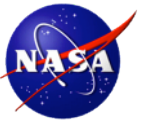
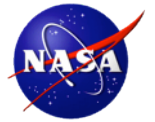


Effects of Tropospheric Spatio-Temporal Correlated Errors on the Analysis of Space Geodetic Data

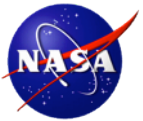
Andres Romero-Wolf,
C.S. Jacobs, and J. T. Ratcliff
Jet Propulsion Laboratory,
California Institute of Technology
March 6, 2012



Motivation



- VLBI error budget (becoming less Gaussian dominated).
- Reduction of thermal errors will require modeling of correlated noise sources.
- Troposphere errors have correlations in both space and time.
 - Kolmogorov frozen flow model.
 - Modeling delay errors Treuhaft & Lanyi (Radio Sci. 1987)
 - Water vapor radiometer (WVR) measurements.
- Instrumental errors introduce station specific temporal correlations.



Thermal Errors

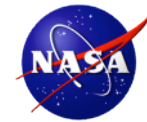


Recording Rate	Improve	
128 Mb/sec	-	Last 10 years
256 Mb/sec	1.4	Recent RDV runs
512 Mb/sec	2	VLBA continuum
2 Gb/sec	4	Mark 5C (R&D fringes)
4 Gb/sec	5.7	Mark 5C dual bank
16-32 Gb/sec	11.3-16	Haystack Mark 6

Data rates are sky-rocketing.
Factors of 10 improvement in the near future.



Future: Trop. Turbulence Dominates

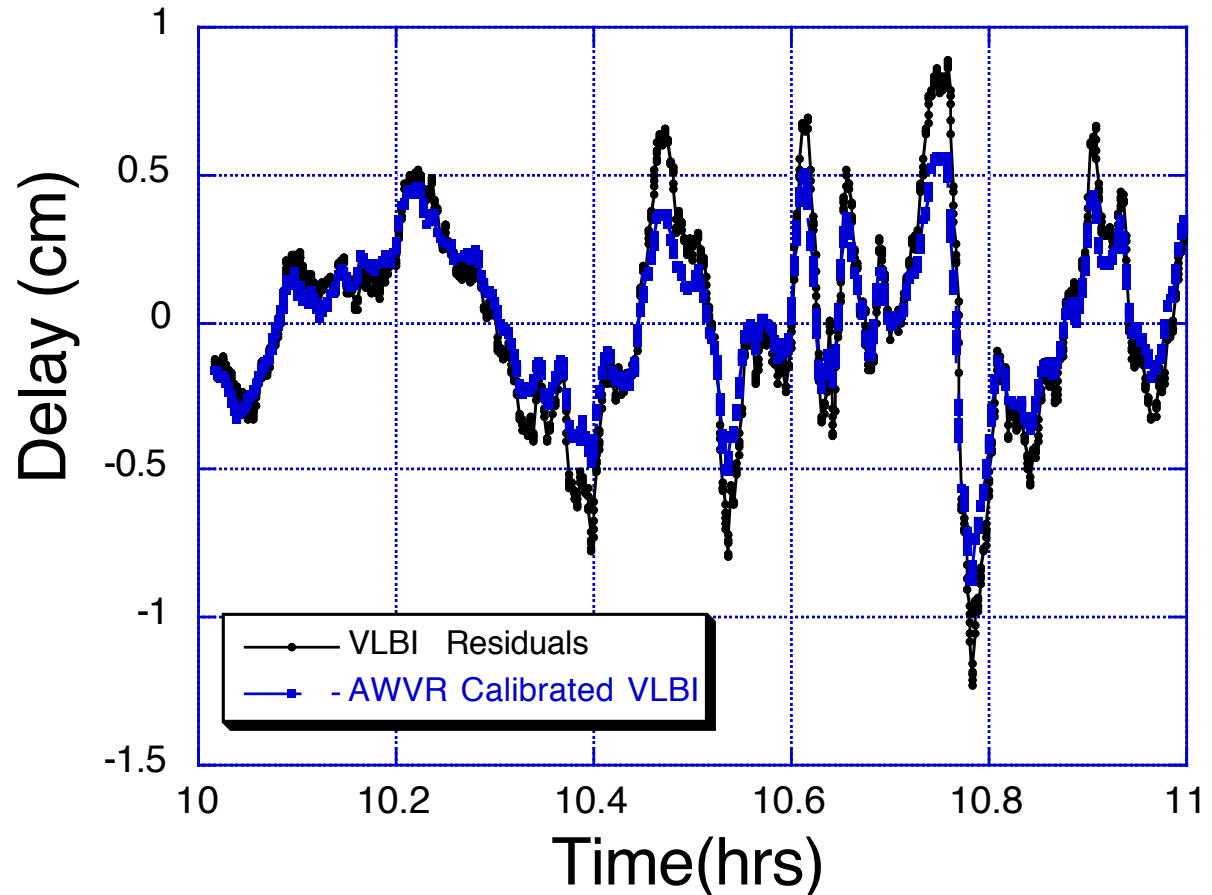


- Monitor 22 GHz/1.3cm water (rotational) line brightness temperature along line-of-sight.
- 3mm scatter reduced to 1mm Goldstone-Madrid 8000 km baseline using X/Ka phase delays.
- As thermal errors average down, the troposphere fluctuations dominate the residuals.

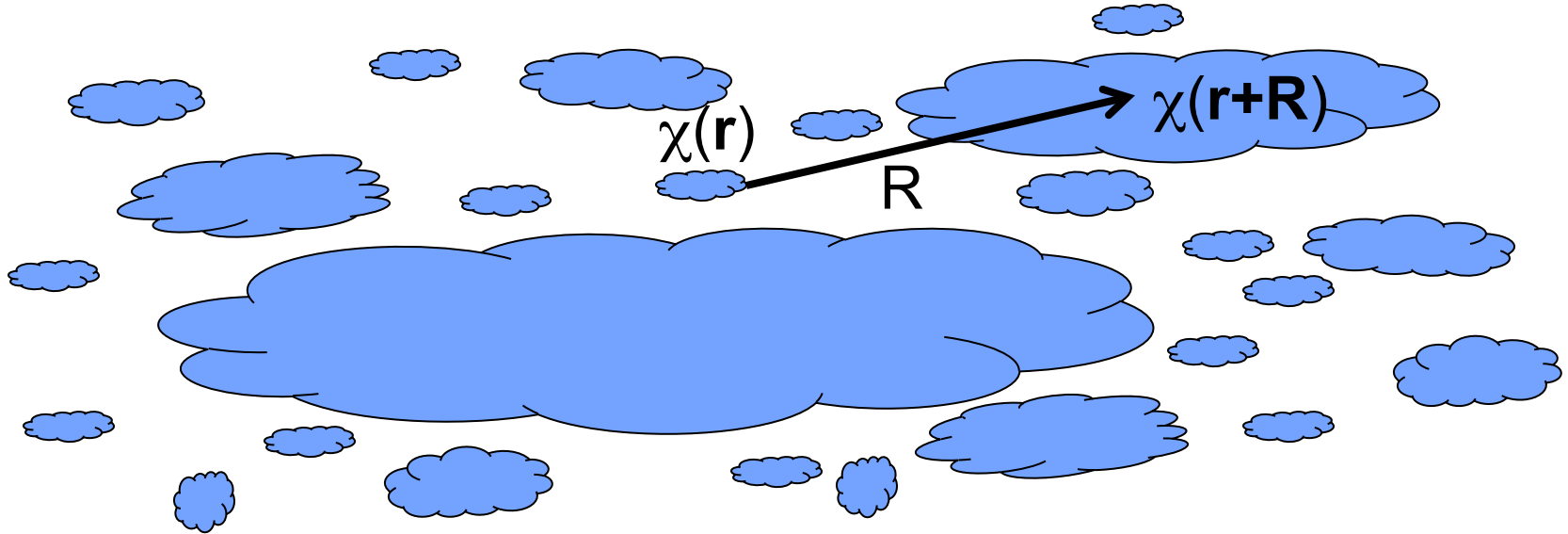


Measured with the JPL Advanced Water Vapor Radiometer

VLBI Delay Residuals DOY 200 Ka-Band DSS26-DSS55



*Jacobs et al, AAS Winter 2005.
Bar Sever et al, IEEE, 2007.*

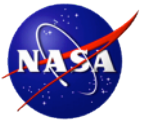


Structure function of refractivity (Tatarskii, 1961)
Based on Kolmogorov turbulence.

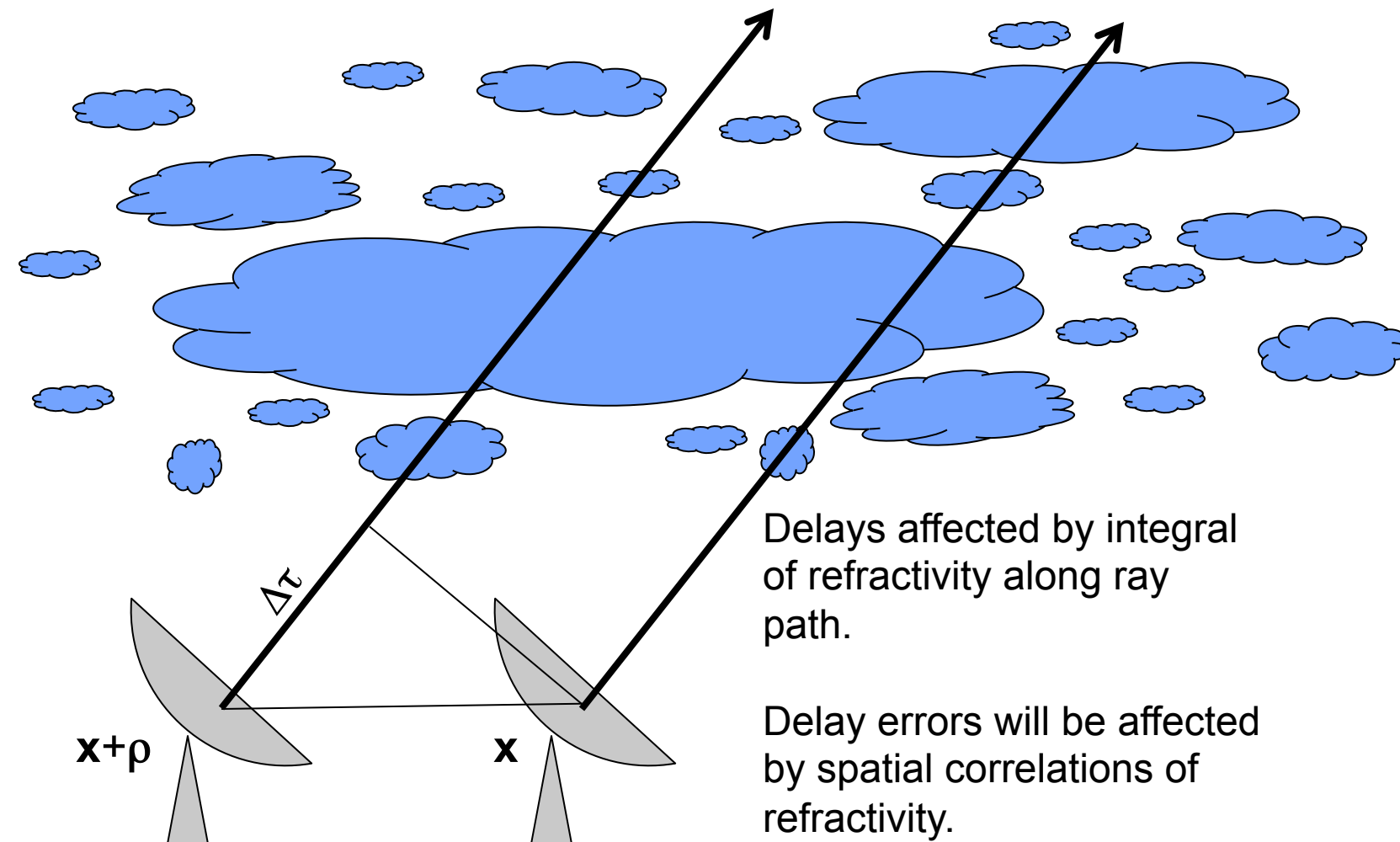
$$D_{\chi}(R) = C^2 R^{2/3}$$

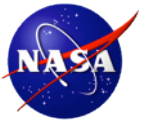
Structure function
contains information
statistical scatter
and covariance.

Constant related
to rate of energy
dissipation and
viscosity.

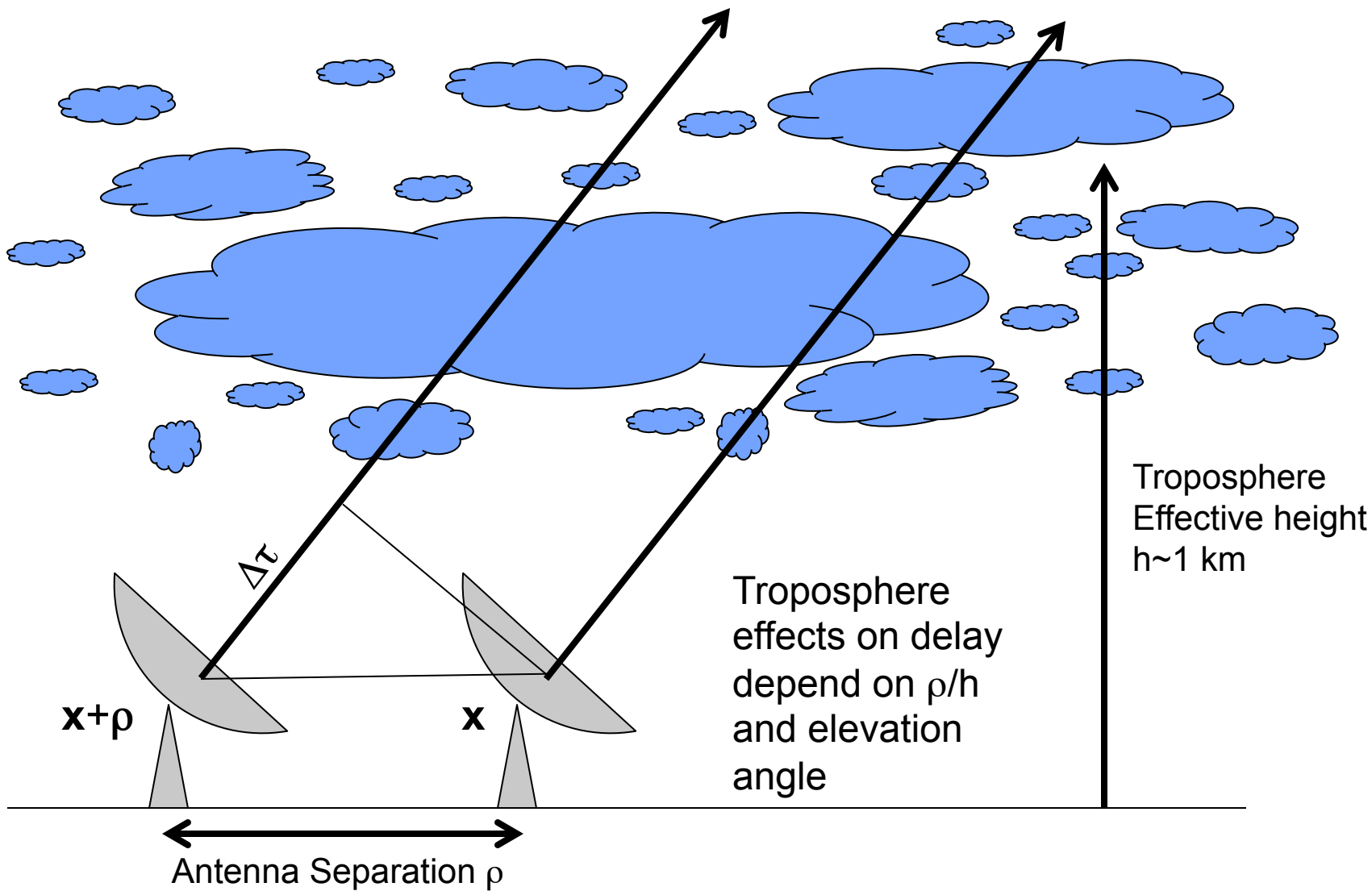


Spatial Correlations on Delay Errors



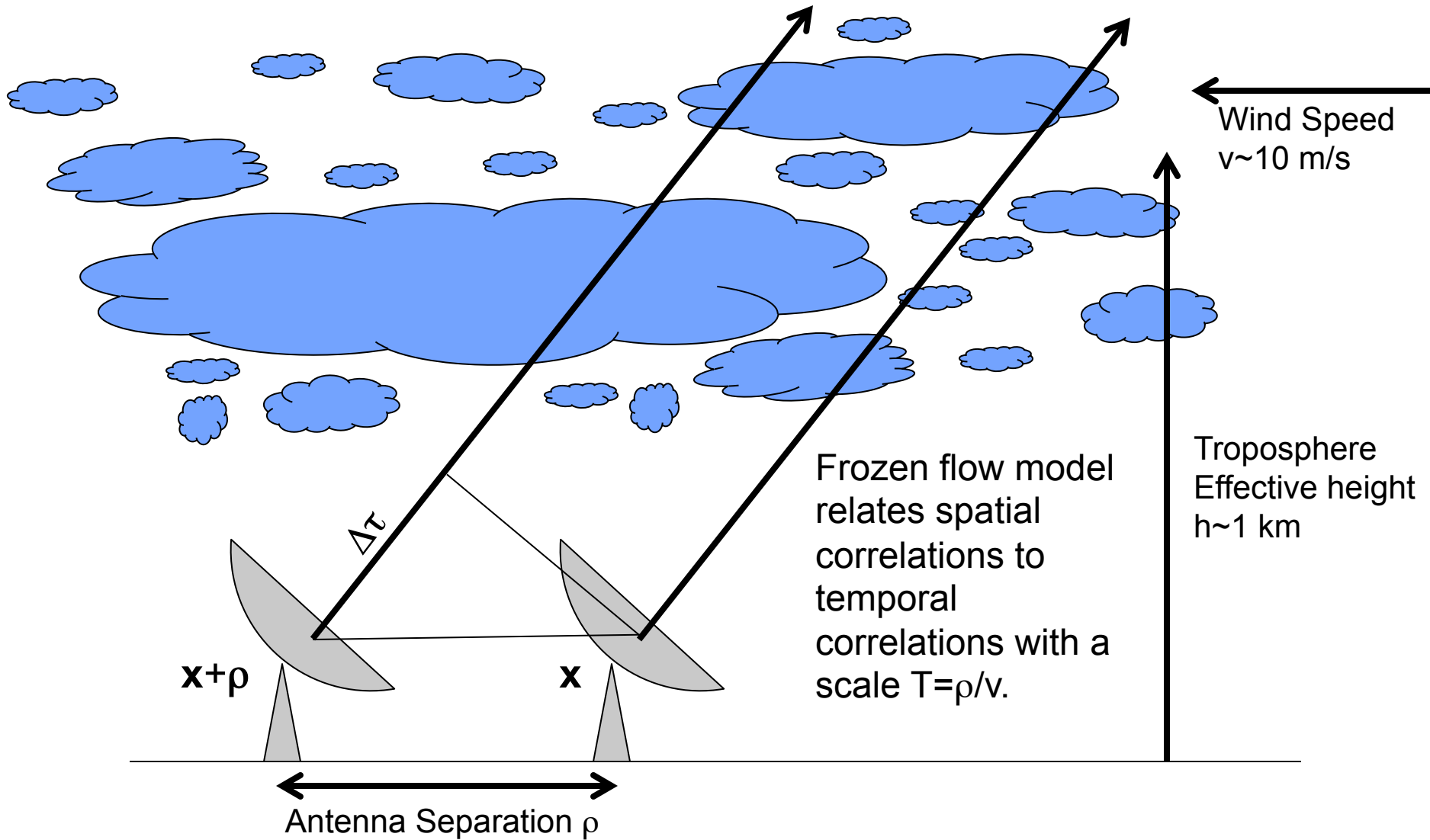


Tropospheric Effects on Delays





Temporal Correlations on Delay

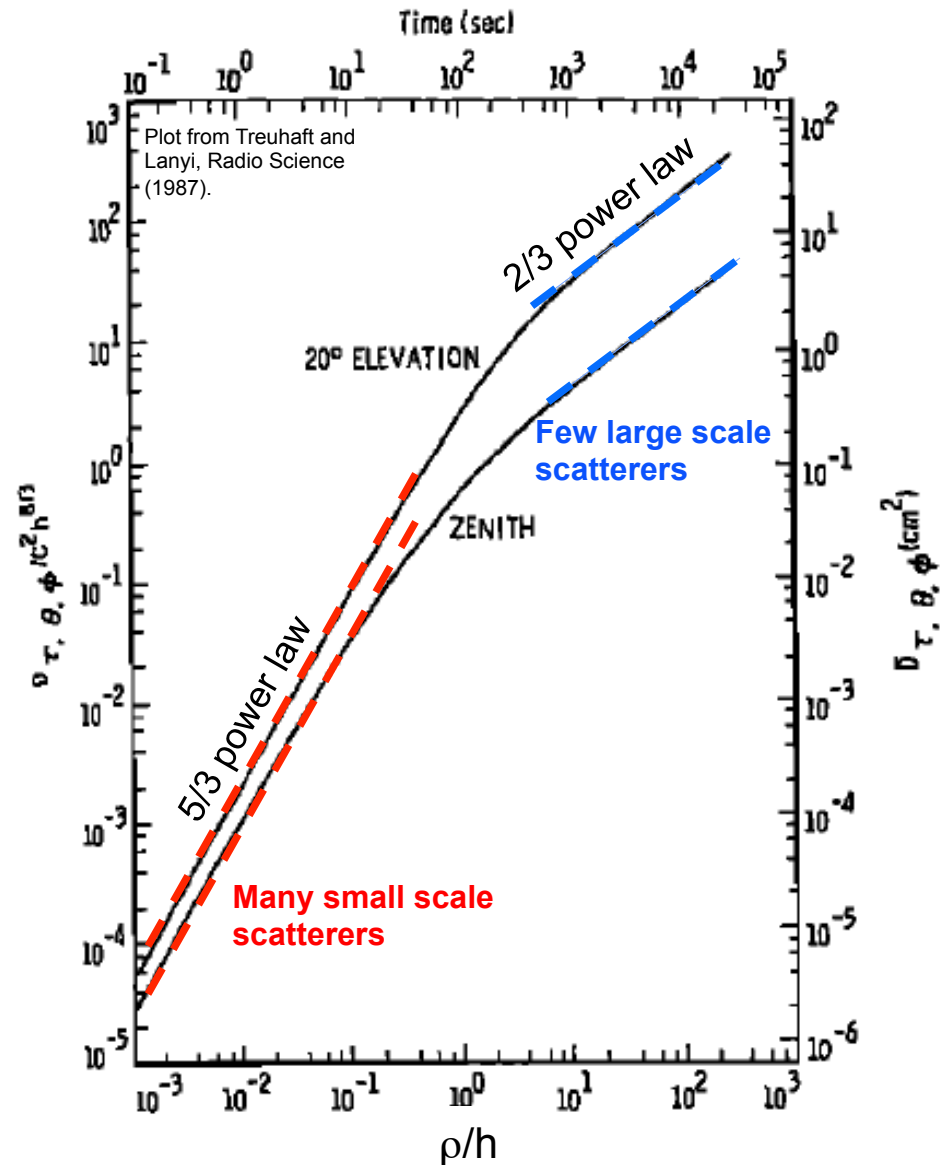


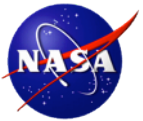
Delay structure function
 $D_{\tau}(\rho) = \langle (\tau(\mathbf{x}+\rho) - \tau(\mathbf{x}))^2 \rangle$
 takes in refractivity variations
 via the media effects on the
 delay.

Delay error correlations depend
 on the ratio of antenna distance
 to troposphere height (ρ/h) and
 elevation angle.

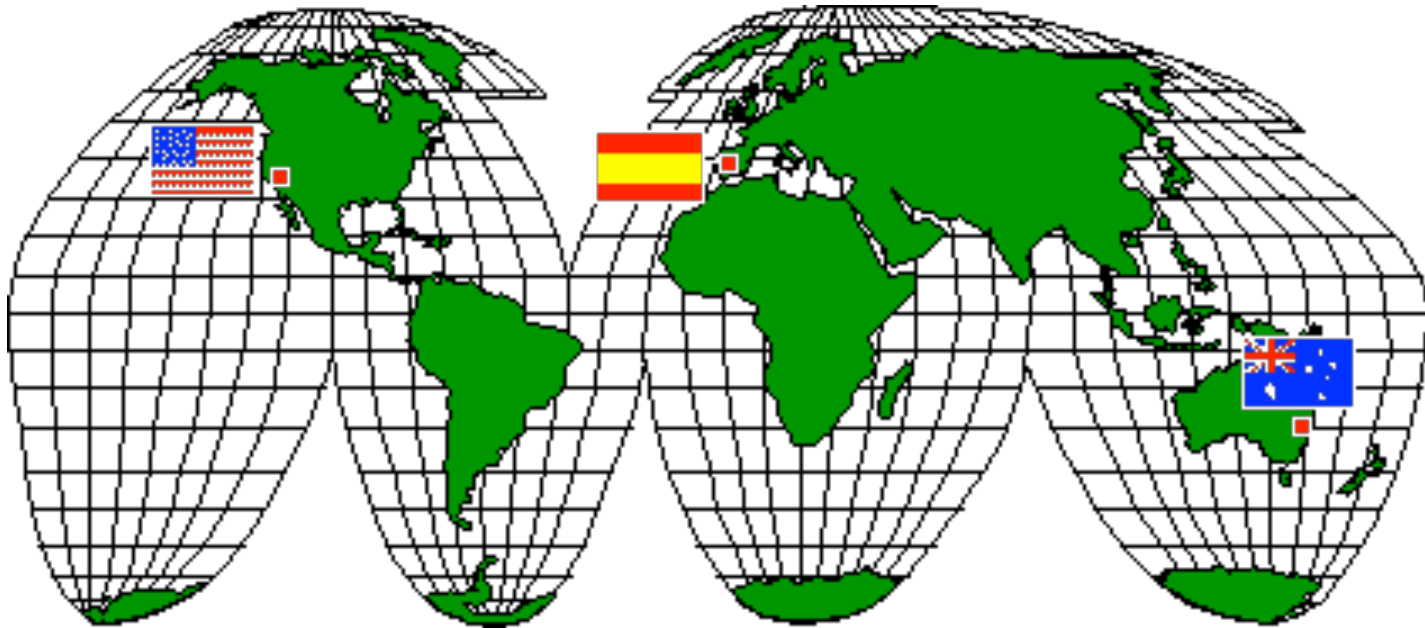
Broken power law with smooth
 switch over at
 $\rho/h \sim 1$.

Correlations in space and time
 are derived from structure
 functions.





The NASA Deep Space Network



Credit: <http://deepspace.jpl.nasa.gov/dsn/antennas/index.html>

VLBI Capabilities:

California-Madrid: 8,400 km

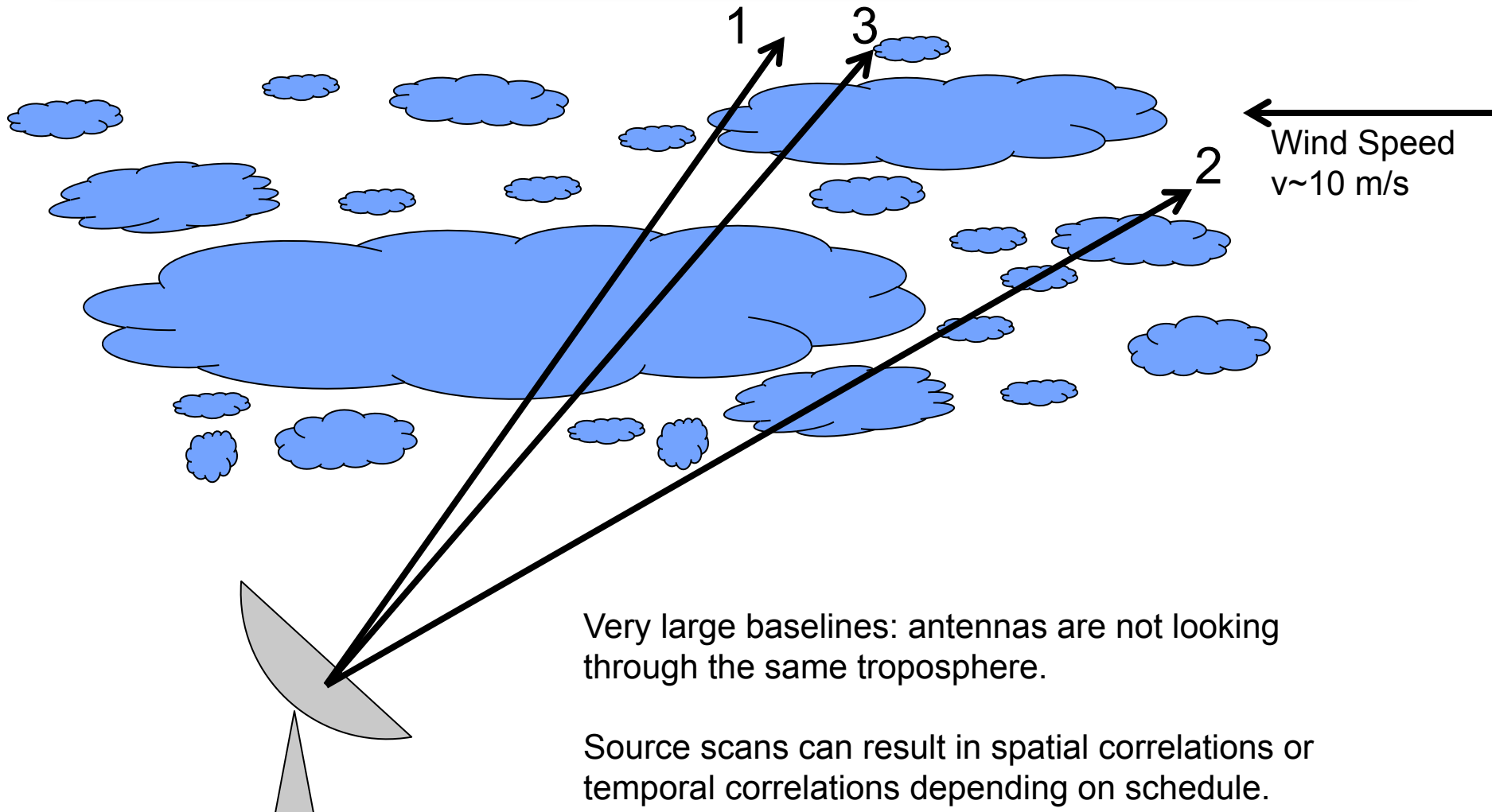
California-Canberra: 10,500 km

X/Ka (34 m antennas)

S/X (34 m and 70 m antennas)



Single Site Spatio-Temporal Correlations





Troposphere Covariance in X/Ka Catalog



X/Ka catalog effort determines celestial coordinates of radio sources, and baseline vectors between DSN stations, for use in spacecraft navigation.

Including trop. cov. in X/Ka catalog improves wRMS agreement with ICRF-2 by 7% (using the exact same data and exact same modeling).

A 7% improvement would take 14% more data to get the same result by pure averaging.

For X/Ka going into its 6th year that would mean almost another year of data to get the improvement we get from trop cov.

wRMS in RA

w/o Trop. Cov. : 215 μs

w/ Trop. Cov. : 198 μs

wRMS in DEC

w/o Trop. Cov. : 300 μs

w/ Trop. Cov. : 283 μs

δDec vs. Dec slope

w/o Trop. Cov. : 1.6 +/- 1.0 $\mu\text{s/deg}$

w/ Trop. Cov. : 1.1 +/- 0.9 $\mu\text{s/deg}$



TEMPO



Time and Earth Motion Precision Observations

Rapid turnaround VLBI measurements earth orientation.

Support of spacecraft navigation, which needs extremely timely and accurate earth rotation information.

Uses the California-Madrid baseline which is mostly East-West directed making it most sensitive to UT1-UTC measurements.



Analysis Comparison

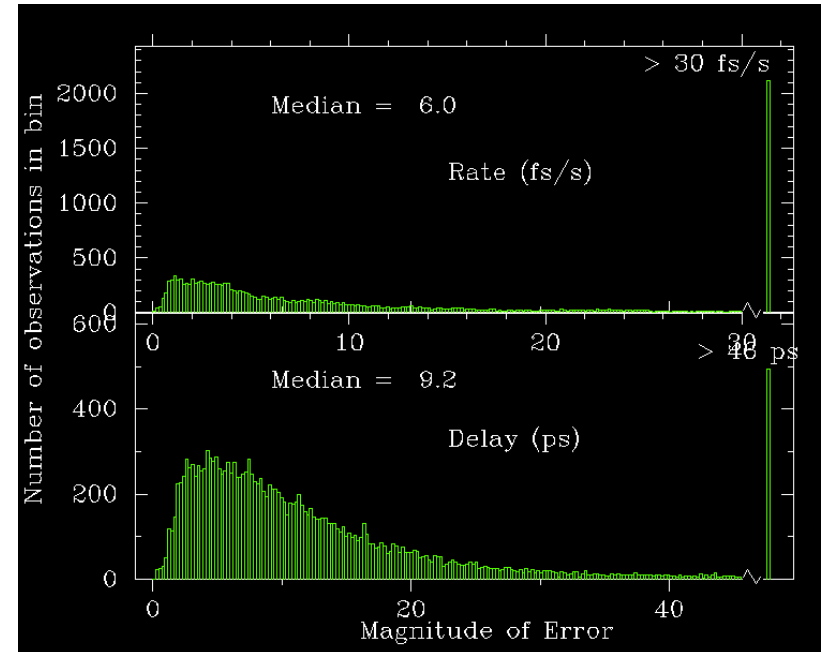


Traditional:

- Gaussian diagonal noise.
- Elevation dependent weighting.
- Troposphere parameter breaks.

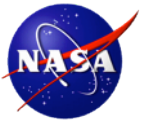
Troposphere Covariance:

- Correlated noise.
- Spatio-temporal dependence.
- Single troposphere estimate per site per session.



Thermal errors:

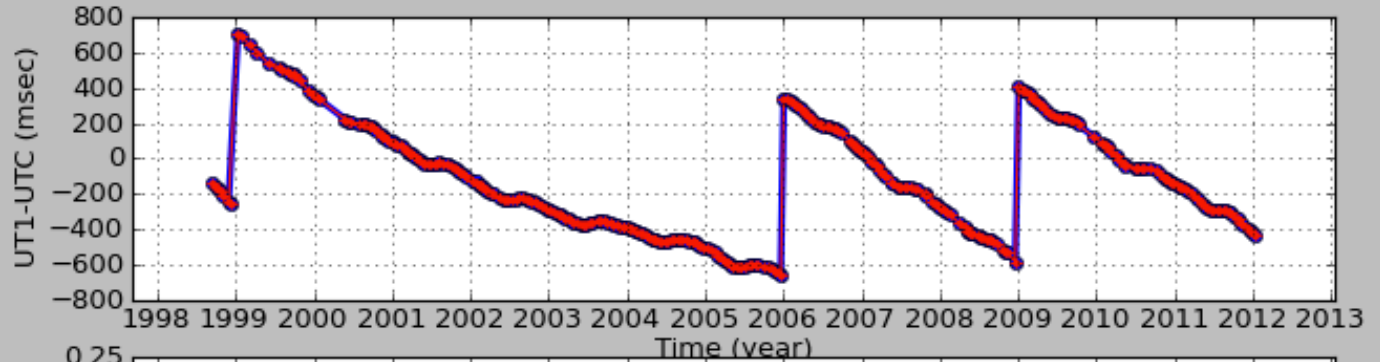
9.2 ps in delay
6.0 fs/s in delay rate



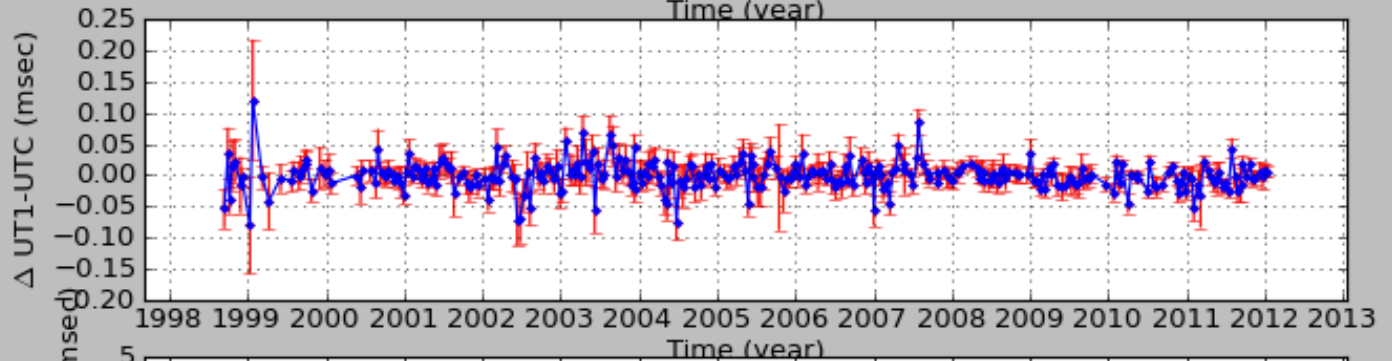
Trop. Cov. Vs. Traditional



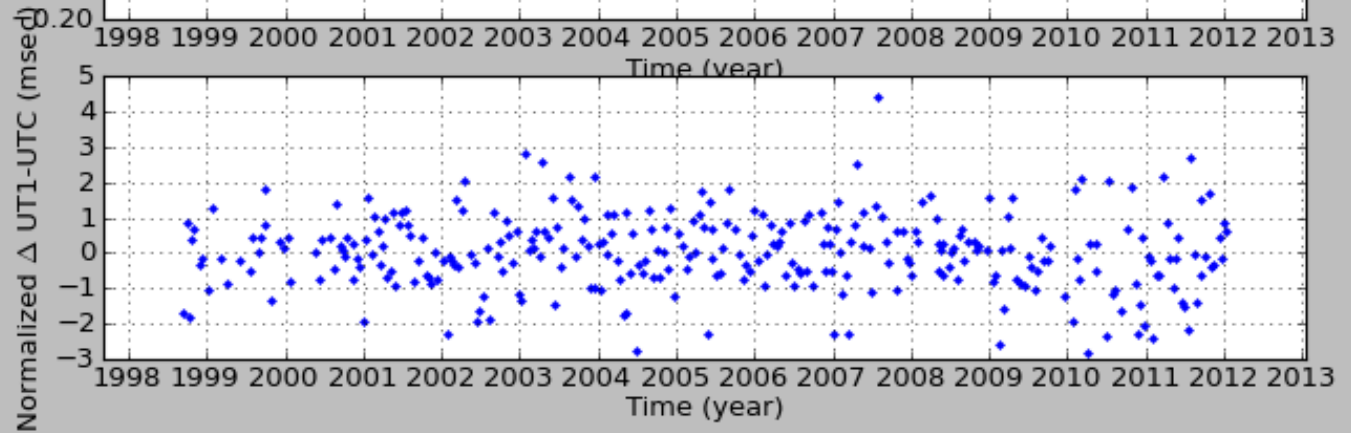
Total UT1-UTC

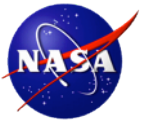


Difference



Normalized Difference





TEMPO Analysis Results



Use JPL Kalman Earth Orientation Filter (KEOF) multi-technique combo as truth.

Subtract KEOF from TEMPO and compare wRMS of residuals.

The wRMS improves by 17% for baseline transverse direction and 27% in the baseline vertical directions.

wRMS in Baseline Transverse

w/o Trop. Cov. : 355 μs

w/ Trop. Cov. : 295 μs

wRMS in Baseline Vertical

w/o Trop. Cov. : 564 μs

w/ Trop. Cov. : 467 μs



Conclusions



As thermal errors are being reduced by higher bit-rates.

Correlated errors e.g. troposphere are becoming increasingly important.

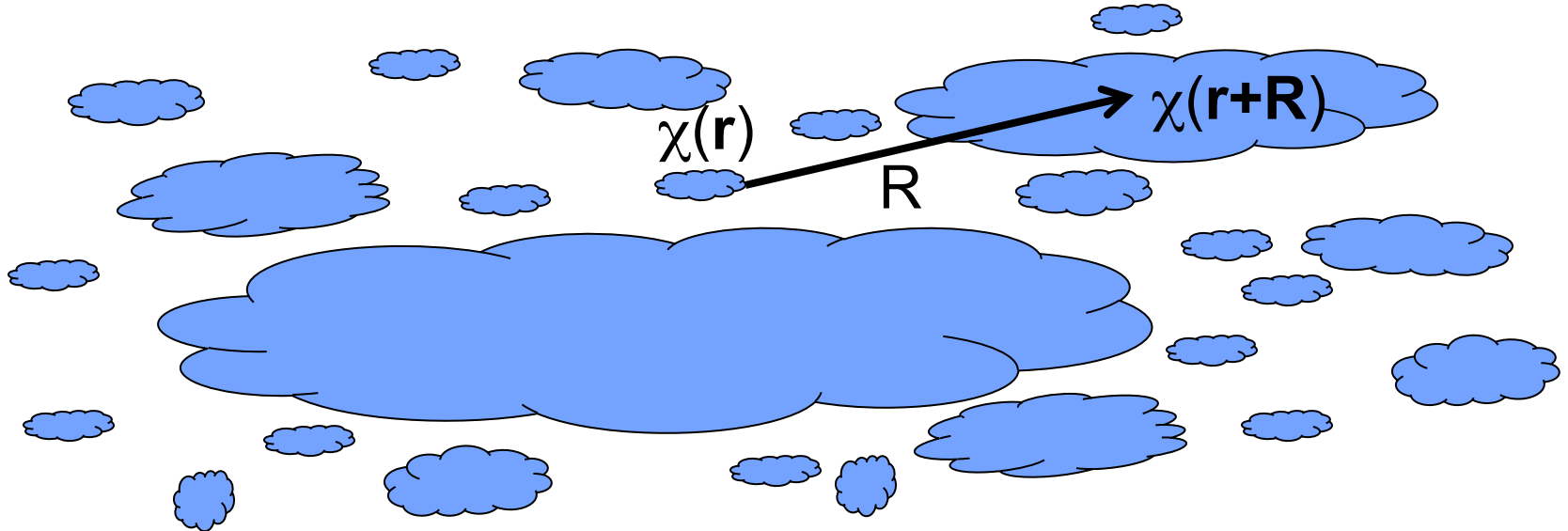
This study shows that troposphere covariance error models improve data analysis results.

Outlook:

We expect to see stronger effects with higher data rates.



BACK UP SLIDES



Structure function of refractivity (Tatarskii, 1961)

$$D_{\chi} = \langle (\chi(\mathbf{r}+\mathbf{R}) - \chi(\mathbf{r}))^2 \rangle$$

$$D_{\chi} = 2\sigma_{\chi}^2 - 2\text{cov}(\chi(\mathbf{r}+\mathbf{R}), \chi(\mathbf{r}))$$

$$D_{\chi} = C^2 R^{2/3} \text{ (Kolmogorov)}$$



The Radio Window of the Atmosphere

