

# EU-VGOS Project

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on behalf of the collaboration

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# Motivation for EU-VGOS

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- VGOS development delayed. Antennas available and ageing.
- Correlator resources at Bonn planned for VGOS ageing.
- Need to acquire know-how for VGOS observing & data reduction.
- Haystack is developing path for observing, correlation and fringe-fitting (FF).
  - ✓ Pseudo Stokes I ignores source polarisation, changes and differences between position of peak in total and polarisation intensity.
  - ✓ Polarisation leakage leads to delay errors
  - ✓ Source structure and its variation, position as function of frequency
- Different FF programs will become available for VGOS (PIMA, CASA, ??)
- Optimal method for VGOS correlation and FF not scientifically evaluated!!

- To reach 1mm accuracy all steps of VGOS observing and data reduction have to be optimised!
- Get community involved! DiFX is excellent example how this works.

# Aims of EU-VGOS

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- ✓ Define Aims of EU-VGOS (some flexibility, ongoing)
- ✓ Find collaborators: Telescopes, experts, correlator, regular meetings
- ✓ Scheduling of VGOS observation, depends on aim of observation
  - ✓ Find fringes, calibration X/Y & AMP, stability, maps, geodetic parameters (ongoing)
- ✓ Observing
- ✓ Data transport → Internet
- ✓ Correlation including QA & feedback
- ✓ Fringe-fitting: initially fourfit, later others
- ✓ Polarisation conversion: X/Y → L/R (includes coarse FF and mapping)

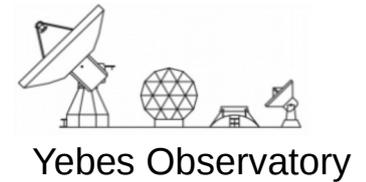
Comparison of methods  
Source structure effects  
Effects of polarisation  
.... more

# Participating Stations

- **ONSA13NE**: 13.2m, antenna, receiver QRFH, VLBI back-end = DBBC3/flexbuff
- **WETTZ13S** (WETTZ13N not equipped with broadband receiver yet): 11-feed, 2xDBBC2/Mark6 units
- **RAEGYEB**: 13.2 m antenna, receiver QRFH, 4xRDBE/1xMark6 units



# Collaborators: see 1<sup>st</sup> slide



# Scheduling of first sessions

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- Used SKED/SCHED: creates VEX, plots for parallactic angle, UV coverage
- Schedule strong calibrators at regular intervals
  - FF, bandpass, pol. cal, phase stability, amp cal
- $\geq 1$  source with a high fractional polarization
- $\geq 1$  source with a low fractional polarization
- vgt050 "imaging" session, observed only strong sources

4 hrs sessions - Standard geodetic schedule (time on source = 30 s) + scans targeting 4 strong calibrators (integration time  $\sim$  120s)

# Frequency set-up

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- Four bands with dual linear polarization:
  - 3000.40 MHz - 3480.40 MHz (band A in Haystack nomenclature; band S in Fourfit)
  - 5240.40 MHz - 5720.40 MHz (band B/ band C)
  - 6360.40 MHz - 6840.40 MHz (band C/ band X)
  - 10200.40 MHz - 10680.40 MHz (band D/ band X)
- In each band are sampled 8 channels of 32 MHz bandwidth (sample\_rate of 64 Ms/s)
- 2-bit sampling, thus meaning a **data rate of about 8 Gbps**
- Recording on Flexbuff (JIVE)/Mark6 modules

# Correlation

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**Onsala** (real data) – 1 VDIF file per scan with 8 threads, 8 channels per thread

**Wetzell** (real data) – 1 VDIF file per scan with 4 threads (1 for each band), 16 channels (8 channels x 2 linear polarizations) per thread.

**Yebe**s (complex data) – 1 VDIF file per scan with 4 threads, 16 channels/thread

Data are e-transferred to Bonn

Issues with file-based correlation in DiFX due to

- 1) clumpy time threads
- 2) too small data buffering in DiFX

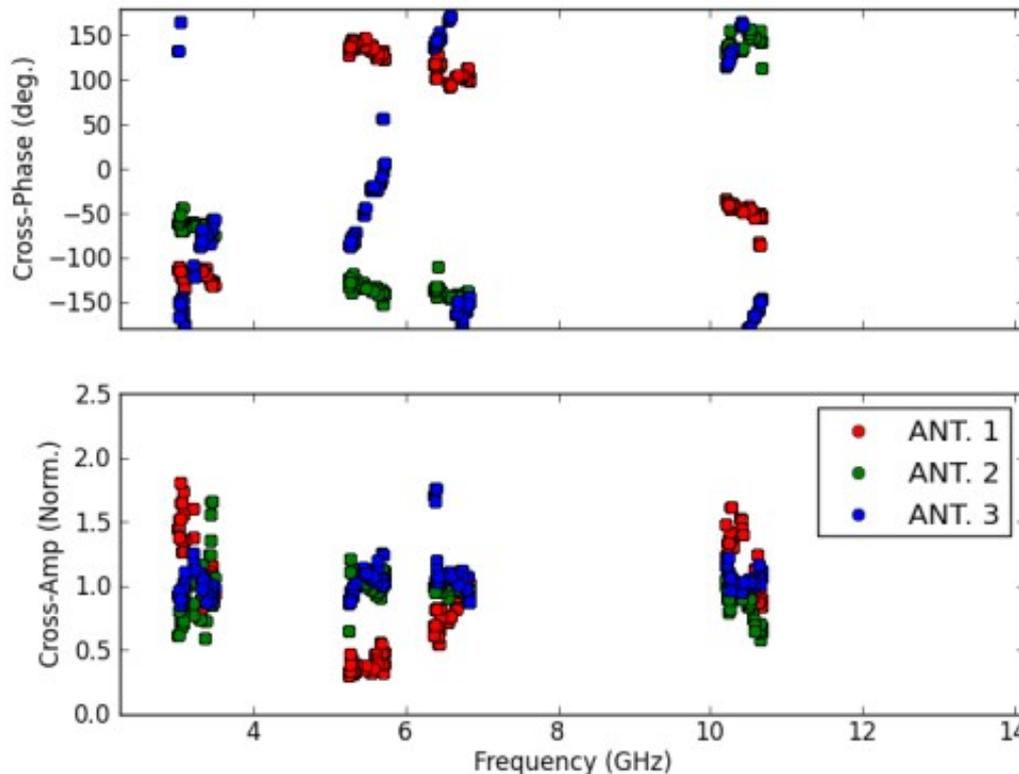
Multi-thread raw data must be converted, else data loss

DiFX produces 4 cross-correlation products (XX,YY,XY,YX) to be fed to the PolConverter

# PolConversion

- X/Y relative phases are tracked using the pcal X-Y phase differences.
- Additional (add-hoc) X/Y phases are applied (due to e.g., pcal cabling).
- Need amplitude calibration to track X/Y relative amplitudes.
- Add-hoc X/Y phases (and amplitudes) can be derived from calibrator observations (Global Cross-Polarization Fringe Fitting, GCPFF, with PolConvert).

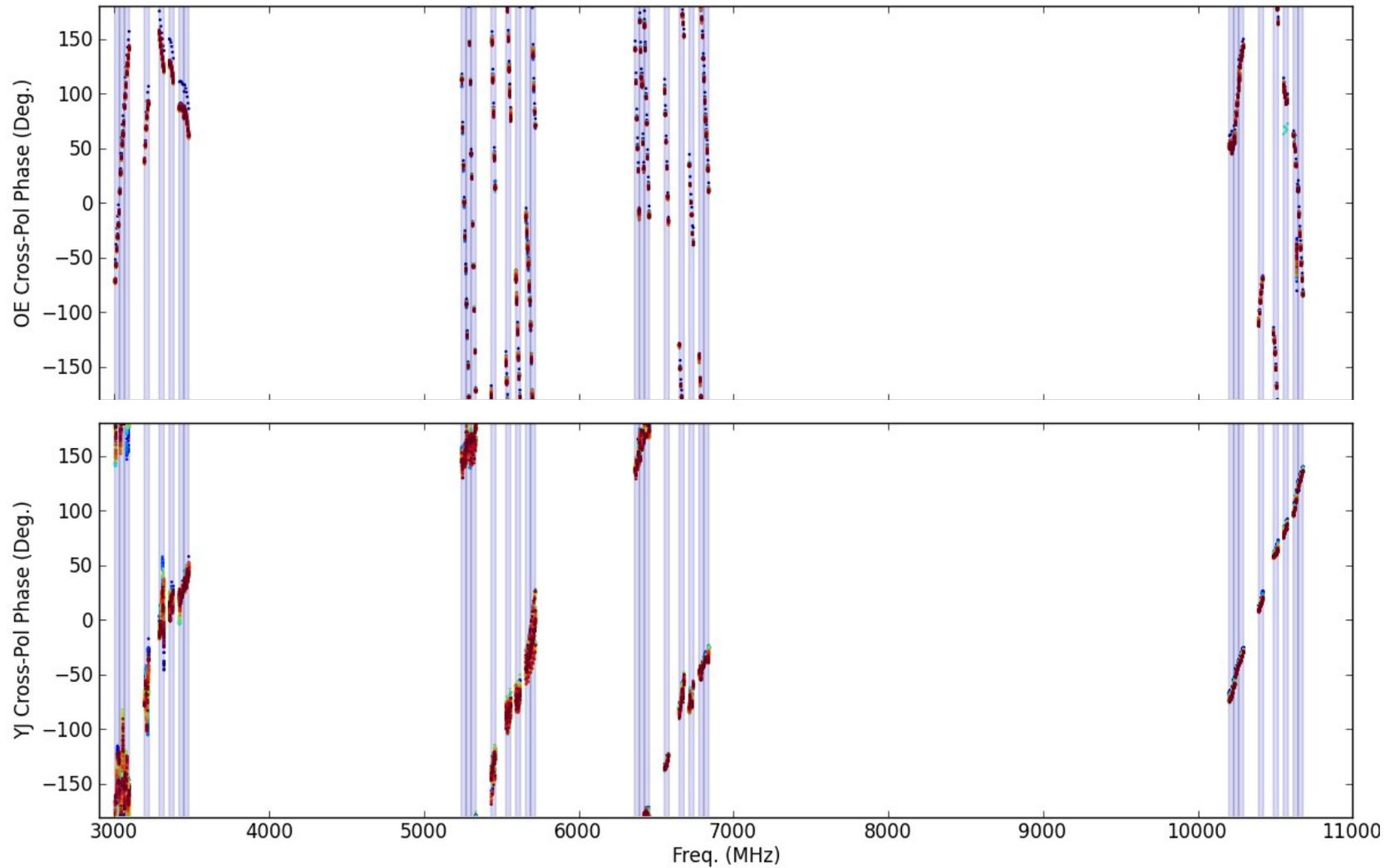
CROSS-POLARIZATION GAINS



X/Y relative phases and amplitudes for **OE** (ANT. 1), **WS** (ANT. 2) and **YJ** (ANT. 3).

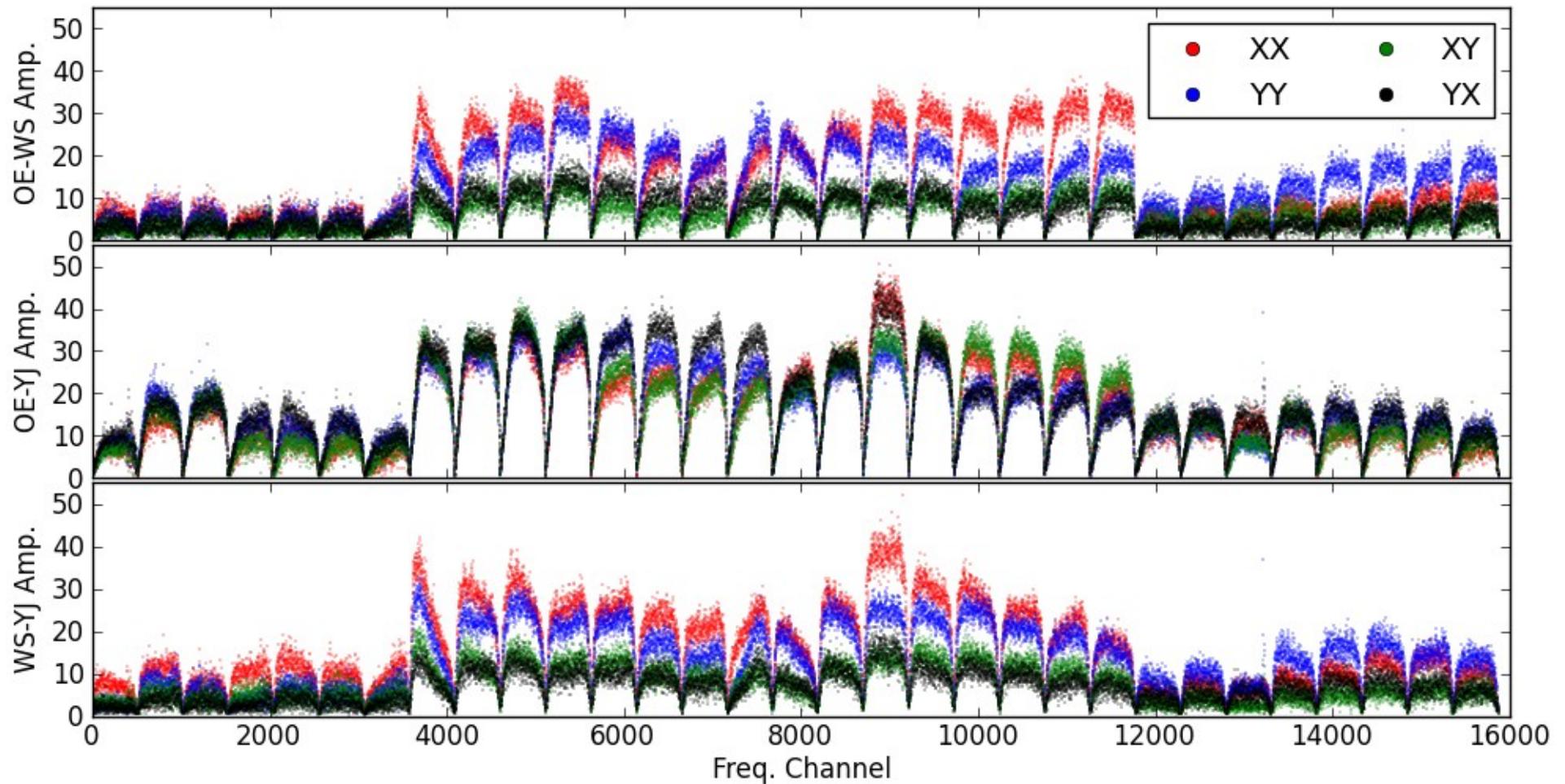
Experiment **VGT274**.

# PolConversion



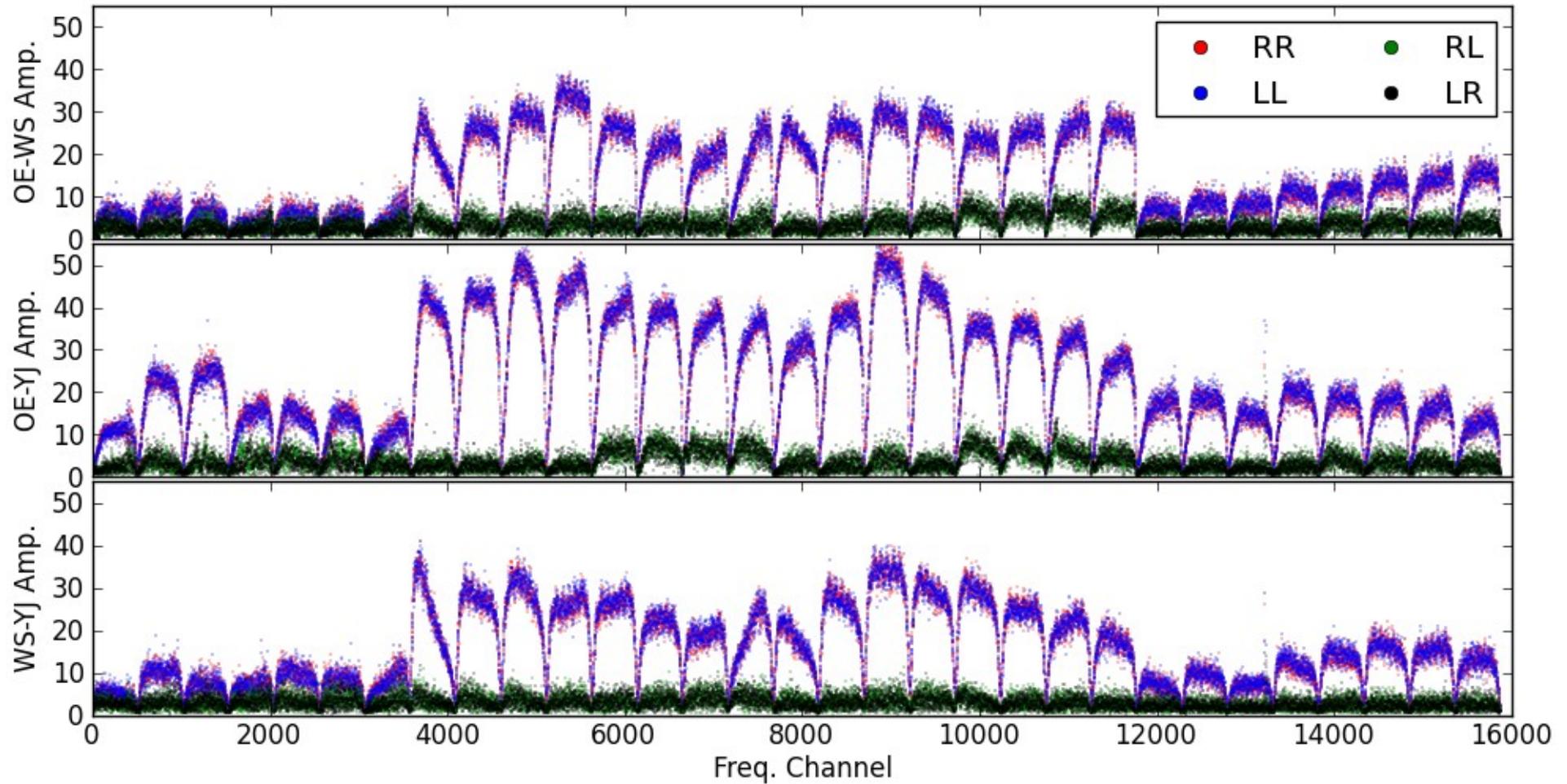
X/Y relative pcal phases for all scans of experiment **VGT274**

# PolConversion



Amplitudes for all IFs (RAW correlation) for a scan on 3C84 (calibrator).

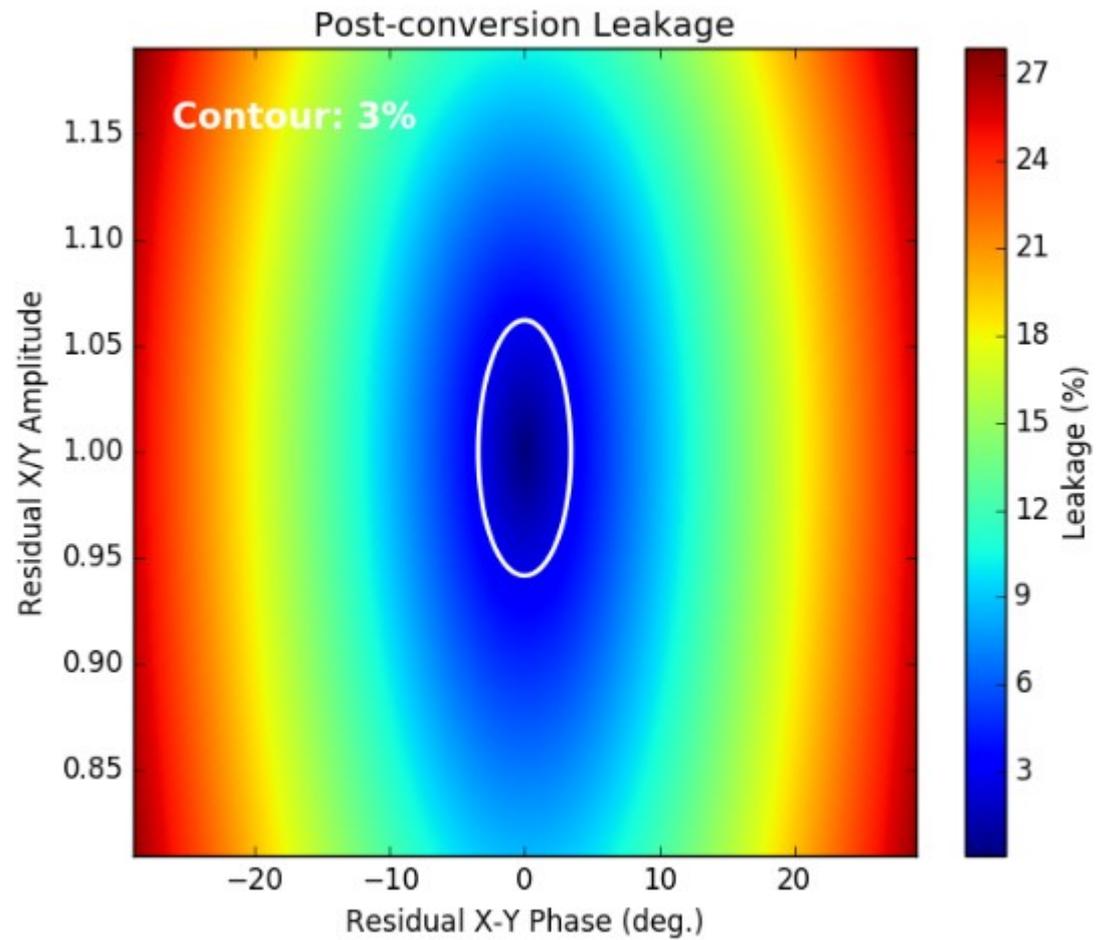
# PolConversion



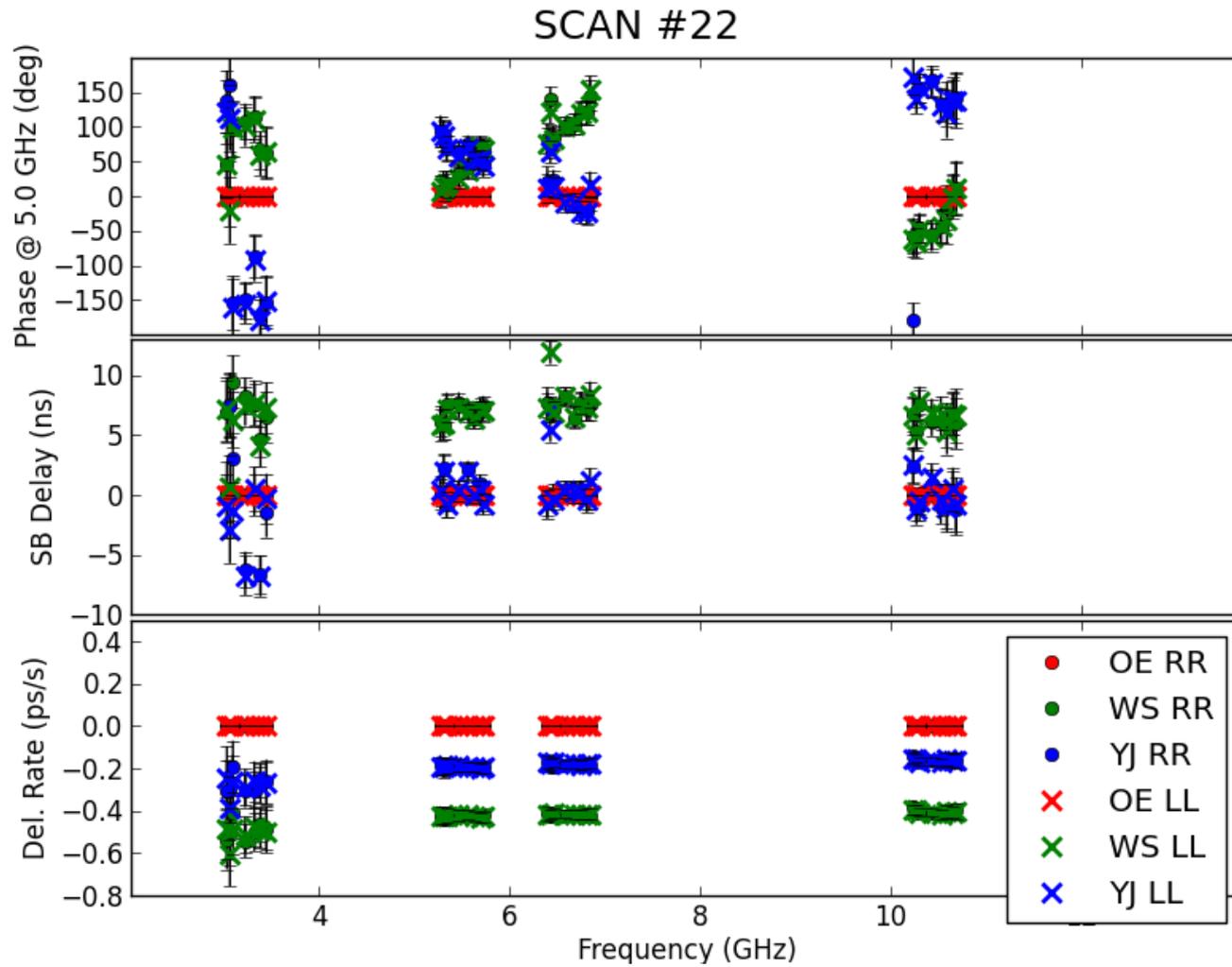
Amplitudes for all IFs (PolConverted) for a scan on 3C84 (calibrator).

# PolConversion

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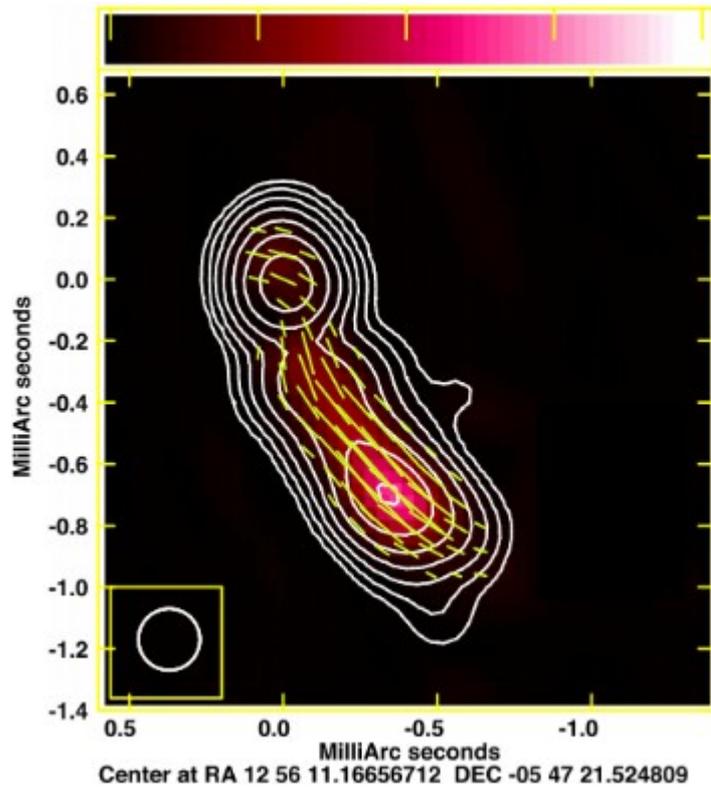


# PolConversion

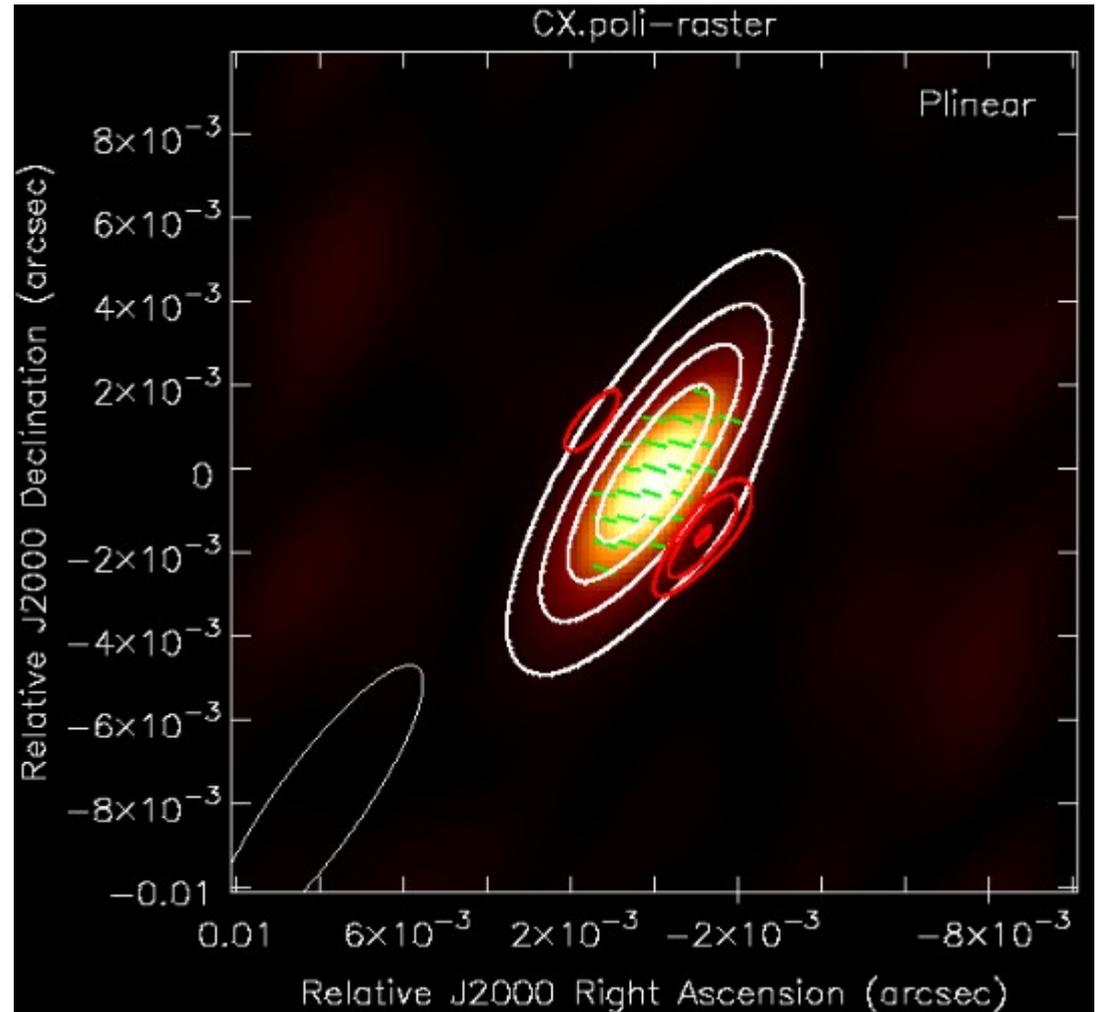


Global Fringe Fitting (per IF) on a PolConverted scan of 3C84.

# Full-Stokes Imaging



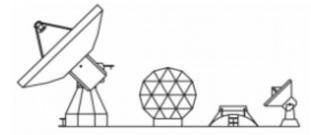
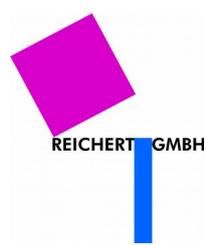
3C279 at 43GHz  
(Rani et al. 2018)



3C279 at 5-11GHz  
(EU-VGOS, Exp. VGT260)



Bundesamt für  
Kartographie und Geodäsie



Yebes Observatory



# Thank you!

