

Progress and Status of e-MERLIN

Ralph Spencer
on behalf of the e-MERLIN team

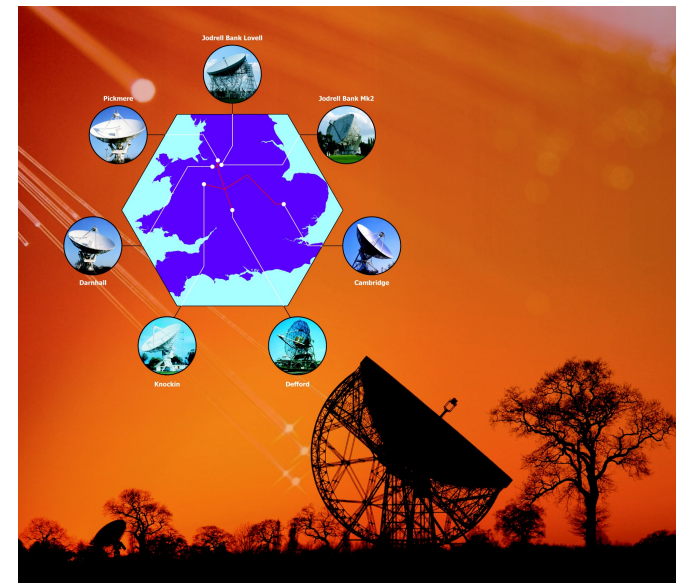
Madrid June 2009



e-MERLIN

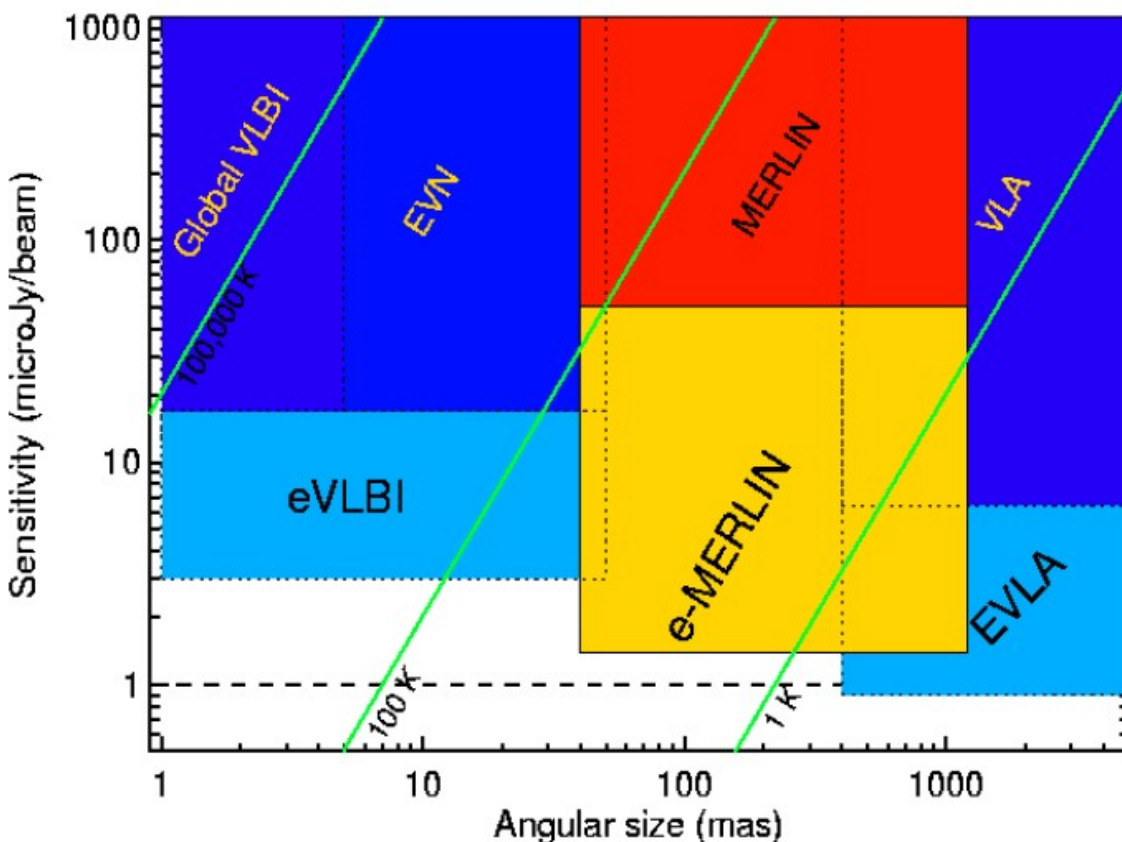
- **Major upgrade to MERLIN array**
- **New optical fibre network connecting 7 telescopes to new correlator at Jodrell Bank Observatory**
- **Unique combination of 50-150 mas resolution and μJy sensitivity**
- **Crucial to resolve AU-scale star-formation processes in Galaxy and kpc-scale regions at $z \sim 1$**

New Receivers
New IF system
New Links
New Science!



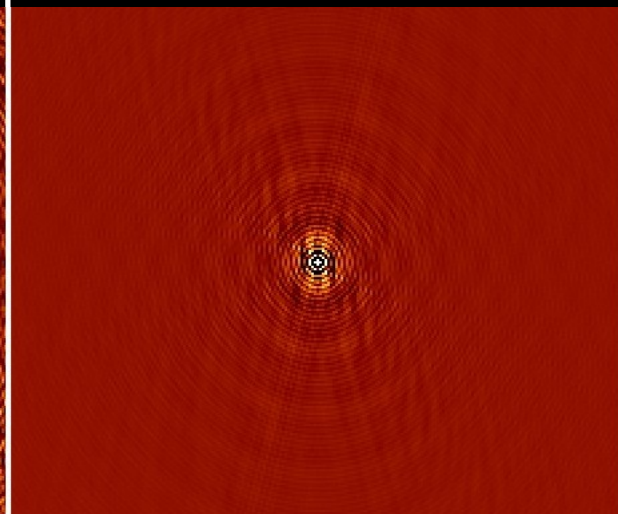
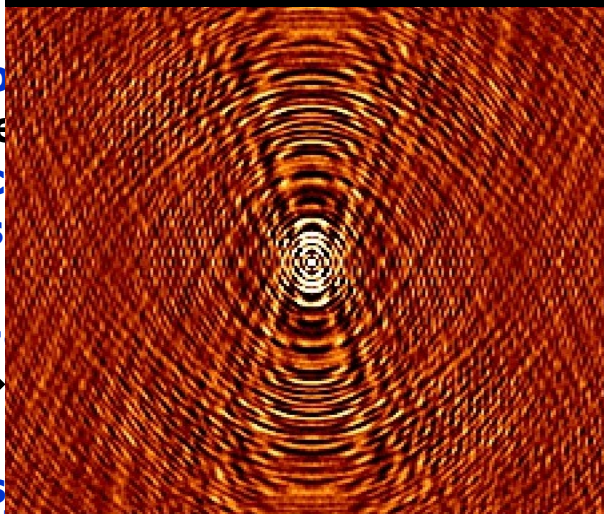
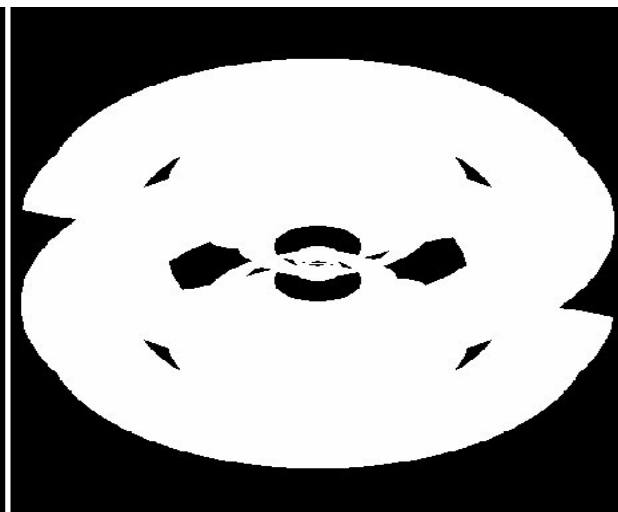
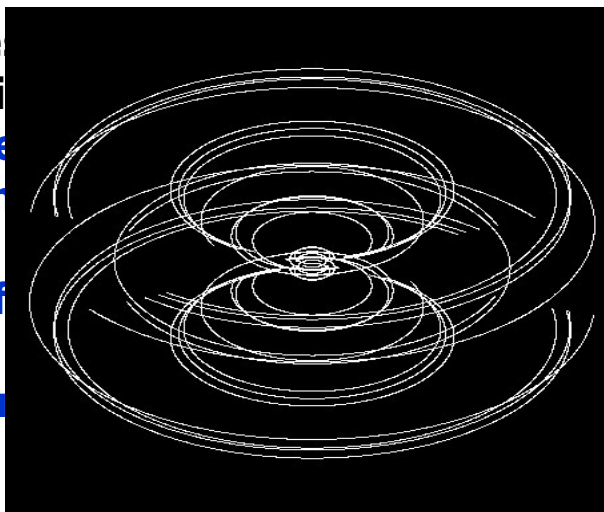
Capabilities

- **150, 40, 10 mas resolution**
- **$\sim 2 \mu\text{Jy}$ sensitivity in 100,000 K**
 - **$< \mu\text{Jy}$ in deep fields**
 - **$\sim 30 \mu\text{Jy}$ in ~ 1 m**
- **Wide fields**
 - **Out to HPBW of 1000 mas**
- **Spectroscopy**
 - **16 placeable subchannels**
 - **Can mix/trade polarisations**
 - **Correlator almost always on**
- **Much improved aperture**
 - **Via frequency comb**
 - **May help snapshots**
- **Spectral mapping**
 - **1.3-1.7; 5-7/4-8 GHz**
- **Polarization ($L, R \rightarrow IQUV$)**
- **Astrometry**
 - **Goal is < 1 mas wrt ICRF: using GPS measurements of troposphere delay (5cm error \rightarrow 5mm); closer calibrators**



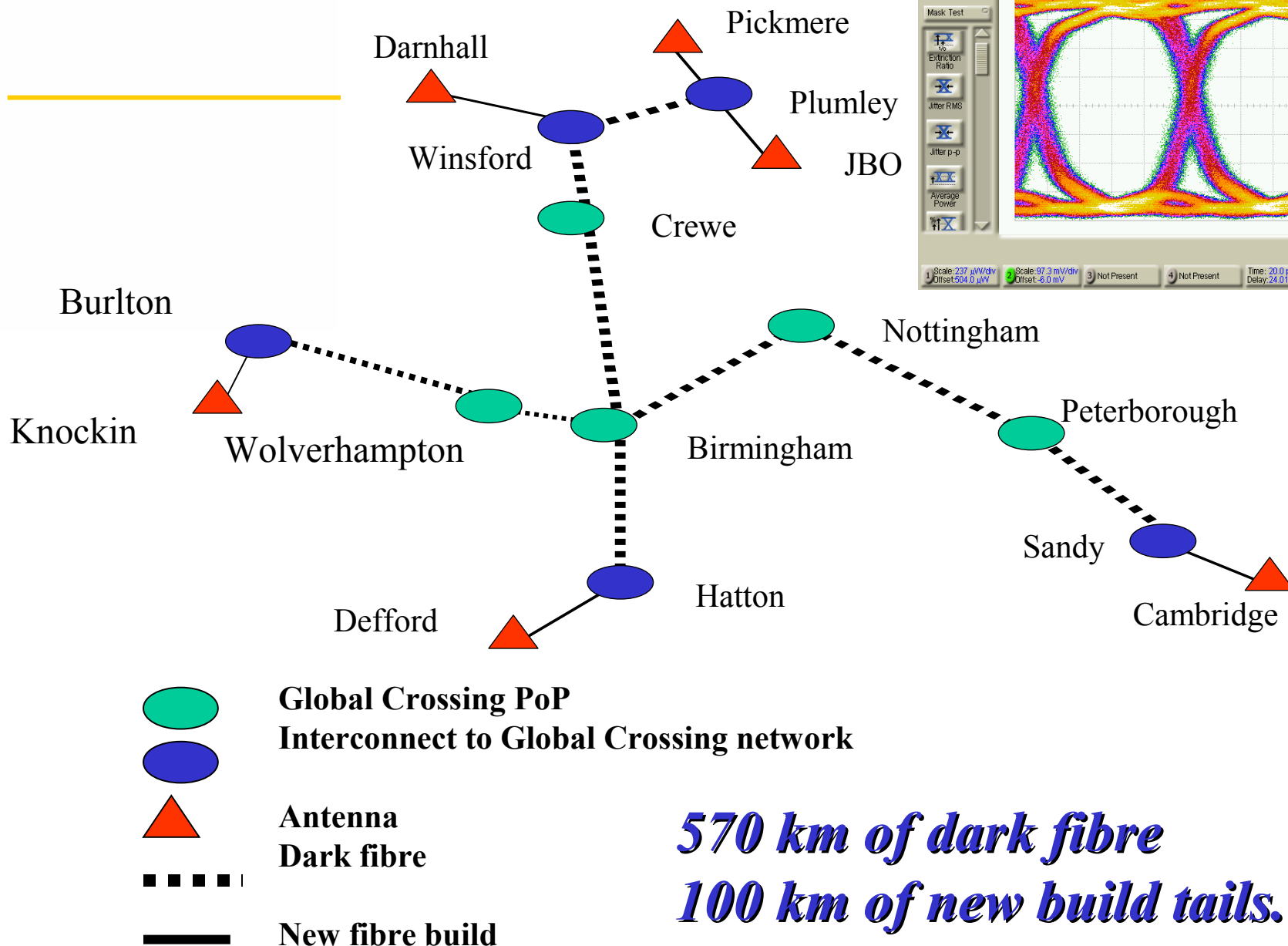
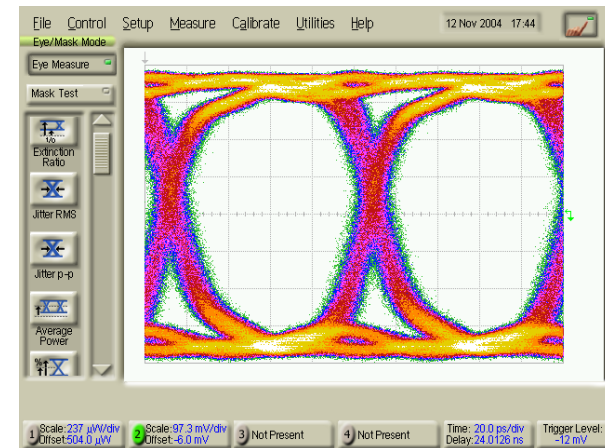
Capabilities

- 150, 40, 10 mas resolution
 - $\sim 2 \mu\text{Jy}$ sensitivity in 100 MHz
 - $< \mu\text{Jy}$ in deep fields
 - $\sim 30 \mu\text{Jy}$ in ~ 1 m
 - Wide fields
 - Out to HPBW of 1000 mas
 - Spectroscopy
 - 16 placeable sub-beams for frequency recirculation
 - Can mix/trade polarisations
 - Correlator almost always on
 - Much improved aperture
 - Via frequency comb
 - May help snapshots
 - Spectral mapping
 - 1.3-1.7; 5-7/4-8
 - Polarization (L,R \rightarrow V, H)
 - Astrometry
 - Goal is < 1 mas
- troposphere delay (5cm error \rightarrow 5mm); closer calibrators





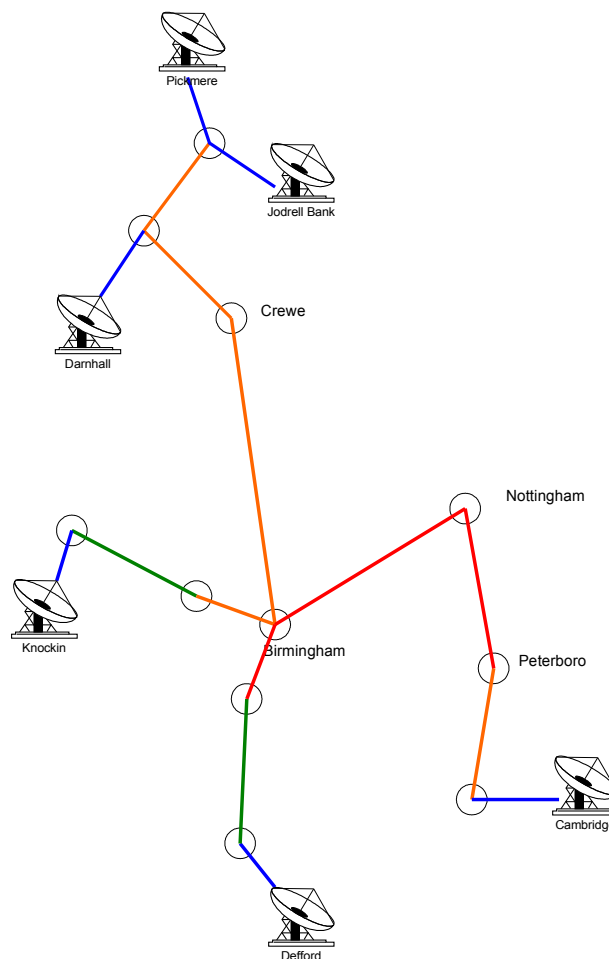
Schematic diagram of the e-MERLIN fibre network

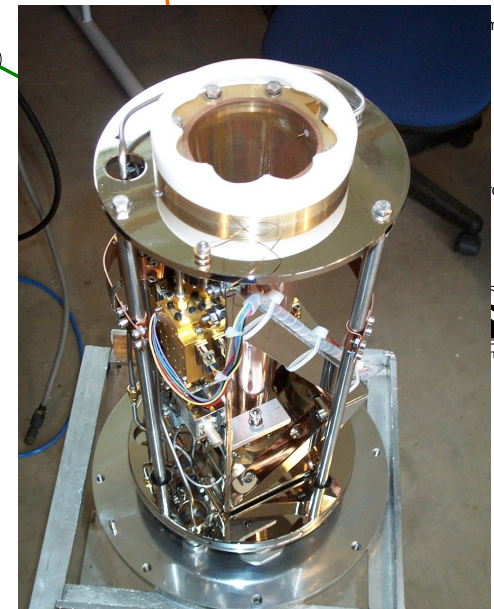
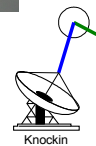
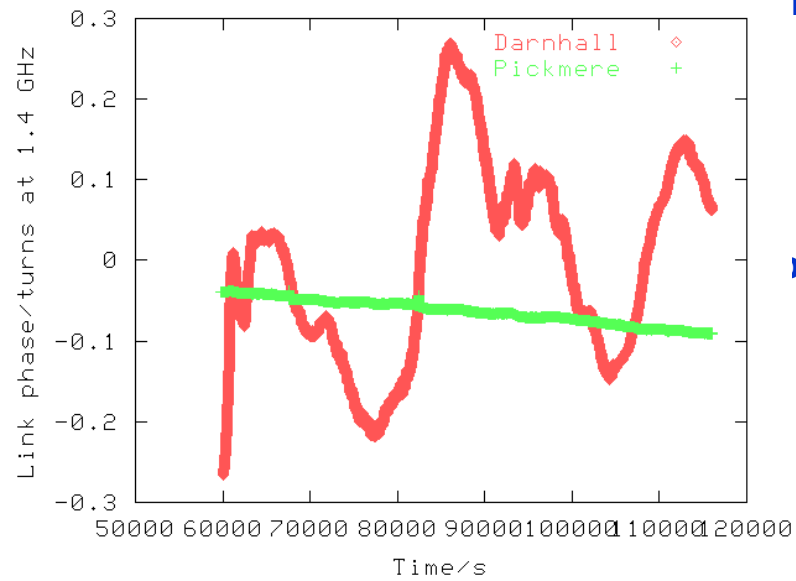


570 km of dark fibre
100 km of new build tails.

Commissioning

- **New receivers/IF installed**
 - wider bandwidth, improved Tsys, quick band-change
- **Samplers & data transmission being installed (2 done)**
 - 0.5 GHz (covers L-band) done, 2 GHz (NRAO design) later this year
- **Optical fibre network in place**
- **New timing system developed and being installed (2 done)**
 - H-maser signal transferred on fibre + GPS at each telescope
- **Prototype correlator (DRAO, Penticton) arrived over last few months**
 - 2 station boards + 1 correlator board + Xbar
 - up and running
- **New array control software**
 - up and running

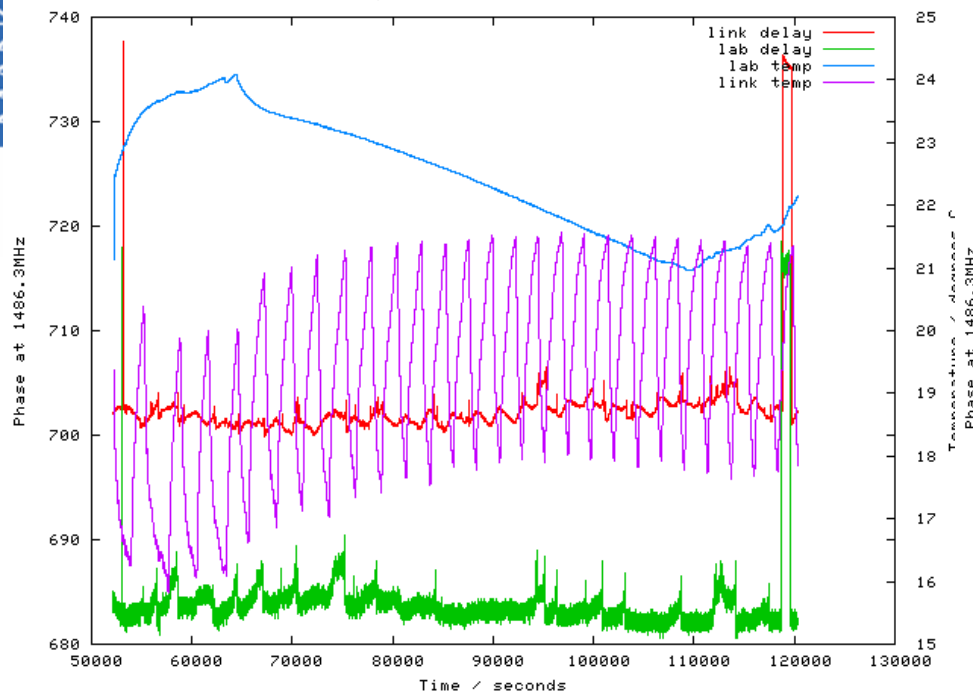




Results

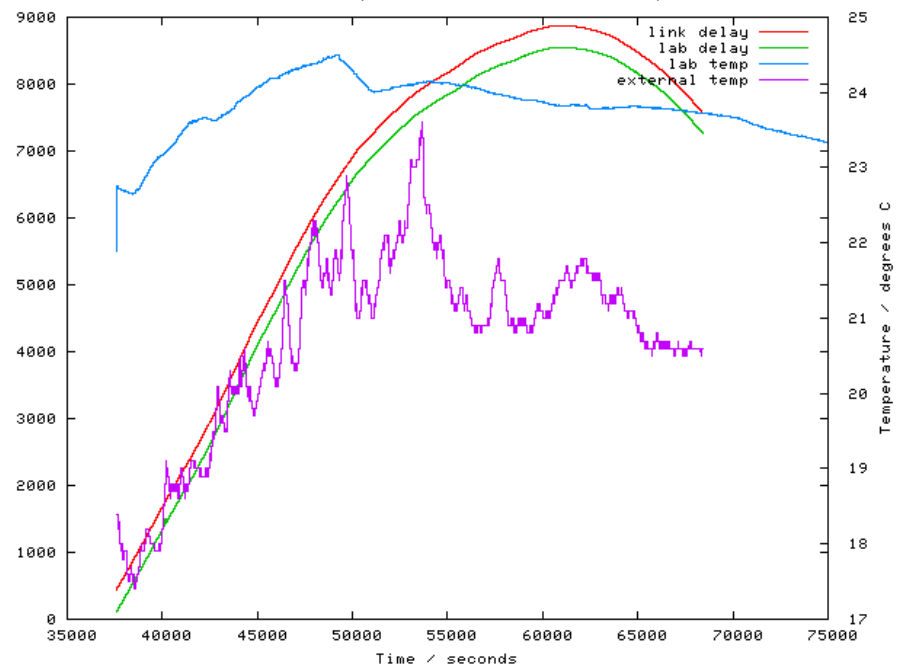
$\Phi_{\text{one way}}$ & $\Phi_{\text{round trip}/2}$

Phase and temperature for back-to-back, no fibre



B2b

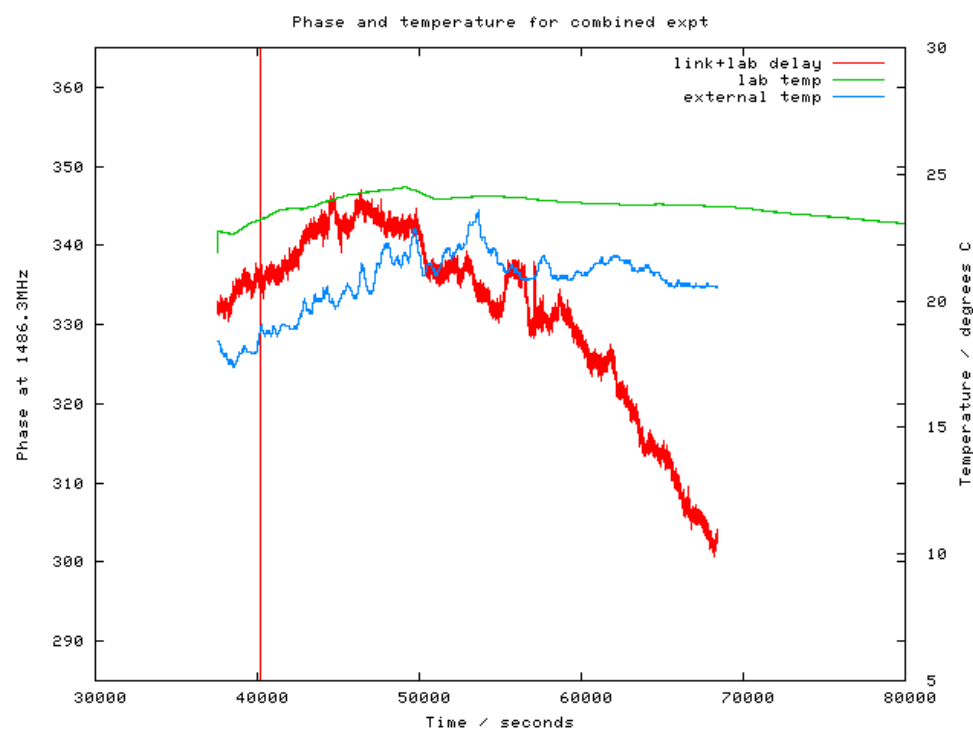
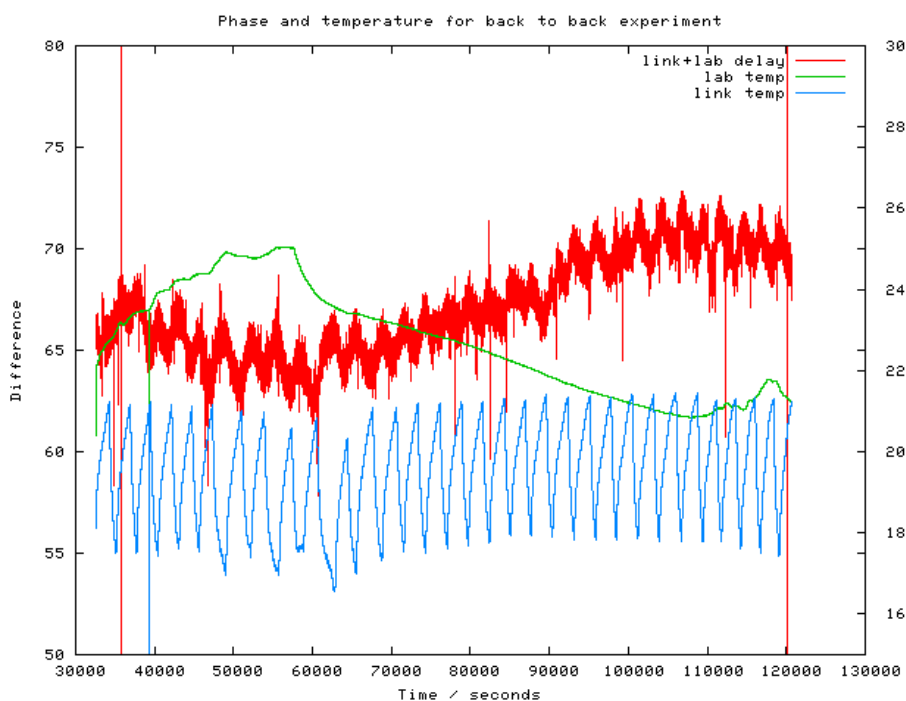
Phase and temperature for 110.8km roundtrip



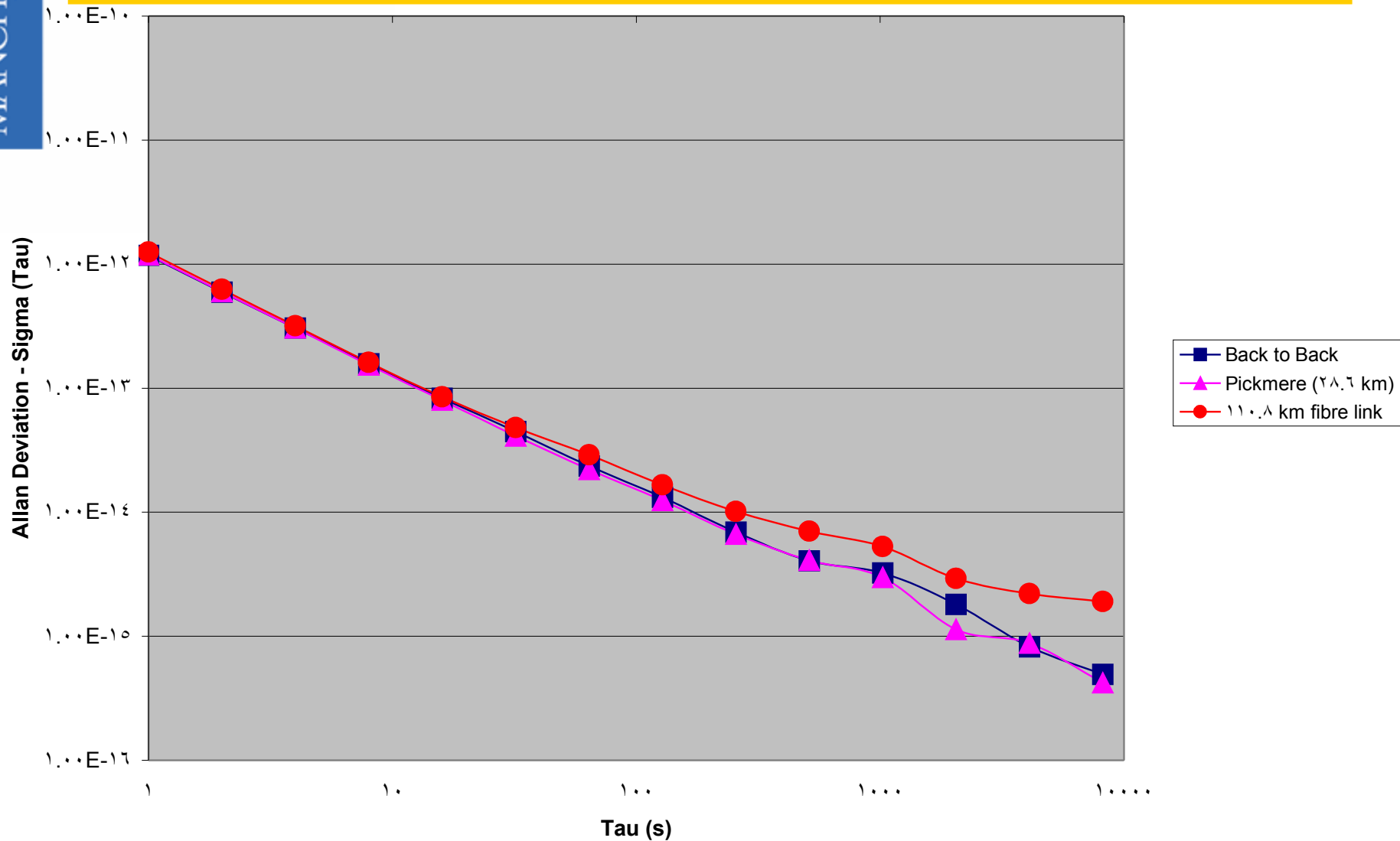
110 km

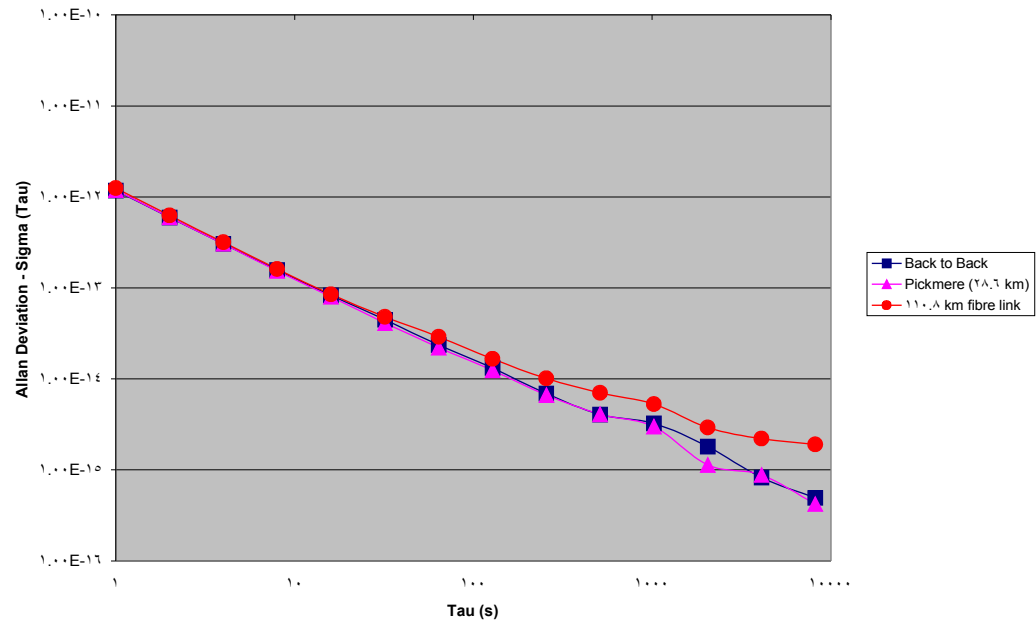
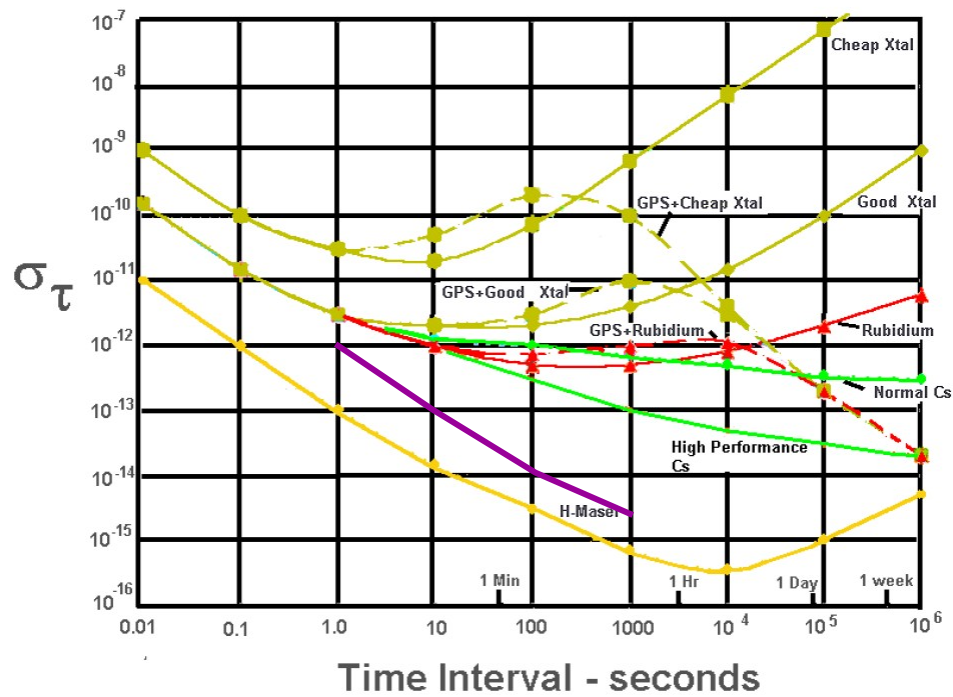
Difference plot

$$(\Phi_{\text{one way}} - \Phi_{\text{round trip}}/2)$$



Allan deviation plot



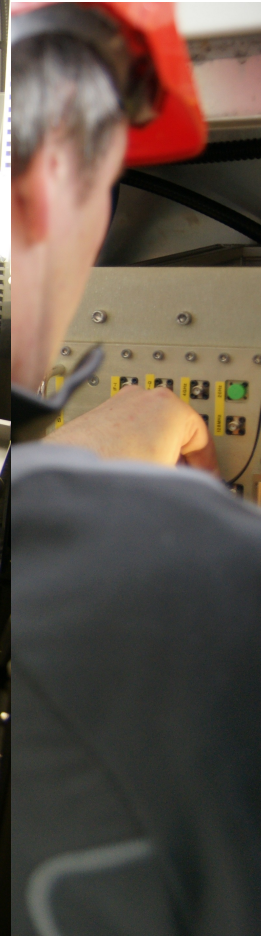


Progress

- Huge effort over



- C



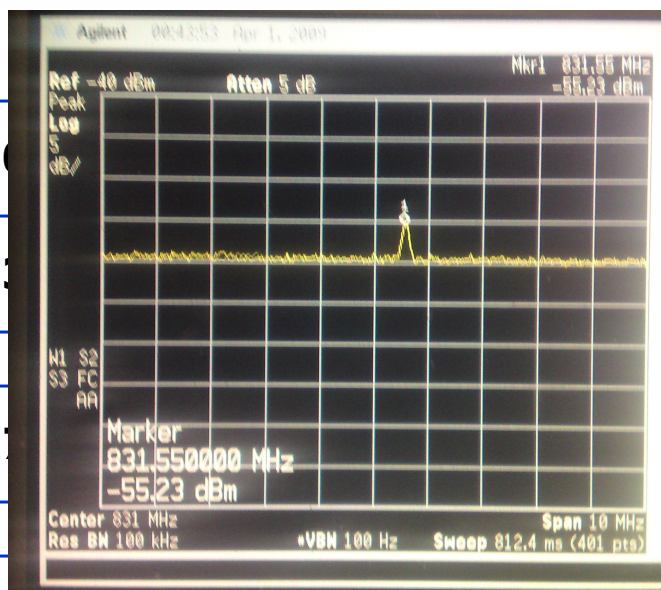
Current Status

- **1 April 09**
 - first 'analogue light' with 4 GHz IF - 6.8 GHz methanol line in IF
- **10 April 09**
 - first 'digital light' - 512 MHz IF through samplers into correlator
- **13 April 09**
 - autocorrelation of methanol line
 - 1024 channels across each of 8 x 512 MHz sub-bands
- **17 April 09**
 - first remote telescope out-fitted and connected
 - auto correlation of methanol line
- **20 April 09**
 - correlator synchronised with two telescopes
 - delay models running; started fringe search...
- **24 April 09**
 - first fringes, 3C84: delay model OK
- **May 09**
 - checking fringe rotation, 0.5 GHz/1 GHz input, synchronisation...
- **June 09**
 - Expres 4 Gbps input from IBOBs



Current Status

- 1 April 09
- 10 April 09
- 11 April 09
- 12 April 09
- 13 April 09
- 20 April 09
 - correlator synchronised with two
 - delay models running; started fr
- 24 April 09
 - first fringes, 3C84: delay model
- May 09
 - checking fringe rotation, 0.5 GHz
- June 09
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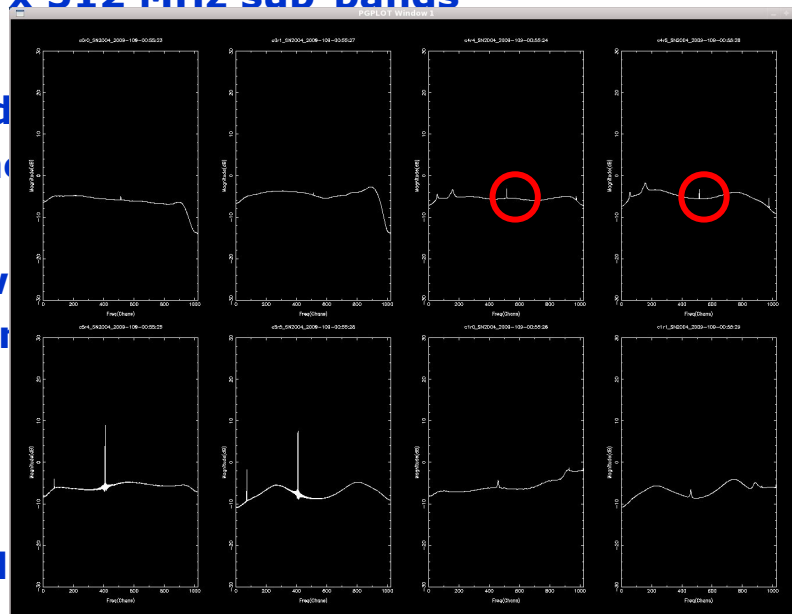
Hz IF - 6.8 GHz methanol line in IF

IF through samplers into correlator

line

f 8 x 512 MHz sub-bands

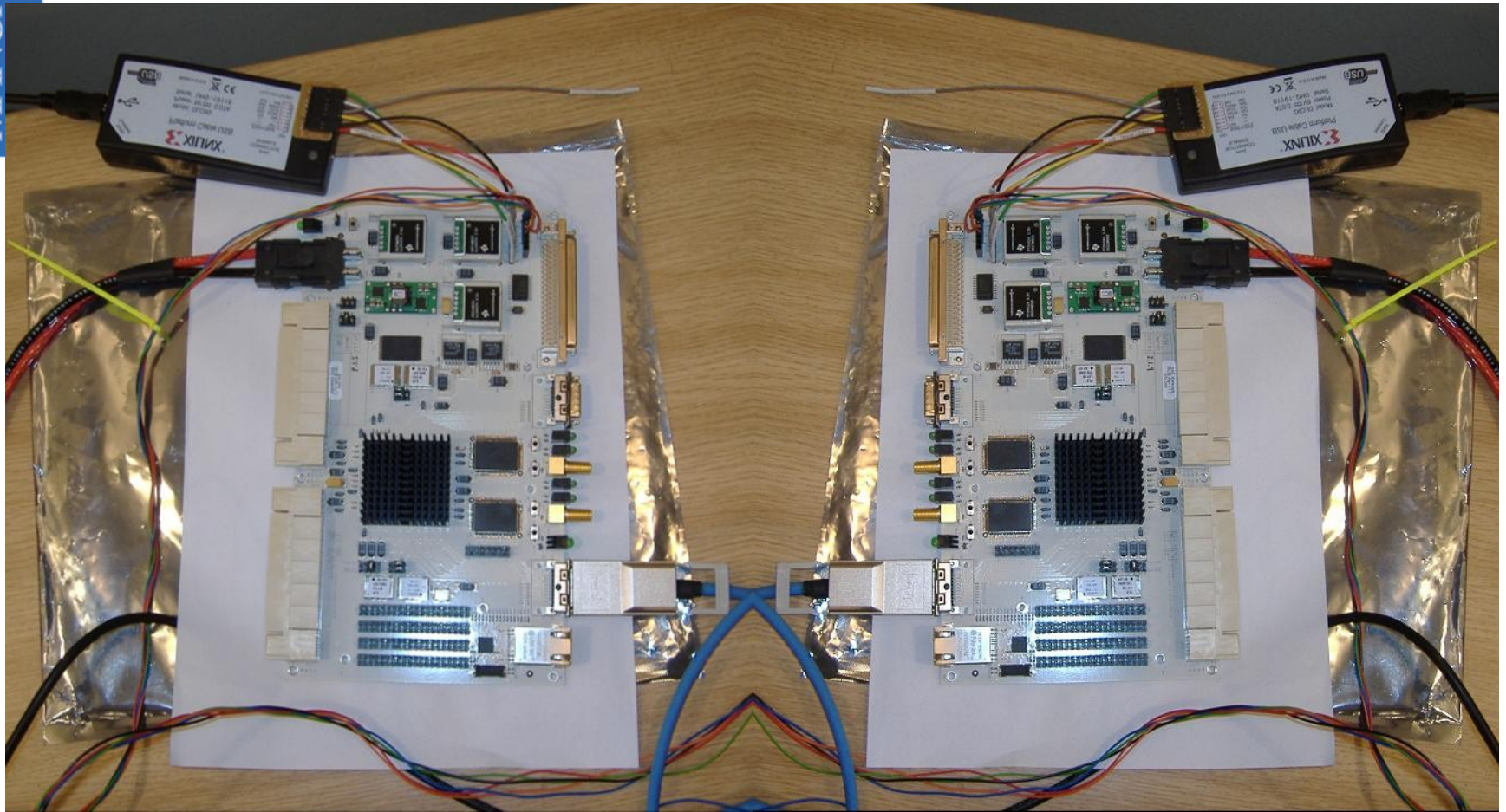
ted
line



Coming months

- **Continue single baseline tests**
- **4 station tests**
 - next 2 station boards now installed
 - next production run of IF, digital boards: Summer 09
- **Full array & full bandwidth**
 - requires further SB, BB (Nov 09), 4 Gs/s samplers
 - regeneration electronic for last, most distant telescope at Cambridge (Dec 09)
- **Engineering/commissioning obs 2009**
- **Limited, shared-risk observations in 2010A**
- **Standard operations 2010B**

e-MERLIN & VLBI



board, 1K5C packet writing already in correlator

e-MERLIN Legacy Programme

- National and international consortia prepared proposals for large, high impact science projects
- 15 proposals submitted from 325 scientists from 119 institutes
 - already reaching a much wider community
- Reviewed by external referees and Legacy Steering Group chaired by Prof R Ivison; rigorous assessment following UK PATT procedure
- 5000 hours allocated to 11 projects, covering topics from the formation of planets to the formation of galaxies in the early universe.
 - important demonstration of community interest & potential science

Legacy programme as a whole: Proposed projects

- Breadth of legacy proposals demonstrate the versatility of e-MERLIN
- Programme covers $\sim 50\%$ of e-MERLIN time in first 5 semesters
- In total $\sim 10,000$ hrs of e-MERLIN time were proposed for
 - **Well oversubscribed**

Allocated programme:

GALACTIC PROJECTS:

• eTT - Pulsar interferometry	- Vlemmings/Stappers et al.	160hrs **
• PEEBLES	- Greaves et al.	72hrs **
• Feedback processes in Massive SF	- Hoare/Vlemmings et al.	450hrs
• Thermal jets from low mass stars	- Rodriguez et al	180hrs
• COBRaS	- Prinja et al.	294hrs

EXTRA-GALACTIC PROJECTS:

• LEMMINGS	- Beswick/McHardy et al.	810hrs
• LIRGI	- Conway/Perez-Torres et al.	353hrs
• Extragalactic Jets	- Laing/Hardcastle et al	375hrs
• AGATE	- Simpson/Smail et al	330hrs
• e-MERGE	- Muxlow/Smail/McHardy++	918hrs
• Gravitational lenses	- Jackson/Serjeant et al	228hrs **

TOTAL : 4170hrs

- ****Up to 5000hrs can be allocated. Remaining 830hrs to be allocated by LSG to pre-identified projects pending initial results**
- DETAILS OF ALL PROJECTS (inc Proposals allocations etc) ARE AVAILABLE AT

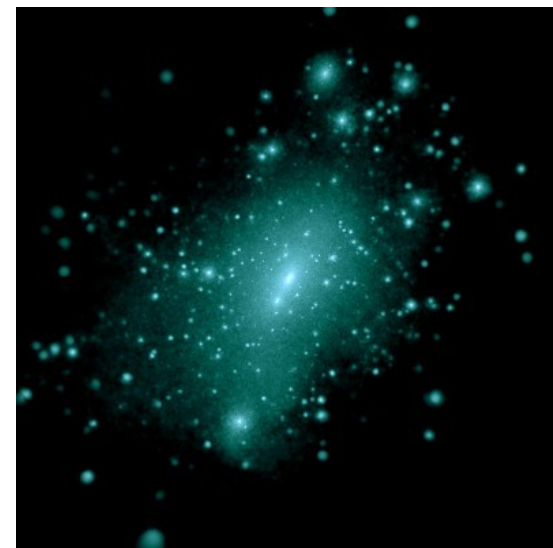
<http://www.merlin.ac.uk/legacy/>

Legacy projects (1)

galaxy formation and evolution

Galaxy formation (Grav lenses)

- **Gravitational lenses - test CDM models (halo substructure) and probe central $\sim 10\text{pc}$**
 - **Requires e-MERLIN sensitivity & resolution/astrometry**
- **New survey (>1000 lenses)**
Herschel - SCUBA2 -e-MERLIN
 - **e-MERLIN vital for confirmation (CLASS experience)**

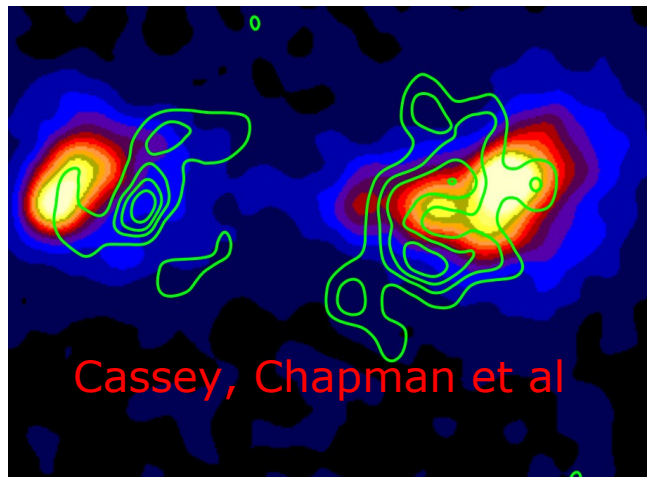
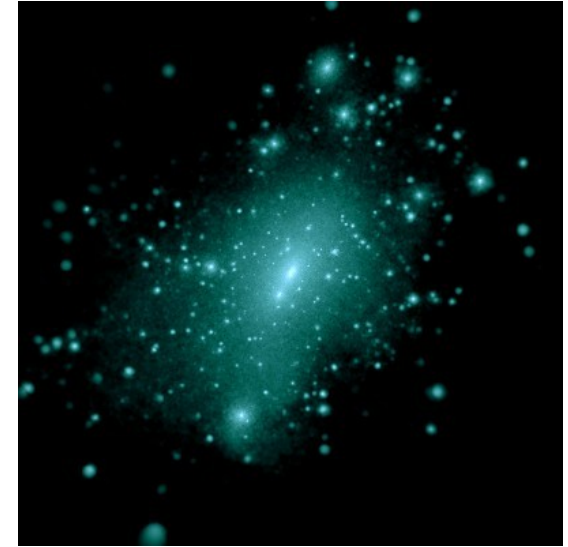


Galaxy evolution (e-MERGE, AGATE)

- **Ultra-deep (cluster lenses): star-formation in normal galaxies at $z \sim 5$**
- **Deep (0.25 dg^2): star-formation history $z \sim 0.5-3$ (extinction-free; AGN; sub-mm quiet SFG)**
- **Environmental impacts on galaxy evolution: clusters;**
 - **e-MERLIN required to resolve typical few-kpc star-forming regions at $z \sim 1$ and identify/separate AGN (HDF experience)**

Legacy projects (1)

galaxy formation and evolution



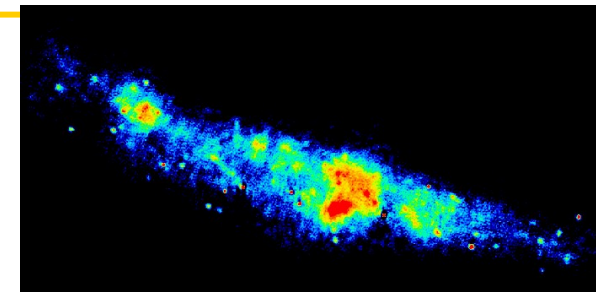
Legacy projects (2)

From galaxies to Star and planet formation

Birth & death of stars in the Local Universe

(LeMMINGS, LIRGI)

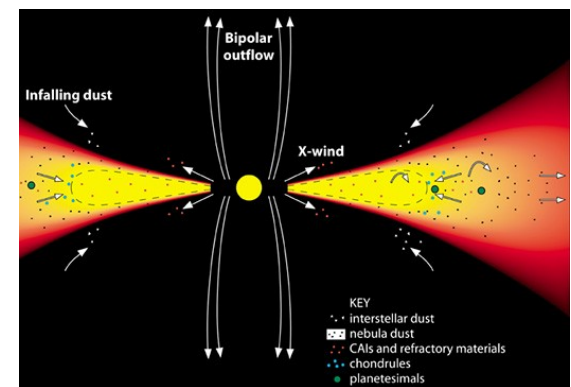
- Add to multi- λ Nearby galaxy samples (SINGS, KINGFISH, GOALS...)
- Super star clusters; determination of IMF via SN rates
- Census of AGN; identify intermediate-mass black holes; ULX
- Magnetic field in LIRGS
- RSNe detection & AGN separation in LIRGS
 - **e-MERLIN resolution essential to resolve SSC, SNR, AGN (M82 experience)**



Physical processes of star-formation

(Massive star-formation, stellar jets, CoBRAS)

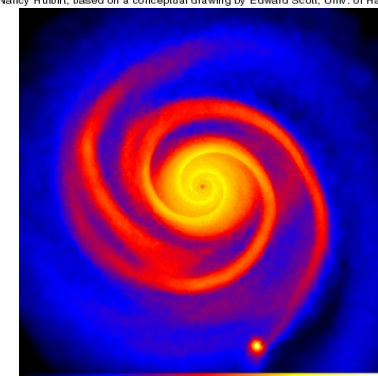
- Trace onset of feedback processes by jets/winds during early stages of massive star formation.
- First direct test of the X-wind model for protostellar jets
 - **e-MERLIN can resolve key physical processes**



Formation of planetary systems

(PEEBLES)

- Detect cm-sized particles in planet-forming disks ~ 1 AU, $\sim 1M_E$
- Astrometric detection of Jupiter-mass planets around low-mass stars
 - **e-MERLIN required for spatial and spectral separation of emission components**

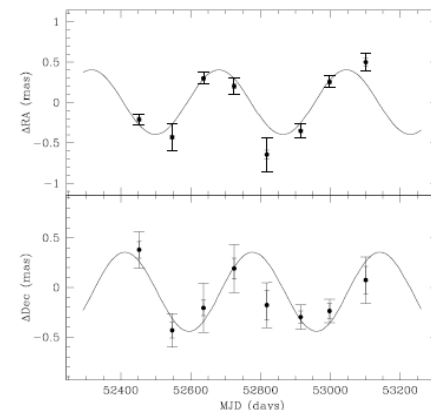


Legacy projects (3)

Extreme environments

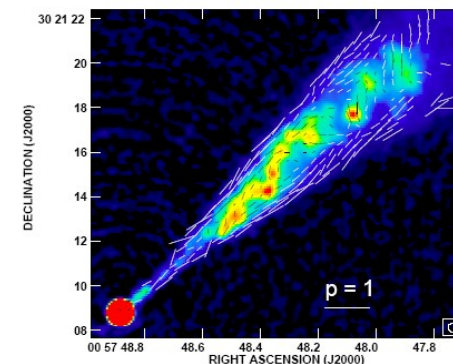
Pulsar proper motions (eTT)

- Proper motions of many pulsars
- Physics of neutron star formation during SNe
- Double the number of pulsars with accurate distances & velocities
 - e-MERLIN resolution essential, multi-epoch proper motions



Extragalactic Jets

- Dynamics of jets, low luminosity radio jet deceleration on sub-kpc scales
- Magnetic field configuration surrounding jets
- Particle acceleration in hotspots
- Polarisation and faraday rotation across e-MERLIN bands
 - e-MERLIN can resolve the key physical processes'



Conclusion

- E-Merlin will have a major impact on high resolution astronomy
- Huge progress in past few months but much to do

- E-Merlin
resolution
- Huge pr
do



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much to