

Recent developments at JIVE

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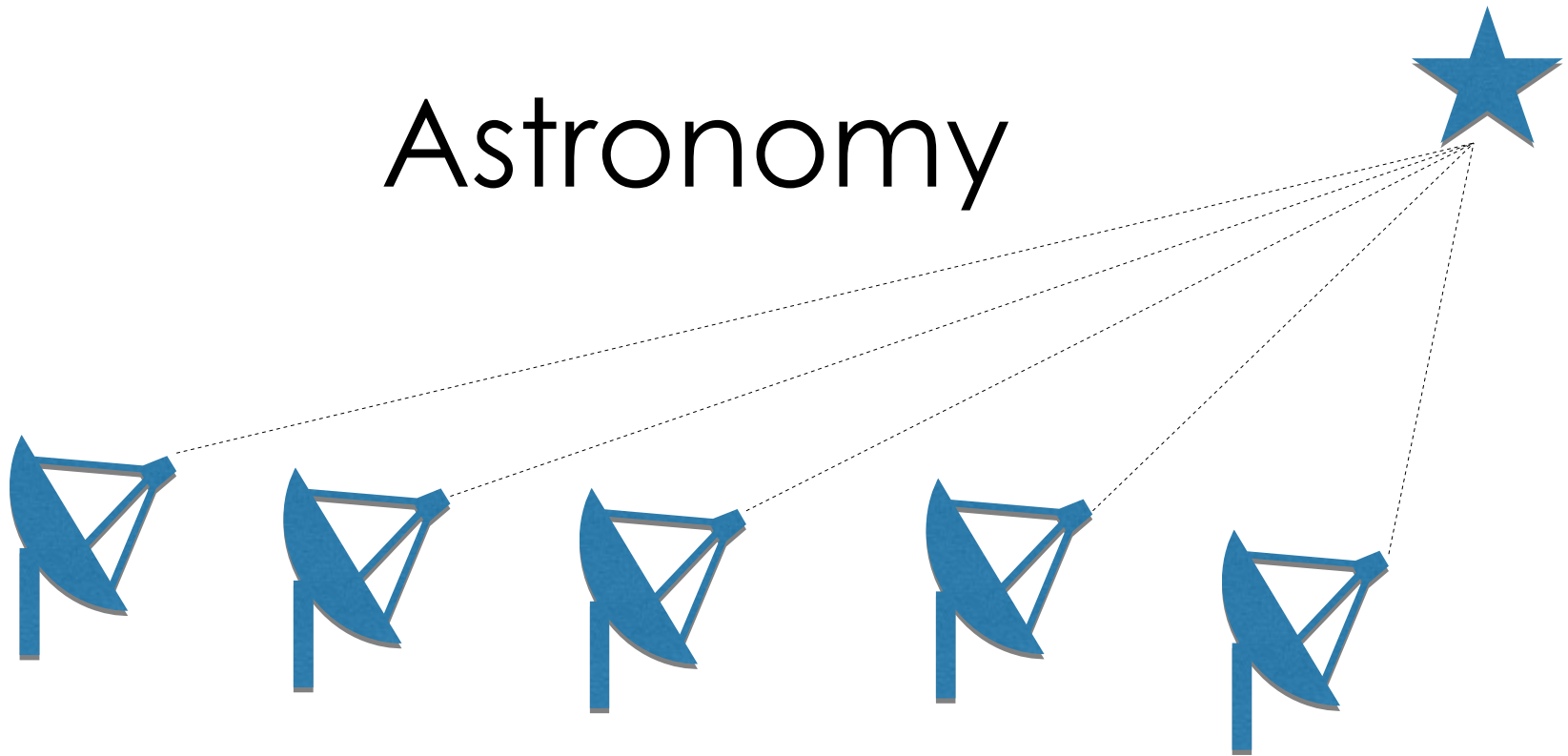
SFXC

- MPI application written in C++
- EVN correlator since 2010, replaced MkIV completely in 2012
- Delay model: CALC10 or external
- Rich set of features
 - Multiple simultaneous phase centers
 - Pulsar binning + coherent dedispersion
 - Mixed-bandwidth correlation
 - **New: Support for Geodetic VLBI**
- Nightly mirror of development repo:
Official: <https://svn.astron.nl/sfxc> Unofficial: <https://github.com/aardk/sfxc>
- Installation instructions:
<http://www.jive.nl/jivewiki/doku.php?id=sfxc>

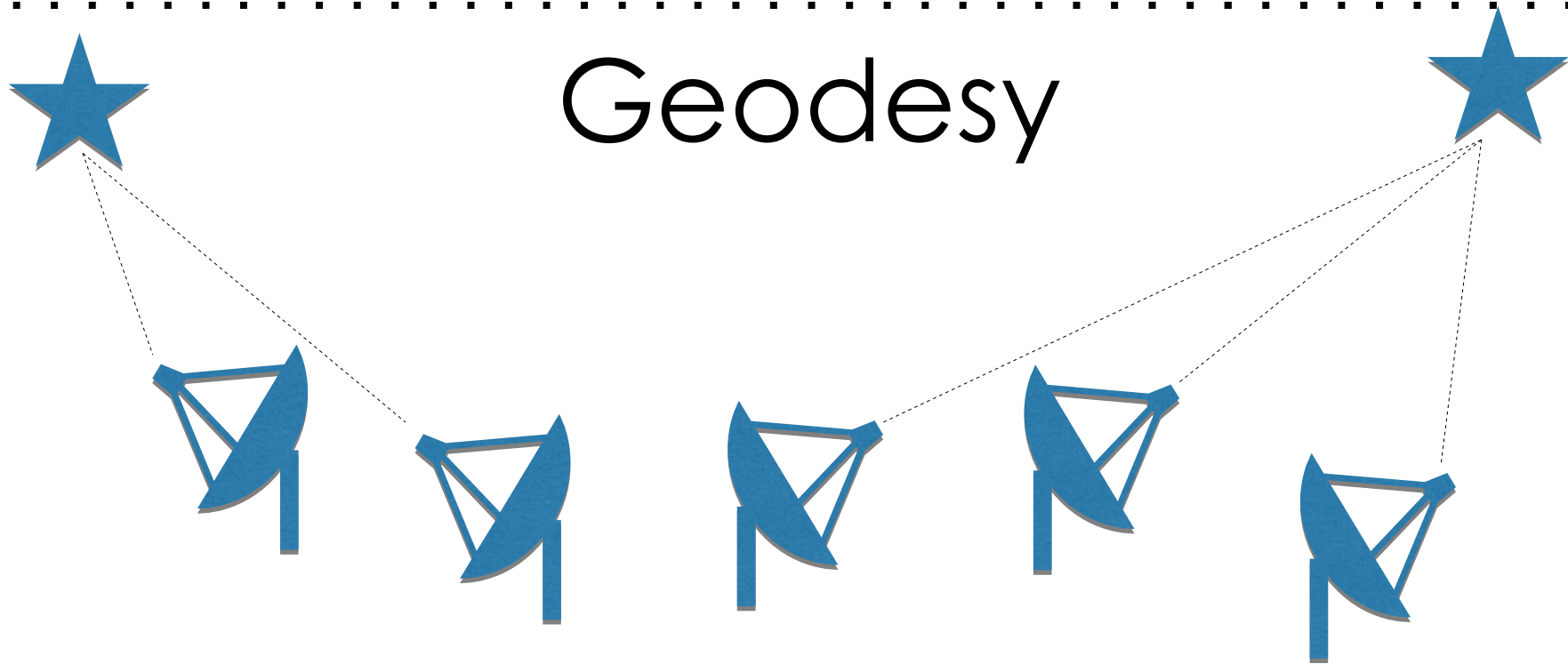
SFXC is Open Source software available under GPL version 2 or later

Keimpema, Kettenis, Pogrebenko, et al., Experimental Astronomy, 39, 259 (2015)
arXiv:1502.00467

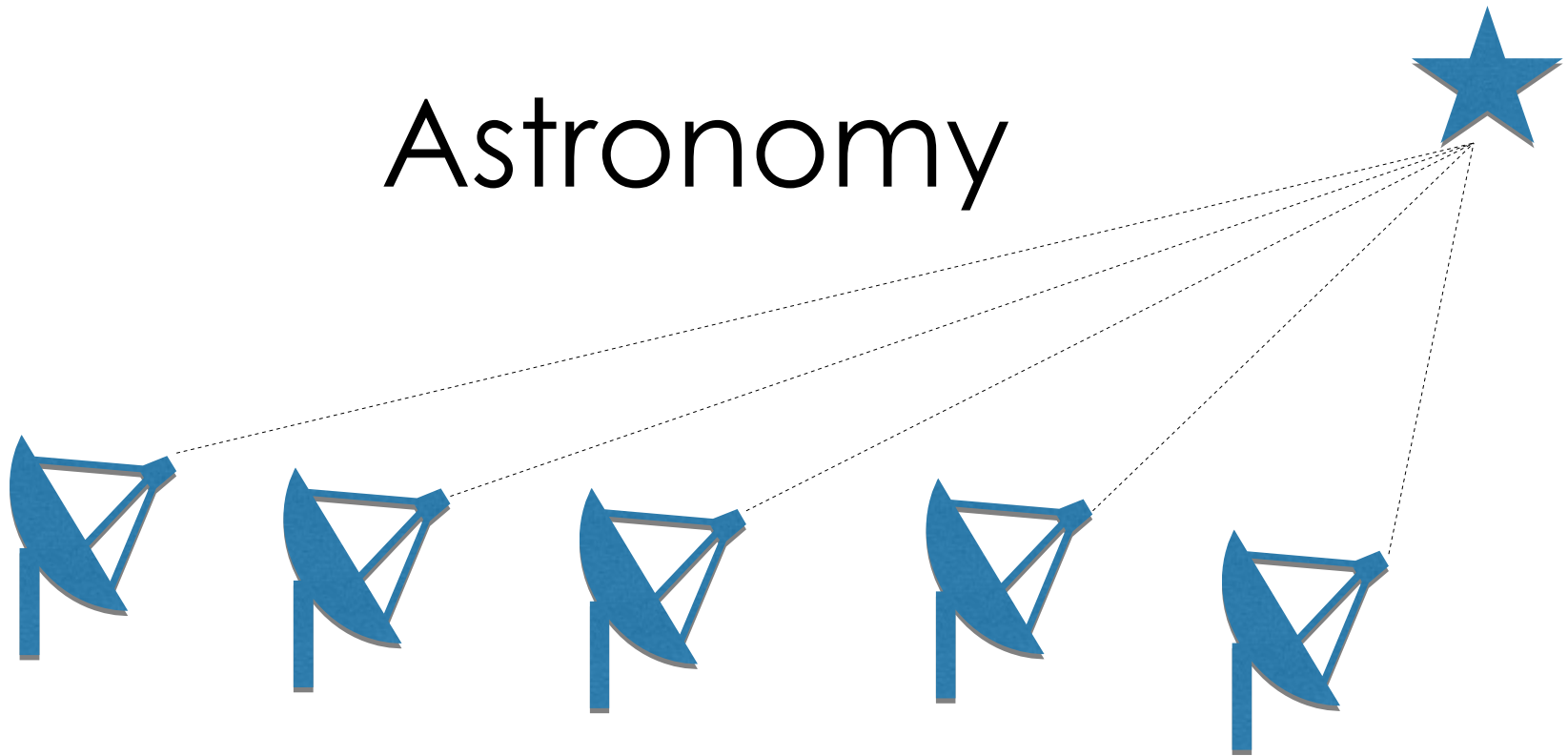
Astronomy



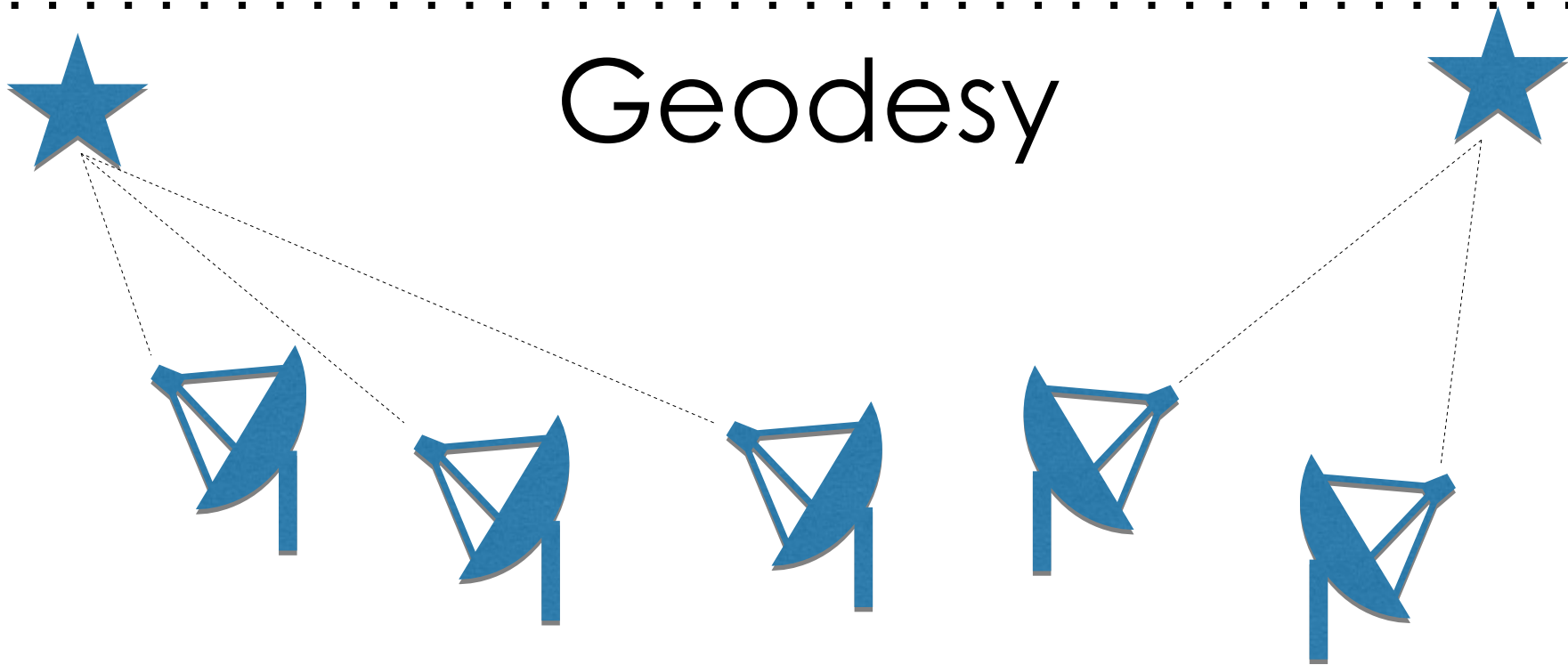
Geodesy



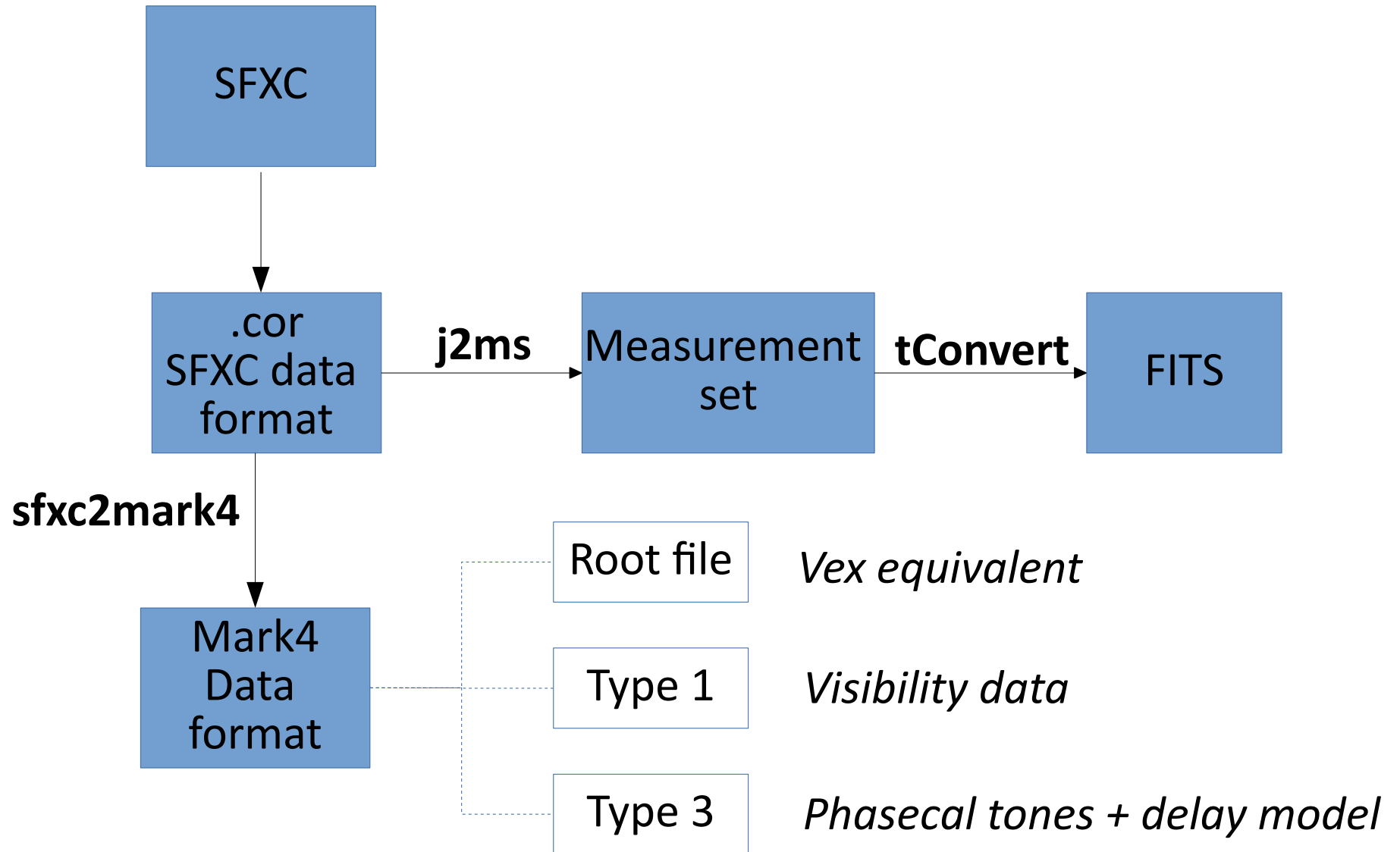
Astronomy



Geodesy



sfxc2mark4

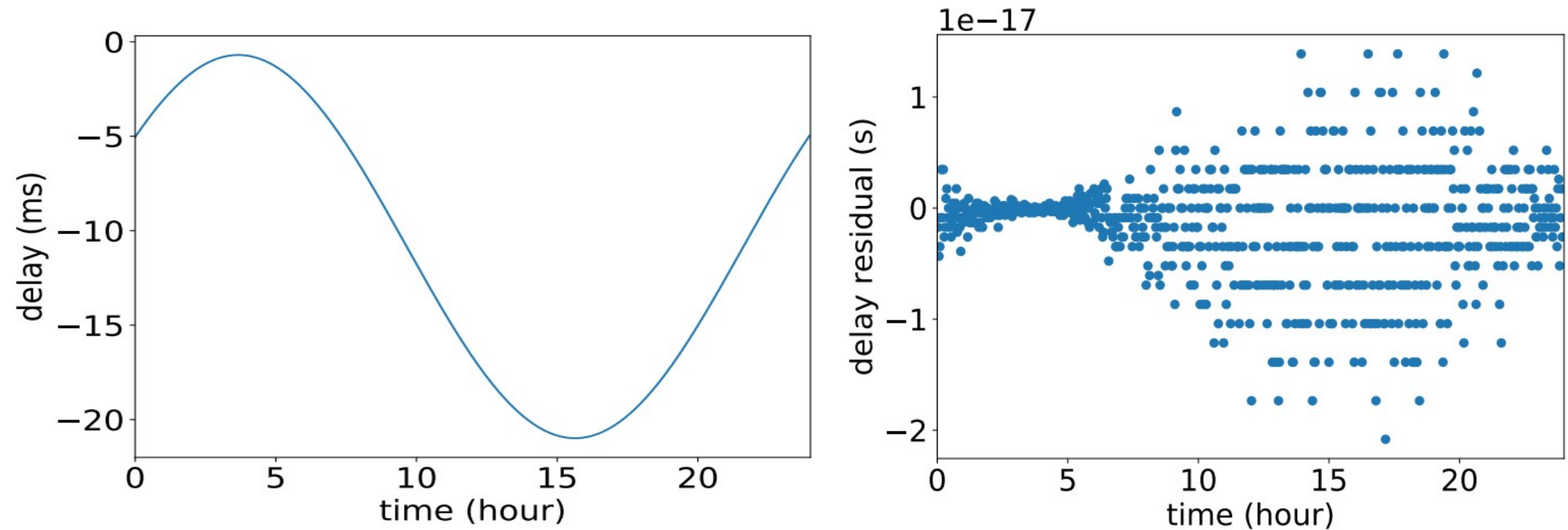


Correlator Model

SFXC	DifX/HOPS/AIPS
sampled every second	spanning a few minutes
interpolated using Akima spline	polynomial of (up to) degree 5

- Sample Akima splines
- Fit polynomial of degree 5 over time intervals, default interval: 120 sec
- Implemented using NumPy / SciPy

Correlator model



More verification in Maria Eugenia Gomez's talk

Building up the cluster

Year	Nodes	CPUs	SIMD	Clockspeed (GHz)	Cores	Stations @1Gb/s
2010	16	Xeon E5520	SSE 4.2	2.3	128	?
2011	16	Xeon E5620	SSE 4.2	2.4	128	7
2012	8	Xeon E5-2670	AVX	2.6	128	13
2015	4	Xeon E5-2630v3	AVX2	2.4	64	16
2016	4	Xeon E5-2630v3	AVX2	2.4	64	19
2018	42	Xeon E5-2630v4	AVX2	2.2	840	12 x 4Gpbs
Total	90				1352	

Building up the cluster

Year	Nodes	CPUs	SIMD	Clockspeed (GHz)	Cores	Stations @1Gb/s
2010	16 10	Xeon E5520	SSE 4.2	2.3	128 80	?
2011	16	Xeon E5620	SSE 4.2	2.4	128	7
2012	8	Xeon E5-2670	AVX	2.6	128	13
2015	4	Xeon E5-2630v3	AVX2	2.4	64	16
2016	4	Xeon E5-2630v3	AVX2	2.4	64	19
2018	42	Xeon E5-2630v4	AVX2	2.2	840	11 x 4 Gbps
Total	84				1304	

Old compute nodes
Flexbufs (512 cores)

Flexbufs

4x
Mellanox
SN2100

HP5412ZL
Switch

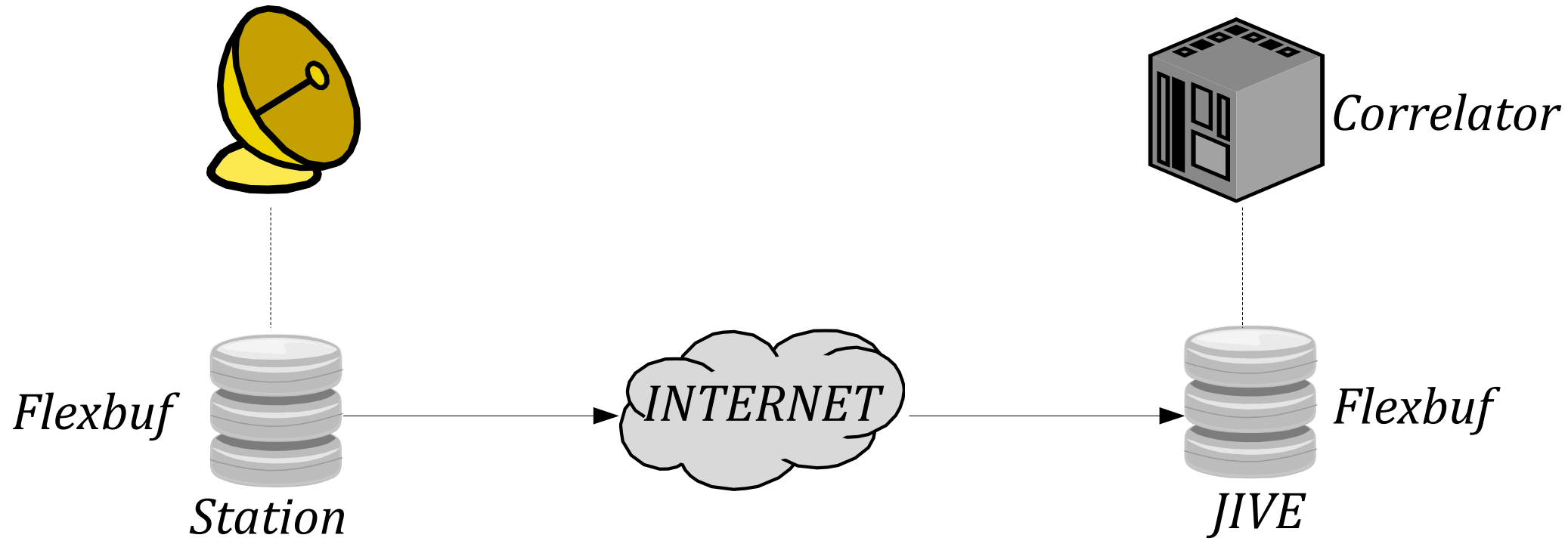
New compute nodes
(840 cores)



Storage

- 17 Mark5A/B units
- 6 Mark5C units
- 2 Mark6; 2 x (master + slave)
- 14 Flexbuf units
 - almost 3 PB RAID-Z storage
 - e-shipping from most EVN stations
 - Wb records directly to Flexbuf @JIVE

Flexbuf transition



- During observation data is recorded on local flexbuff at the station
- Data is transferred to flexbuff at JIVE for correlation
- For each station there is a corresponding flexbuff at JIVE
- Raw data is kept for one additional session after correlation
- Currently 14 flexbuffs at JIVE

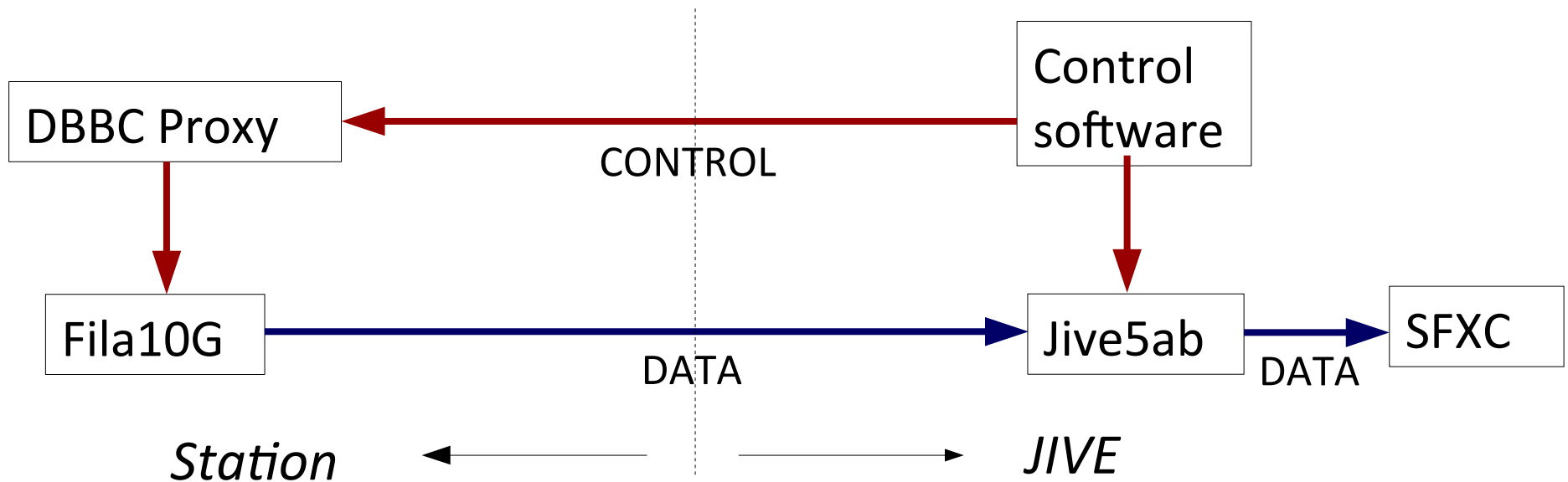
Realtime eVLBI at JIVE

- About 23-30% of EVN observing time
- Advantages of eVLBI
 - Reliability, online monitoring
 - Low latency from observation to distribution to PI
 - Rapid response
 - (Automatic) triggered observations
 - Target of opportunity
- Typical observation size is 10-11 stations
- EVN now defaults to 2 Gbps eVLBI (C band and higher), 4 Gbps after network upgrade (30 Gbps → 100 Gbps)
- Simultaneous correlation and recording on Flexbuff

e-EVN

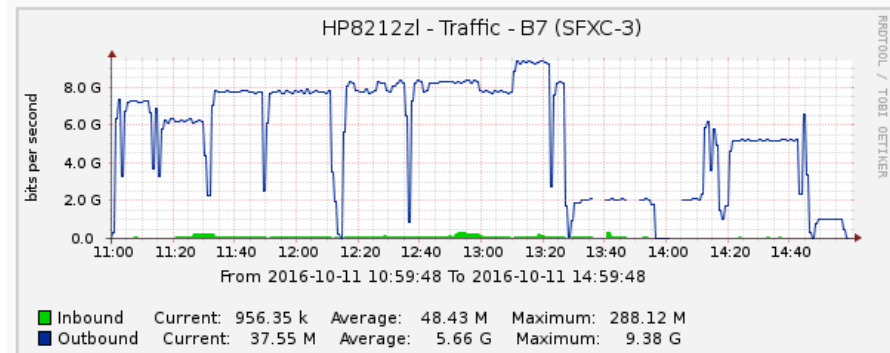
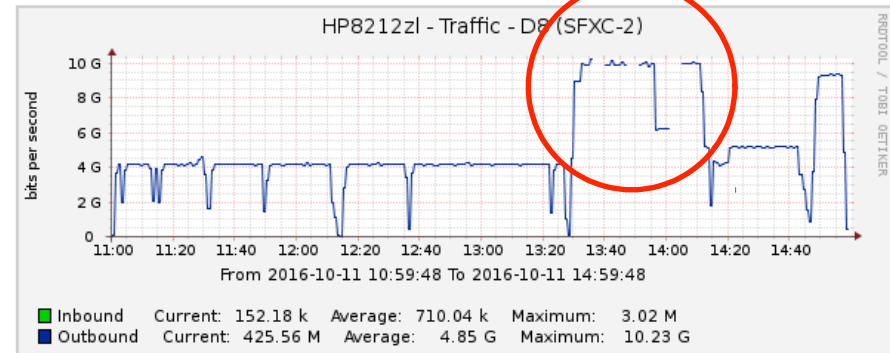
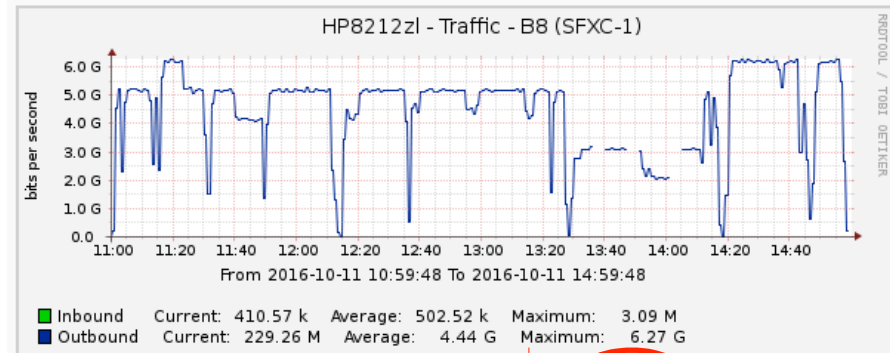
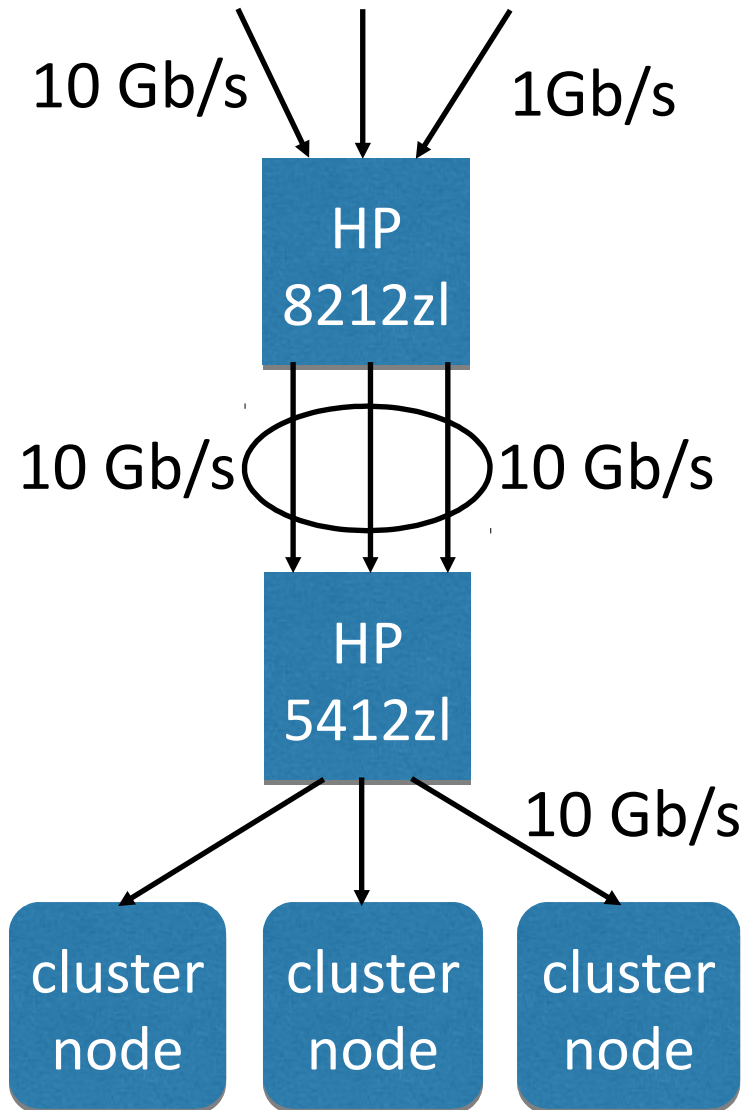
Station	Bandwidth	Type	Fila10G
Arecibo	512 Mbps	Routed	No
Effelsberg	4 Gbps	Lightpath	Yes
HartRAO	2 Gbps	Lightpath	Yes
Irbene	4 Gbps	Lightpath	Yes
Jodrell Bank	4 Gbps	Dedicated	Yes
Metsahovi	4 Gbps	Routed	Yes (?)
Medicina	4 Gbps	Routed	Yes
Noto	4 Gbps	Routed	Yes
Onsala	4 Gbps	Lightpath	Yes
Shanghai / Tianma65	1 Gbps	Routed	Yes
Torun	1 Gbps	Routed	No
Westerbork	2 Gbps	Dedicated	Yes
Yebes	4 Gbps	Routed	Yes

eVLBI data flow



- DBBC proxy APU allows communication with Fila10G
- Fila10G sends UDP data stream (typical 8032 byte packet size)
- One UDP packet = one VDIF frame
- Jive5ab detaches the data stream from the correlation, correlation continues if data stream stops
- **4 Gbps maximum data rate**

Internal Network Bottleneck



Scalability DBBC3 (32 Gbps)

- EVN personality: 2 x 4 GHz IF (=2 x 16 Gbps)
- How to divide data?

4 GHz

4 GHz

VS

1 GHz

1 GHz

1 GHz

1 GHz

1 GHz

1 GHz

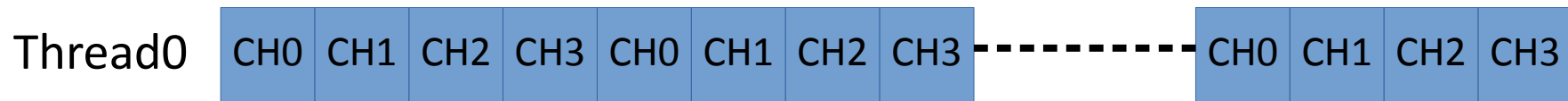
1 GHz

1 GHz

More is better!

Scalability DBBC3 (32 Gbps)

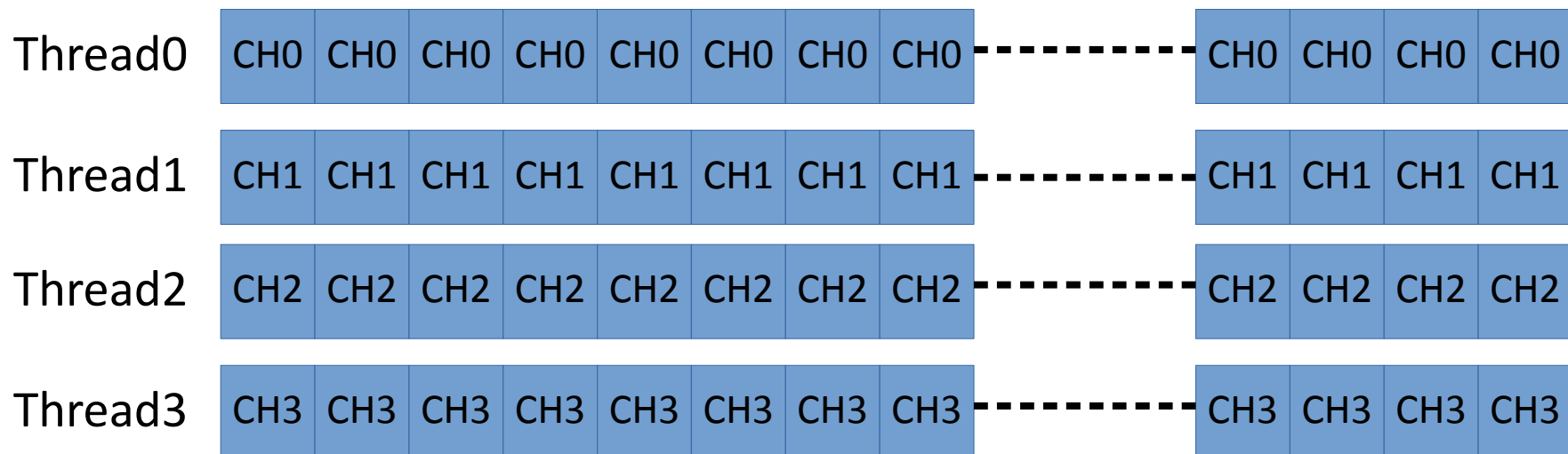
- Single threaded VDIF



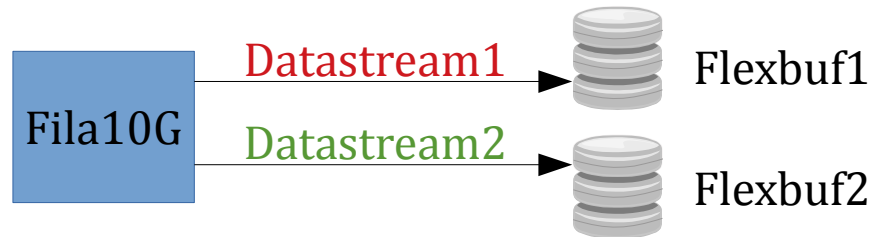
Requires expensive corner turning

VS

- Multi threaded VDIF

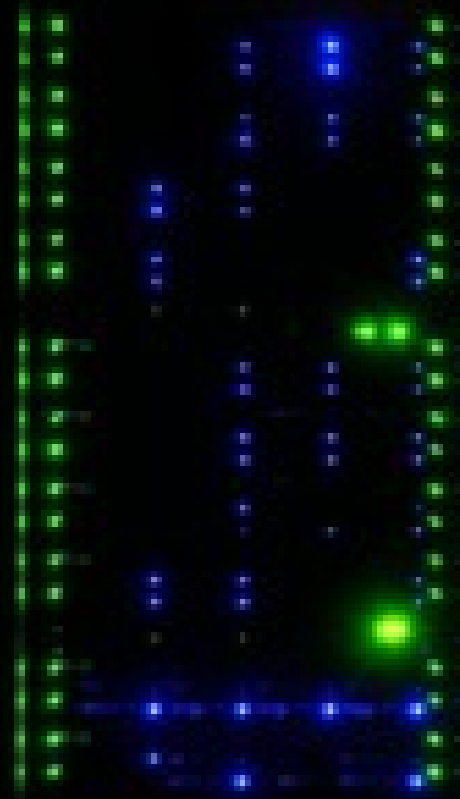


Multiple Data Streams



- To prevent networking / computation bottlenecks the data is split into multiple (typically 2 Gbps) datastreams
- Each data stream is handled by a separate input node
- Supported in VEX2:

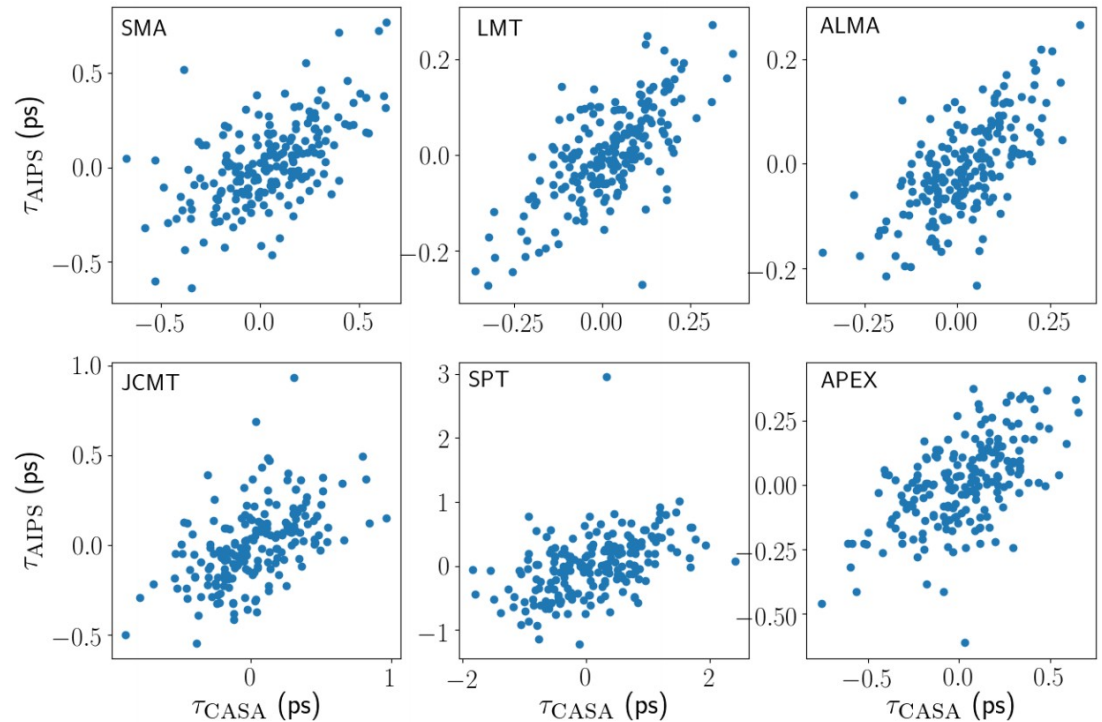
```
$DATASTREAMS;  
*  
def DVP;  
  datastream = &DS1 : VDIF_legacy;  
  thread = &DS1 : &THR1 : 0 : 32 : 8 Ms/sec : 2 : real : 8000;  
  channel = &DS1 : &THR1 : &CH00 : 0;  
  channel = &DS1 : &THR1 : &CH01 : 1;  
  ...  
  channel = &DS1 : &THR1 : &CH15 : 15;  
  datastream = &DS2 : VDIF_legacy;  
  thread = &DS2 : &THR2 : 0 : 32 : 8 Ms/sec : 2 : real : 8000;  
  channel = &DS2 : &THR2 : &CH16 : 16;  
  channel = &DS2 : &THR2 : &CH17 : 17;  
  ...  
  channel = &DS2 : &THR2 : &CH31 : 31;  
enddef;
```



CASA transition

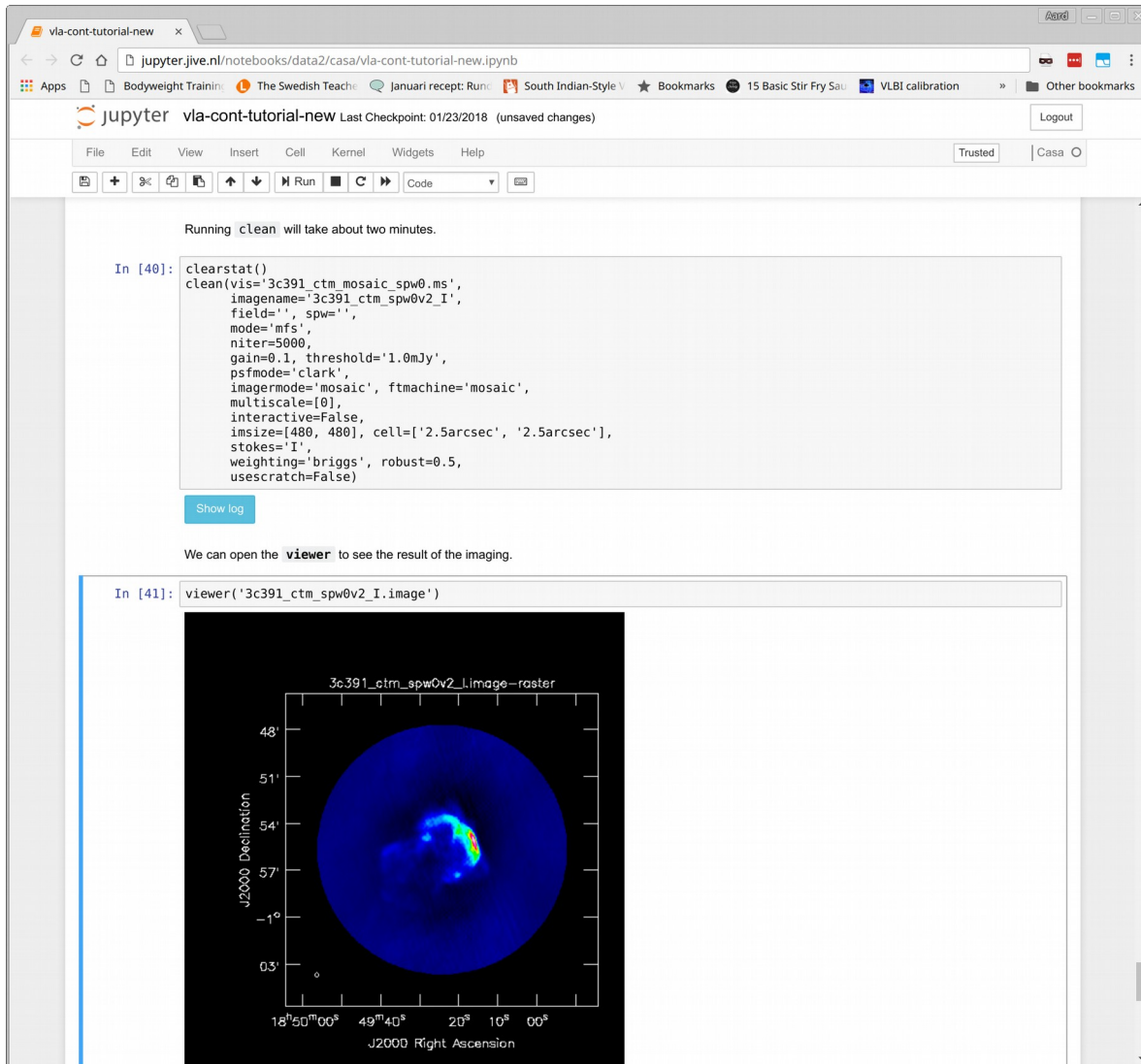
- AIPS is now legacy software
- CASA very widely used:
 - VLA, ALMA, Meerkat, SKA
- ANTAB a priori amplitude calibration
- Fringe fitter for CASA
- Developed at JIVE by:
*Des Small, Mark Kettenis,
and Ilse van Bemmelen*

Residual delay CASA vs AIPS
for simulated EHT data



*Iniyana Natarajan (Rhodes University),
MNRAS (2019), in preparation*

Remote pipelining using Jupyter and CASA



The screenshot shows a Jupyter Notebook interface in a web browser. The browser address bar shows the URL `jupyter.jive.nl/notebooks/data2/casa/vla-cont-tutorial-new.ipynb`. The notebook title is `vla-cont-tutorial-new` and it shows the last checkpoint as `01/23/2018`. The interface includes a menu bar (File, Edit, View, Insert, Cell, Kernel, Widgets, Help) and a toolbar with icons for file operations, running, and code execution. The main content area shows two code cells. The first cell, labeled `In [40]:`, contains a `clean` command with various parameters: `clearstat()`, `clean(vis='3c391_ctm_mosaic_spw0.ms', imagename='3c391_ctm_spw0v2_I', field='', spw='', mode='mfs', niter=5000, gain=0.1, threshold='1.0mJy', psfmode='clark', imagermode='mosaic', ftmachine='mosaic', multiscale=0, interactive=False, imsize=[480, 480], cell=['2.5arcsec', '2.5arcsec'], stokes='I', weighting='briggs', robust=0.5, usescratch=False)`. Below the code cell is a `Show log` button. The second cell, labeled `In [41]:`, contains the command `viewer('3c391_ctm_spw0v2_I.image')`. Below this command is a plot of a radio astronomy image. The plot is titled `3c391_ctm_spw0v2_I.image--raster` and shows a circular field of view with a central bright source. The axes are labeled `J2000 Declination` (ranging from 48' to 03') and `J2000 Right Ascension` (ranging from 18^h50^m00^s to 20^h10^m00^s).

- Jupyter kernel for CASA
- Docker and Singularity containers
- Documentation and download <https://github.com/aardk/jupyter-casa>
- Live demo: <http://jupyter.jive.nl>

Architecture

*Data are corner-turned,
subbands are sent to
different correlator
nodes*

