



Long Baseline Interferometry with the Square Kilometre Array

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Presentation Outline

- ▶ SKA – A description
 - Outline Specifications
- ▶ Long Baselines in the SKA
 - Science drivers
 - Delivery of signal transport networks
- ▶ SKA Programme development office
 - Organisation
 - Turning science requirements into instrument specifications

What is the SKA for?

Five Key Science Projects (KSPs)

2.Probing the Dark Ages

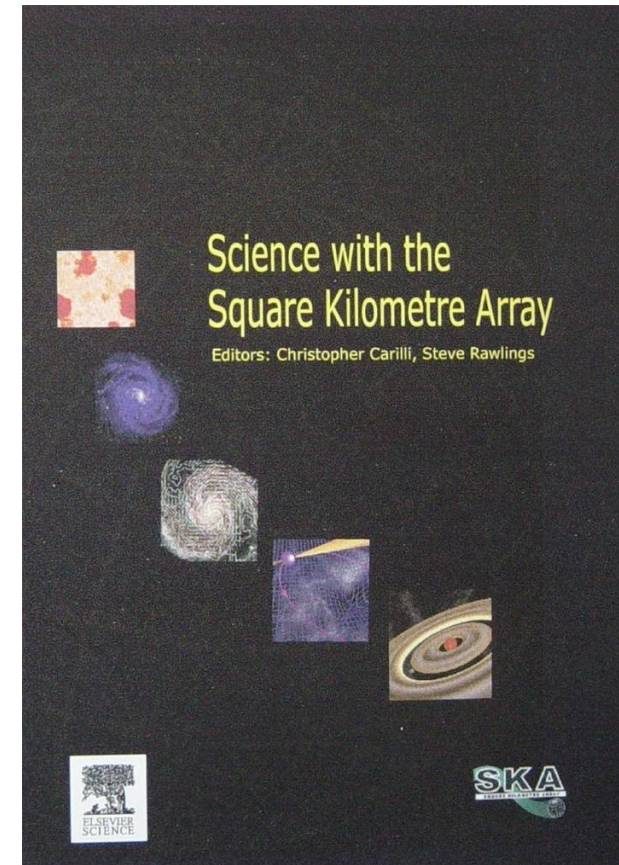
3.Galaxy Evolution, Cosmology, & Dark Energy

4.The Origin & Evolution of Cosmic Magnetism

5.Strong Field Tests of Gravity Using Pulsars and Black Holes

6.The Cradle of Life/Astrobiology

... plus **The Exploration of the Unknown** as an underlying philosophy for design of the instrument



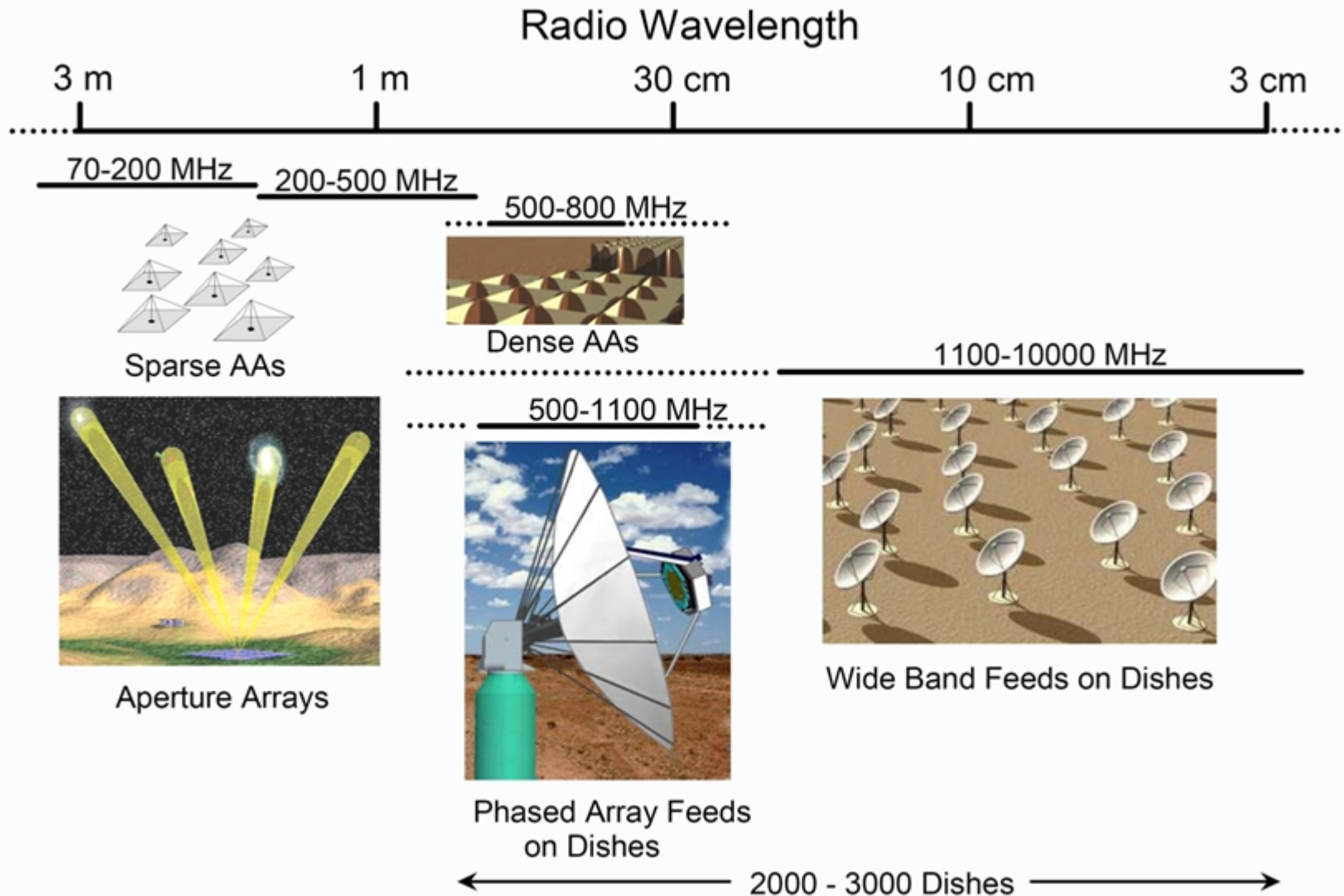


The Square Kilometre Array

4 prime characteristics

- **very large collecting area (km^2)** → sensitivity to detect and image hydrogen in the early universe
 - *sensitivity 40 x EVLA, 50 x LOFAR*
- **very-large-angle field of view** → fast surveying capability over the whole sky
 - *survey speed ~ 10000 x EVLA with FoV=1 sq. deg.*
- **wide frequency range required for the Science Reference Mission**
 - **low : 70-300 MHz**
 - **mid: 300 MHz-10 GHz**
 - **high: 10-25+ GHz**
- **large physical extent (3000+ km)** → capability for detailed imaging of compact objects and astrometry with milli-arcsec resolution

More Concise Picture of S



Numbers of dishes (2000-3000) depends on whether Phased Array Feeds and/or Aperture Arrays are used in the SKA.

Artist's impression of SKA central region

1000- 1500 dishes in the central 5 km; (50% of collecting area)

25% of collecting area at baselines > 180 km

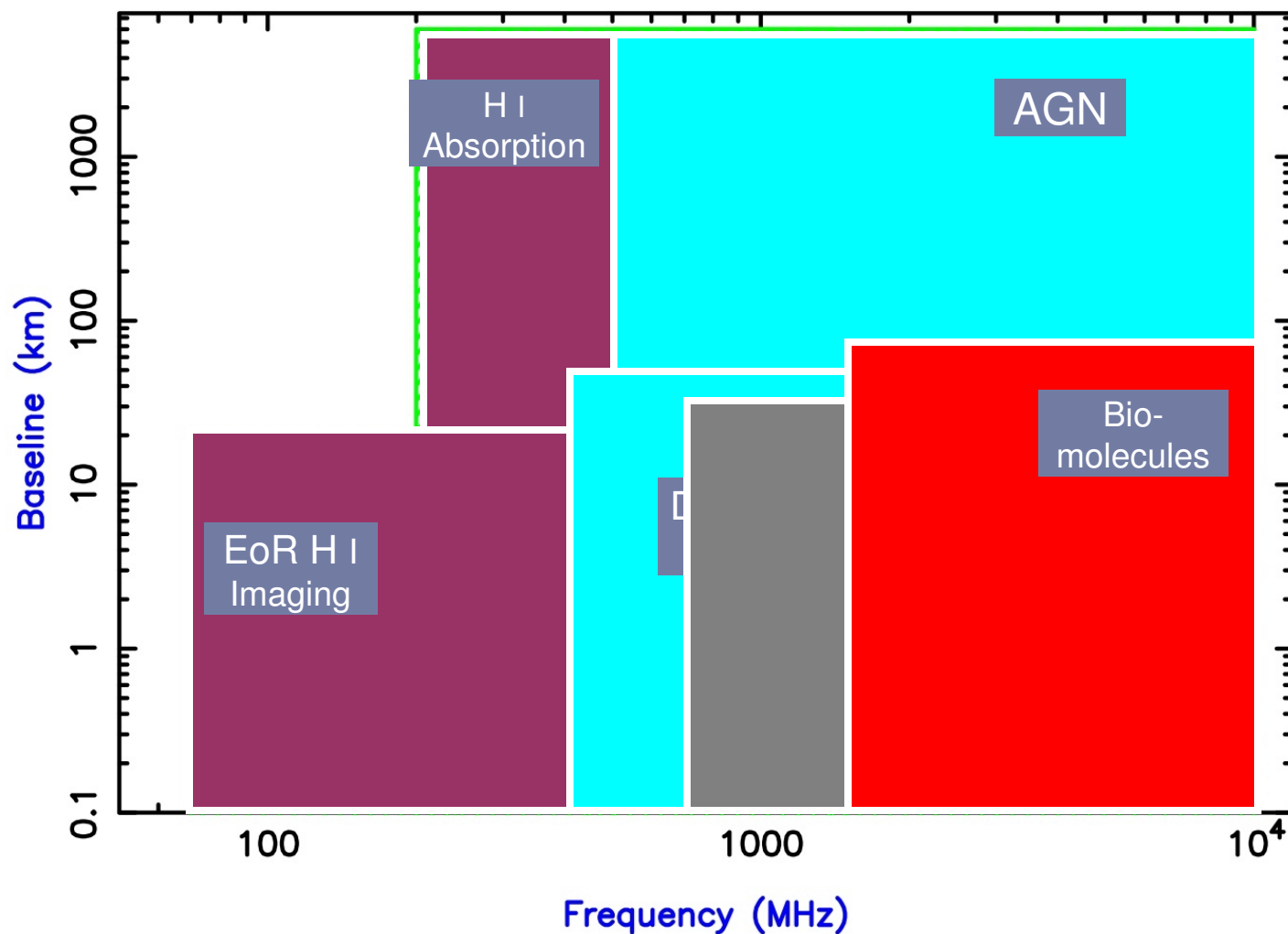
+ sparse aperture arrays

+ wide FoV technologies, dense aperture arrays and/or phased array feeds.

Located in Western Australia or South Africa

Baseline Distribution

Imaging Use Case Requirements



Long Baselines – Science Drivers

- ▶ Resolving Active Galactic Nuclei (AGN) and Star Formation in Galaxies
 - Specific goal of tracking and distinguishing the contribution of star formation and AGN to the evolution of galaxies.
 - Maximum baseline $> 3,000$ km
for Flux densities of $> 30\mu\text{J}$; Brightness temperatures of 10^6 K ; Redshifts 2 to 7
 - Frequency 0.5-8 GHz

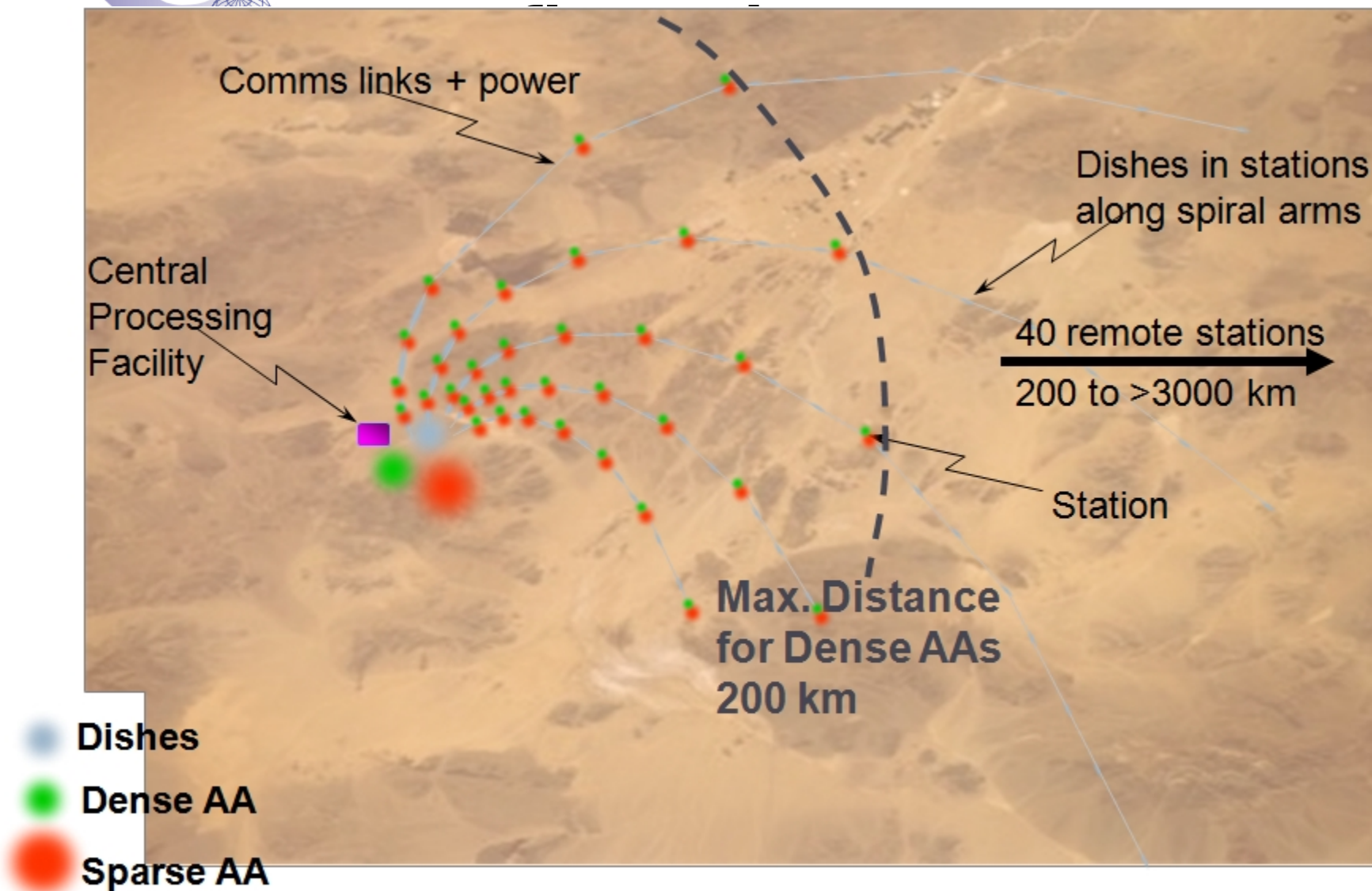
Long Baselines – Science Drivers

► H1 Absorption

- Probe the circum-nuclear environment of Active Galactic Nuclei.
- Maximum baseline $> 3,000$ km

Resolutions of the order of sub kpc to probe the innermost regions of galactic nuclei at high red-shift

Possible telescope



Adapted from A. Faulkner

SKA Signal Transport Infrastructure

The Tour de France is a 3,000 km race



SKA antenna
will be spread
along 3,000
km arms



Signal Transport & Networks for the SKA

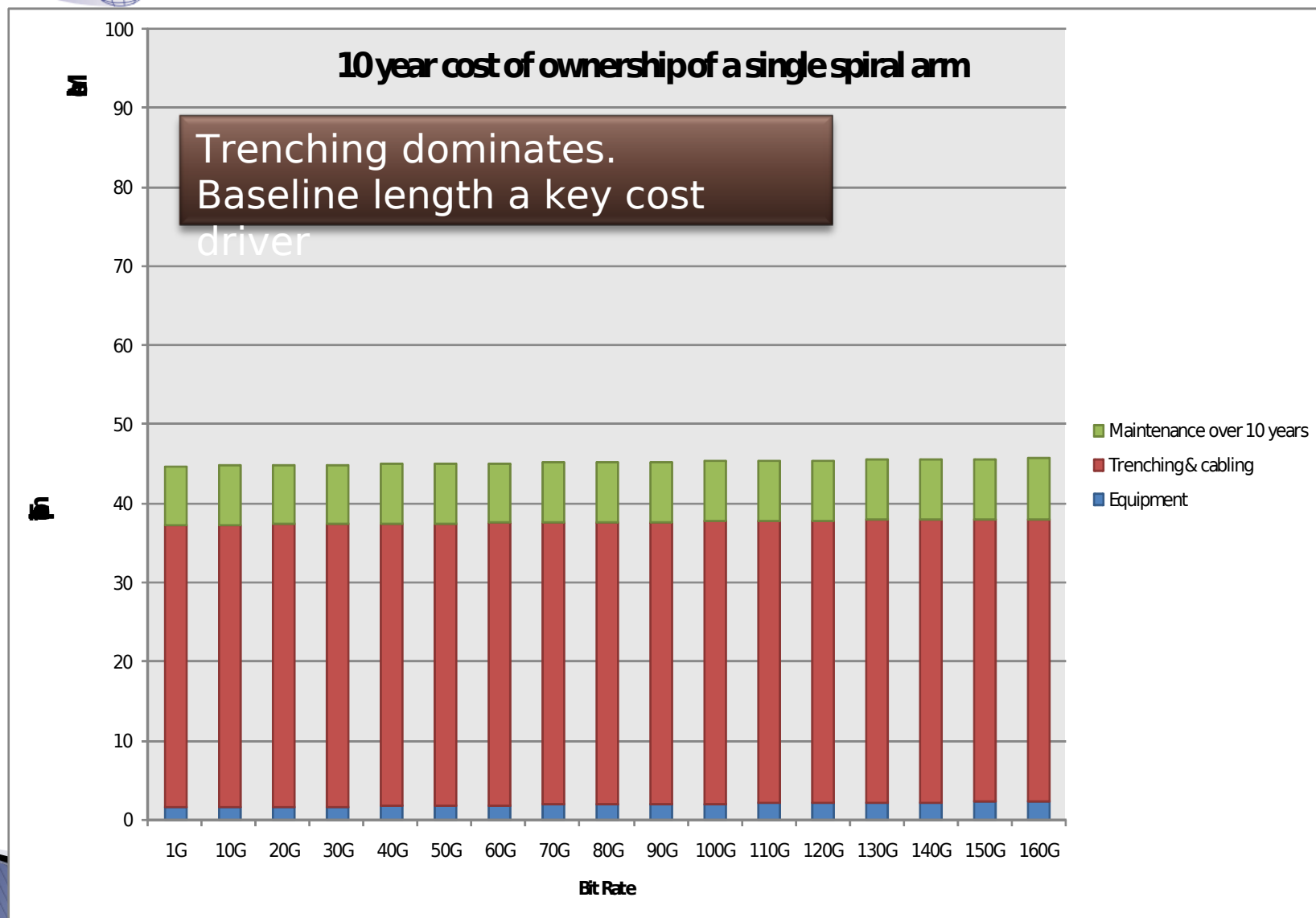
- ▶ Data Network
 - For transporting astronomical signals to a central processing facility (CPF)
- ▶ Timing Network
 - For the distribution of local oscillator signals for clocks and down converters.
- ▶ A Monitor & Control Network (M&C)
- ▶ Connections from the CPF to the outside world



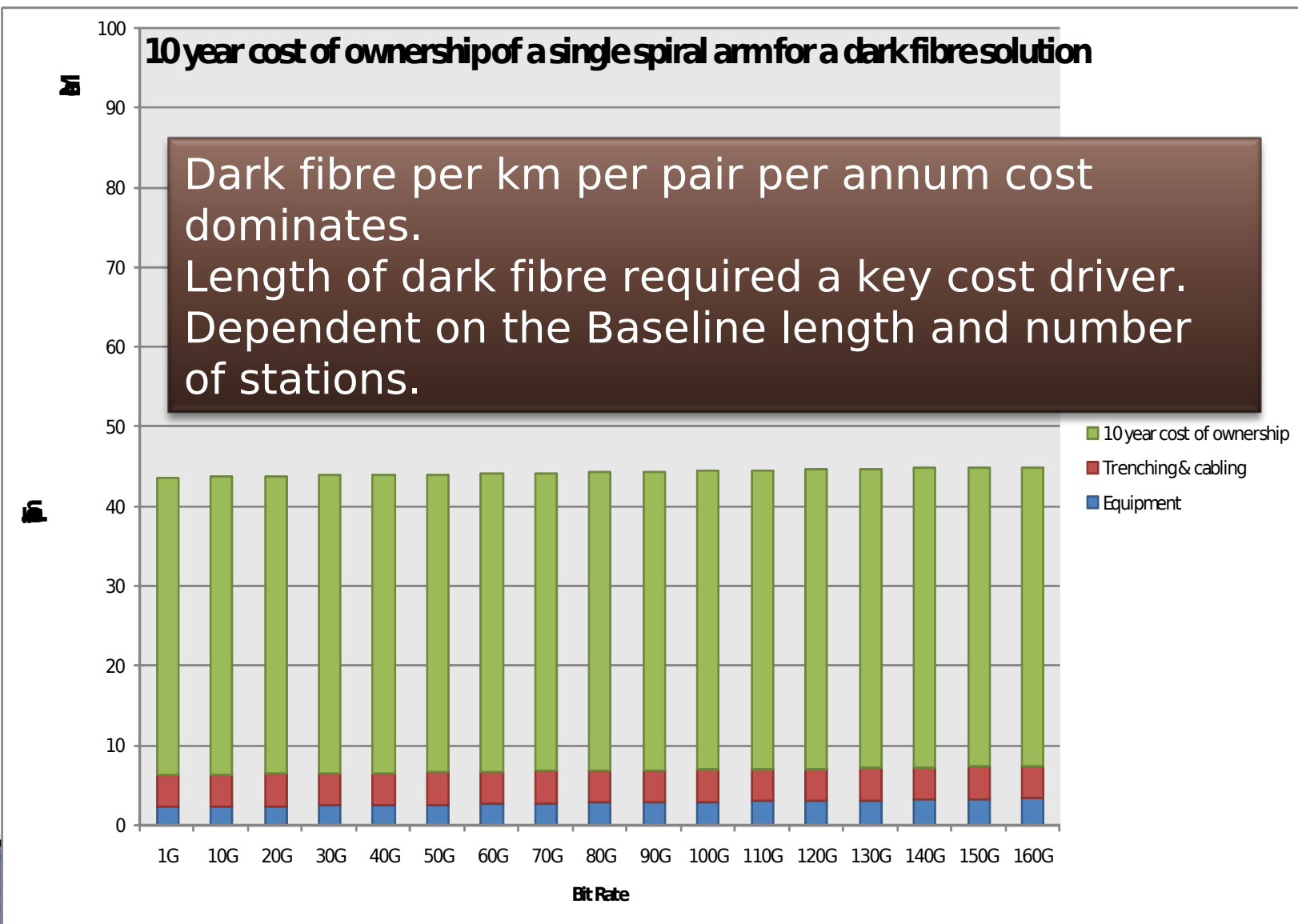
Implementation strategies for Signal Transport on long baselines

- ▶ Construction of a private fibre network along the baseline lengths
- ▶ Dark fibre network provision to long baseline stations
 - Assuming capacity in the locations required
- ▶ Bandwidth provision over commercial networks to long baseline stations
 - Assuming capacity in the locations required

Network Construction

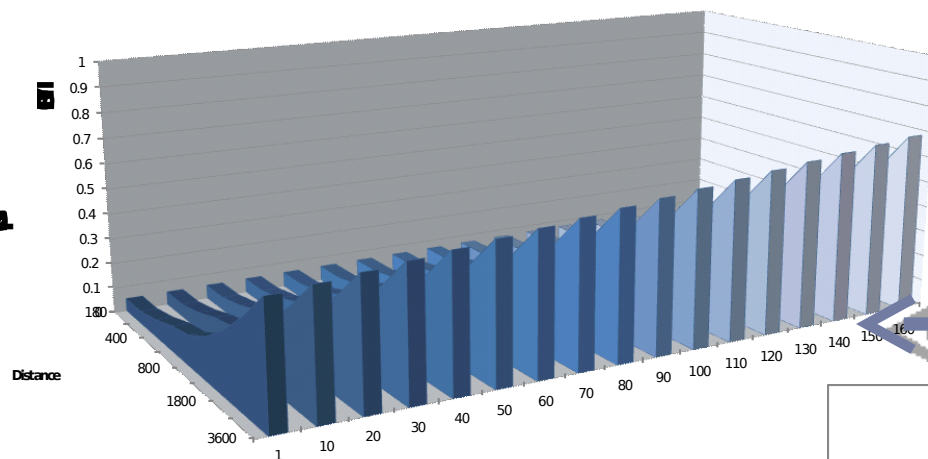


Dark fibre



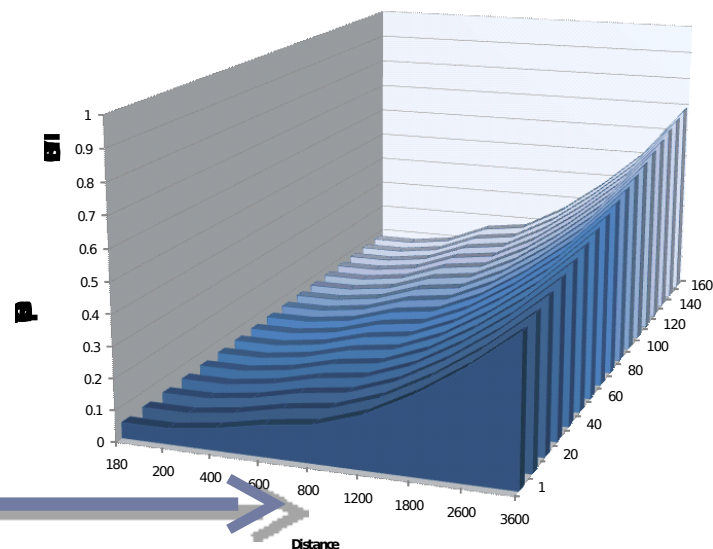
Increased bandwidth
not a significant cost
driver. Provide
multiple beams at
long baselines for a
small incremental

Estimated cost per station of data transmission,
for stations at distances > 180 km



Distance main cost driver
for transmission.

Estimated cost per station of data transmission,
for stations at distances > 180 km



Data Transport

Very Large Bit Rates:

R = Total Bit Rate

B = Bandwidth;

A_{tot} = Total Collecting Area

Ω = Field of View

λ = wavelength

$$R \propto \frac{BA_{\text{tot}}\Omega}{\lambda^2}$$

In summary this leads to bit rates of

- ▶ 160 Gbps per WBSPP dish
- ▶ 20 Tbps per Aperture Array Station
- ▶ 420 Gbps per Phased Array Feed



Commercial Bandwidth

- ▶ For 40 stations on long baselines, the total bandwidth requirements are:
 - 6.4 Tbps for an 8GHz bandwidth
 - 1.6 Tbps for a 2 GHz bandwidth
 - 0.4 Tbps for a 500 MHz bandwidth
- ▶ SKA Bandwidth requirements are static, whilst IP traffic and networks are growing.



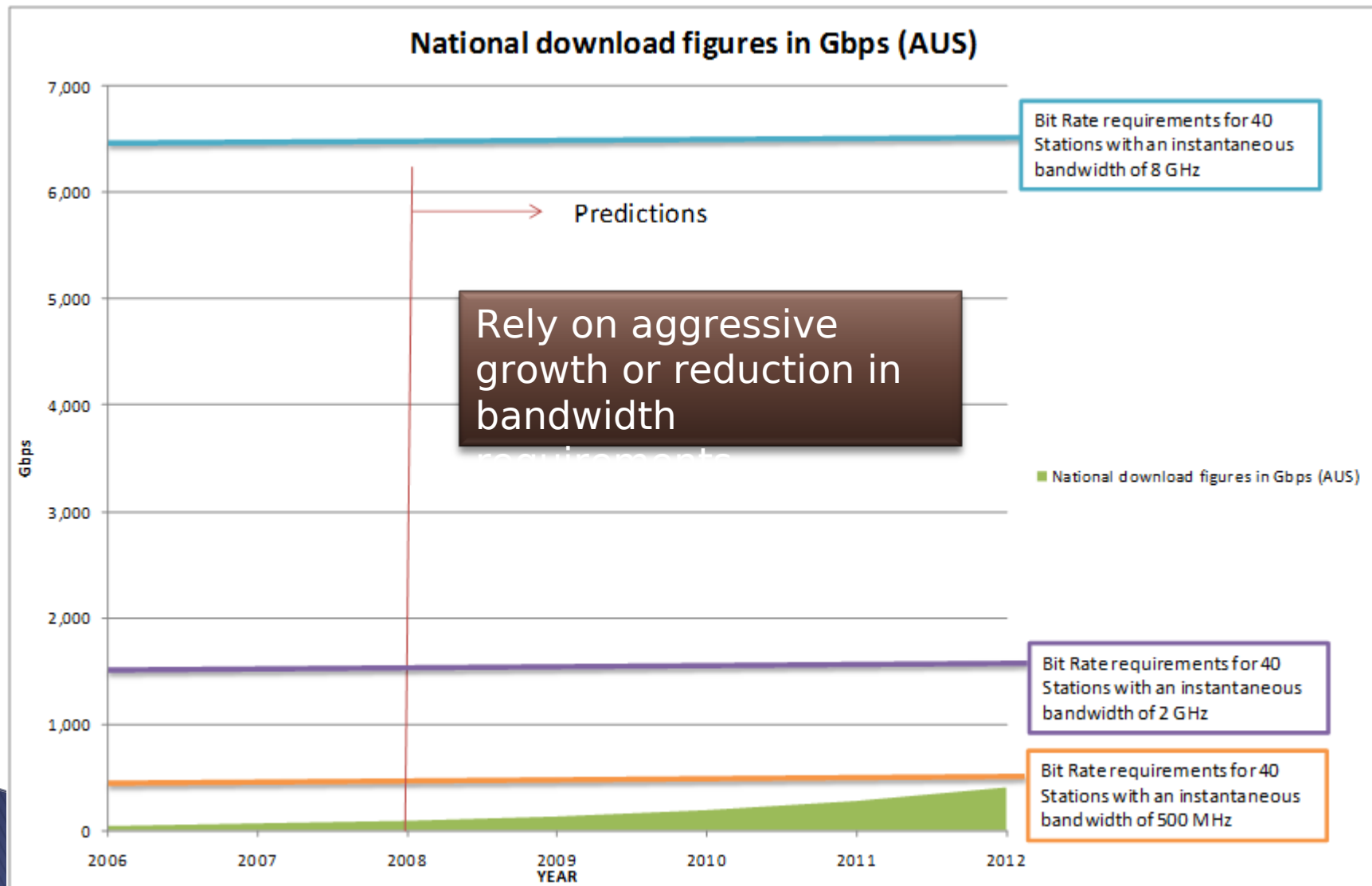
Putting SKA Bandwidths into perspective....

- ▶ Cisco* estimated that in 2007 the total, global IP traffic was the equivalent of **22 Tbps**.
- ▶ Cisco predict a **46% CAGR** of IP traffic upto 2012
- ▶ John Stankey** of AT&T showed AT&T's IP traffic at **1 Tbps** in 2007 (or 5% of the global total).
- ▶ Australian Bureau of statistics show national download figures for 2007 are the equivalent of **65 Gbps** (or < 1% of the global total)
- ▶ No equivalent statistics are published for South

*Cisco White Paper, Visual Networking Index – Forecast and Methodology, 2007–2012

**Presentation by John Stankey, Group President-Telecom Operations, AT&T Inc., delivered to 2007 Analyst Conference, December 11, 2007

SKA Bandwidths compared to AUS download traffic



Cost Drivers

Design Solution	Dominant Cost drivers	Key Cost assumptions used
Station solution	<ul style="list-style-type: none"> • Trenching cost/km • Length of trenching <ul style="list-style-type: none"> ◦ Baseline 	€ 10,000 per km
Dark Fibre solution	<ul style="list-style-type: none"> • Cost of the dark fibre/annum • Length of dark fibre required which depends upon: <ul style="list-style-type: none"> ◦ Baseline 	€ 500 per fibre pair per annum
Bandwidth solution	<ul style="list-style-type: none"> • Cost per lambda or Gbps /annum <ul style="list-style-type: none"> ◦ Number of Stations • Bandwidth Requirements <ul style="list-style-type: none"> ◦ Number of Stations 	unknown



SPDO Team

Project Director
Project Engineer
Executive Officer
Project Scientist
System Engineer
Domain Specialist Receptors
Domain Specialist Signal Transport
Domain Specialist Computing & Software
Domain Specialist Signal Processing
Site Engineer
Project Management Officer
Industry Relations Manager
Office Manager

Richard Schilizzi
Peter Dewdney
Colin Greenwood
Joe Lazio
Kobus Cloete
Neil Roddis
Roshene McCool
Duncan Hall
Wallace Turner
Rob Millenaar
Billy Adams
Phil Crosby
Lisa Bell

www.skatelescope.org

Goals

“A costed system design and deployment plan for the SKA”

PrepSKA WP2 Work

Flow

Science
Requirements

DRM Case Studies

DRM =
Design Reference
Mission

Eng.
Simulations
Prototypes

Pathfinders

Technical R&D

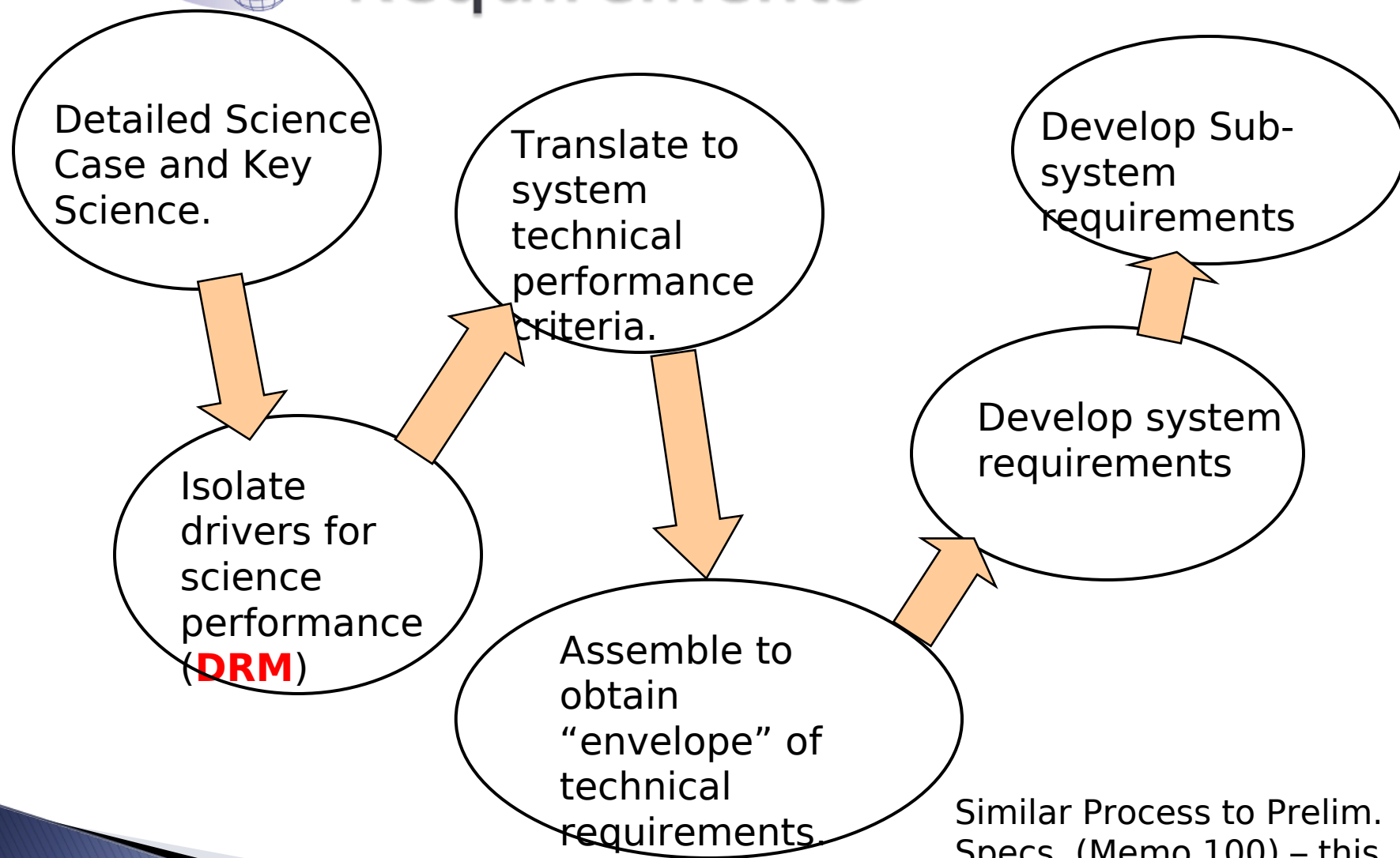
SKACost

Four
Years

Readiness Assessments

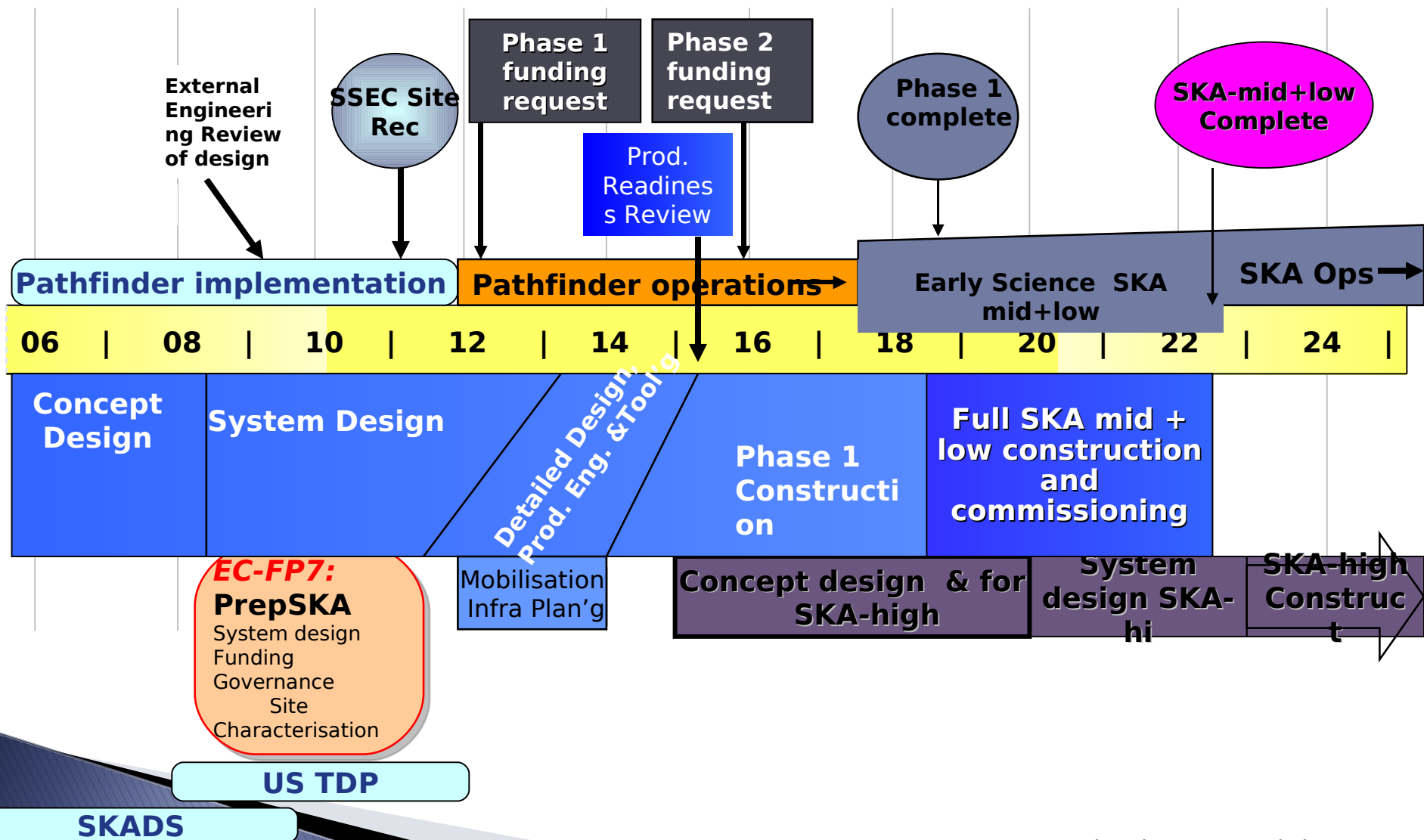
Engineering
Design & Cost

Developing System Requirements



Similar Process to Prelim. Specs. (Memo 100) – this time leading to “design action”.

SKA timeline



Questions

