

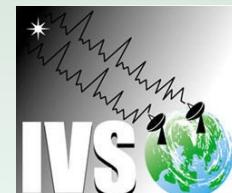


EOP determination from observations of Russian VLBI-network «QUASAR»

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Institute of Applied Astronomy RAS

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«QUASAR» NETWORK

2015×4282×4405 км



Центр корреляционной
обработки,
Санкт-Петербург



Обсерватория «Светлое»,
Ленинградская область



Обсерватория «Зеленчукская»,
Карачаево-Черкесская Республика



Обсерватория «Бадары»,
Республика Бурятия

“There is no national science, as there is no national multiplication table.”
(Anton Chekhov)

This work is of very importance for:

- GLONASS support
- contribution to station positions appointment
- contribution to sources posinions monitoring for ICRF
- contribution to IVS EOP time series for densification

“QUASAR” COMPLEX COMPONENTS:

- Three observatories:
«Svetloe», «Zelenuchukskaya», «Badary»
- IAA correlator center (schedule, correlation data treatment)
- IAA technical Consulate – obs. plan for year, month
- IAA technology center
- IAA AC

STATIONS EQUIPMENT



Svetloe
operable :1999,
IVS obs:2003,
QUASARobs:2006

- 32 m radio telescope, equipped with low noise receivers,
- frequency and time keeping system with H-masers (VCH-1003M),
- local geodetic network,
- control computers, local computer
- recording terminals Mark5B+,
- DAS R1002M (digital back end)



Zelenchukskaya
operable :2001,
IVS obs: 2005 (end),
QUASAR-obs: 2006



Badary
operable :2005,
IVS obs:2007,
QUASAR-obs : 2006

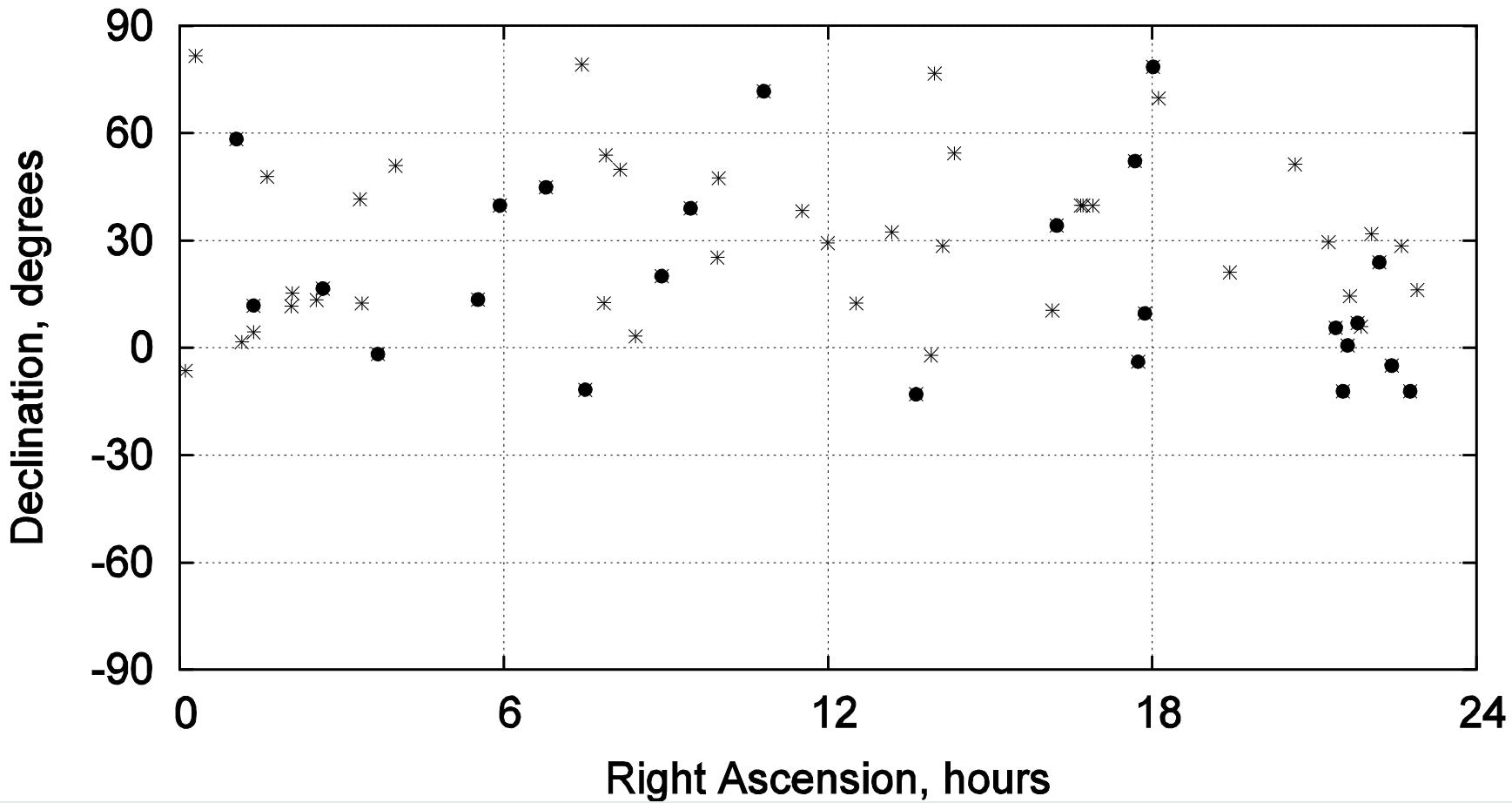
DOMESTIC VLBI PROGRAMS

	Ru-U	Ru-E
stations	Bd Zc(Sv)	BdZcSv
Duration, hours	1	24
aim	dUT1	EOP (Xp, Yp, dUT1, Xc, Yc)
Turn-around time	2 hours	2-5 days
Data transfer	e-vlbi	shipping
Frequency, start time	weekly, Friday 20 UT	weekly, Friday 22UT
scheduling	SKED (IAA), X/S, 1min 159 sources, > 0.25 J	SKED (IAA), X/S, 1min 63 sources, > 0.5 J
bandwidth	8 MGz	16 MHz
Data sampling	1 bit	1 bit
Data rate	256 Mb/s	512 Mb/c
Number of scans	20	330-350
Number of obs.	20	~1000

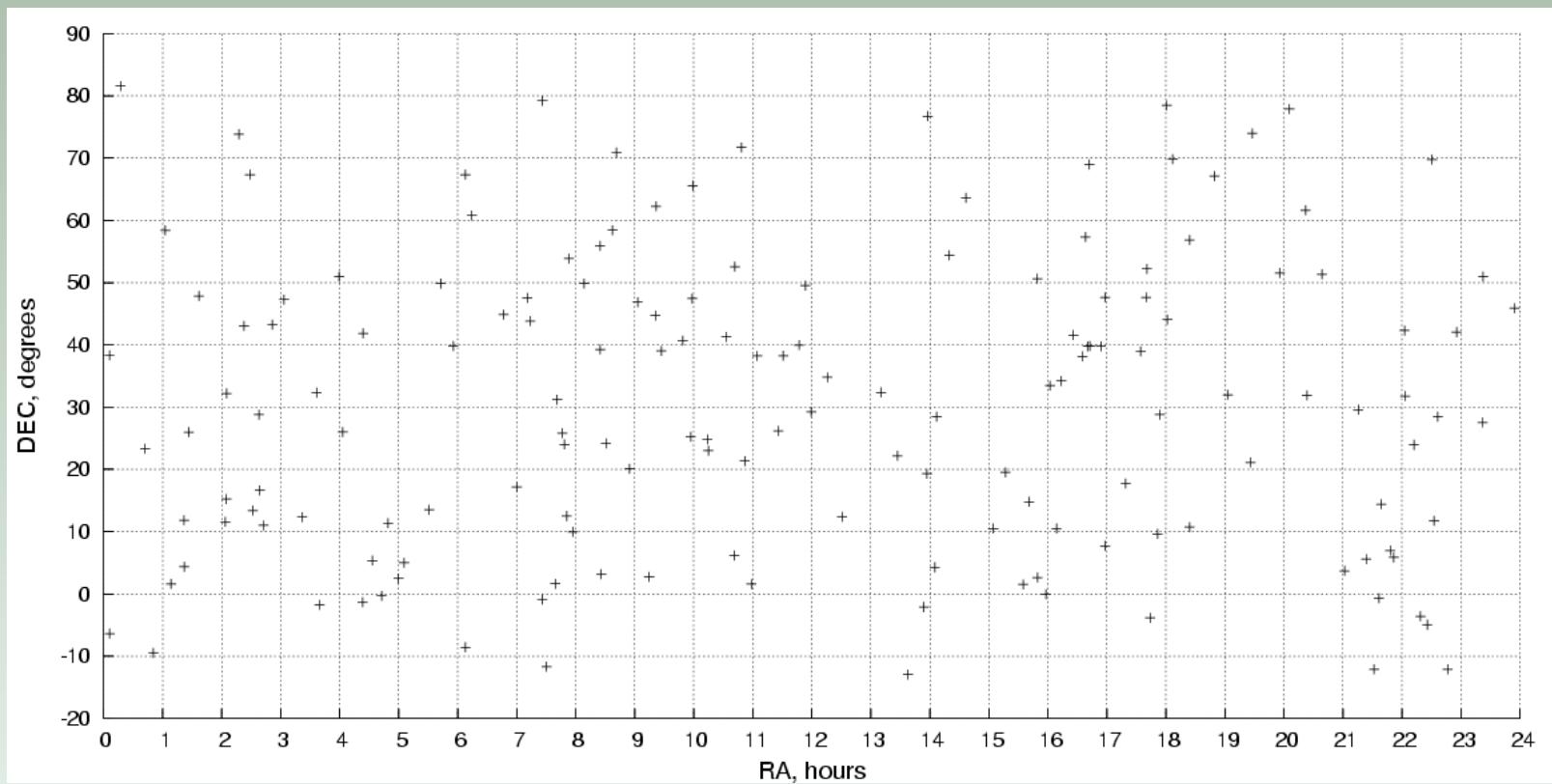
DOMESTIC SESSIONS

	Ru-U			Ru-E		
	Sv	Zc	Bd	Sv	Zc	Bd
2006		6	6	9	9	9
2007	10	12	17	9	9	9
2008	18	15	18	14	14	14
2009	13	26	30	23	23	23
2010	3	50	50	20	20	20
2011	5	65	65	49	49	49
2012	51(1)	51(8)	51(7)	51(5)	51(5)	51(5)

new radio sources set Ru-E (63 sources, > 0.5 J)



new radio sources set Ru-U (159 sources, > 0.25 J)



The main steps at the development Russian Domestic VLBI observations:

- 2005 first VLBI experiments on X-band, ionosferic delay account - from TEC.
- Since August 2006 - start regular QUASAR network observations, biweekly with S2 registrator, MicroPARSEC 3-station correlator
- 2009 - Mark5b since February, weekly observations - first attempts were made, ARC correlator Mark5B
- April 2009: first e-vlbi experiment, since September 2009 - e-vlbi data transfer on regular base for 1-hour Ru-U sessions. Weekly Ru-E and Ru-U sessions on regular base since July 2010 (after Svetloe repair works).
- New DAS - first experiment (May 2011), IVS-R4 sess.
- Oct.2011 DAS P1002M – Sv, Bd, Zc
- Since Nov. 2011 – observations with “RadioAstron”

IAA correlator ARC

- 2006-2008
3-station IAA correlator
MicroPARSEC , S2
- 2009
2 st. Correlator IAA
ARC, Mark5B
- 2010: 6-station
correlator IAA ARC

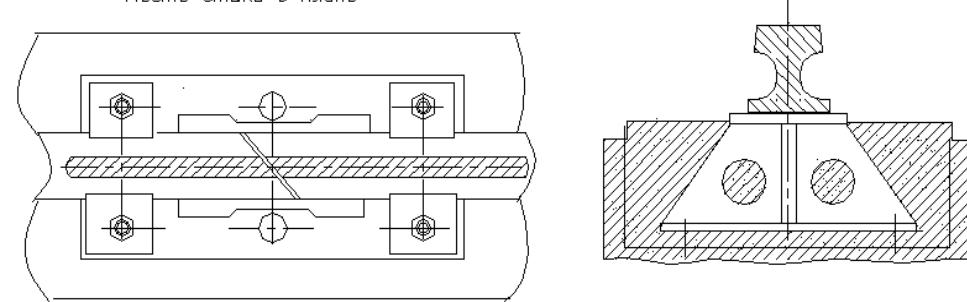
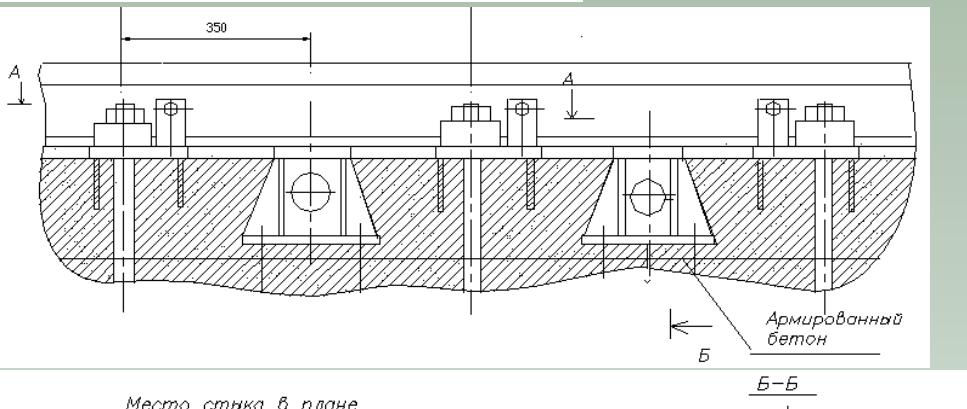
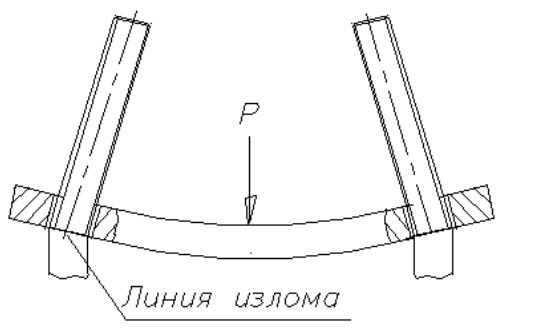


Svetloe: 2006: reconstruction



1. rail reconstruction

Sv: 2006, Zc, Bd – 2007



The antenna rail was reconstructed by adding a steel supporting construction and rebuilding the concrete under it.

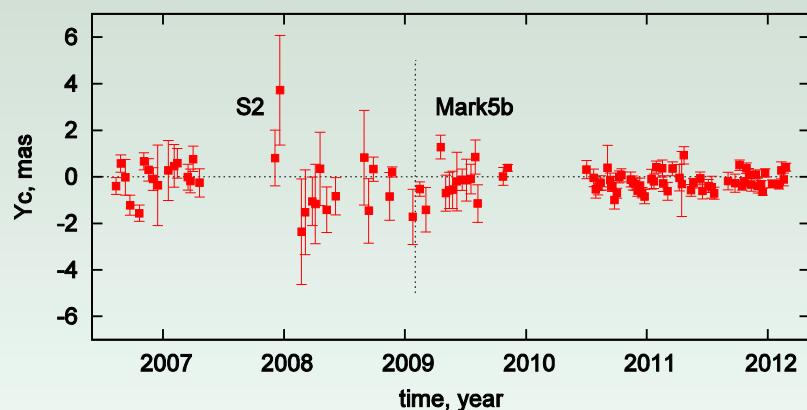
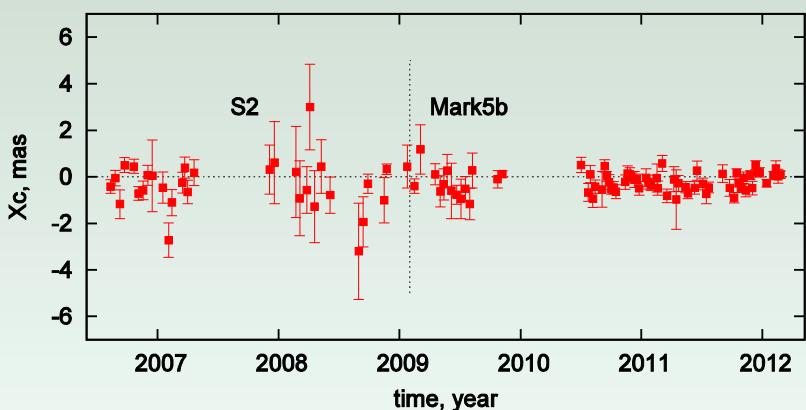
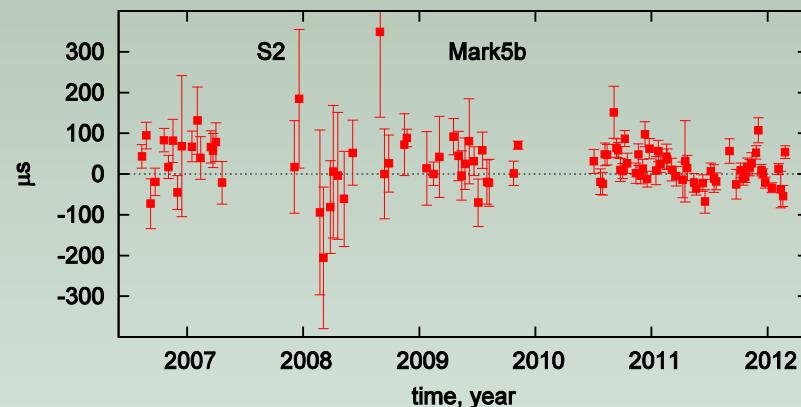
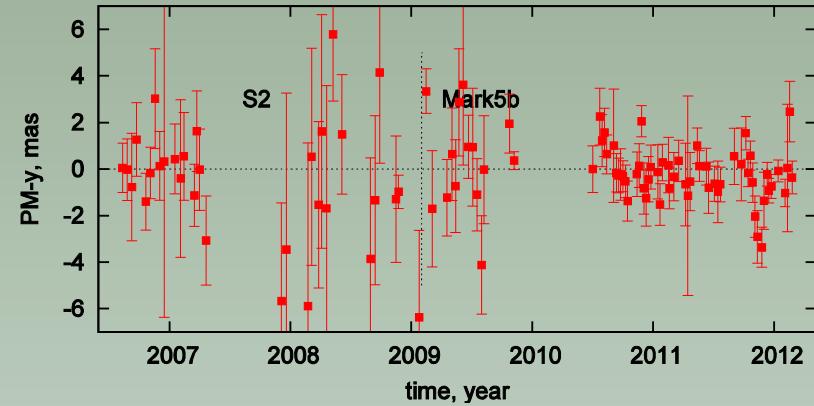
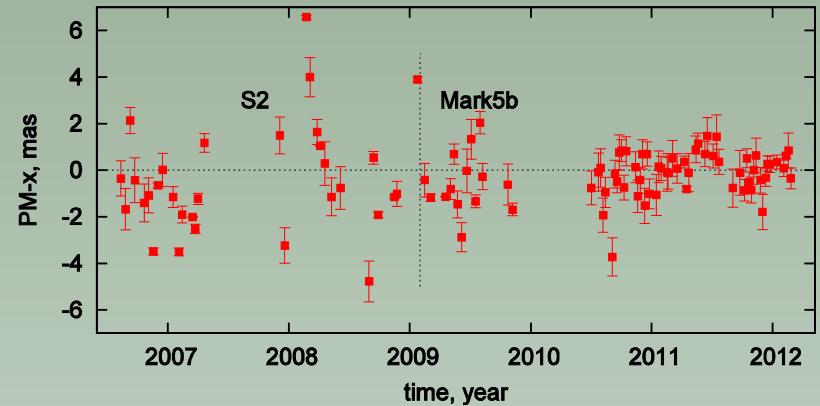


**Nov.2009 – Jun. 2010 – SVETLOE no operable
antenna repair works for steel construction
supporting,
summer 2010 – prophylactic works for Badary,
Zelenchukskaya**

VLBI data analysis

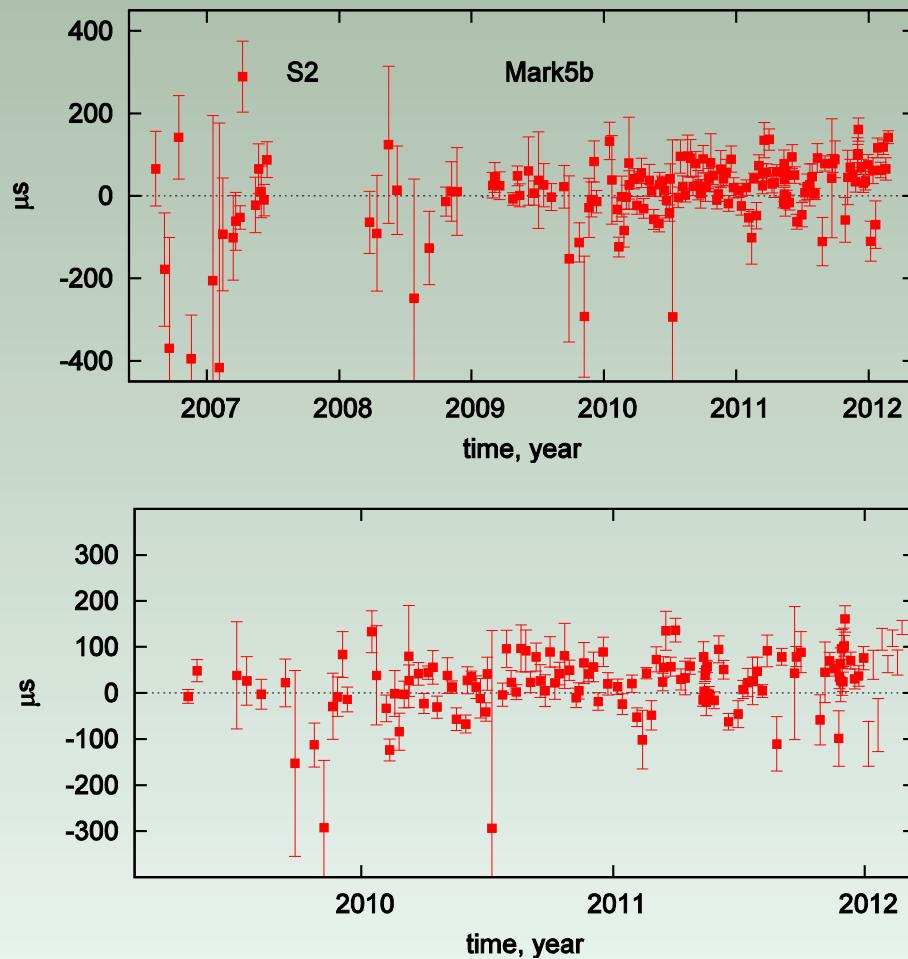
- QUASAR software
- OCCAM/GROSS (for results verification)
- Reduction – IERS Conventions
- catalogues: from QUASAR global solution
- standard parametrization
- tropospheric gradients are not estimated
- Ru-U sessions - in automatic mode

RuE-IERS 08 C04



Ru-U-IERS 08 C04

e-vlbi



TEST FOR EOP ACCURACY COMPARISON

For Ru-E, we extract observations for BdSvZc from IVS-R4 sessions

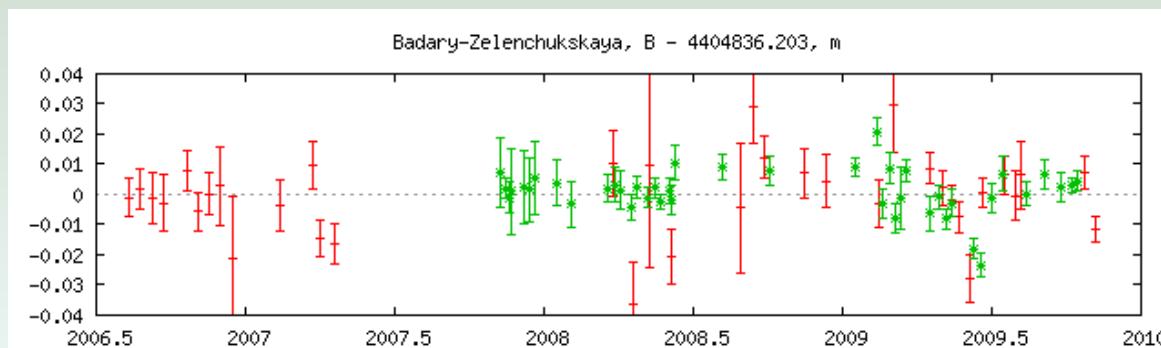
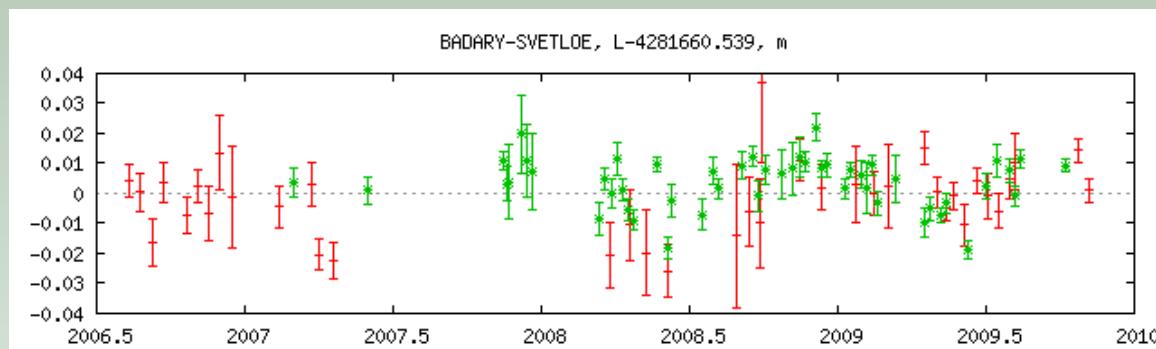
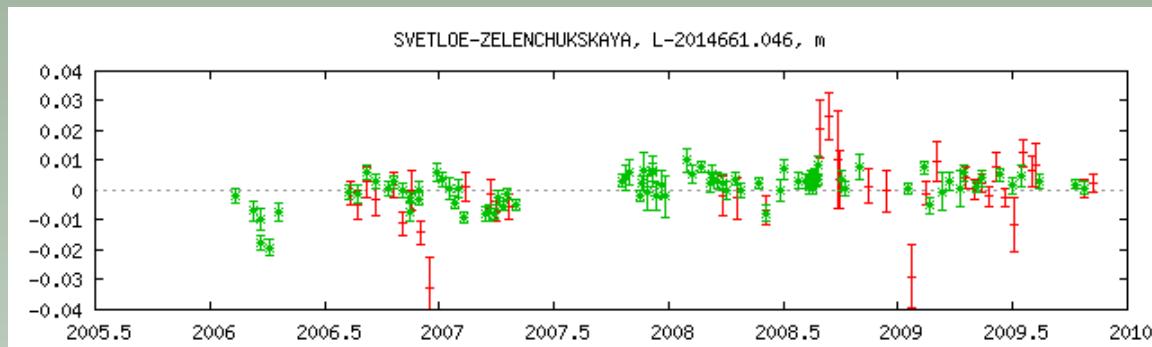
For Ru-U comparison we used IVS-Int2 sessions for 2011 WzTs

EOP time series accuracy

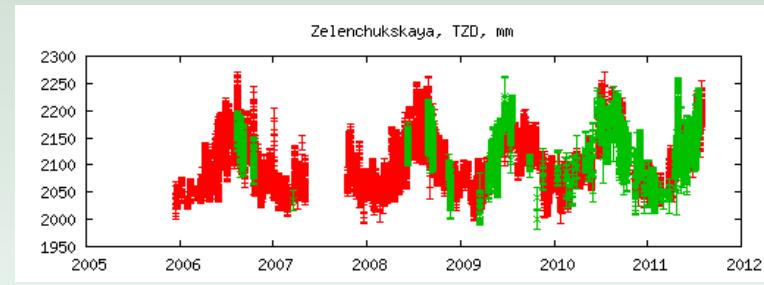
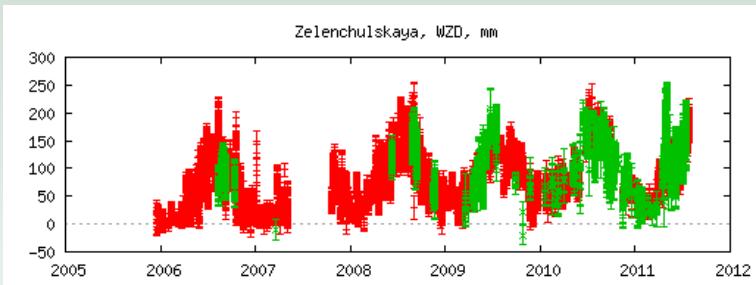
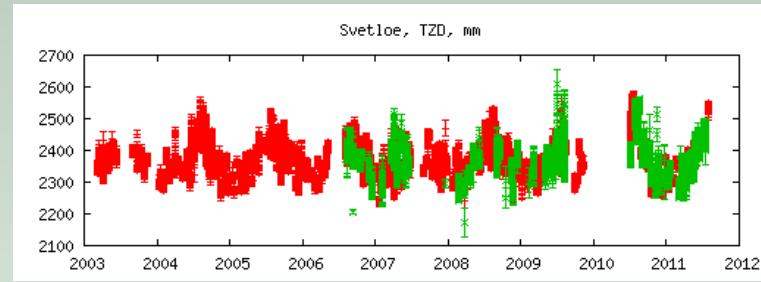
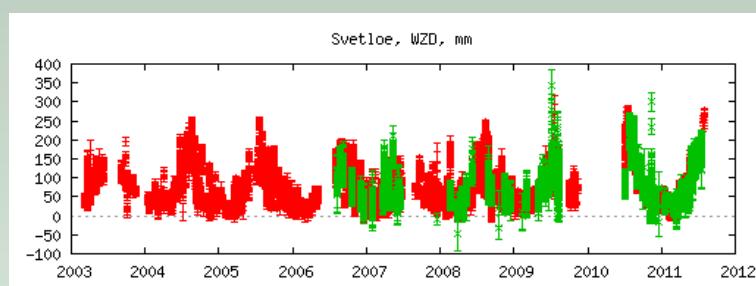
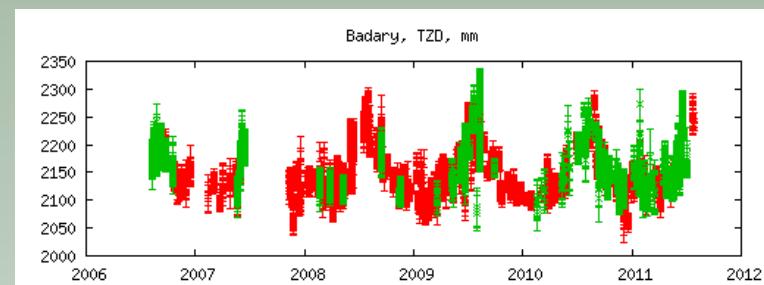
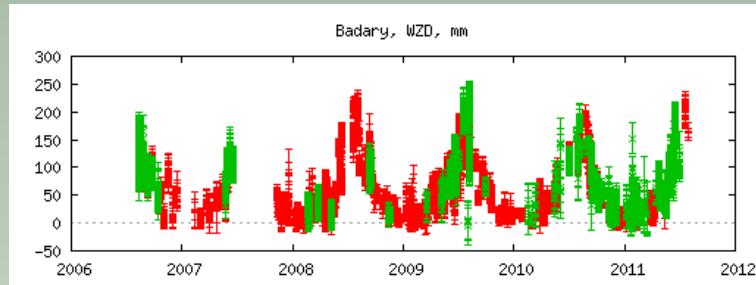
(rms (EOP(IAA) – IERS 08 C04))

	Domestic sessions		IVS-R4 (SvZcBd)	
	N sessions 2011.2- 2012.2	rms	N sessions 2007-2011	rms
dUT1, Int, μ s	53	59	125	37
Xp, mas	30	0.72	34	0.73
Yp, mas	30	1.13	34	1.13
dUt, μ s	30	35	34	37
Xc, mas	30	0.41	34	0.29
Yc, mas	30	0.39	34	0.34

TIME SERIES OF BASELINES LENGTH



TROPOSPHERIC PARAMETERS WZD, TZD



CONCLUSION REMARKS

- Problem: ~ only 70% observations correlated (? – SNR filtering, source set, other reasons – under investigation) – possible reserve for EOP accuracy improving
- Further improvements:
 - software correlator
 - WVR
 - 2010 VLBI antennas (Zelenchukskaya, Badary)
- Colocation with GNSS, SLR (see poster)

THANK YOU!