

# ON THE ESTIMATION OF A CELESTIAL REFERENCE FRAME IN THE PRESENCE OF SOURCE STRUCTURE

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# GEODESY MEETS ASTRONOMY

We know that,

- quasars are NOT ideal sources  
and that
- quasars change with time

**Source structure can corrupt  
our geodetic/astrometric  
VLBI measurements!**

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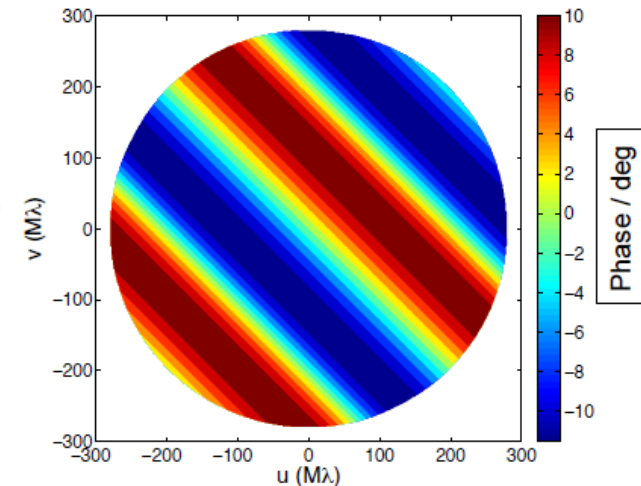
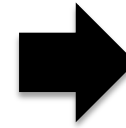
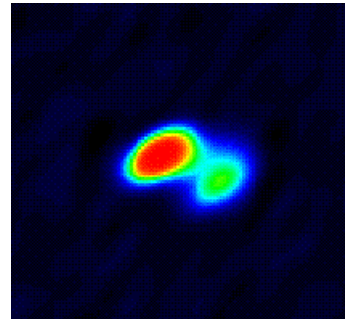
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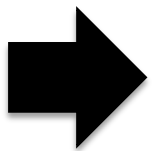
# SOURCE STRUCTURE IN GEODESY

Additional Phase in dependence of the

- frequency,
- length &
- orientation of the baseline



- Effect is different at each of the 8 sub-bands at X-band, resp. different at S-band



**group delay (= slope across band) changes due to structure**

# SOURCE STRUCTURE SIMULATIONS

- **VieVS structure simulator** (Shabala et al. JoG 2015)
  - Calculate delay due to source structure based on multi-component source models.
  - Apply in analysis and/or use it in simulations.

# THOROUGH INVESTIGATION

Plank et al. MNRAS 2015 (soon)

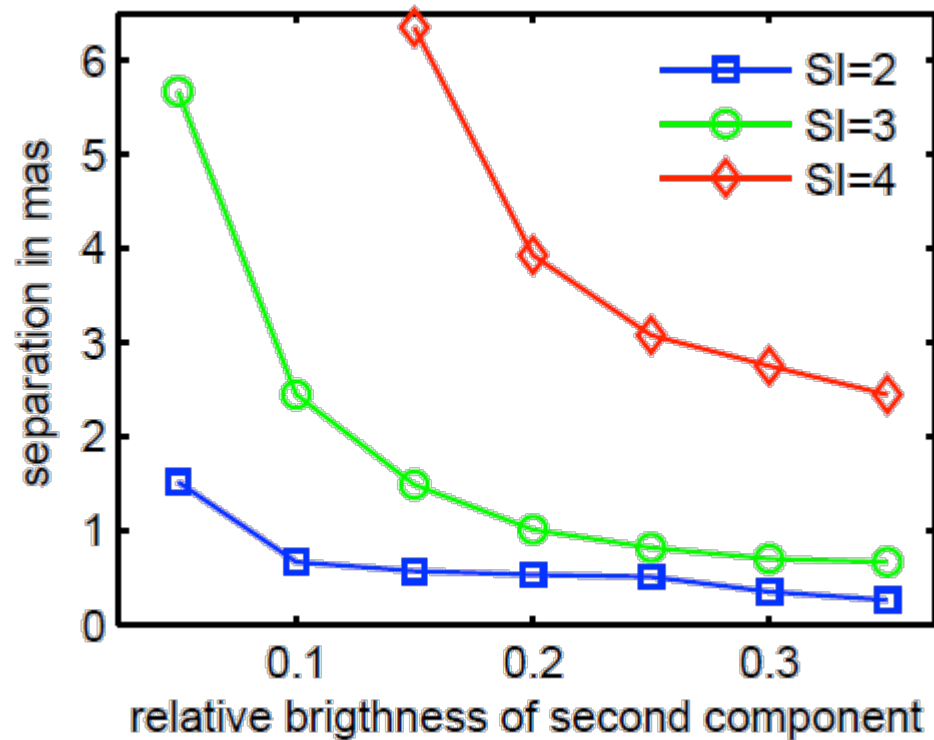
- Mock two-component sources
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- Mock two-component sources
- Apply to one year of R1/R4 schedules
  - Identical model for all sources, random jet directions
- Structure-only
- structure+troposphere+noise  
(VGOS Simulations)
- Globally estimated source positions  
(loose constraints)

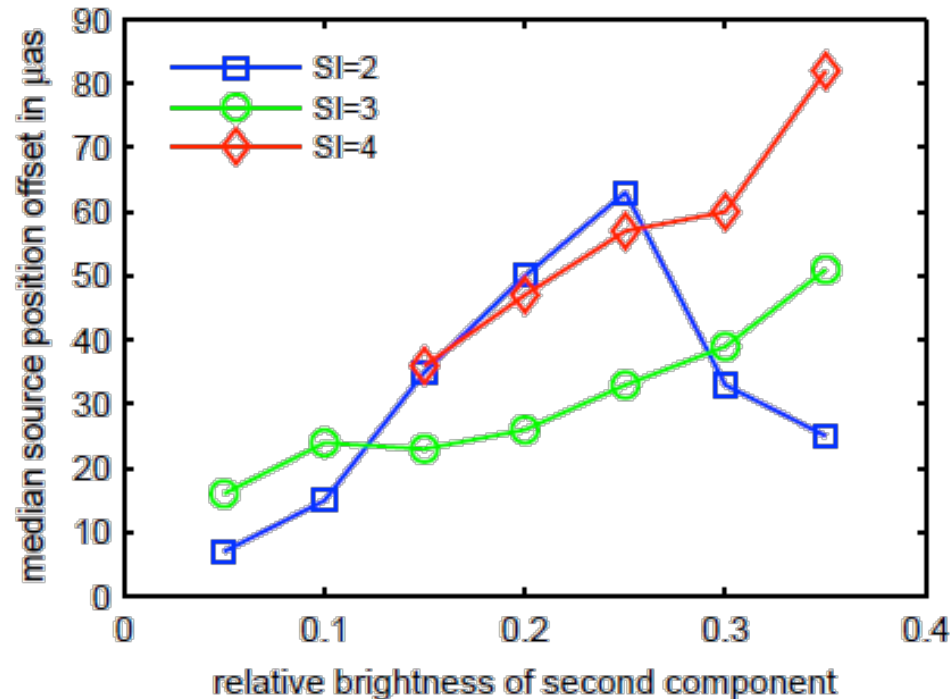
# DIFFERENT STRUCTURE MODELS



- Apply models of various structure indices (SI=2,3,4).



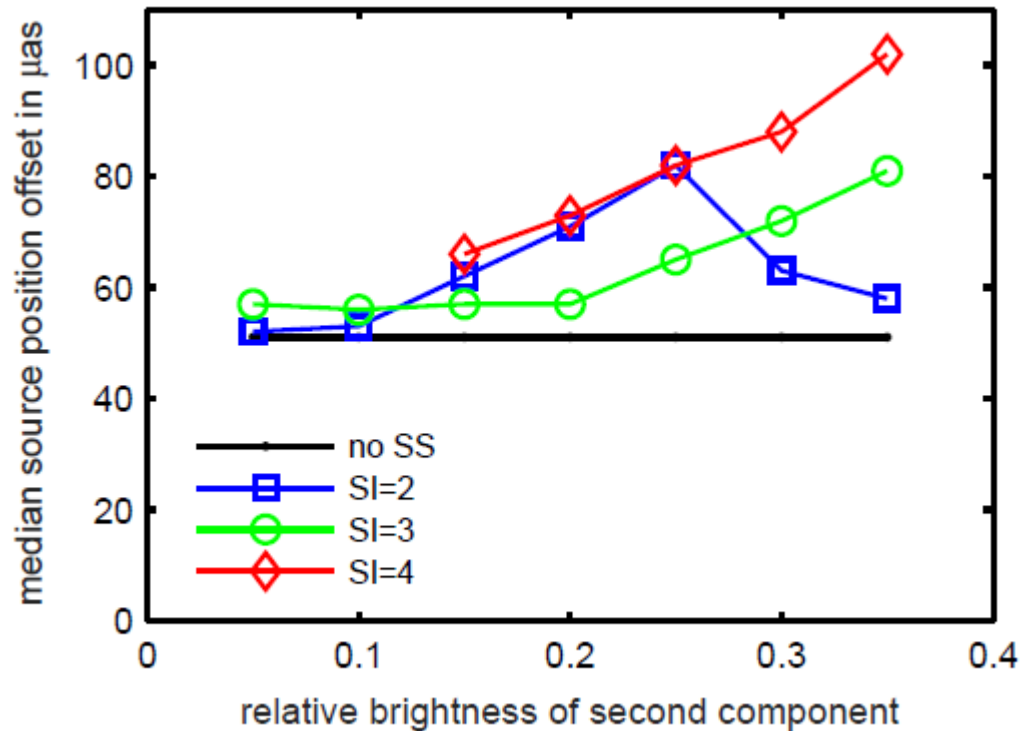
# MEDIAN SOURCE DISPLACEMENT



- Systematic displacements of several tens of  $\mu\text{as}$

Simulated median source position offsets  $d$  due to source structure using various two-component source models with nominal structure indices SI=2, 3, and 4 and a relative brightness of the second component between 0.05 and 0.35 of the main component. These are structure-only simulations.

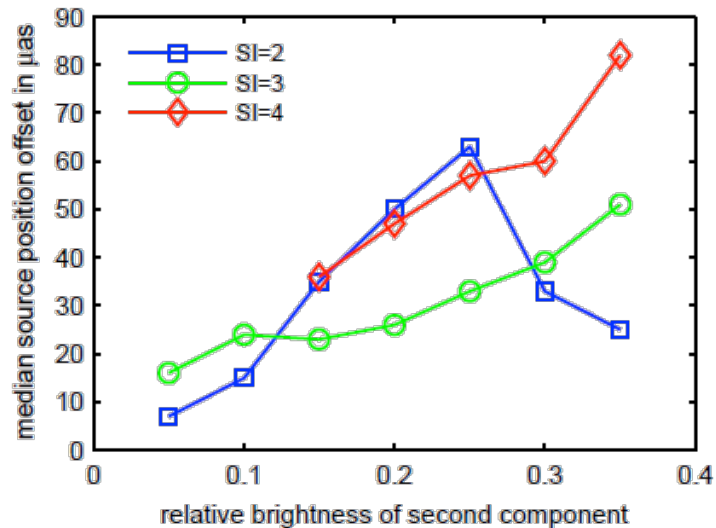
# EXCEEDS THE TROPOSPHERE



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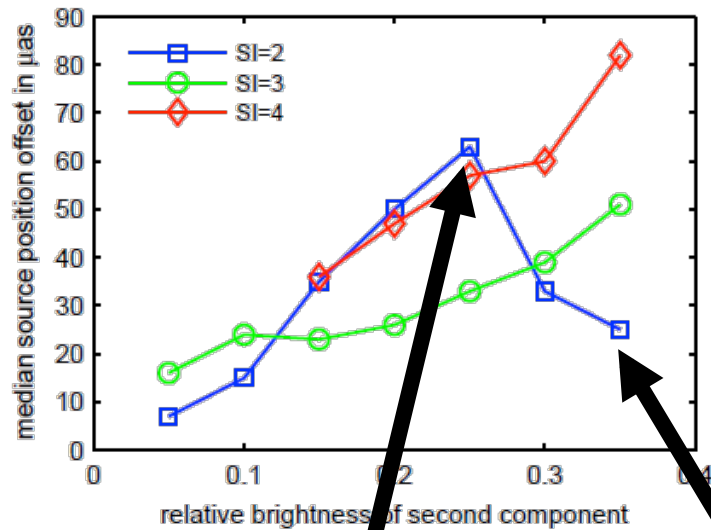
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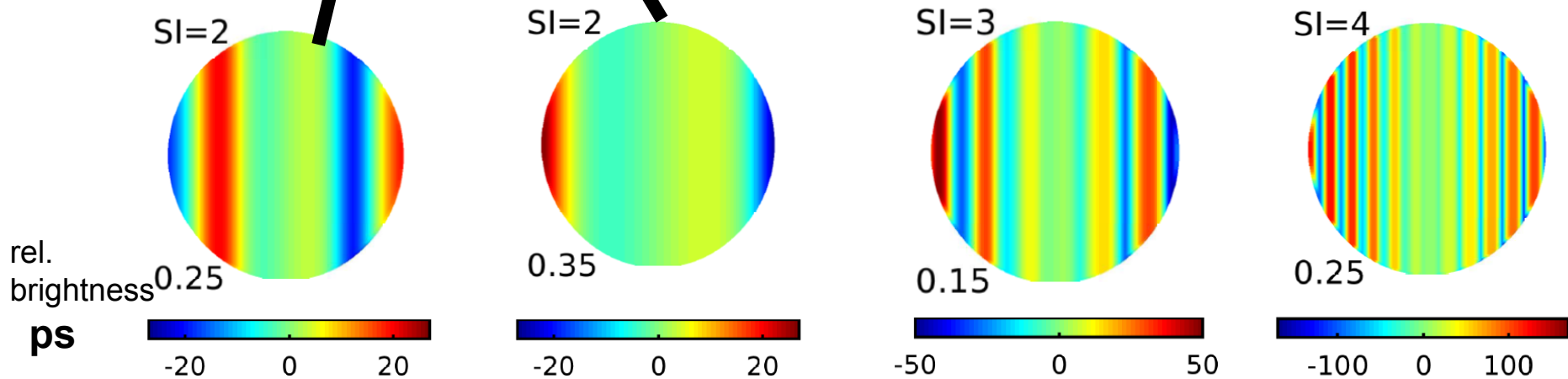


- SI=3, 4:  
Larger displacement for stronger secondary component close by.

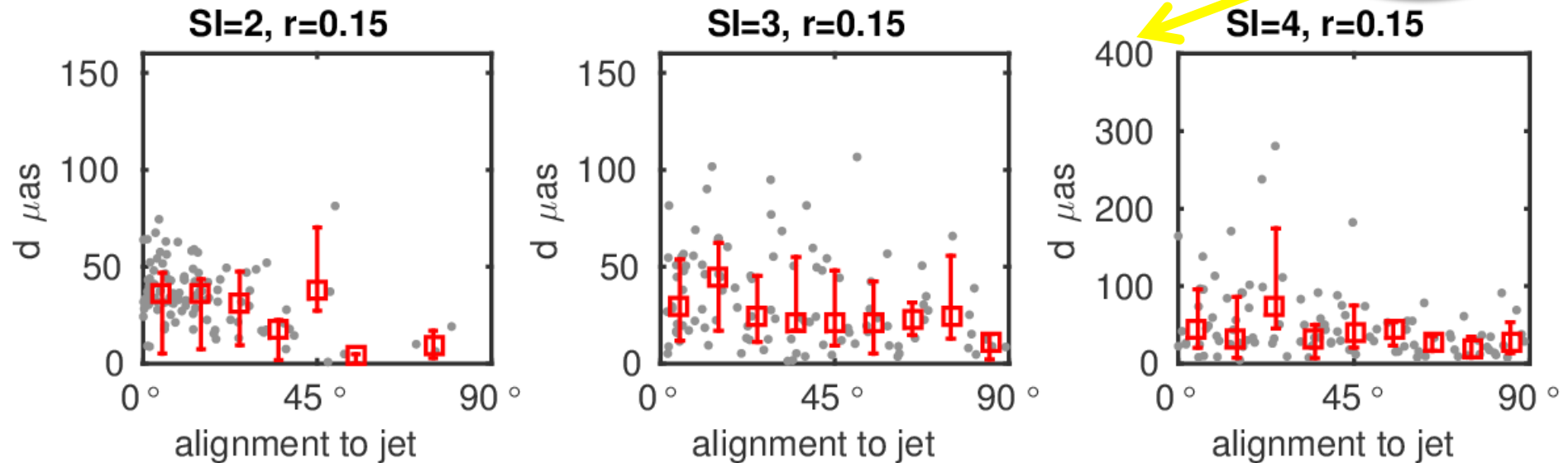
# DIFFERENT STRUCTURE MODELS



- SI=3, 4:  
Larger displacement for stronger secondary component close by.
- SI=2:  
smaller effect, but more systematic



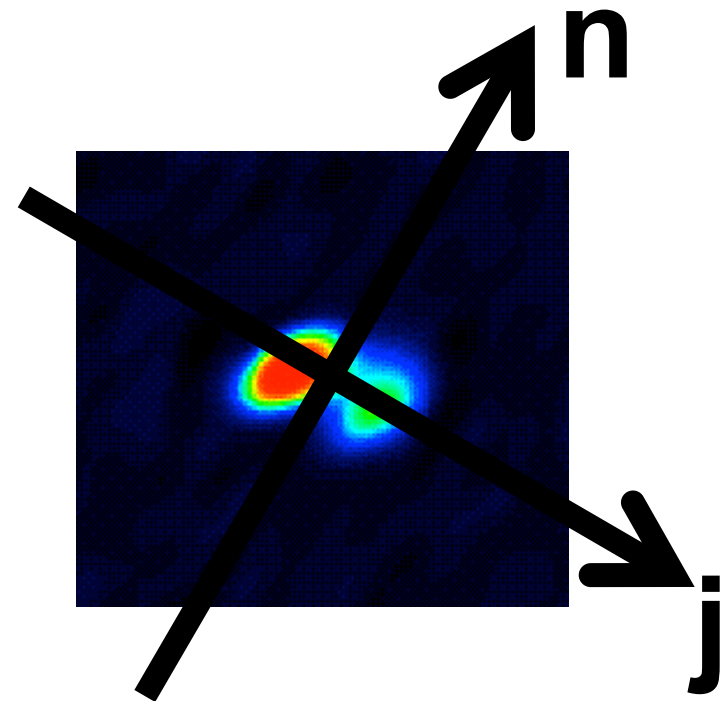
# SYSTEMATIC VERSUS NOISE



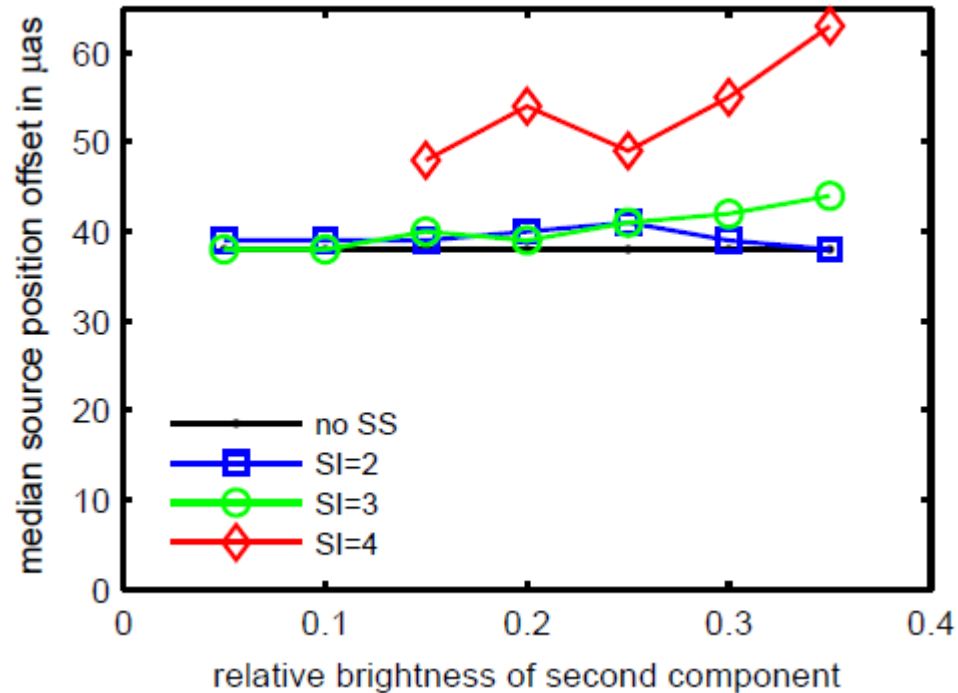
- Sources of lower structure indices (SI=2) tend to displace the sources along the jet direction
- The displacement for higher SI (3,4) are more noise-like

## IDEA

- Parameterise the source position in components along the jet ( $j$ ) / cross jet ( $n$ )
- Reduce the  $j$ -component session-wise and
- Estimate the  $n$ -component in a global solution of all sessions (1 year)

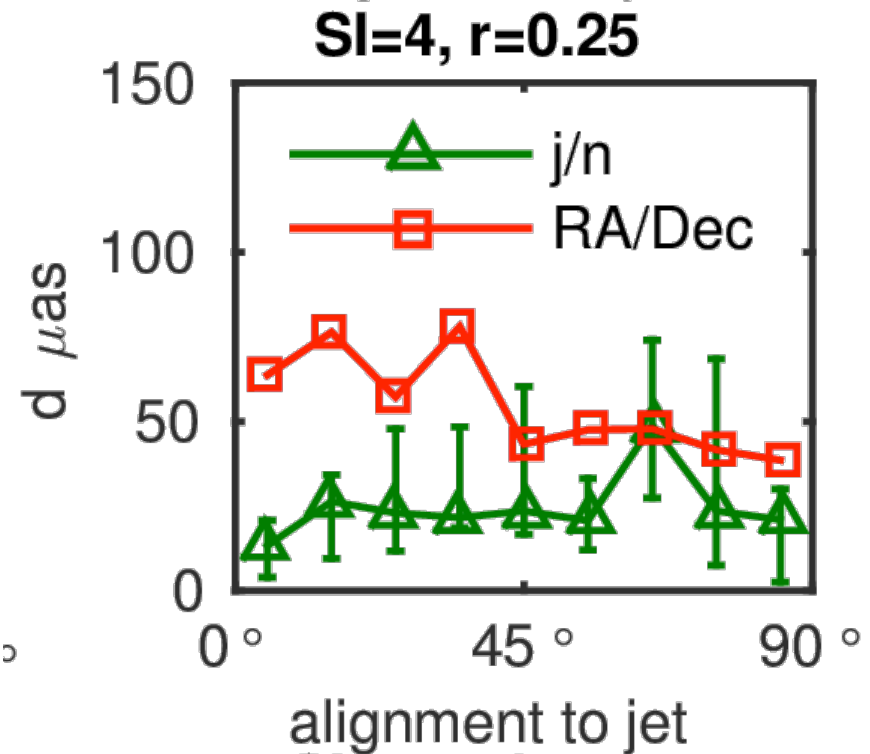
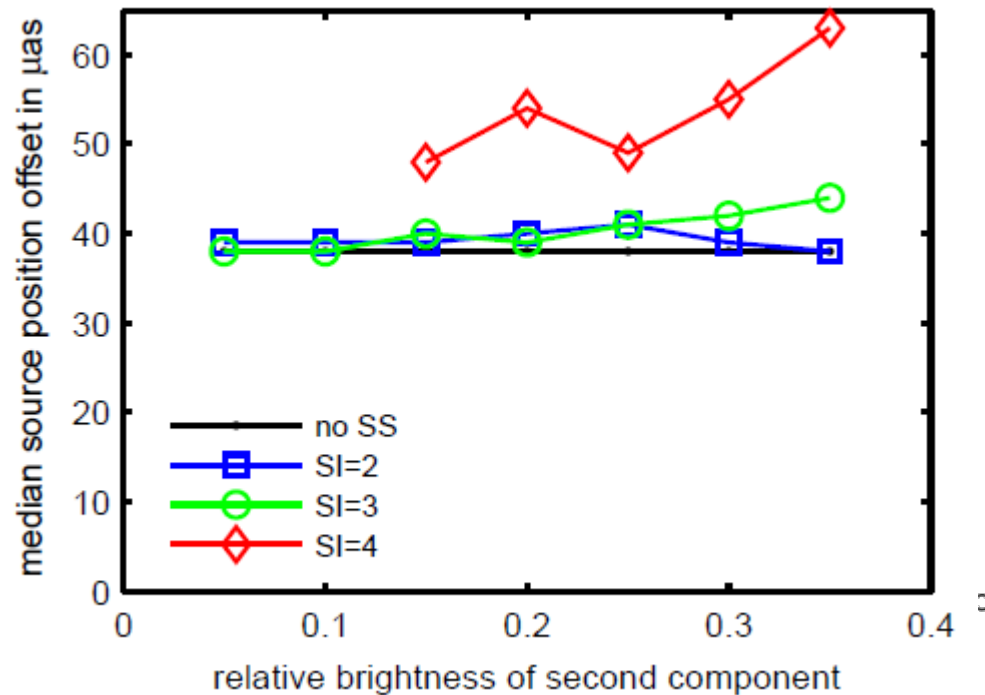


# NEW METHOD



- For SI=2,3 the new  $j/n$  parameterisation reduces the median effect down to the level of tropospheric errors.

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- For  $SI=2,3$  the new  $j/n$  parameterisation reduces the median effect down to the level of tropospheric errors.
- However, it does not perfectly work for structure  $SI=4$ .



# SUMMARY

- In VieVS, we can **apply source structure corrections** (based on real or mock source models) **in VLBI analysis and simulations**.
- Source structure can **systematically displace source positions**
  - Two component models, stable with time.
  - which is not necessarily connected to the nominal SI of the source.
- Sources of lower structure indices (SI=2) tend to **move the source along the jet direction**, while the displacement for higher SI (3,4) are more noise-like.
- A **new modelling strategy (along jet and cross jet)** could be a way to absorb this effect .

# THANK YOU FOR YOUR ATTENTION!

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