

# Ionospheric calibration for K-band celestial reference frames

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# Introduction

- CRF based on K-band (24 GHz) part of ICRF3
  - Very Long Baseline Array (VLBA) + Hobart-HartRAO
  - Mostly 2 Gbps data rate, testing 4 Gbps
  - Posters by De Witt et al, Krasna et al, and talk by Le Bail et al
- + Advantage: less source structure compared to S/X
  - S/X CRFs limited by a 20-30  $\mu$ as source structure noise floor [Le Bail et al]
- Disadvantage: single-frequency observations
  - About 95% of the K-band ionospheric effect needs to be corrected to get below S/X noise level

Accuracy depends on external ionospheric calibration

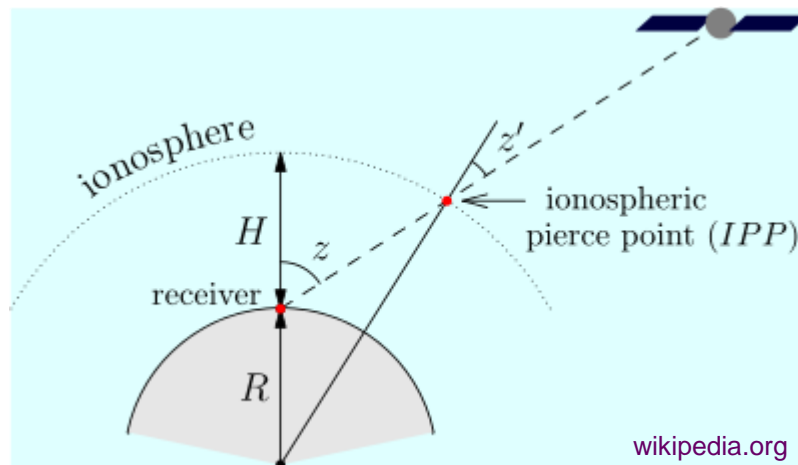
# How to improve the ionospheric calibration?

*Ongoing investigation*

1. Improvements in ionosphere models
2. Improvements from better model coverage of VLBI stations

# Background – ionospheric corrections

- Dual-frequency observations: ionosphere-free linear combination
- Single-frequency observations: ionospheric delay determined from ionosphere models
- Simplest models: two-dimensional, single thin-layer approximation



Ionospheric delay:

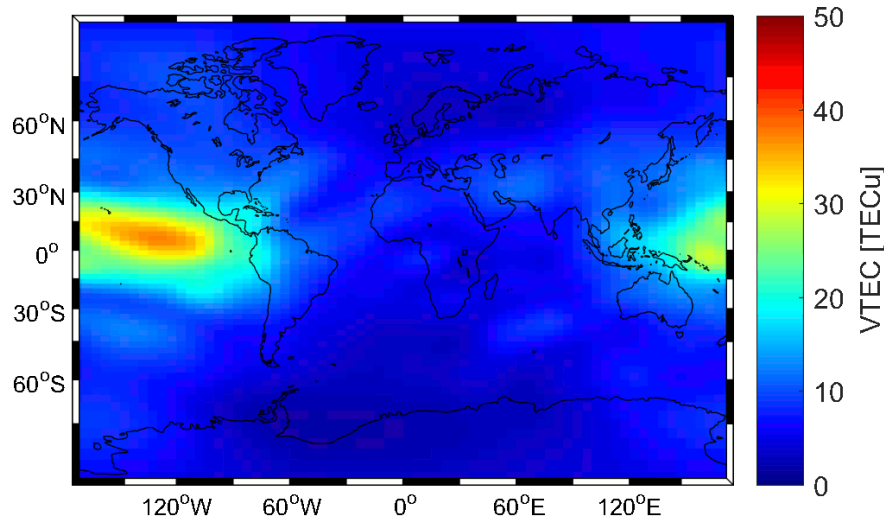
$$\tau_{\text{iono}} = \frac{40.3}{cf^2} \text{ STEC}$$

STEC ... slant total  
electron content

# Global Ionospheric Maps (GIMs)

- International GNSS Service (IGS) operationally provides 2-dimensional maps of vertical total electron content (VTEC)
  - Based on a combination of GIMs from individual analysis centers
  - Temporal resolution: 2 hours
  - Spatial resolution: 5 deg (lon) x 2.5 deg (lat)
  - Stated uncertainty 2-8 TEC units (TECu)

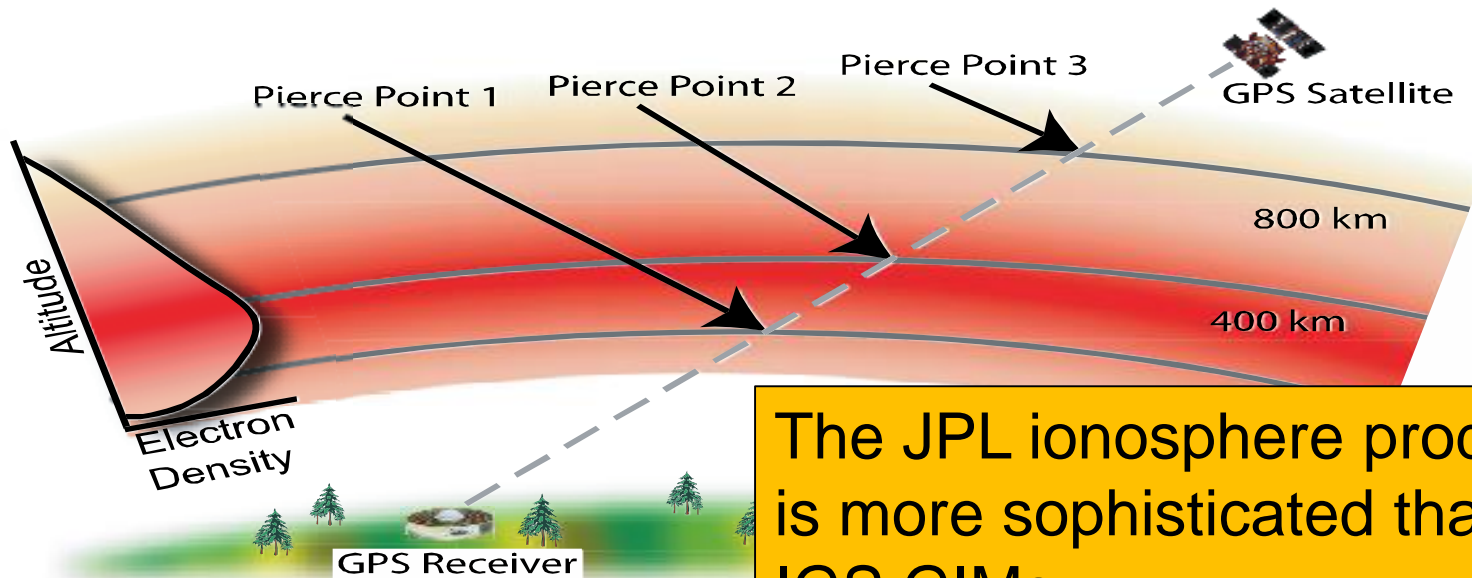
1 TECu  $\approx$  0.7 mm delay in K-band



2-h GIMs by JPL  
were used for the  
K-band ICRF3

# JPL ionosphere product

- JPL computes ionosphere models from GNSS data assuming 3 horizontal layers instead of a single one
  - Temporal resolution: 15 minutes
  - Spatial resolution: 330 base functions per layer instead of grid



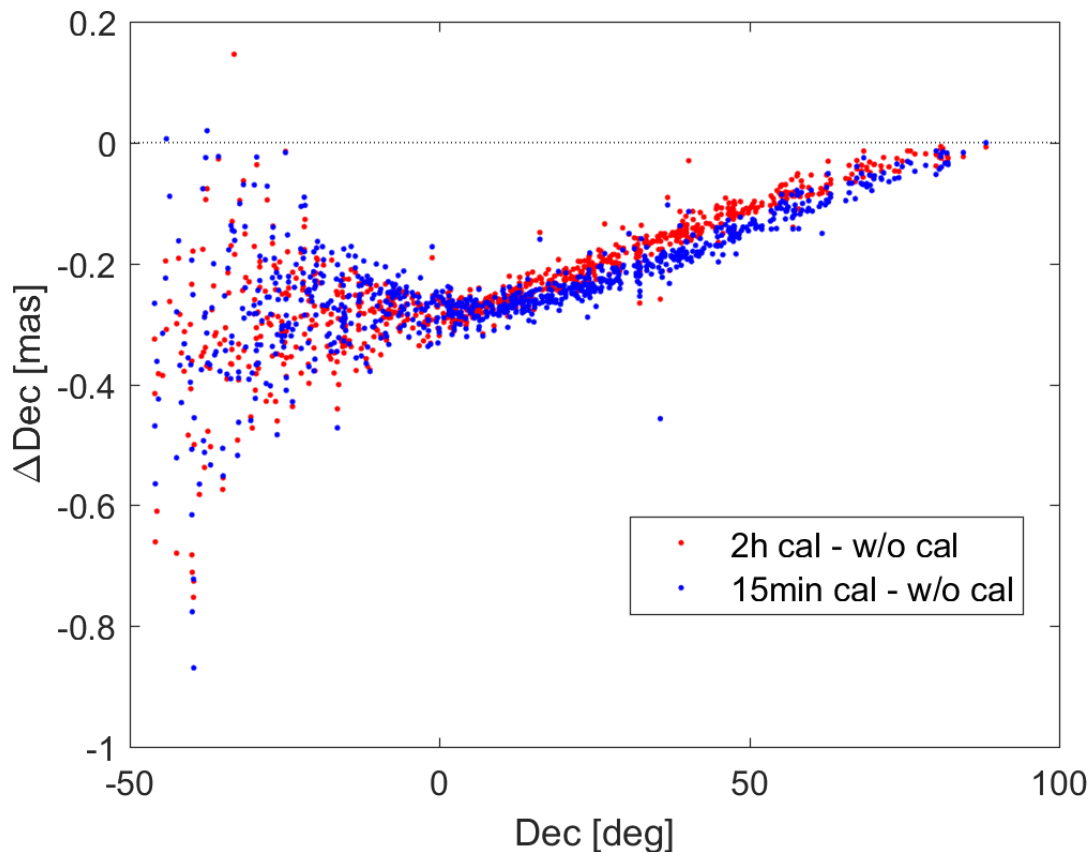
The JPL ionosphere product is more sophisticated than IGS GIMs

# K-CRF comparisons

- Preliminary results: VLBA sessions only (no HartRAO/Hobart)
- Comparison of K-CRF solutions
  1. without calibration
  2. 2h calibration
  3. 15min calibration

Difference w.r.t. w/o cal:

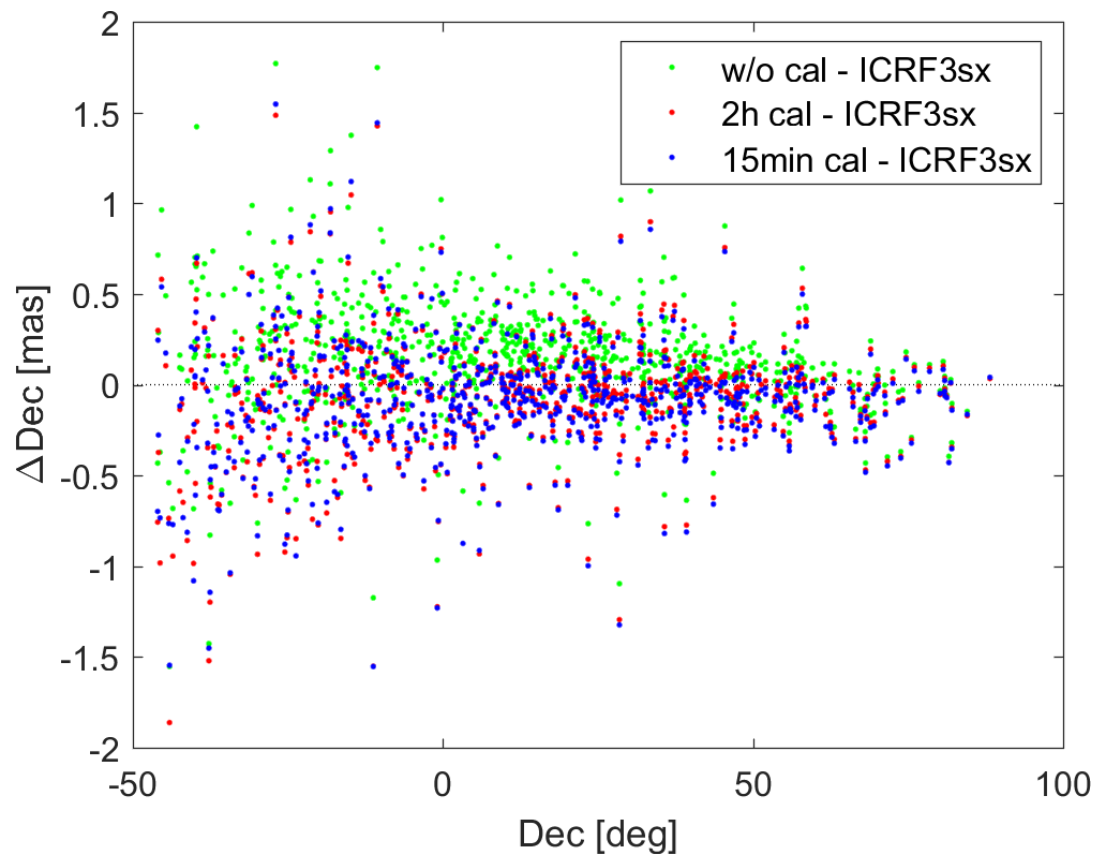
RMS [ $\mu$ as]	RA*cos Dec	Dec
2h cal	20	248
15min cal	24	252



# K-CRF comparisons vs. ICRF3 S/X

Difference w.r.t. ICRF3:

RMS [ $\mu$ as]	RA*cos Dec	Dec
w/o cal	220	360
2h cal	220	320
15min cal	218	316





# How to improve the ionospheric calibration?

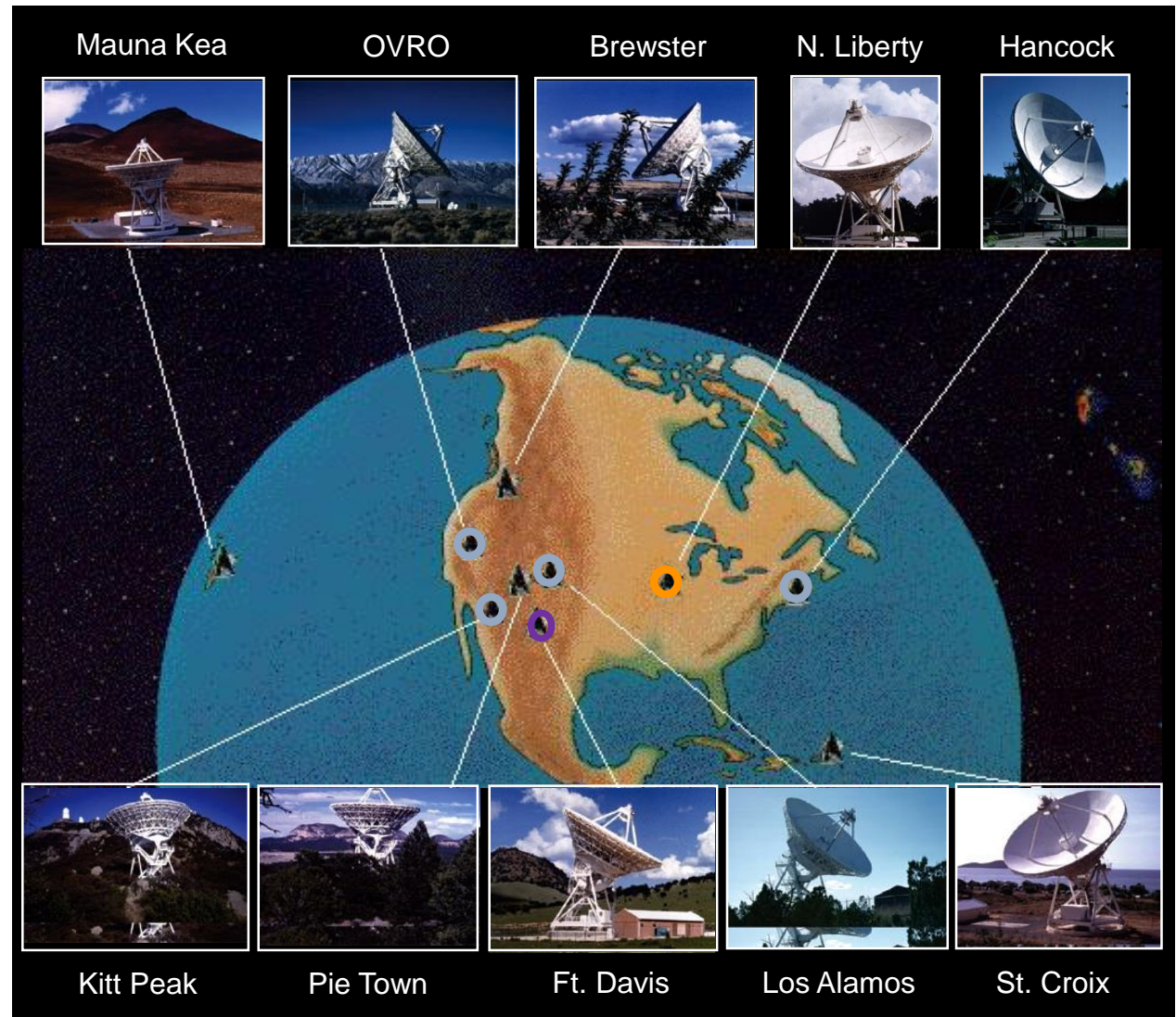
1. Improvements in ionosphere models

2. Improvements from better model coverage of VLBI stations

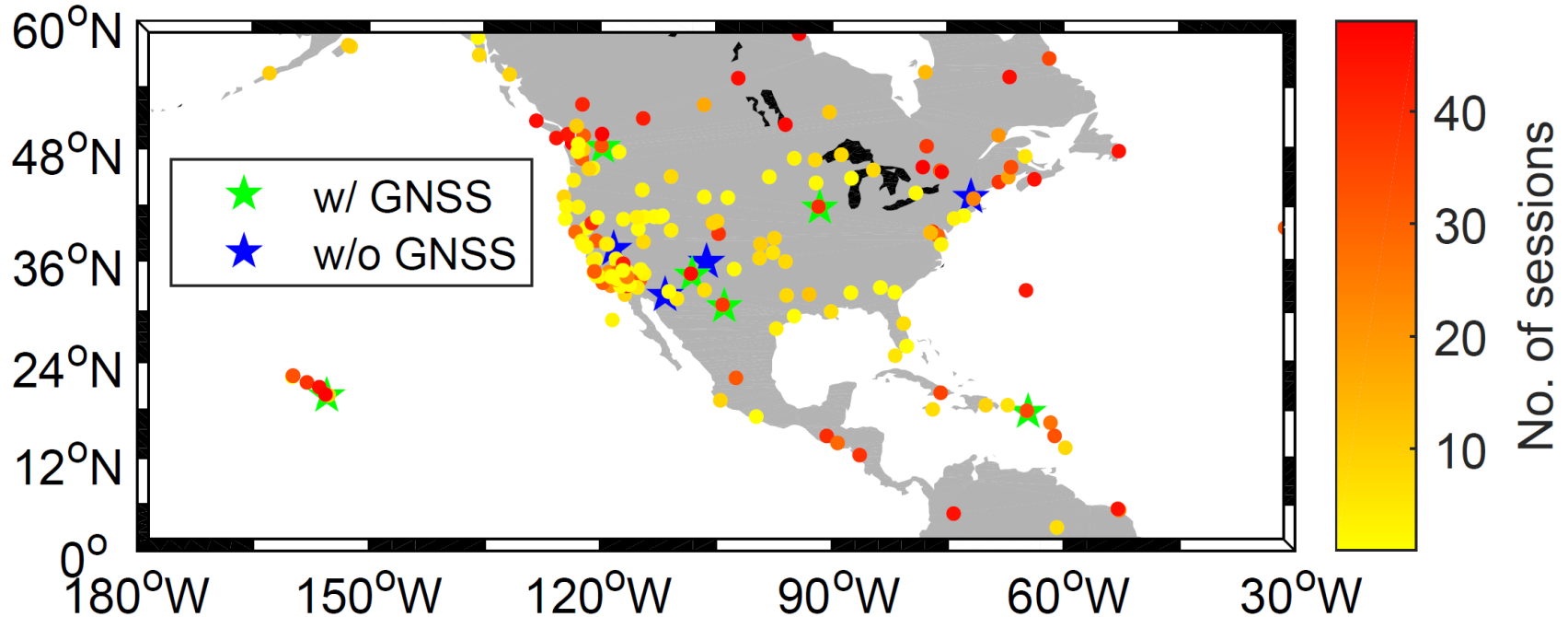
# GNSS antennas at VLBA sites

Very Large Baseline Array <http://www.vlba.nrao.edu/>

- Coverage gaps:
  - HN missing
  - OV missing
  - KP missing
  - LA missing
- NL **broken** since Nov 2018
- FD **offset 8 km**



# GNSS station contributions to IGS GIM

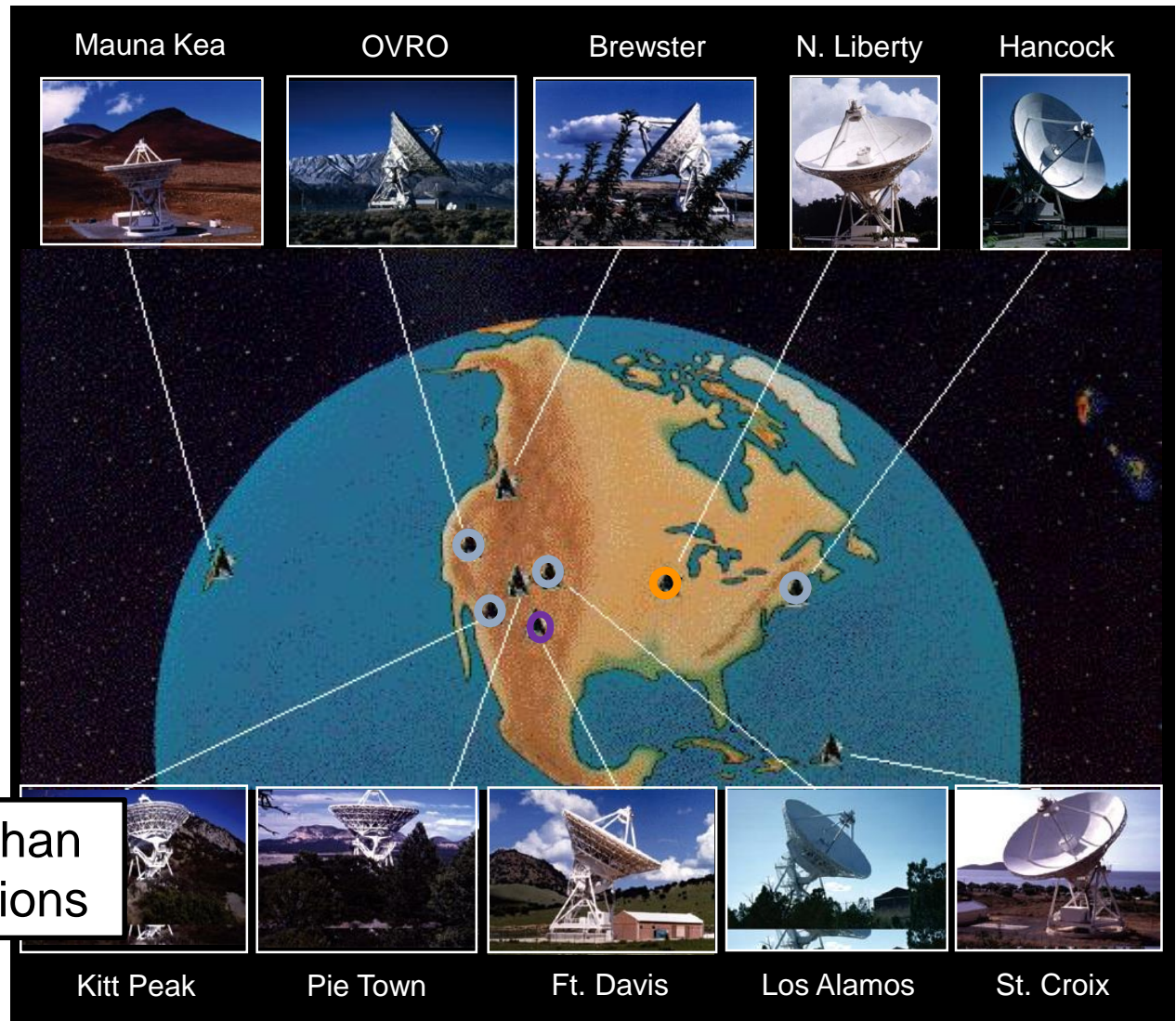


Colors indicate how often GNSS stations contributed to the IGS GIMs during VLBA sessions

# Distances to active\* IGS GIM stations

Very Large Baseline Array <http://www.vlba.nrao.edu/>

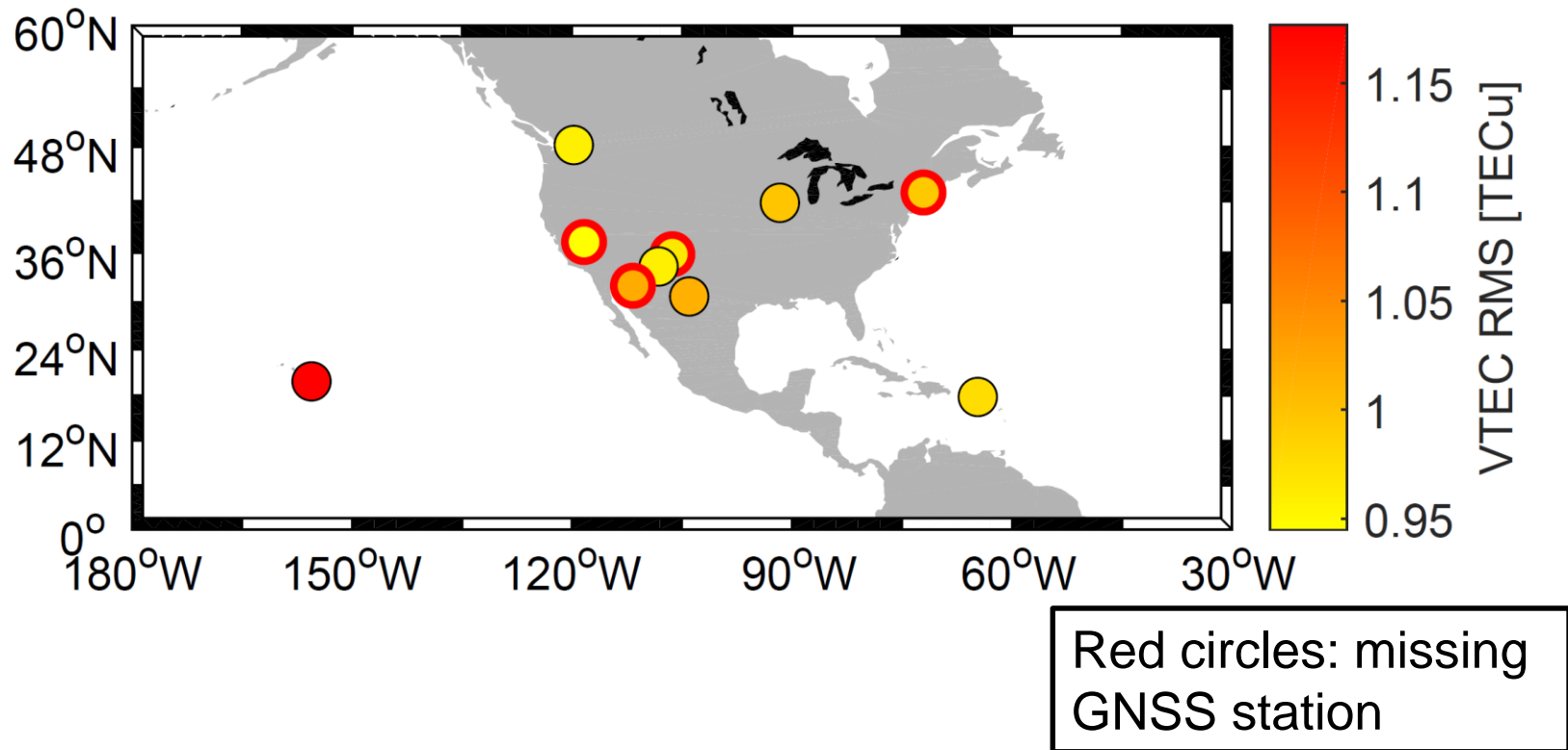
- HN 344 km
- OV 206 km
- KP 343 km
- LA 236 km
  
- NL 999 km  
since Nov 2018
- FD 8 km



\*contributed to more than half of the VLBA sessions

# Precision of ionospheric model for VLBA sites

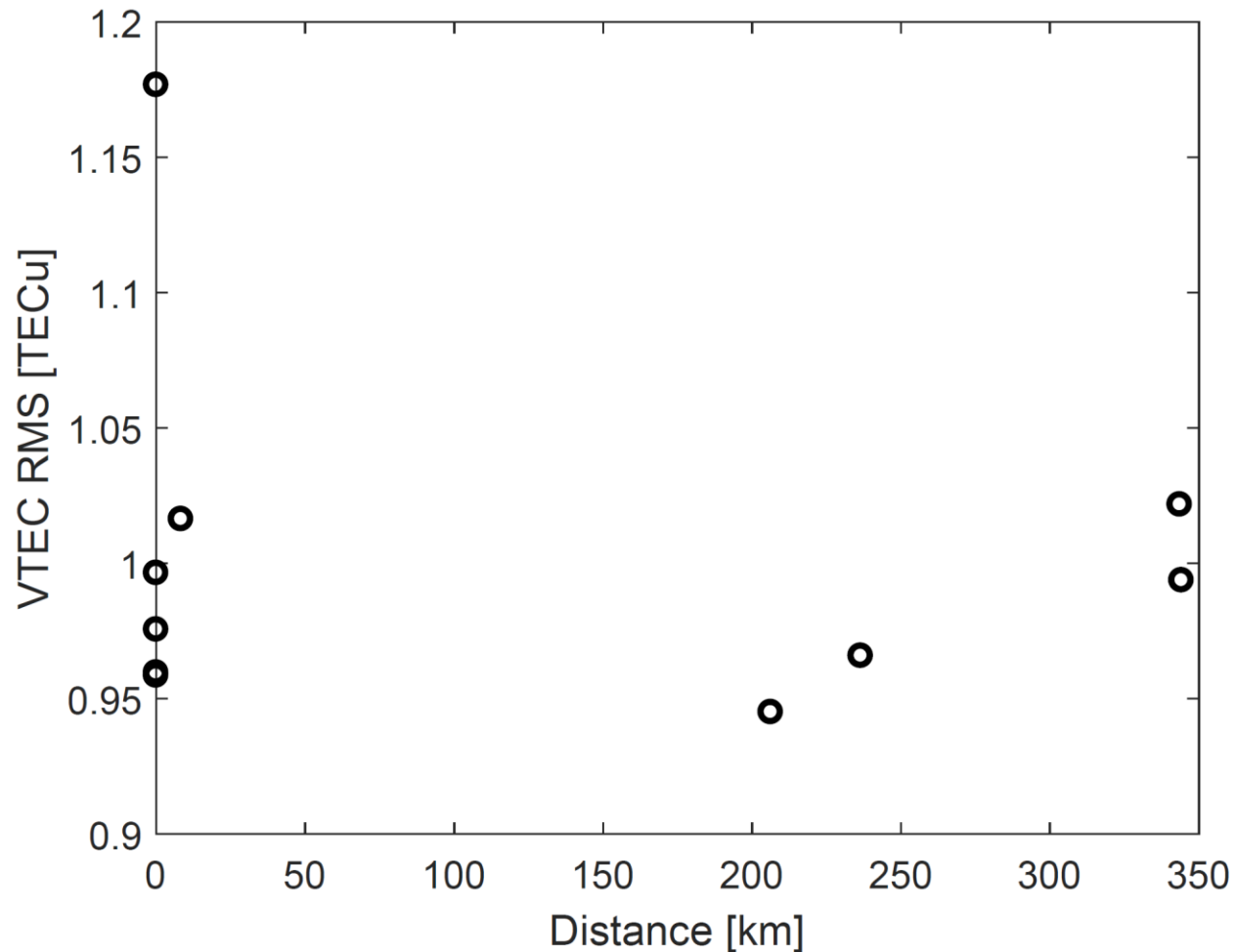
- IGS GIMs contain VTEC RMS as a measure of precision
- VTEC RMS averaged over K-band VLBA sessions



# Precision of ionospheric model for VLBA sites

Station name	GNSS on site	GNSS distance	VTEC RMS [TECu]	K-band delay RMS [mm]
HN-VLBA	0	344	0.99	0.72
LA-VLBA	0	237	0.97	0.70
NL-VLBA	0 / 1	0.068 / 999	1.00	0.73
OV-VLBA	0	206	0.94	0.69
PIETOWN	1	0.062	0.96	0.70
MK-VLBA	1	0.089	1.18	0.86
SC-VLBA	1	0.083	0.98	0.71
BR-VLBA	1	0.059	0.96	0.70
KP-VLBA	0	343	1.02	0.74
FD-VLBA	1	8.4	1.02	0.74

# VTEC RMS vs. GNSS station distance



# Conclusions

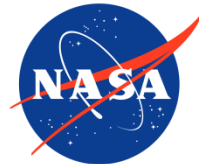
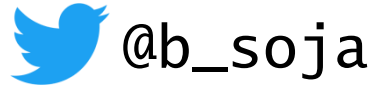
- **K-band** VLBI data has the potential to improve the CRF due to less source structure than S/X
- K-CRF accuracy depends on **ionospheric calibration**
- Implementation of ionospheric calibrations with **15 min** temporal resolution (vs. 2h previously – factor 8!)
  - Preliminary results: slightly better agreement with ICRF3 S/X
- Four VLBA sites **without GNSS stations** actively contributing to IGS GIMs
  - No significant degradation detected in the VTEC precision for these sites
  - Could be worse for N. Liberty (1000 km from active GNSS station)



# Thanks for your attention!



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