



Challenges and perspectives for CRF and TRF determination

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• ITRF2008 from VLBI/GNSS/SLR/DORIS normal equations



- ITRF2008 from VLBI/GNSS/SLR/DORIS normal equations
- ICRF2 from global VLBI solution



- This situation causes a complicated mutual impact of ITRF and ICRF which should be carefully investigated
 - to improve the accuracy of both, ICRF and ITRF,
 - and the consistency between them.



- Many groups working on that topic
 - DFG Forschergruppe on Reference Systems
 - IAG SC 1.4
 - see other presentations



Interaction of Celestial and Terrestrial Reference Frames

- IAG SC 1.4 (2011-2015) (Chair: J. Böhm)
- WG 1: Geophysical and Astronomical Effects and the Consistent Determination of CRF and TRF (Chair: Z. Malkin)
- WG 2: Co-location on Earth and in Space for the Determination of the CRF (Chair: S. Lambert)
- WG 3: Maintenance of CRF and the link to the new GAIA frame (Chair: C. Ma)

- Insufficient models
 - Tropospheric delays, in particular gradients



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Constraints on Gradients

- 6 h piecewise linear offsets as gradients
- 0.5 mm relative and 1 mm absolute constraints
- Zero vs. DAO a priori gradients



mean bias in declination $\approx 40 \ \mu as$

Constraints on Gradients



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- Insufficient models
- Terrestrial reference frame
 - Poor geometry in the southern hemisphere
 - Modeling of non-linear station motions
 - Dependence on ITRF datum
 - Choice of datum stations for NNR/NNT

Selection of Datum Stations

• 6 instead of 22 datum stations



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- Insufficient models
- Terrestrial reference frame
- Celestial reference frame
 - Uneven distribution of sources over sky



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 - Uneven distribution of sources over sky
 - Impact on Earth rotation (presentation by Nilsson)

- Insufficient models
- Terrestrial reference frame
- Celestial reference frame
- Astronomical effects
 - Apparent source motions
 - Source structure effects / wavelength dependency
 - Galactic rotation

• CRF velocity (correlated with galactic rotation)



- ITRF2008 is determined by a VLBI/GNSS/SLR/DORIS combination (intra-, then inter-technique)
- ICRF2 is VLBI-only solution (once AC)
- Next ICRF should be estimated together with ITRF
 - Connection via local ties
 - see presentation by M. Seitz

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 - Connection via local ties
 - see presentation by M. Seitz
- more possibilities
 - Troposphere ties

- Space ties
 - Observing GNSS satellites with VLBI (Tornatore and Haas)



- Space ties
 - Observing GNSS satellites with VLBI
 - Adding differential observations to quasars

440000;

• mitigates atmospheric influences



- Space ties
 - Observing GNSS satellites with VLBI
 - Adding differential observations to quasars
 - Or even mixing both with twin telescopes



- Effect on slant wet delay at 30° elevation
 - when alternately observing source and satellite every 15 seconds (turbulent atmosphere)



- Space ties
 - Observing GNSS satellites with VLBI
 - Adding differential observations to quasars
 - Or even mixing both with twin telescopes
 - GRASP like satellites

- GRASP (Geodetic Reference Antenna in Space)
- E.g., h = 1350 km, i = 99.9°, e = 0.0334



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Number of observing sites



Maintenance of CRF and the link to the new GAIA frame

- GAIA mission scheduled for launch in 2013
- Optical realization of the CRF with similar precision
- At least an order of magnitude more objects
- For geodetic use the CRF realization must be accessible from the ground
- Search for optical/radio counterparts

Conclusions and Outlook

- A combined ITRF/ICRF/EOP solution from VLBI, GNSS, SLR, and DORIS should be striven for.
- Additionally, systematic errors need to be mitigated, not only by improved models but also by better observing strategies.
- VLBI2010 will open new possibilities.
- Space ties will greatly enhance the consistency between TRF and CRF.

Thanks for your attention!







FШF

D-VLBI Integrated VLBI