

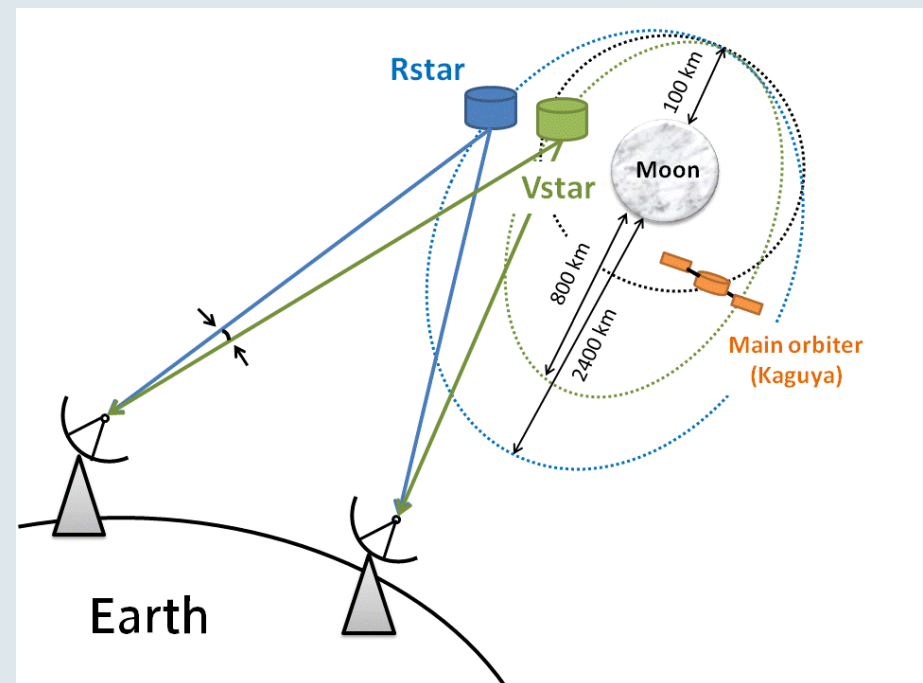
Session 3

Processing SELENE Differential VLBI Data

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J. Böhm, M. Madzak, C. Tierno Ros, H. Schuh

- SELenological and ENgineering Explorer
 - JAXA, 2007-2009
 - 3 lunar satellites
- same-beam differential VLBI (D-VLBI)
 - improved orbit consistency from several hundreds to several tens of metres [Goossens et al., 2010]
- nominal accuracy
 (differential phase delay rms):
 - 3.44 ps (1 mm) S-band [Kikuchi et al., 2009]



$$\Delta\tau = \tau_{\text{RSTAR}} - \tau_{\text{VSTAR}}$$

- 2 datasets (from NAOJ Mizusawa):
 - October 2008, 17-22: 4 japanese antennas
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- Implementation into VieVS
- Challenges:
 - *Moving sources at finite distance*
 - Relativistic time and coordinate transformations (TDB, TT, TCB, TCG resp. BDRS, TRS, BCRS, GCRS)
 - Consistent application of corresponding scaling factors (L_c , L_g , L_b)
- Good agreement of 3 models (< 1 ps resp. < 3 ps)
 - Sekido & Fukushima 2006
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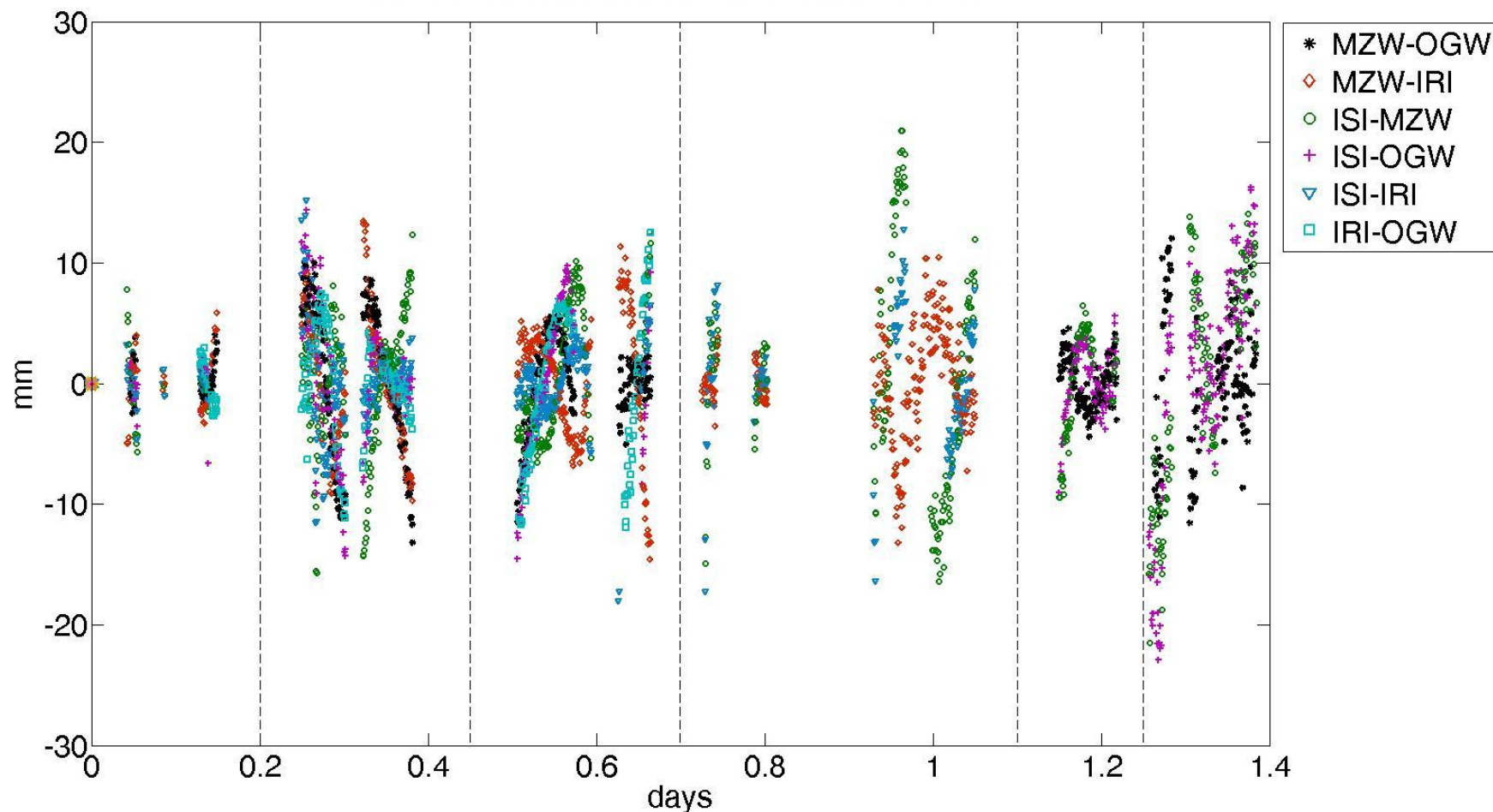
➡ **Differences correspond to the theoretical values**

➡ **Verification through comparison of 3 models**

10 mm $\leftrightarrow$ 33 ps

residuals o-c

OCT08 residuals obs-com after baseline offset

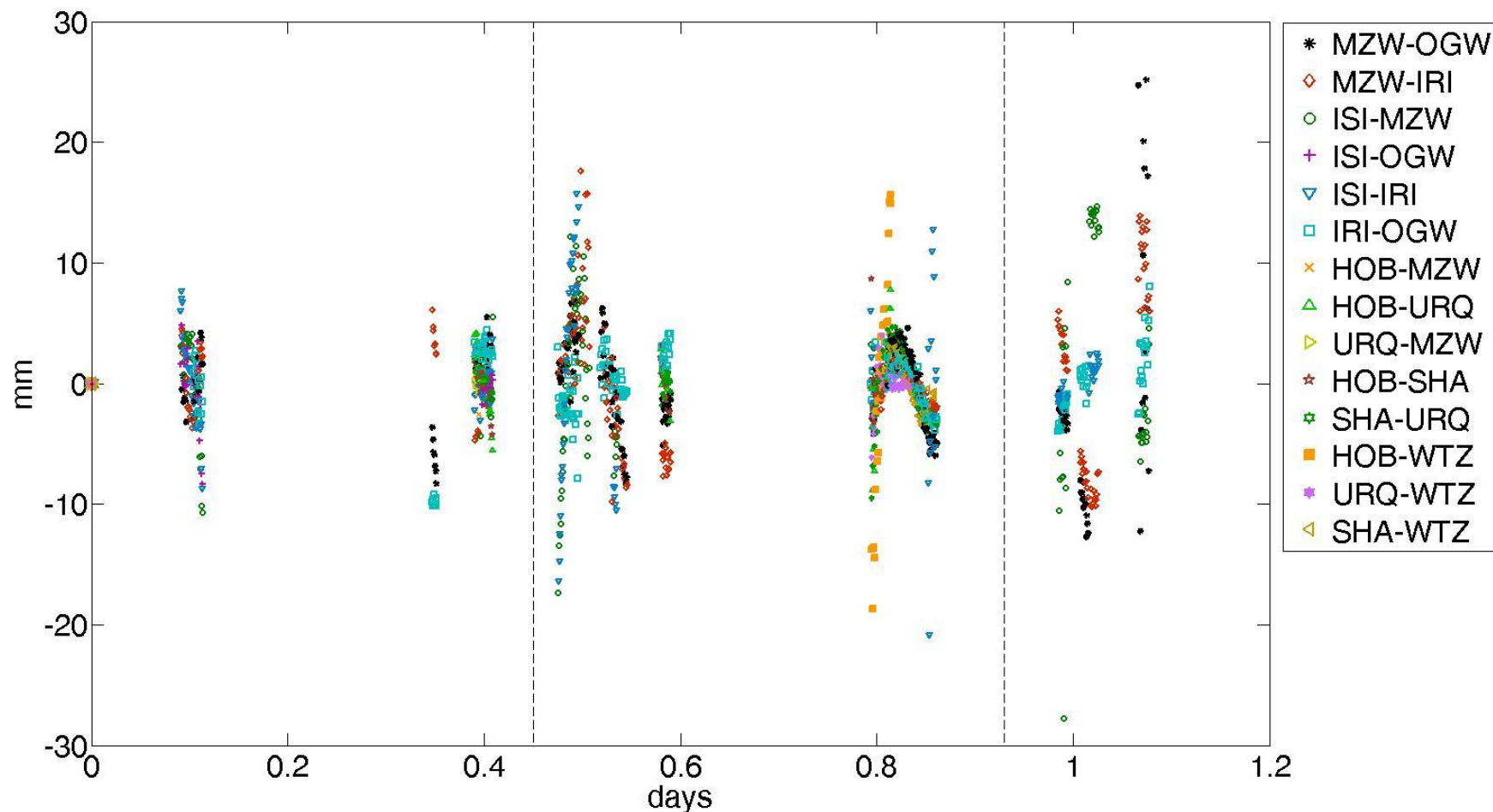


$$\Delta\tau = \tau_{\text{RSTAR}} - \tau_{\text{VSTAR}}$$

10 mm $\leftrightarrow$ 33 ps

residuals o-c

JAN08 residuals obs-com after baseline offset



$$\Delta\tau = \tau_{\text{RSTAR}} - \tau_{\text{VSTAR}}$$

- Residuals are at the same level as those from NAOJ
 ± 10 mm (30 ps)
- NAOJ explains residuals with
*„probably mismodelling of solar radiation
pressure“*

(long baselines)

τ

$\Delta\tau$

GEOMETRY:

Antenna $\pm$ 5cm

300 ps

1-2 ps

Orbit $\pm$ 10 m

150 -1000 ps

2-8 ps

EOP

5 ps (60 ps)

< 0.05 (0.1) ps

dUT1: 5 ms

xp, yp: 200 mas

dX, dY: 300 mas

(long baselines)

τ

$\Delta\tau$

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usual elev > 20°
min. elev. 10° (5°)

ATMOSPHERE:

Trop_hydr

2-20 (10-60) ns

30-300 (50-1000) ps

a priori

Trop_wet

1-3 (4) ns

4-40 (10-60) ps

ECMWF

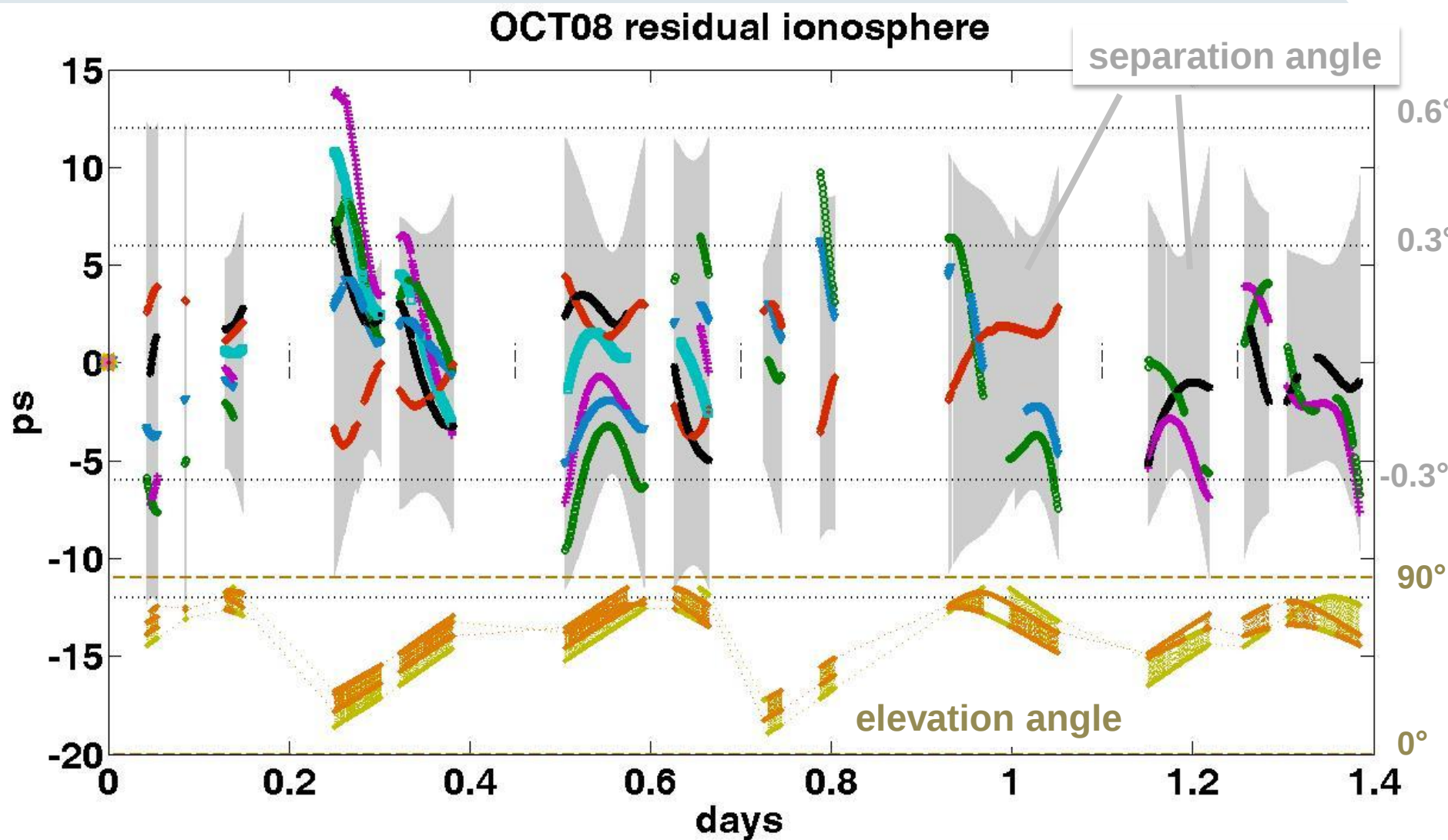
Ionosphere

1.5 (2-10) ns

10 (80) ps

TEC-maps

Peak at noon; nearly all observations during the night

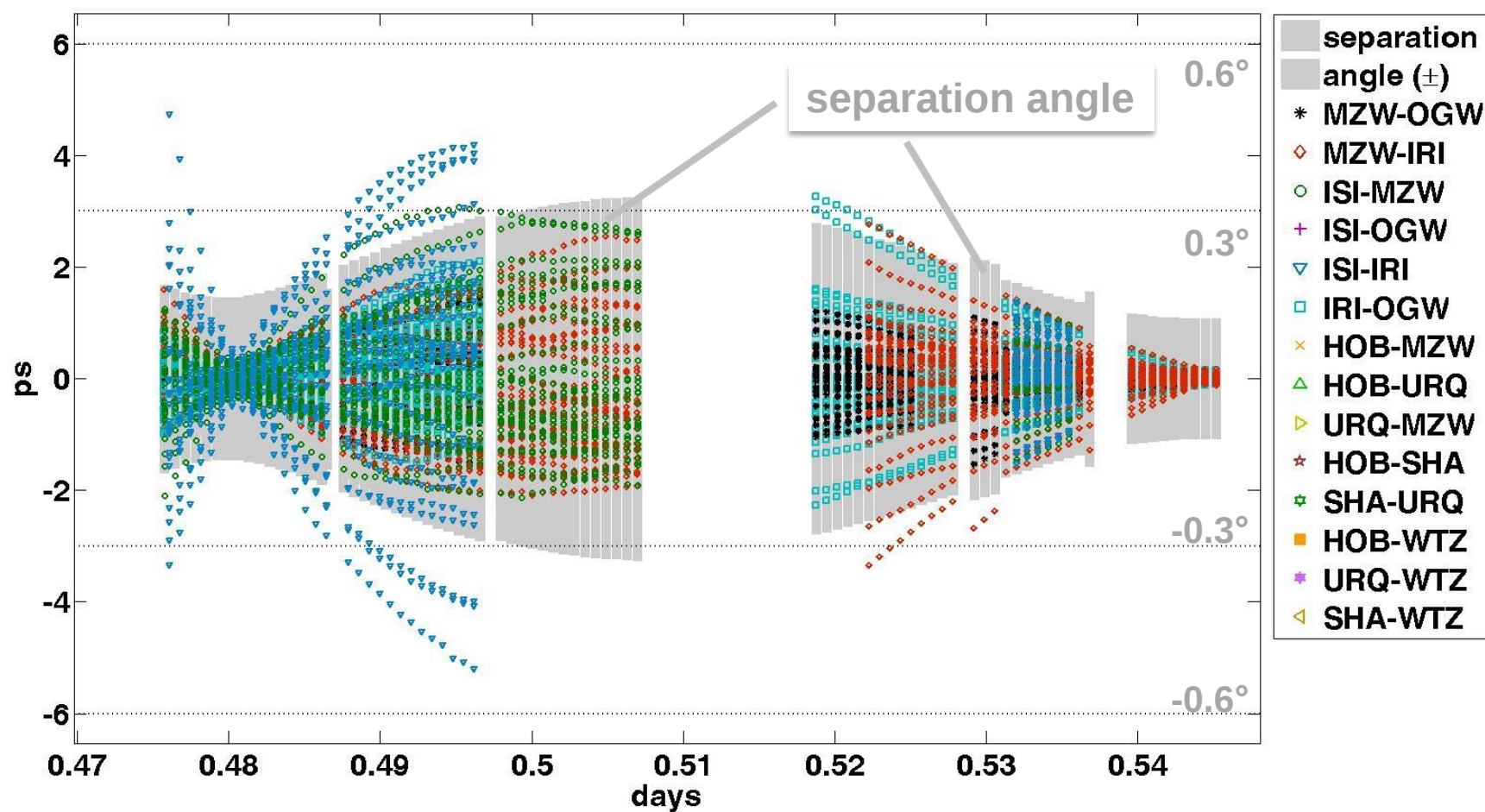




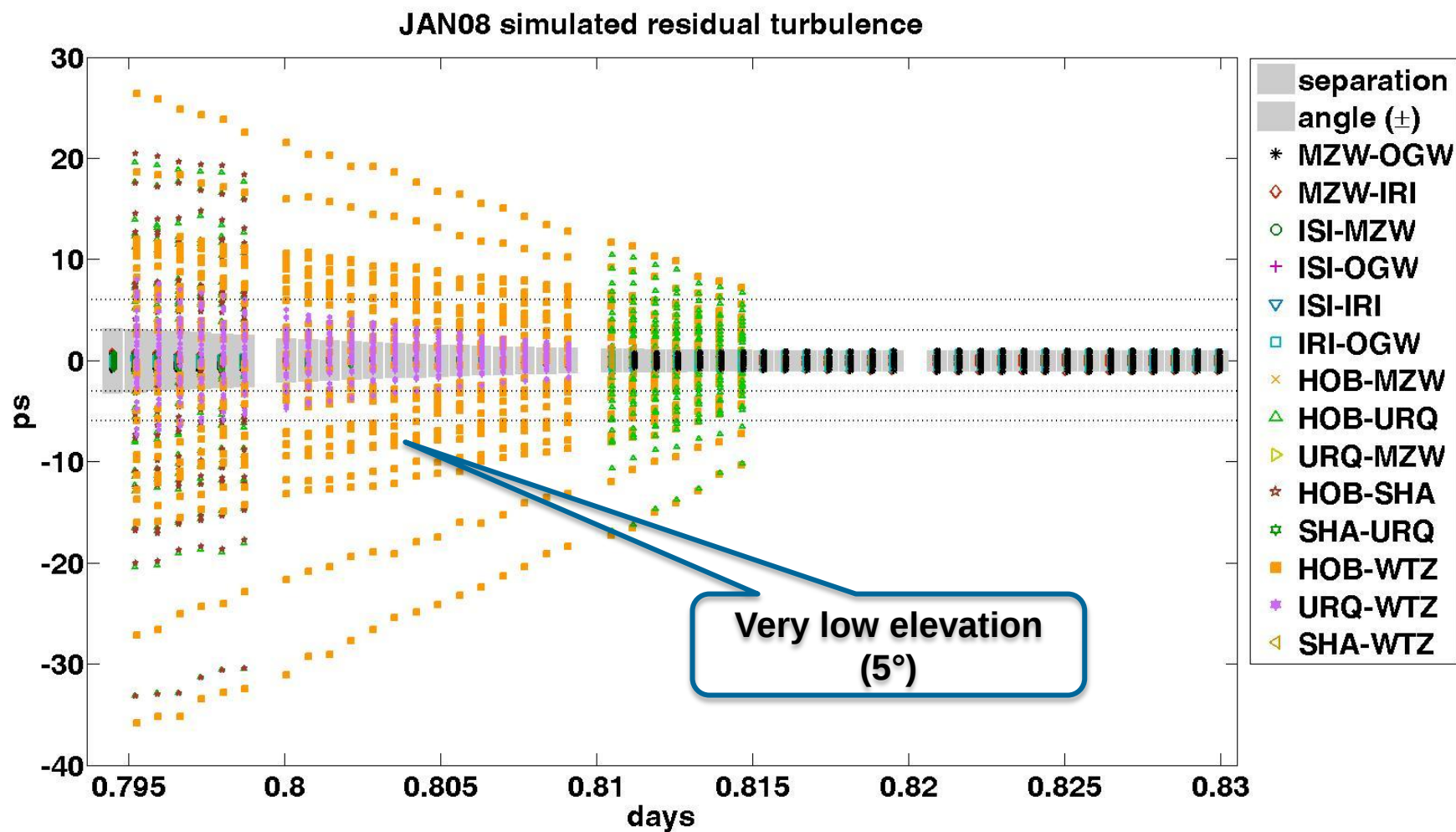
VieVS-simulator [Pany et al., 2010]; $N = 30$

➔ VieVS-simulator [Pany et al., 2010]; N = 30

JAN08 simulated residual turbulence



➔ VieVS-simulator [Pany et al., 2010]; N = 30



Results / Comparison (2)



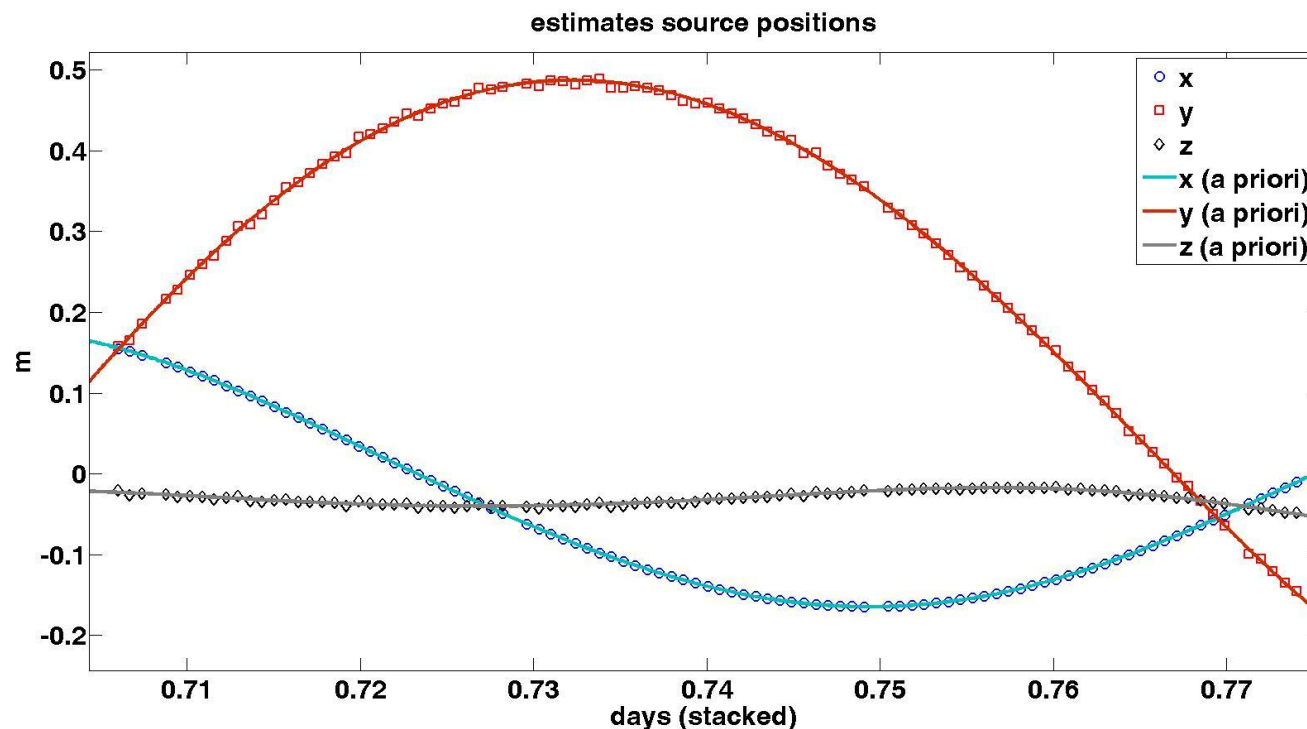
- Geometric effects (station coordinates, source coordinates, EOP) are nearly cancelled out.
- Residual tropospheric effects can reach the level of significance.
 - *Particularly for intercontinental baselines.*
- But: Applying wet troposphere and ionospheric corrections doesn't improve the result significantly.
 - *Some residuals decrease but some others increase.*

$$\begin{aligned}\Delta\tau &= \tau_{RSTAR} - \tau_{VSTAR} = \\ &= [(RSTAR - St1) - (RSTAR - St2)] - [(VSTAR - St1) - (VSTAR - St2)]\end{aligned}$$

$$\frac{\partial\Delta\tau}{\partial VSTAR} = -\frac{VSTAR - St1}{r_{VSTAR-St1}} + \frac{VSTAR - St2}{r_{VSTAR-St2}}$$

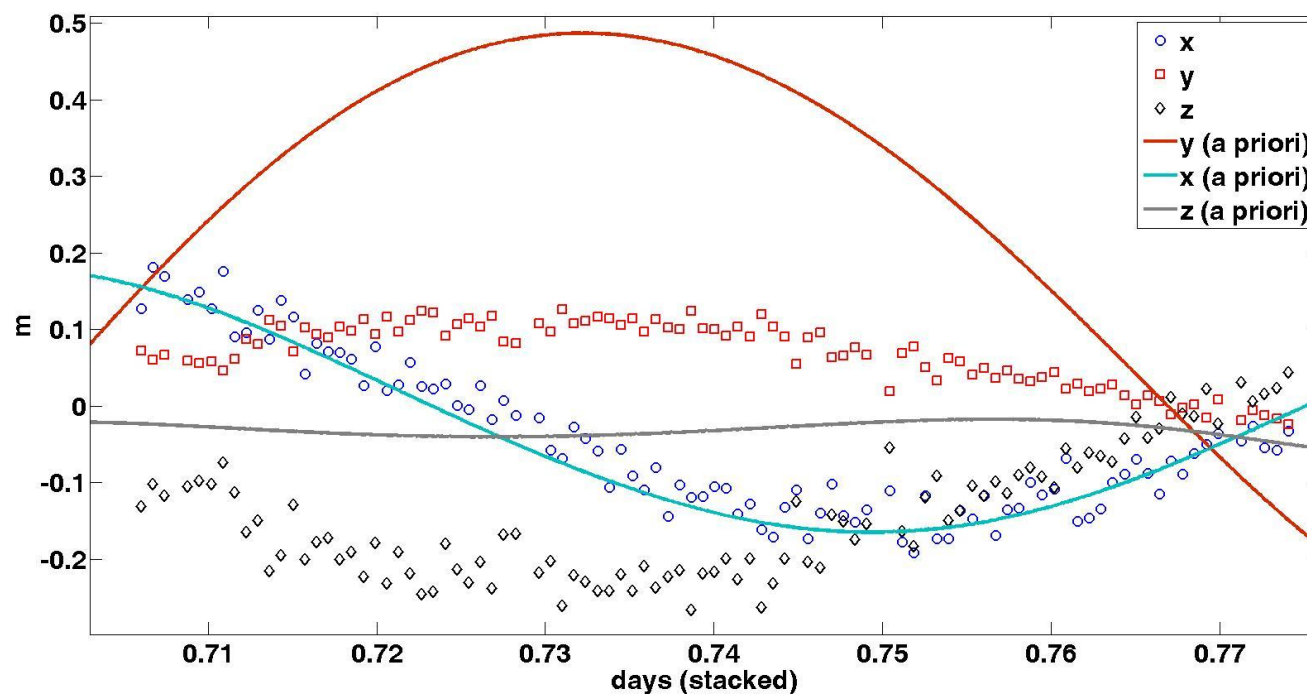
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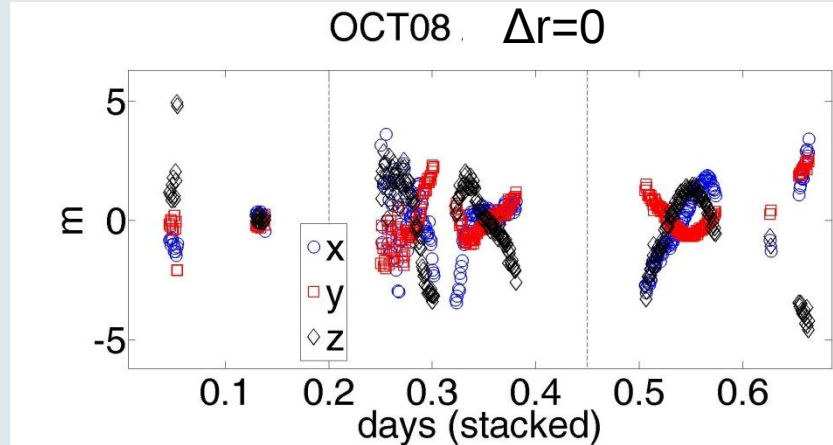
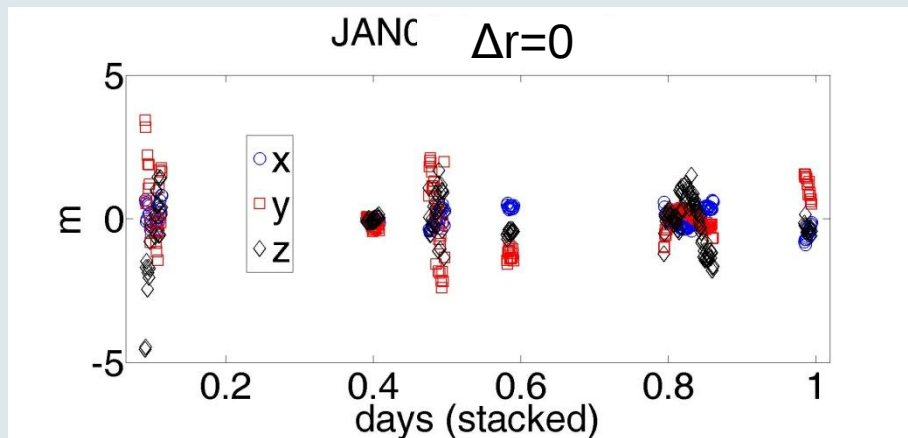
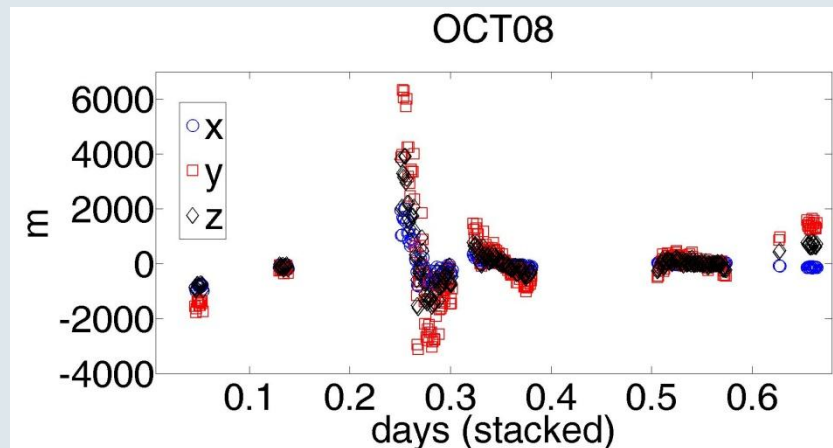
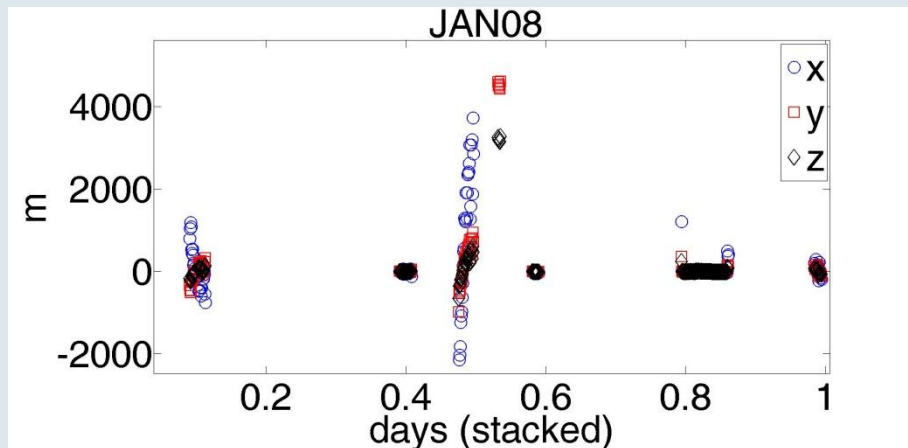
$$\frac{\partial\Delta\tau}{\partial VSTAR} = -\frac{VSTAR - St1}{r_{VSTAR-St1}} + \frac{VSTAR - St2}{r_{VSTAR-St2}}$$



➔ 1 ps noise: OCT08: ± 100 m
 JAN08: ± 50 m

➔ implement condition $\Delta r = 0$





- Model for orbit estimation (xyz) is correct, but highly correlated and unstable.
- *Possible solution: $\Delta r = 0$*
- NAOJ: *total orbit errors at an average level of 18m*
- Estimates: several km $\rightarrow$ some m
- Goal: find better strategies
 - *Line of sight*
 - *estimate orbit parameters*
- **Upcoming: Studying estimation strategy with VLBI observations to GNSS-satellites.**

THE END

Thank You for listening!

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References:

- Goossens et al., 2010: *Lunar gravity field determination using SELENE same-beam differential VLBI tracking data*, J. of Geodesy, DOI: 10.1007/s00190-010-0430-2.
- Kikuchi et al., 2009: *Pico-second accuracy VLBI of the two sub-satellites of SELENE (KAGUYA) using multi-frequency and same-beam methods*, Radio Sci. 44(RS2008).
- Klioner, 1991: *General relativistic model of VLBI observables*. In: Carter WE (eds) Proceedings of AGU Chapman conference on geodetic VLBI: monitoring of global change. NOAA Technical Report NOS 137 NGS 49, AGU, Washington DC, pp 188-202.
- Sekido & Fukushima, 2006: *A VLBI delay model for radio sources at a finite distance*, J. of Geodesy, 80:137-149, DOI 10.1007/s00190-006-0035-y.
- Pany et al., 2010: *Monte Carlo simulations of the impact of troposphere, clocks and measurement errors on the repeatability of VLBI positions*. J. of Geodesy, DOI: 10.1007/s00190-010-0415-1.