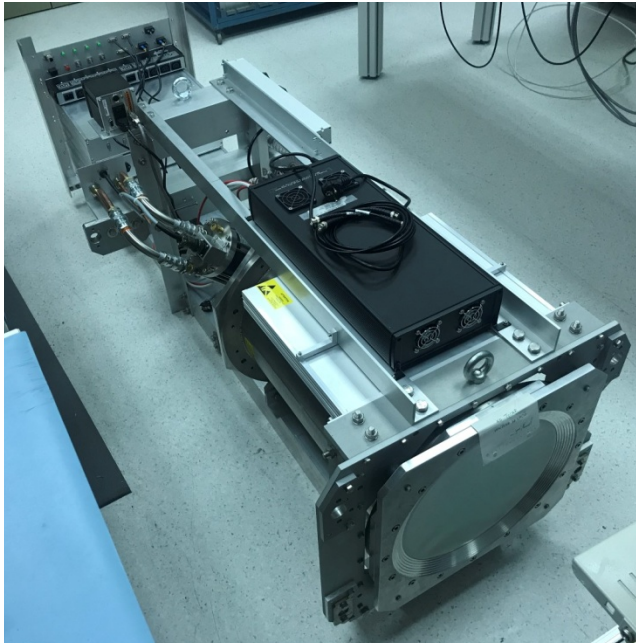


Instrumentation Developments for VGOS at IGN-Yebes Observatory



J. A. López-Pérez, F. Tercero-Martínez, J.M. Serna-Puente, B. Vaquero-Jiménez, M. Patino-Esteban, P. García-Carreño, J.D. Gallego-Puyol, I. Malo-Gómez, I. López-Fernández, M. Díez-González, J. González-García, A. García-Castellano, C. Albo-Castaño, J. López-Ramasco, L. Barbas-Calvo, M. Bautista-Durán, O. García-Pérez, F. J. Beltrán-Martínez, A. Baldominos-Delgado, R. Amils, P. de Vicente-Abad, J. A. López-Fernández

VGOS receivers: NMA & FGI



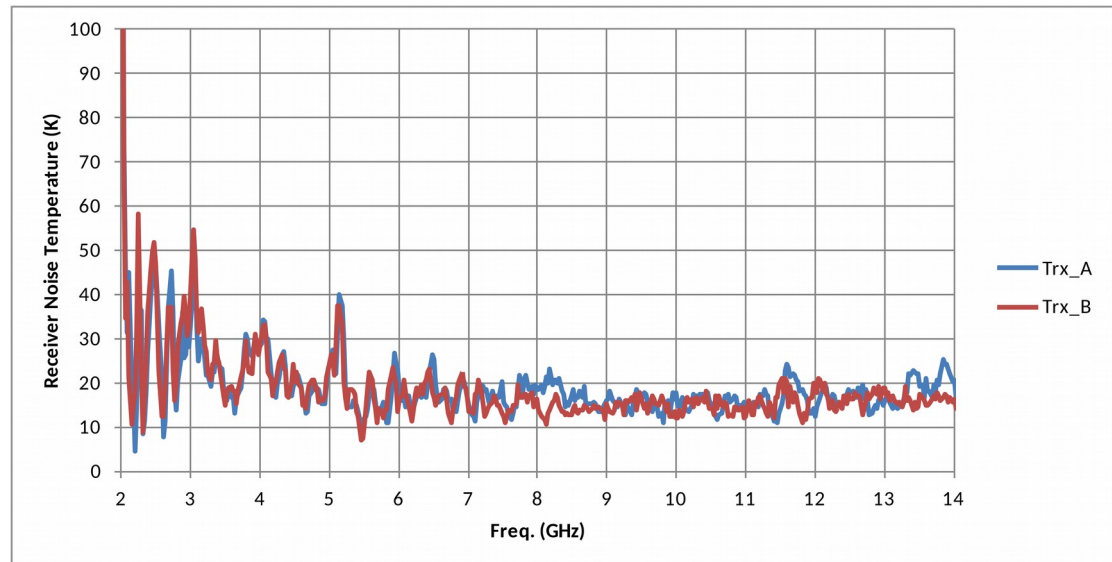
NMA (2 BB receivers)

- 1st delivery scheduled for June 2019

FGI (1 BB receiver):

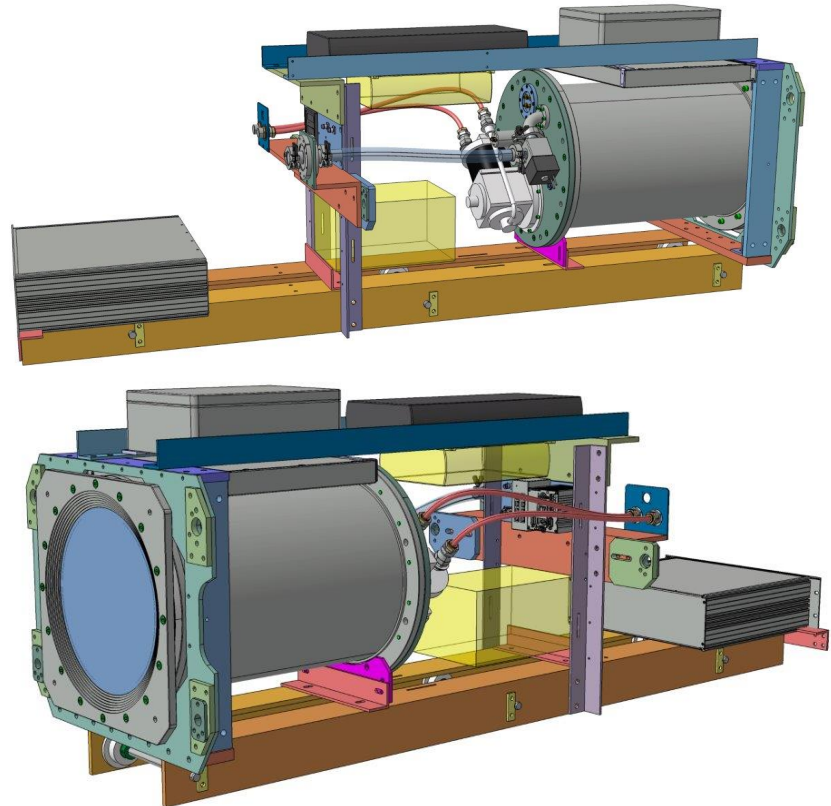
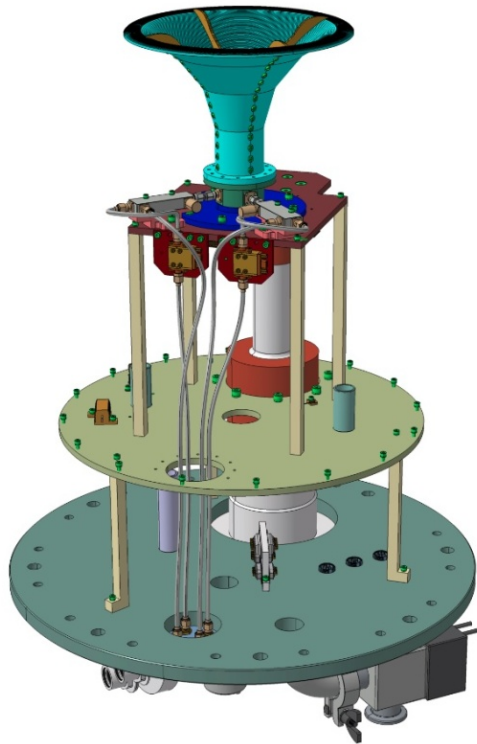
- Delivery September 2019

Trx < 25 Kelvin where no RFI



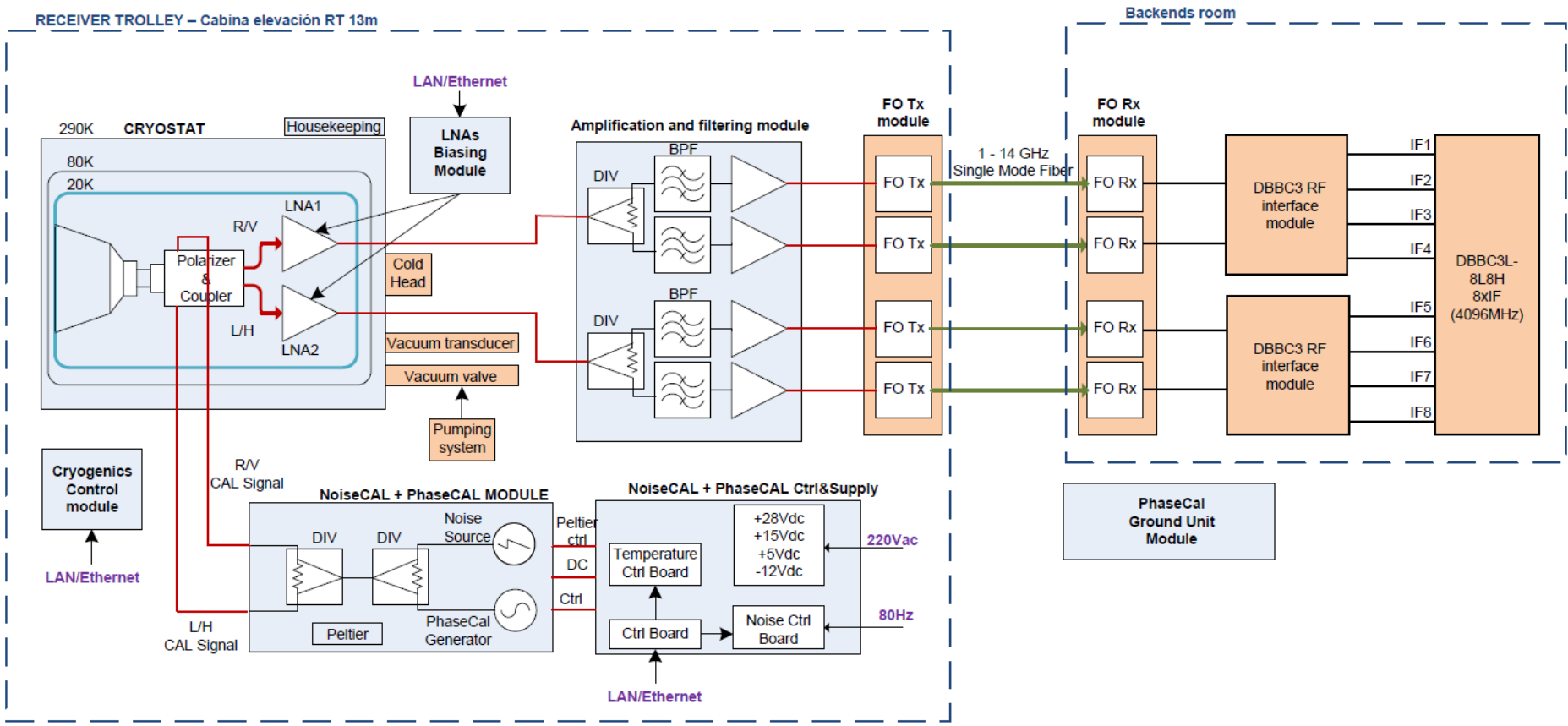
VGOS receivers: Santa María

- Contracted with Spanish company: TTI-Norte
- Delivery: summer 2019



VGOS receivers

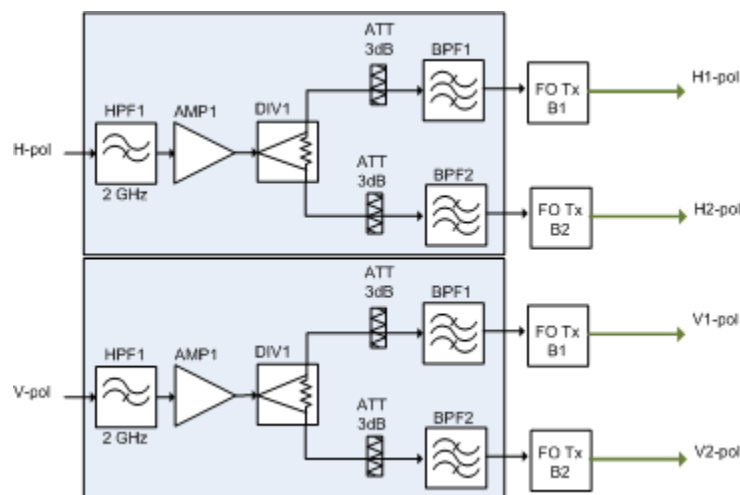
Block diagram



VGOS receivers

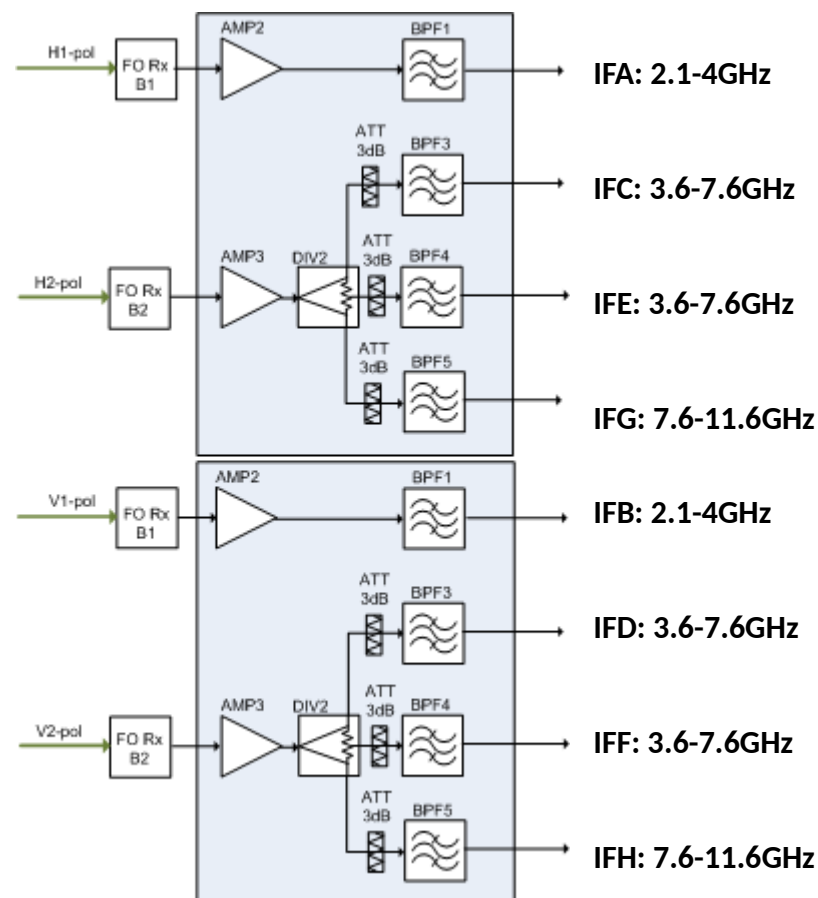
Filtering and pre-amplification units

- Signal conditioning for DBBC3
- Under construction for NMA and FGI



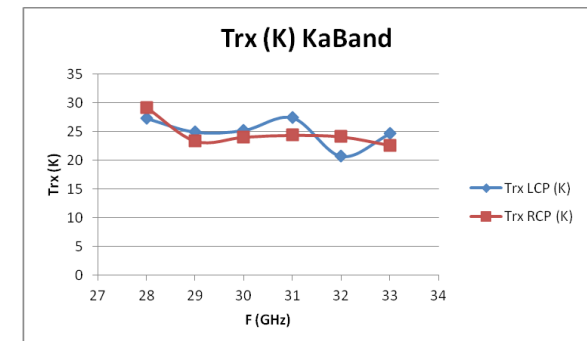
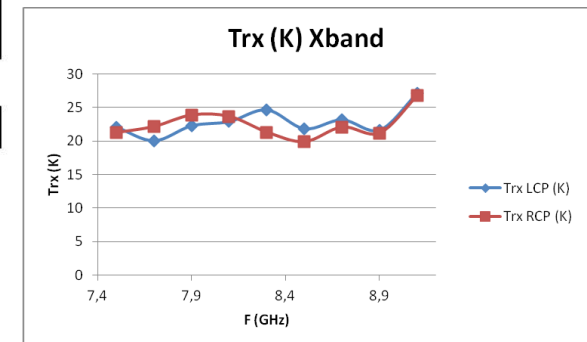
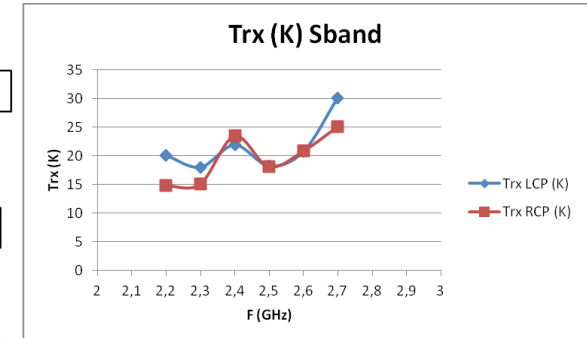
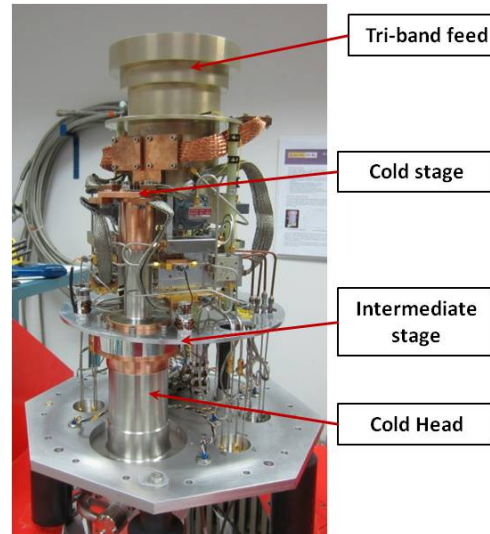
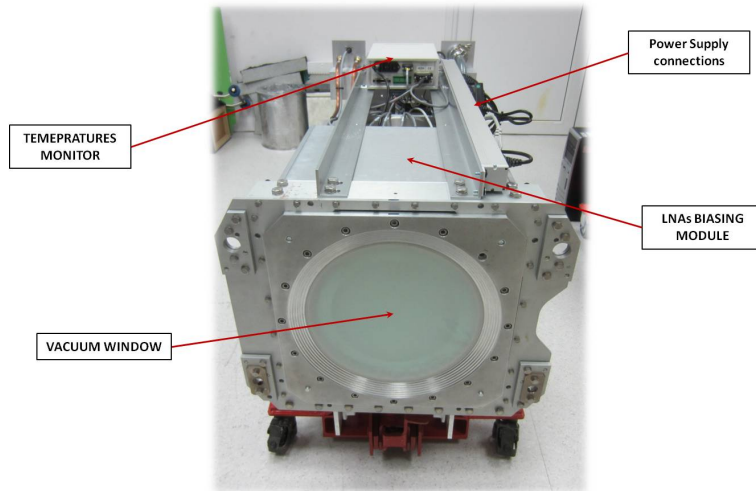
Similar to Hasytack scheme:

- Low band: 2.1 – 5.6 GHz
- High band: 3.6 – 11.6 GHz



Tri-band receivers

FRONT VIEW



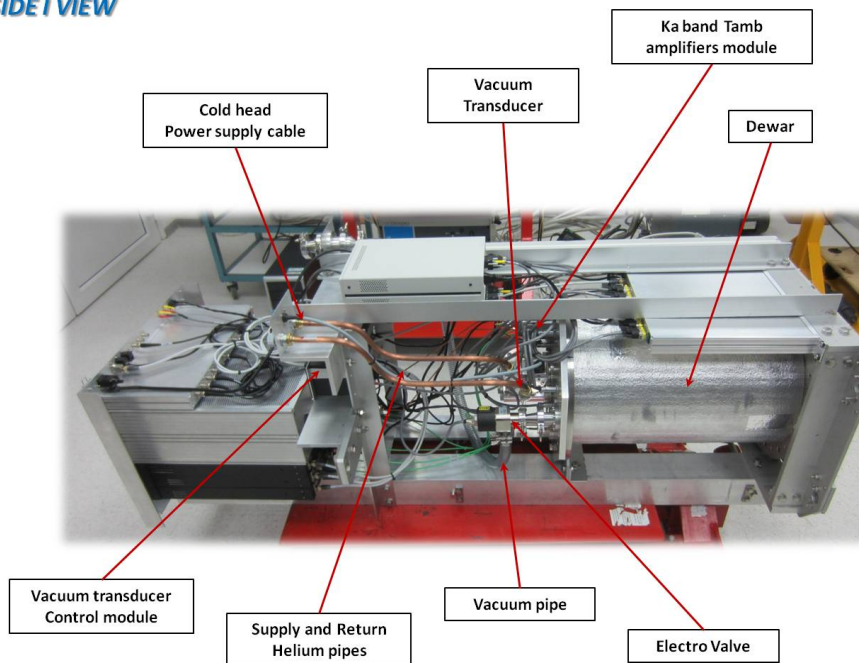
Three tri-band receivers built at Yebes:

- First receiver for Yebes VGOS antenna (2015 - 2016)
- Tri-band receiver for Ishioka station
- Tri-band receiver for Santa Maria VGOS station

Tri-band receiver support to NMA



SIDE VIEW



Installation in VLBI2 antenna: April 2017
Installation in VLBI1 antenna: March 2018

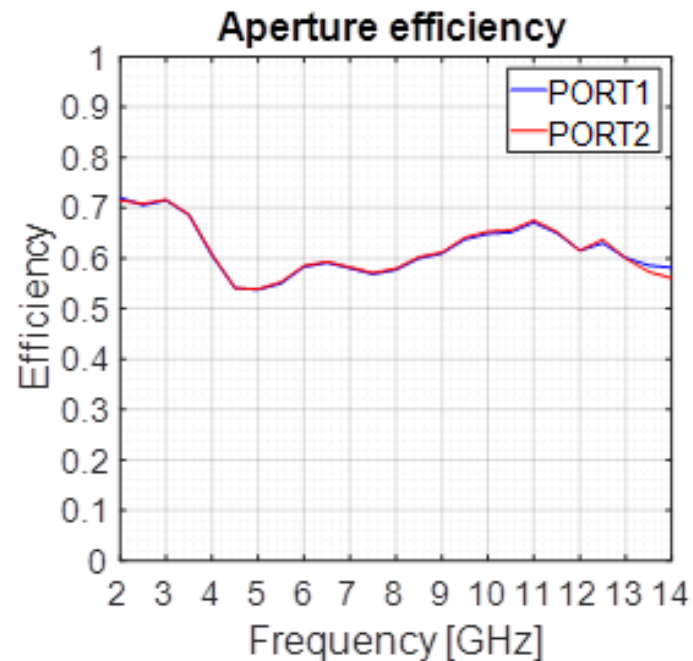
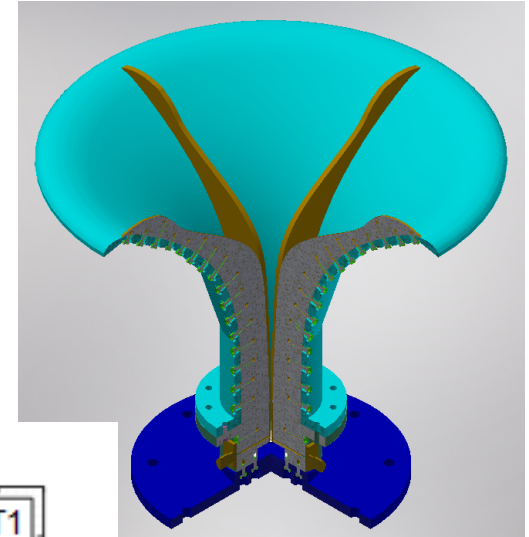
Control software & observation support

VGOS receiver installation in VLBI2 antenna around june/july 2019.

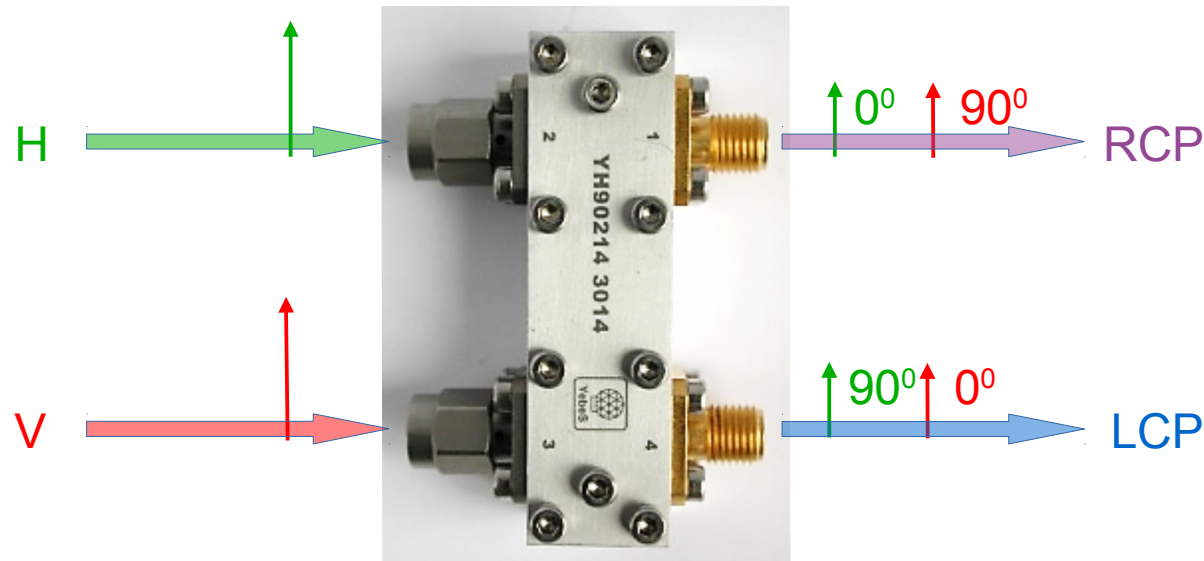
Special care with RFI from ship radars from the fyord !!! Two X-band LNAs were destroyed.

QRFH revisited

- Based in Calech design. Modified profile & interface
- Wideband antenna 2.3-14 GHz.
- Dual linear polarization
- Manufacturing easiness
- 2 output coaxial connectors



Linear to circular polarization conversion using hardware (1)



- 3dB/90° microwave hybrid
- Degradation of LNA average noise by 1.5 K along the 2-14GHz band
- XP < -25 dB. Axial Ratio < 1 dB

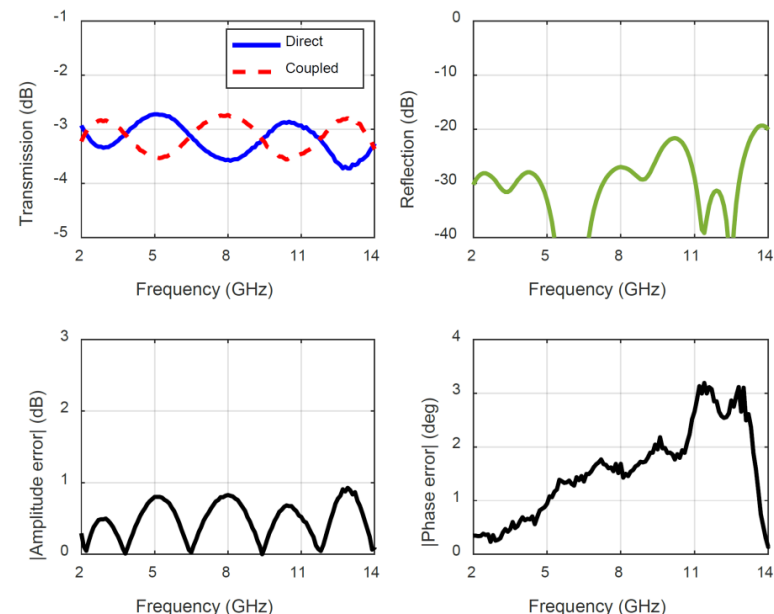
Linear to circular polarization conversion using hardware (2)

Multi-octave stripline hybrids developed at Yebes:

- Bandwidth: 2-14 GHz (1.5 – 15.5 GHz also available)
- Amplitude unbalance < 0.9 dB
- Phase unbalance < 3 deg
- Optimized for cryogenic operation



Hybrid performance



Length of lines is critic, but can be controlled

Linear to circular polarization conversion using software

PolConvert & global fringe fitting efforts by Iván Martí Vidal

Part of EU-VGOS effort (collaboration with European partners)

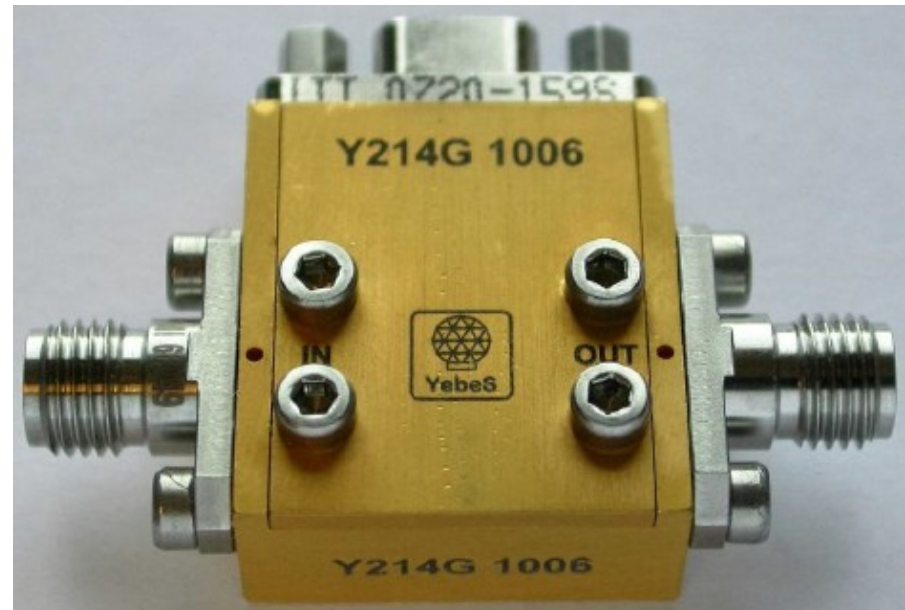
See Walter Alef's on EU-VGOS talk for further details

Yebes ultra-wide band LNAs (1)

Single-ended version

- Hybrid
- Small/compact size
- Usable in 0.5-18 GHz

Band	2-14 GHz
Tn	6.1 K (ETH)
Gain	33.9 dB
IRL	-1.5 dB
ORL	-16.9 dB
Power	36 mW



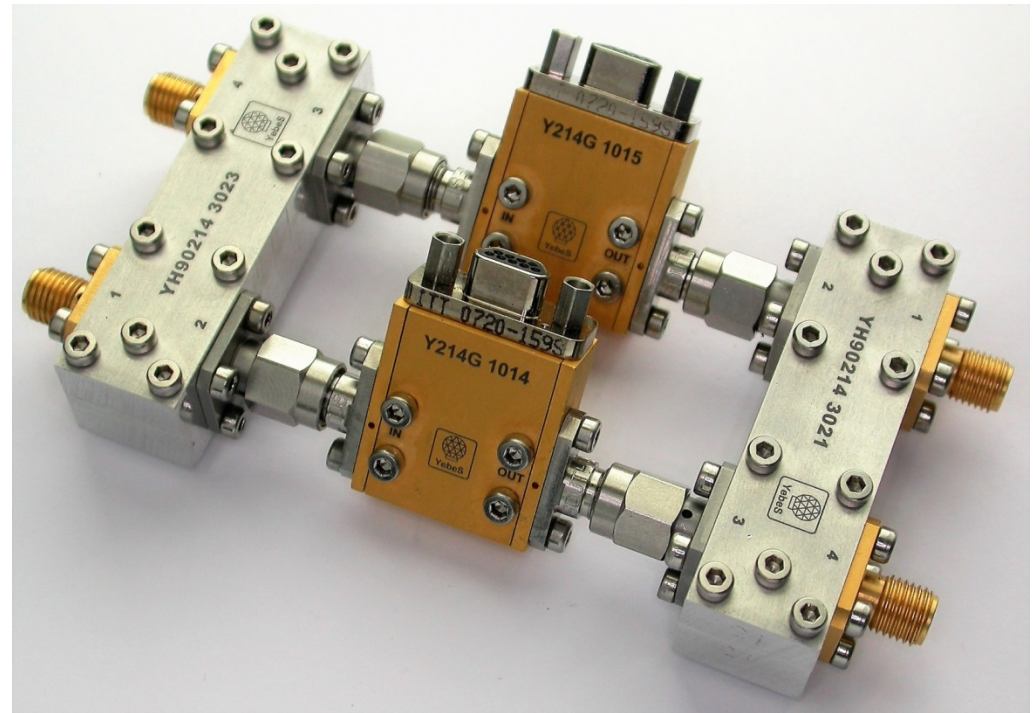
20 x 22 x 9 mm (excl. conn.)

Yebes ultra-wide band LNAs (2)

Balanced version

- Using 3dB/90° hybrids
- Very low noise penalty
- Drastic improvement of input reflection
- Balanced versions for 2-14 GHz and 1.5-15.5 GHz available

Band	2-14 GHz
Tn	7.6 K (ETH)
Gain	33.8 dB
IRL	-21 dB
ORL	-23 dB



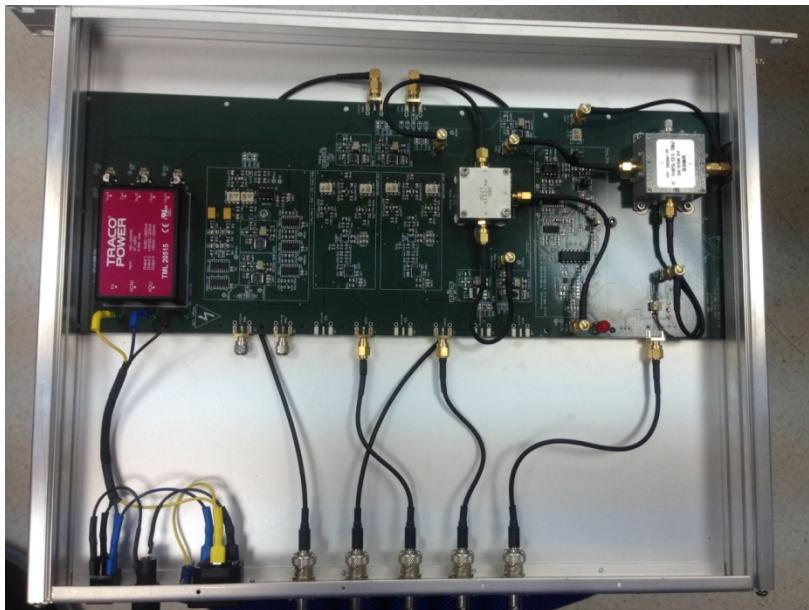
PhaseCal Developments (1)

Cable Delay Measurement System (CDMS)

Follows Haystack legacy approach (**A.E.E. Rogers**) integrated in a single board inside a 19"/1U rack

Subsystems:

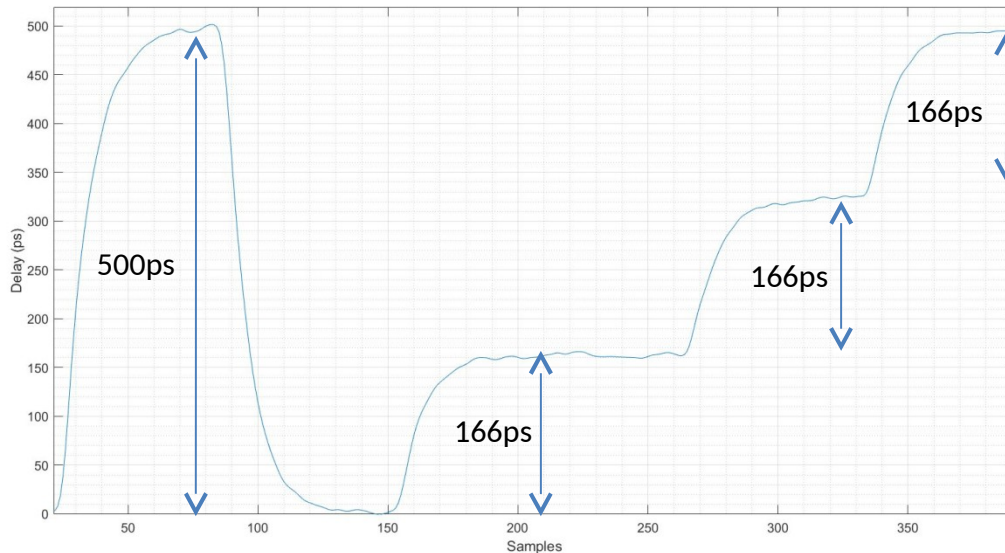
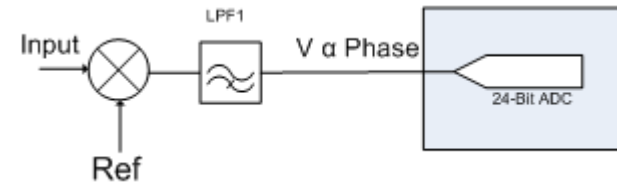
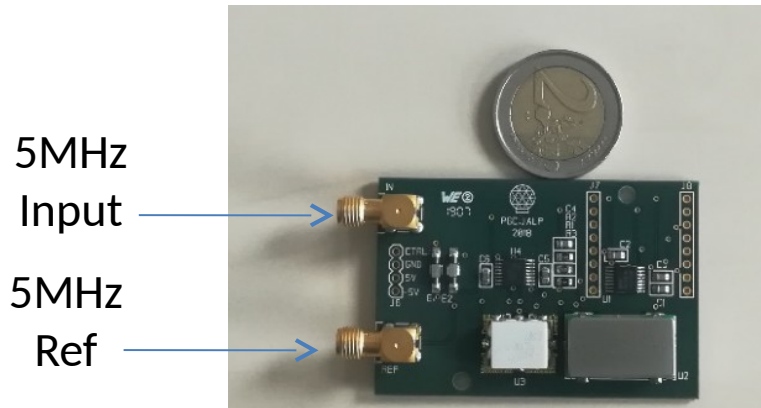
- Phase Detector / Modulating Generator
 - 5MHz – 25Hz Generator
 - Start / Stop Generator
 - Cable Multiplexer
-
- Units developed at Yebes labs for:
BKG, AGGO, NMA and FGI



IT REQUIRES A FREQUENCY COUNTER

PhaseCal Developments (2)

New CDMS under test



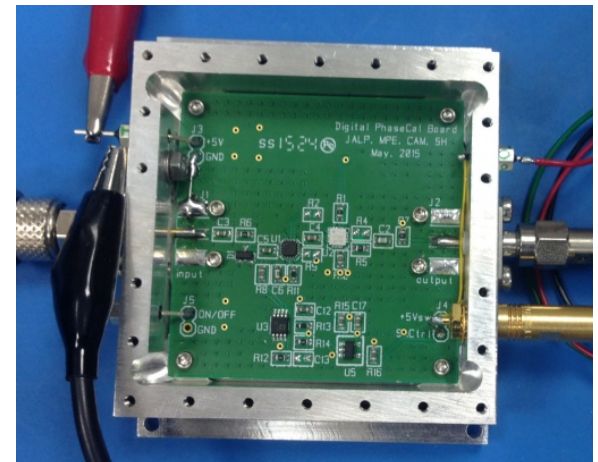
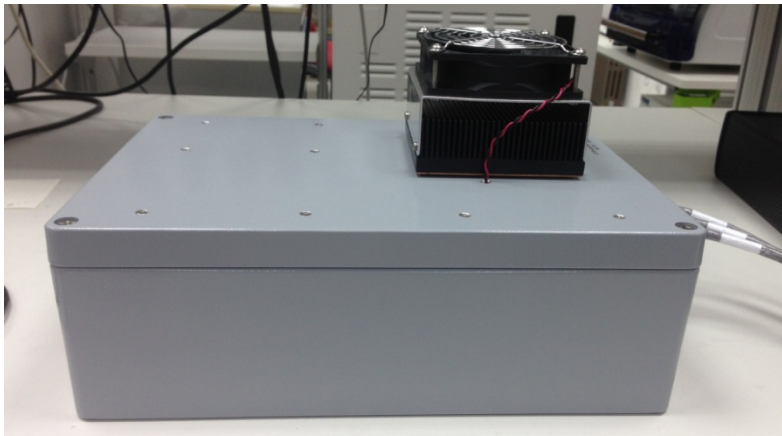
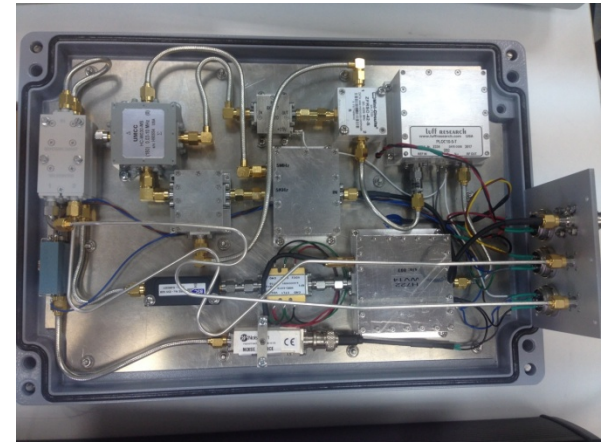
- Phase detector at 5MHz
- 24-bit ADC
- RMS: 5 ps (0.003 deg)

**NO FREQUENCY COUNTER
ONLY A RASPBERRY PI**

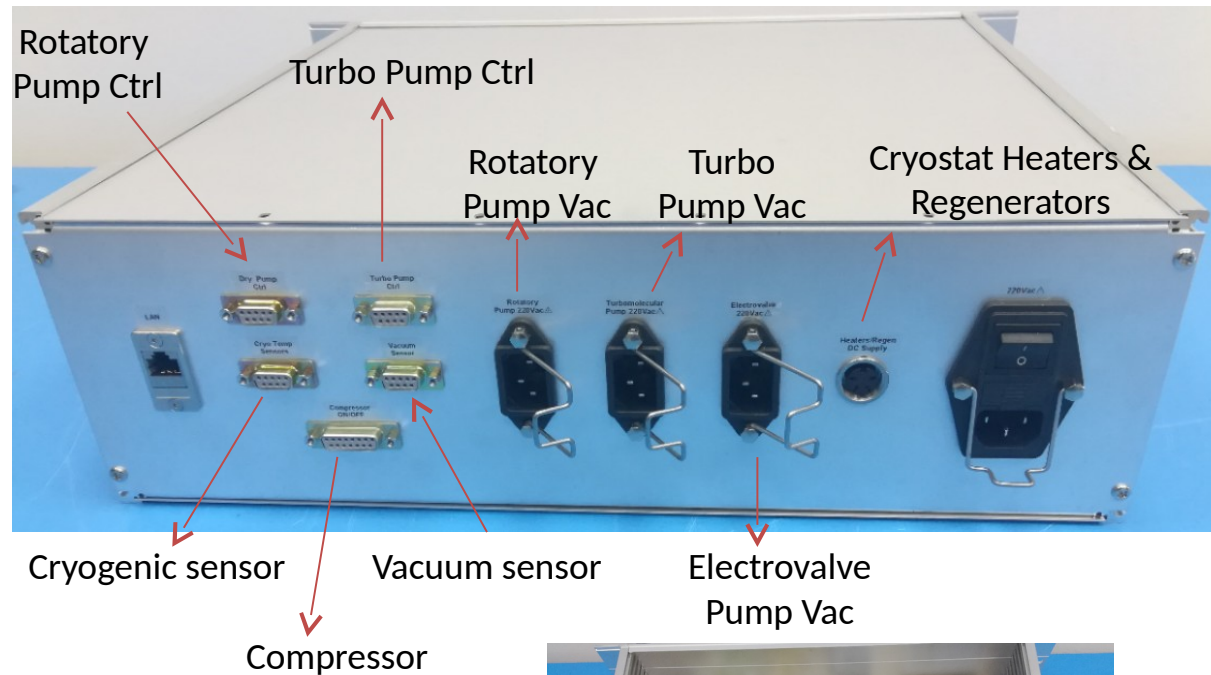
PhaseCal Developments (3)

Noise & Phase Cal Unit

- At the receiver trolley (Front-end). Shorter cables
- 2-14 GHz
- Pulse spacing: 10MHz
- Pulse Generator: based on Hitite Ultrafast logic gates (**Haystack approach**)
- NoiseCal can be switched at 80 Hz rate
- Temperature stabilization: Peltier cooler and passive insulation.
- Units developed at Yebes labs for: BKG, AGGO, NMA and FGI

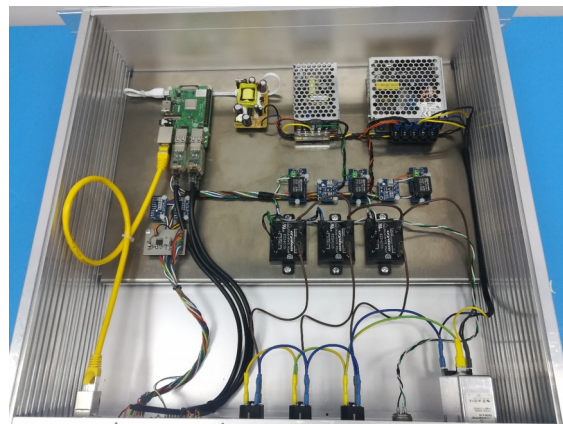


Cryogenics & vacuum control System

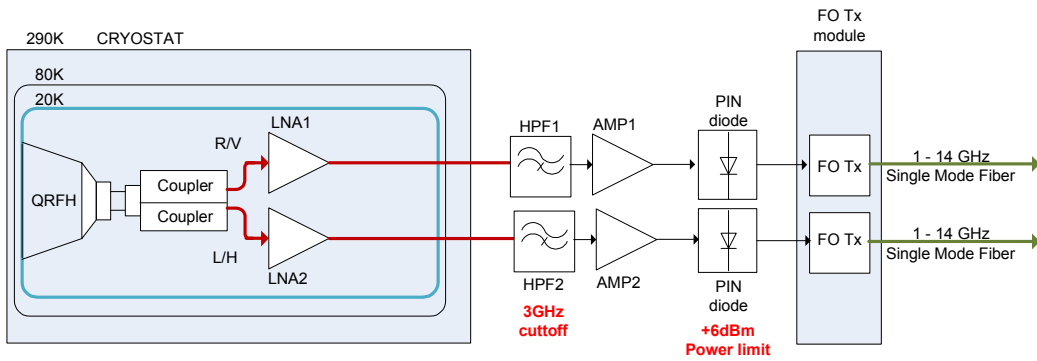


LAN remote monitor and control of:

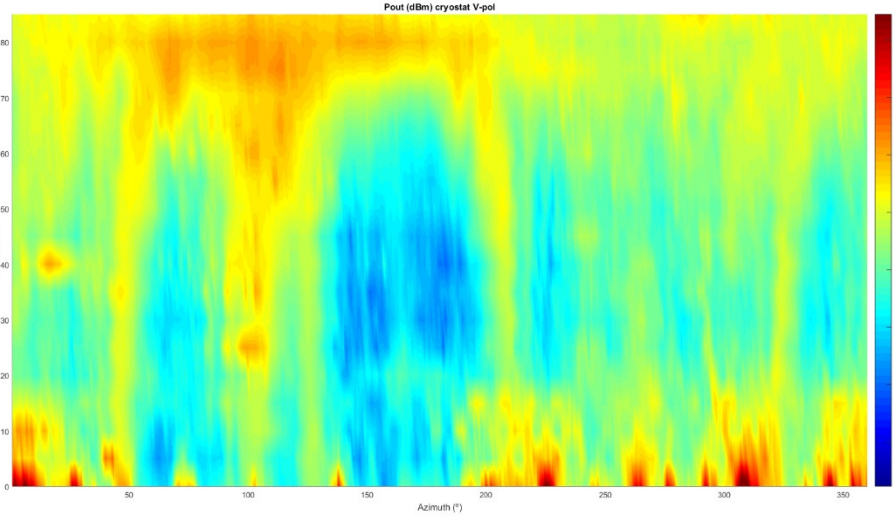
- Rotatory Pump
- Turbo Pump
- Cryogenic temperature sensors
- Vacuum sensor
- Compressor
- Electrovalve
- Heat Resistors
- Regenerators



RFI at VGOS band in Yebes

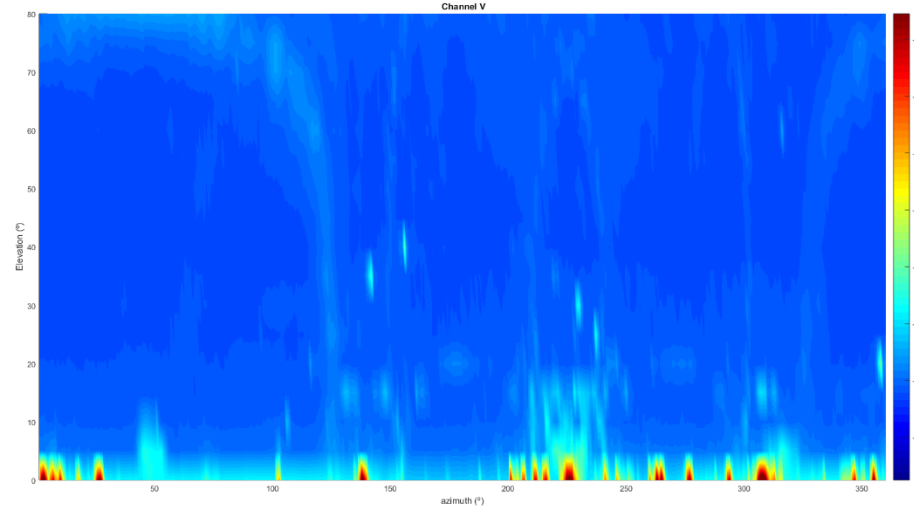


Elevation over Azimuth RFI power maps



a) Original situation without RFI filters

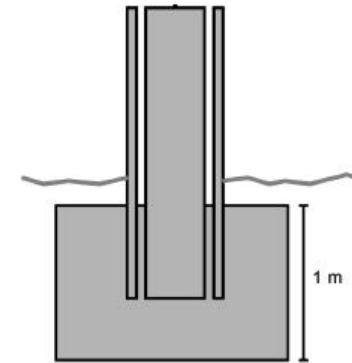
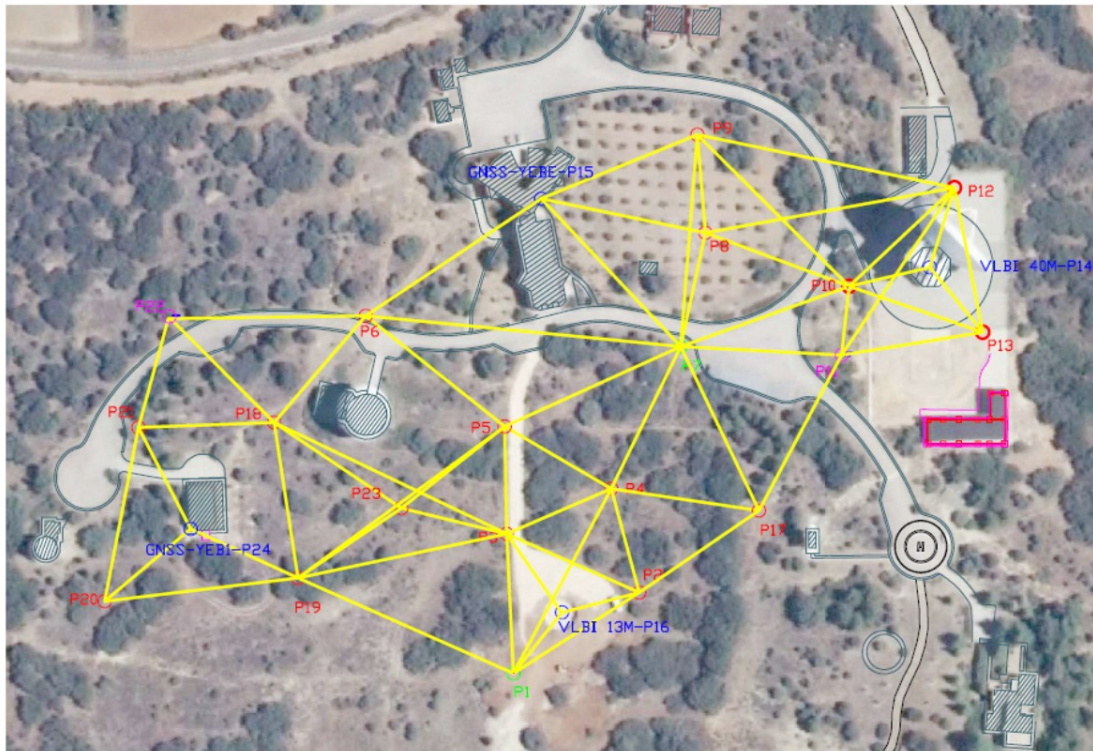
Tsys ≈ 70K



b) New situation with filters and pre-amplifier

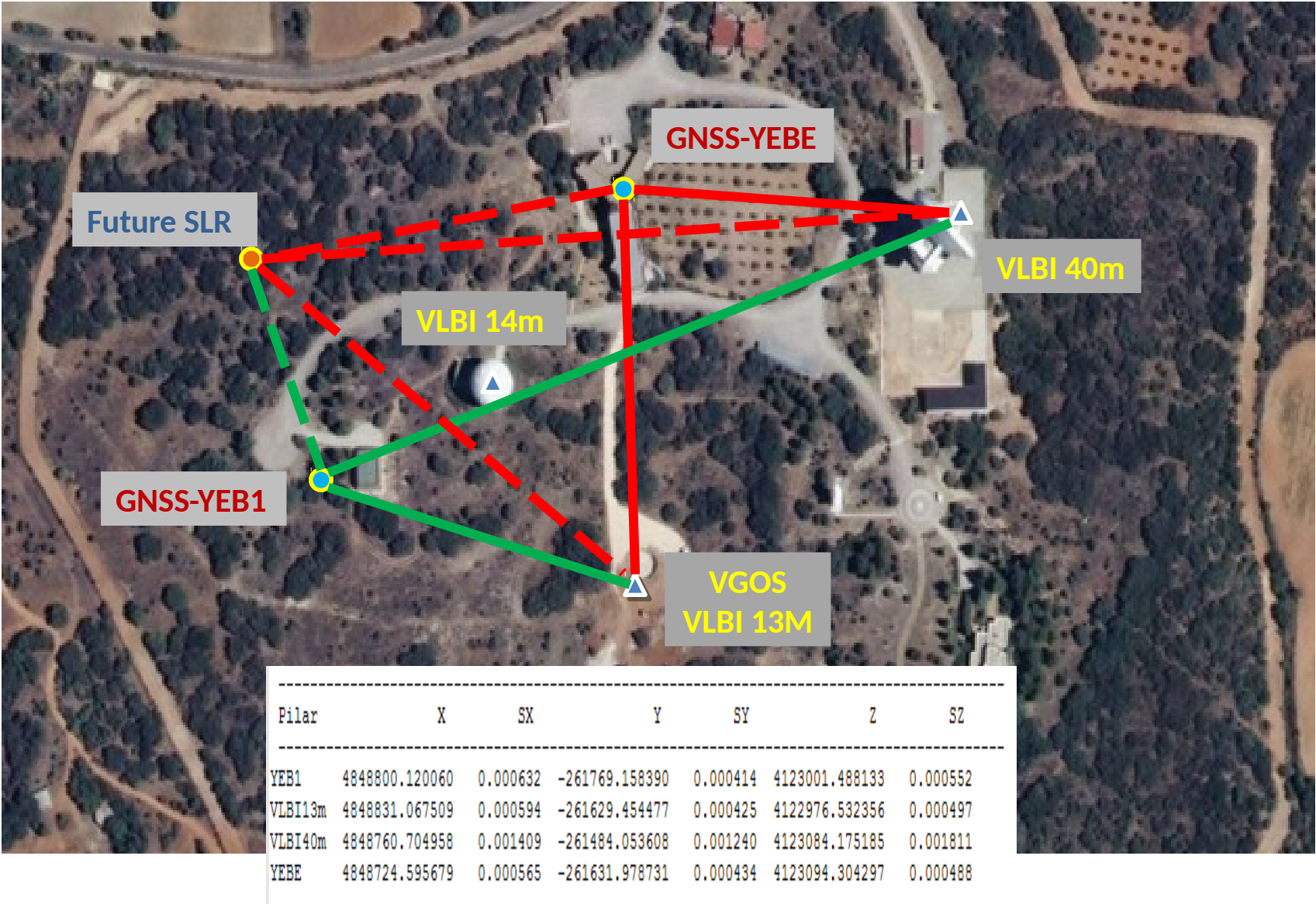
Tsys ≤ 50K

Local Tie. Network of pillars



- 20 Reinforced concrete pillars of 30 cm in diameter and 1.30 m high.
- Isolated 5 cm of a protective tube which provides thermal and climatic stability.

Local Tie (1)

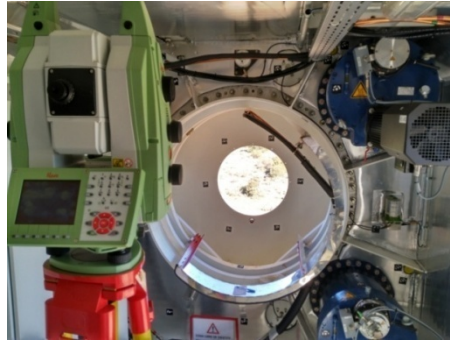


Thank for your attention



Invariant Reference point of the RT measurements

Inside the cabin



Outside the cabin

