



Herschel mission: Overview and observing opportunities

Observing opportunities with *Herschel*: A window to the far-infrared universe – an ASTROCAM workshop

ESAC, Villafranca, 14-15 December 2006

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Astrophysics Missions Division

Research and Scientific Support Department

Herschel in a nutshell



- **ESA cornerstone observatory**
 - instruments 'nationally' funded, int'l - NASA, CSA, Poland – collaboration
 - ~1/3 guaranteed time, ~2/3 open time
- **FIR (57 - 670 μm) space facility**
 - large (3.5 m) monolithic low emissivity passively cooled telescope
 - 3 focal plane science instruments
 - 3 years routine operational lifetime
 - full spectral access
 - low and stable background
- **Unique and complementary**
 - for $\lambda < 200 \mu\text{m}$ larger aperture than cryogenically cooled telescopes (IRAS, ISO, Spitzer, Akari,...)
 - more observing time than balloon- and/or air-borne instruments
 - larger field of view than interferometers
- **Launch in 2008**
 - initial observing AO on 1 Feb 2007



HERSCHEL SPACE OBSERVATORY

The Cool Universe



- **Herschel spectral coverage**

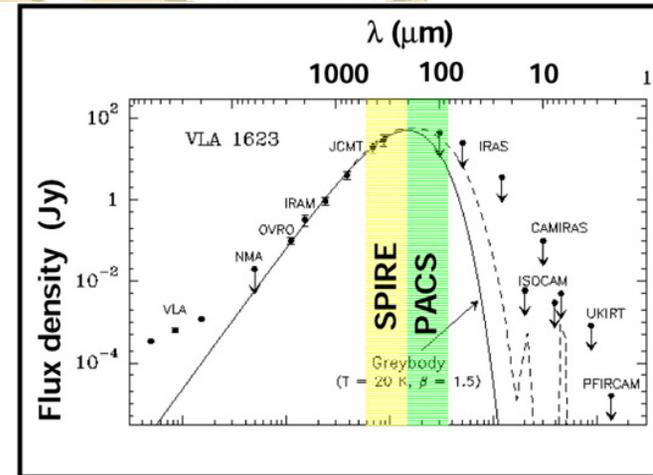
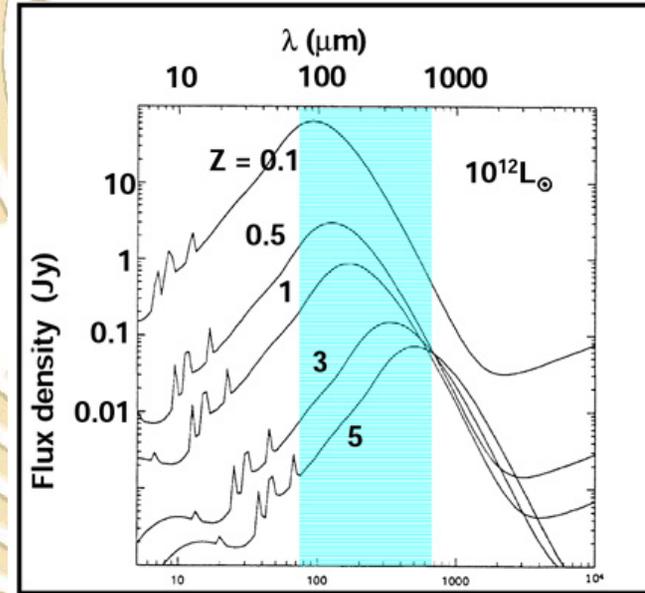
- black-bodies 5-50 K
 - continuum radiation
 - dust grains (re-)radiating
- gases 10-few100 K
 - brightest atomic/molecular lines

- **Herschel strengths**

- covers IR dominated galaxies & protostar SED peaks
- wide area mapping
- full coverage spectral scans & particular (water) lines

- **Herschel emphasis**

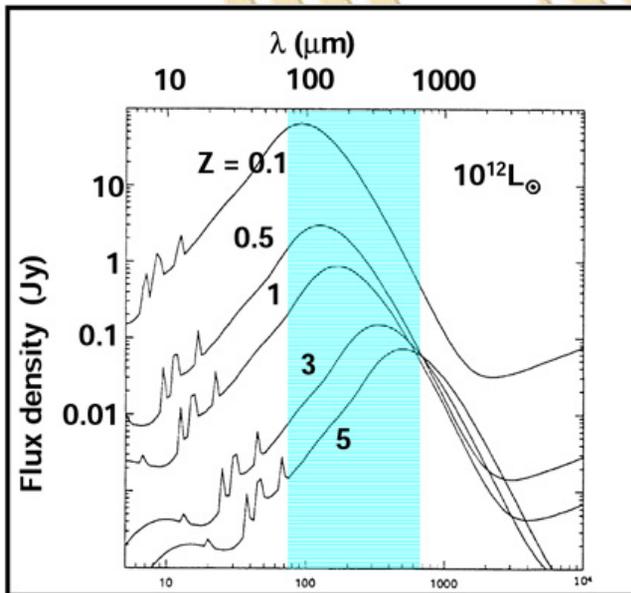
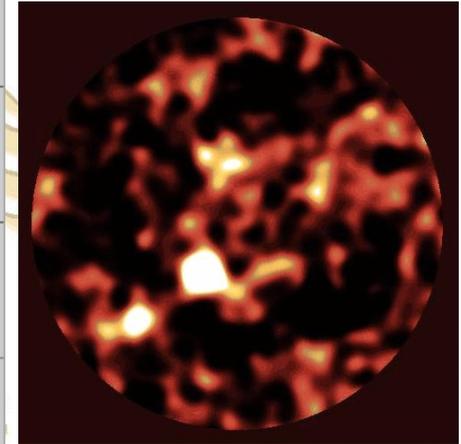
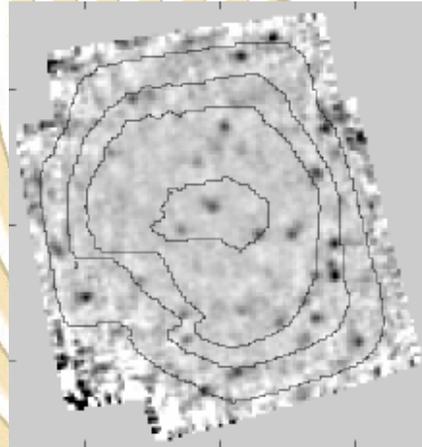
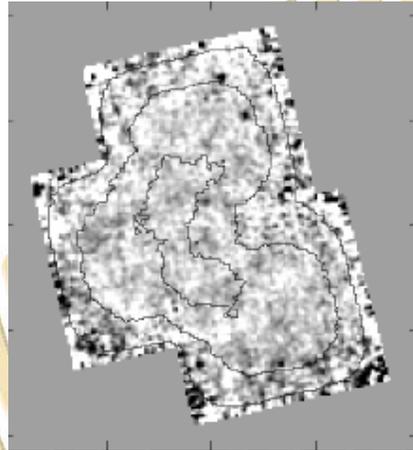
- formation and evolution of galaxies & stars
- ISM physics & chemistry
- solar system bodies



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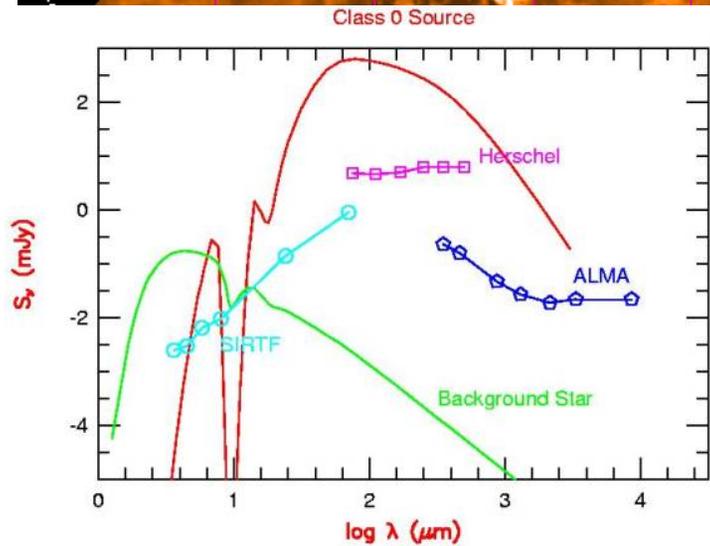
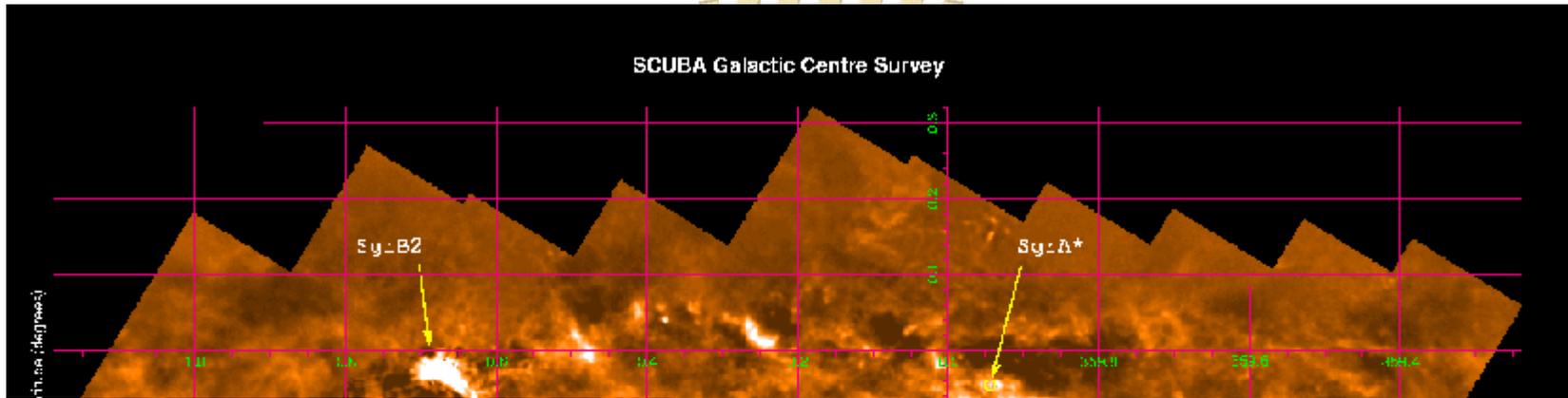
Galaxy formation & evolution



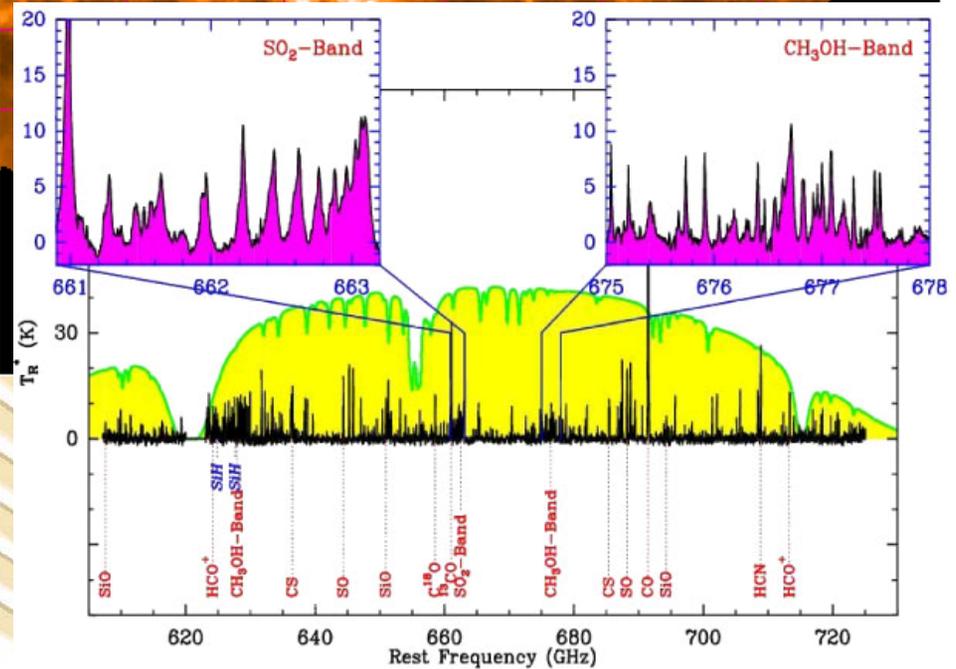
Early Spitzer results:

- 24 μm – 70%
- 70 μm – 23%
- 160 μm – 7%
- but now claiming all resolved!

Surveys & follow-up



0.1 L_{sun} protostar at 300 pc



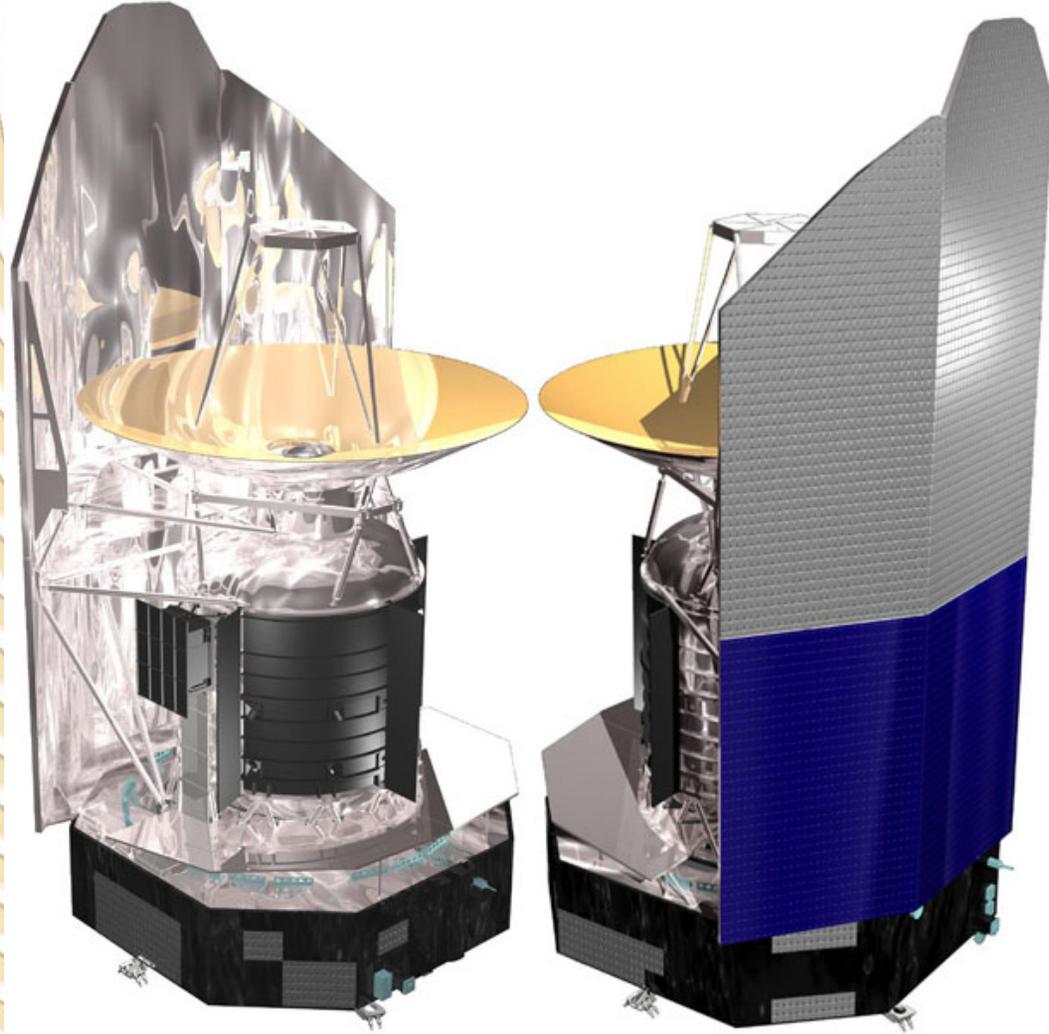
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Herschel spacecraft specs



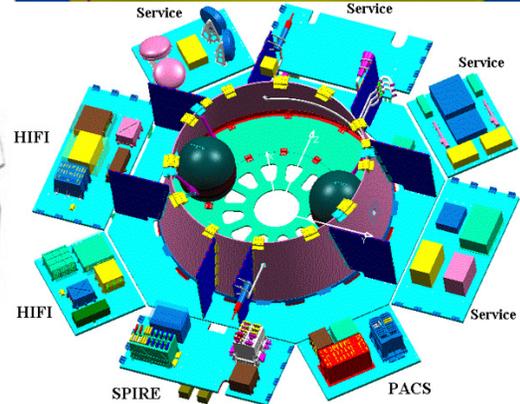
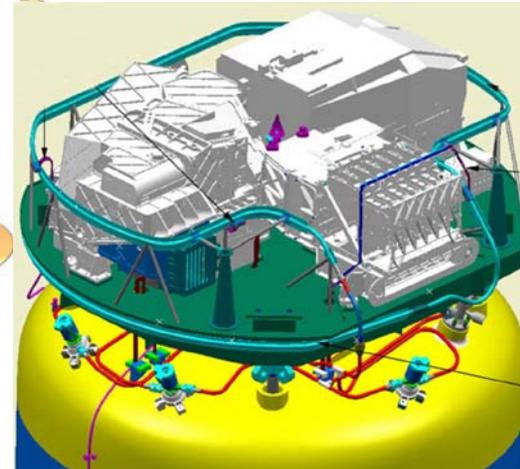
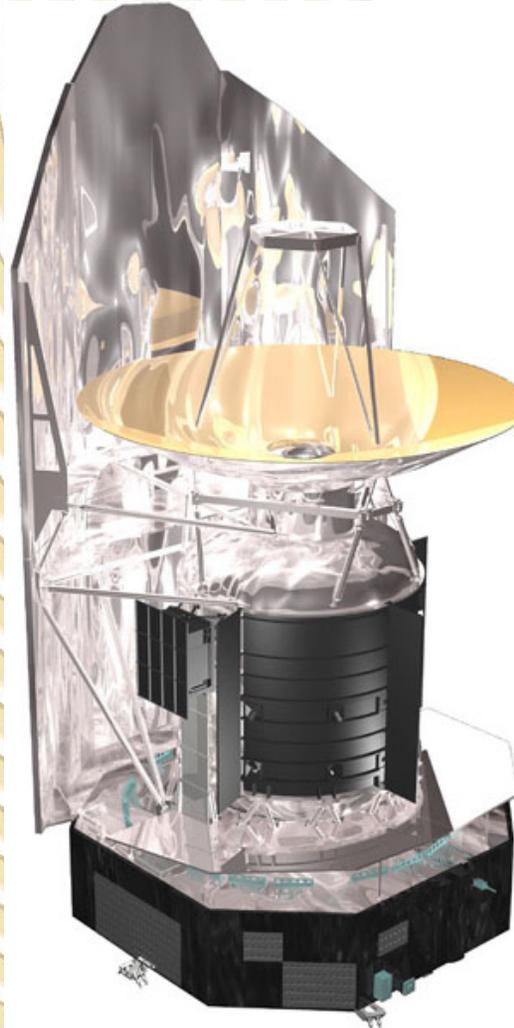
- telescope (eff) diam (3.3) 3.5 m
- *telescope WFE* < 6 μm
- *telescope temp* < 90 K
- *telescope emissivity* < 4%
- *abs/rel pointg (68%)* < 3.7" / 0.3"
- science instruments 3
- science data rate 130 kbps
- *cryostat lifetime* > 3.5 years
- height / width ~ 7.5 / 4 m
- launch mass ~ 3200 kg
- power ~ 1500 W
- orbit 'large' Lissajous around L2
- solar aspect angle 60-120 deg
- launcher (w Planck) Ariane 5 ECA



Herschel spacecraft specs

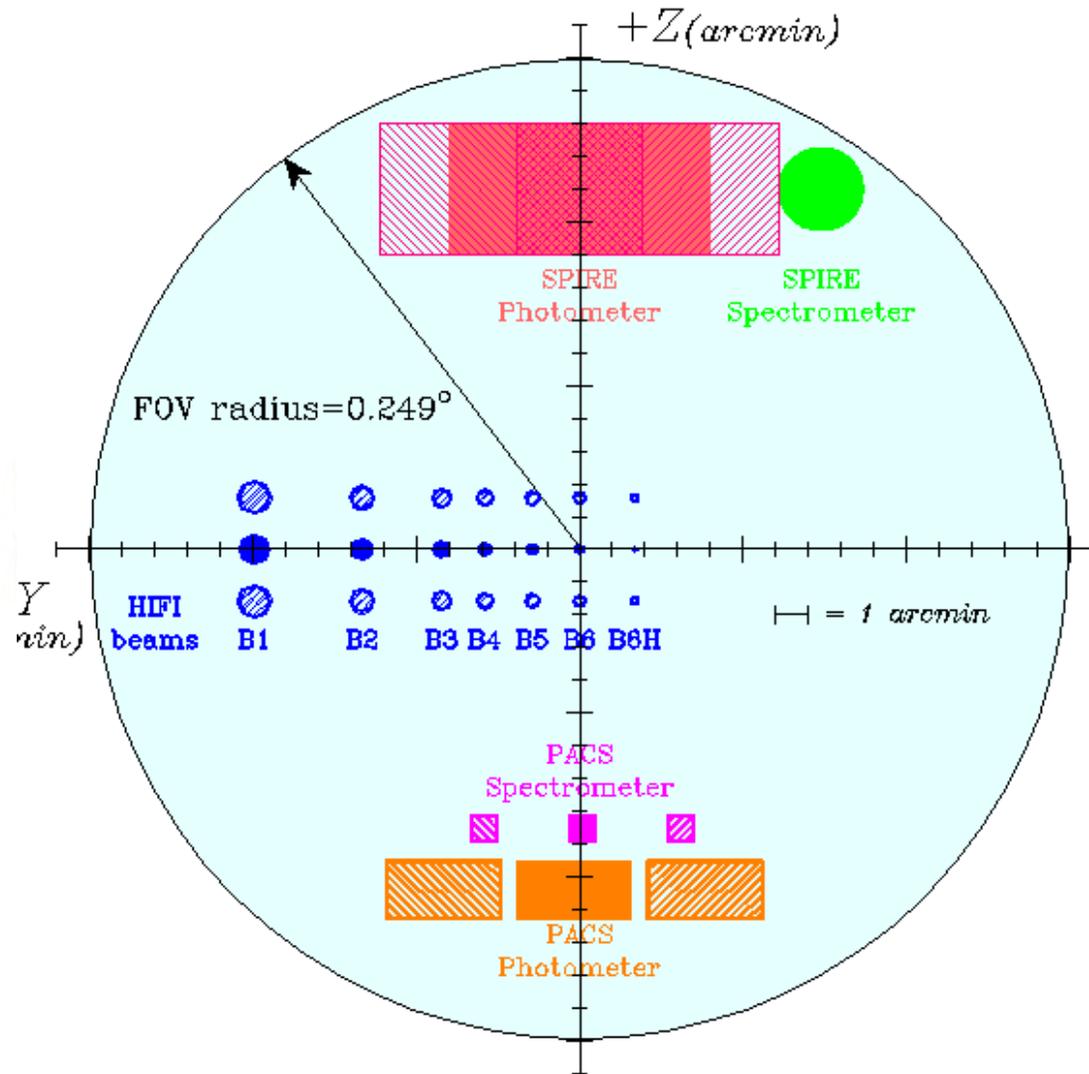


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Herschel focal plane



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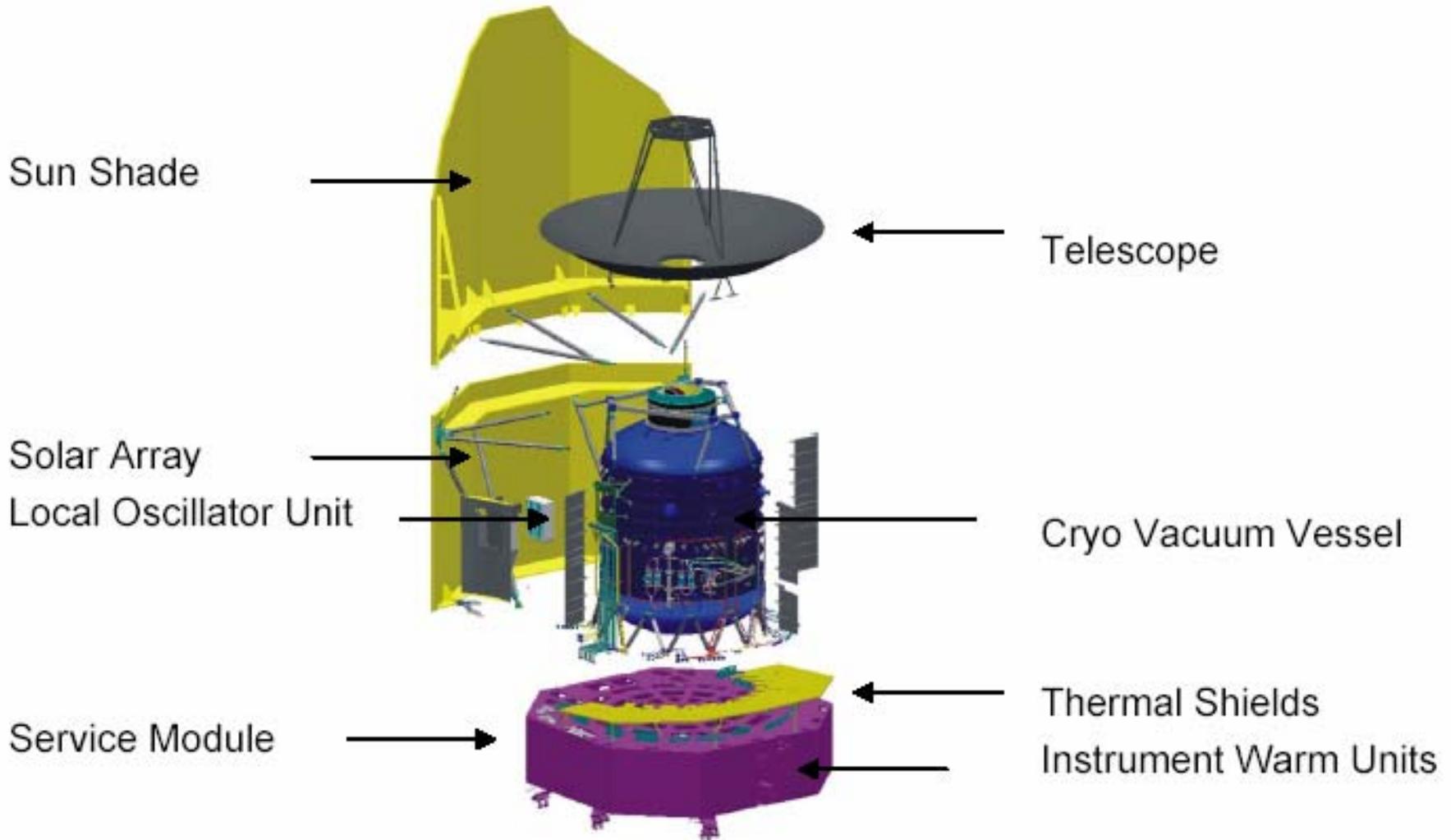


Herschel observatory capabilities

- **Photometry - imaging, 6 broad bands in 75-500 μm range**
 - **PACS** – simultaneous 2 colour fully-sampled ($0.5F\lambda$) imaging with FOV 1.75×3.5 arcmin and $R \sim 2.5$ centred at 75/110 and 170 μm
 - **SPIRE** – simultaneous 3 colour $2F\lambda$ imaging with FOV 4×3 arcmin and $R \sim 3$ centred at 250, 363, and 517 μm
 - for larger fields 'on-the-fly' mapping, mosaicing
 - sensitivity is somewhat wavelength and observing mode dependant, very roughly for point sources $1 \text{ mJy} - 1\sigma - 1$ hour; for mapping confusion limit is important
- **Spectroscopy - in 57-670 μm range, varying R in 20- 10^7 range**
 - **PACS** – grating spectrometer, 5×5 spatial $\times 16$ spectral pixels, FOV 0.8 arcmin, $R \sim 1500-4000$, $\lambda \sim 57-210 \mu\text{m}$
 - **SPIRE** – MTS spectrometer, $R \sim 20-100+$, FOV 2.6 arcmin, $\lambda \sim 200-670 \mu\text{m}$
 - **HIFI** – heterodyne spectroscopy with R up to 10^7 , $\lambda \sim 157-212$ and $240-625 \mu\text{m}$, 2 orthogonal polarisations, 4000 spectral channels per polarisation, single pixel on the sky, mapping by 'on-the fly' or mosaicing observations

Following talks!

Herschel spacecraft



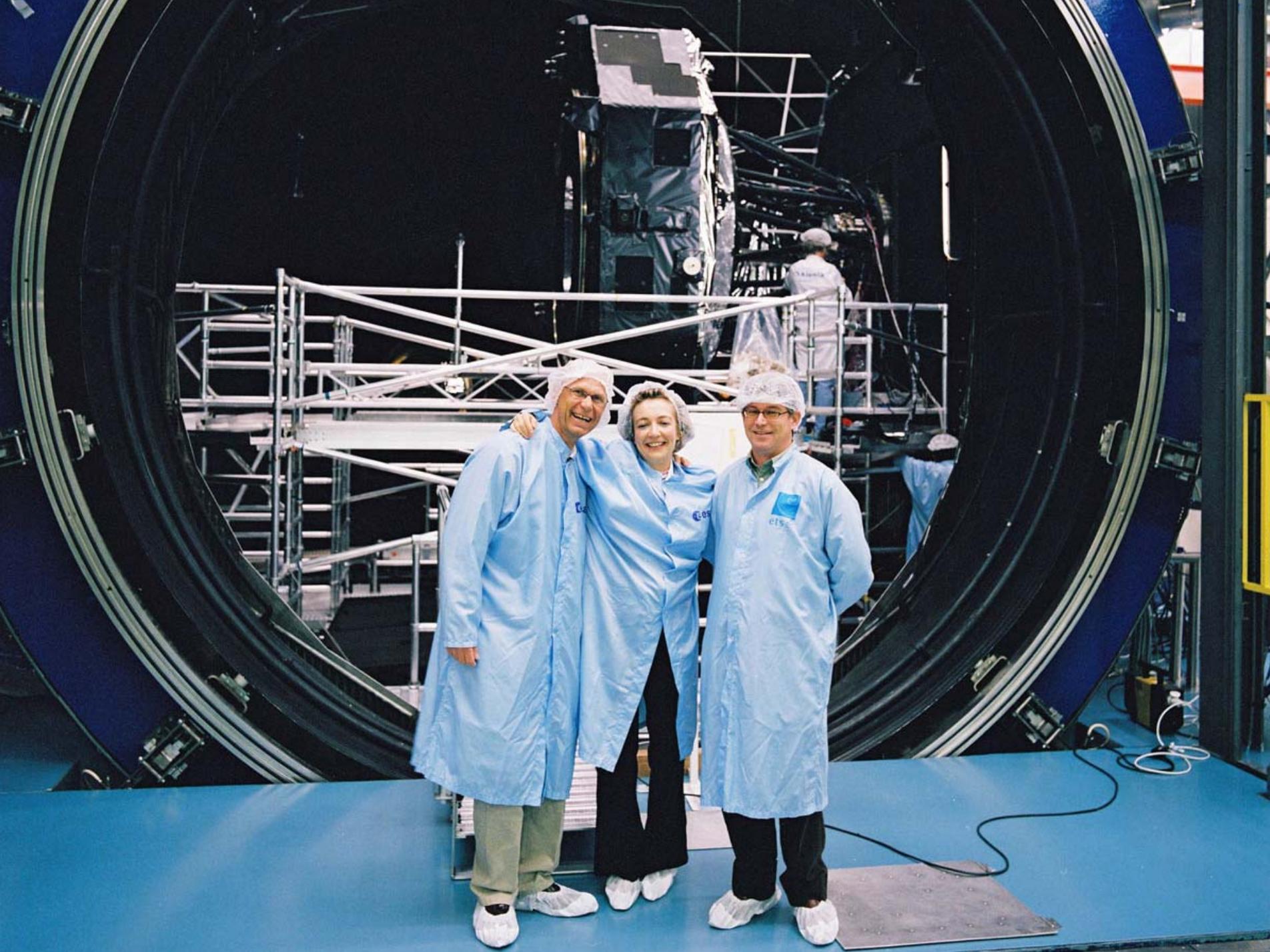


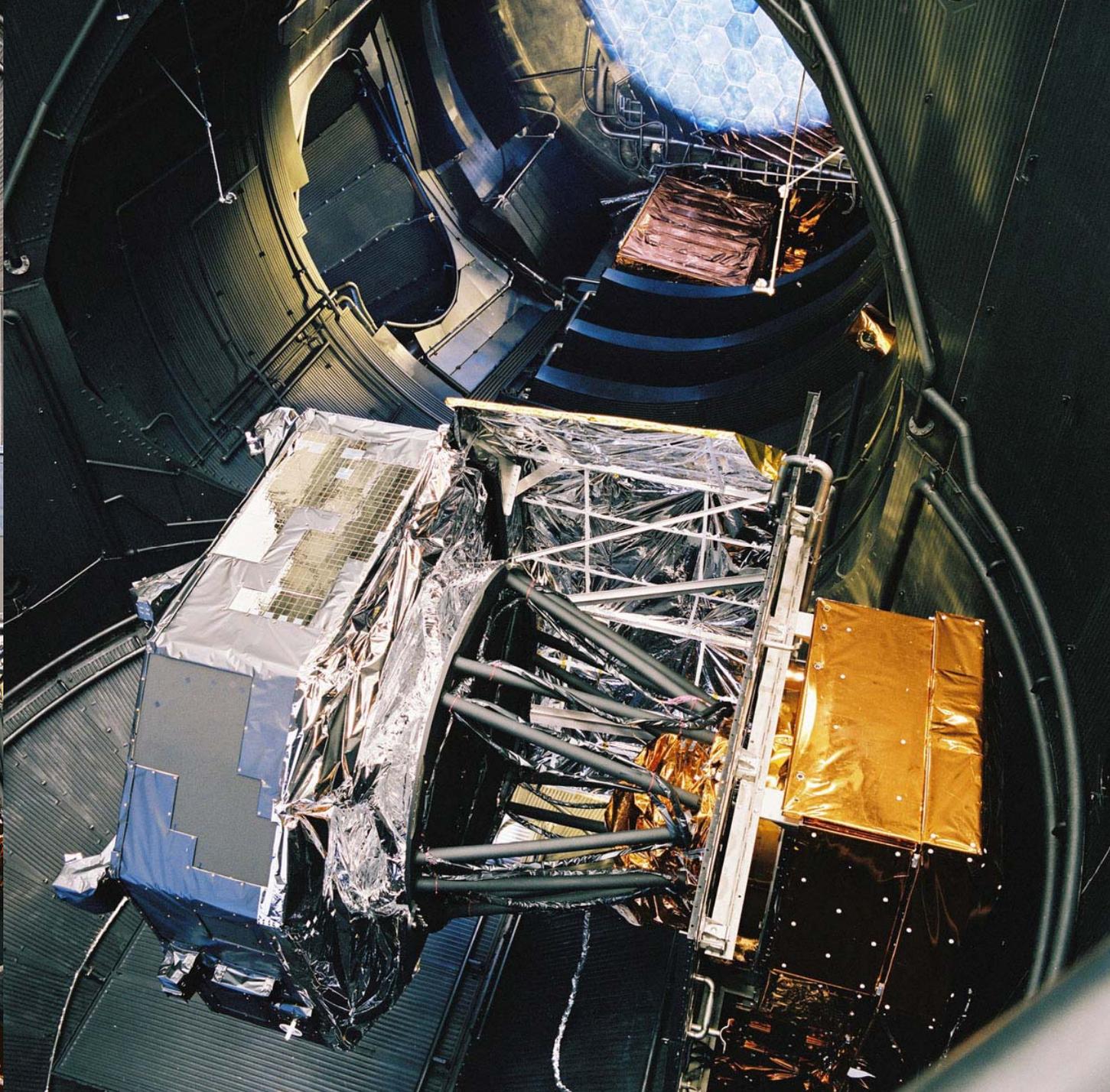
Herschel spacecraft programmes

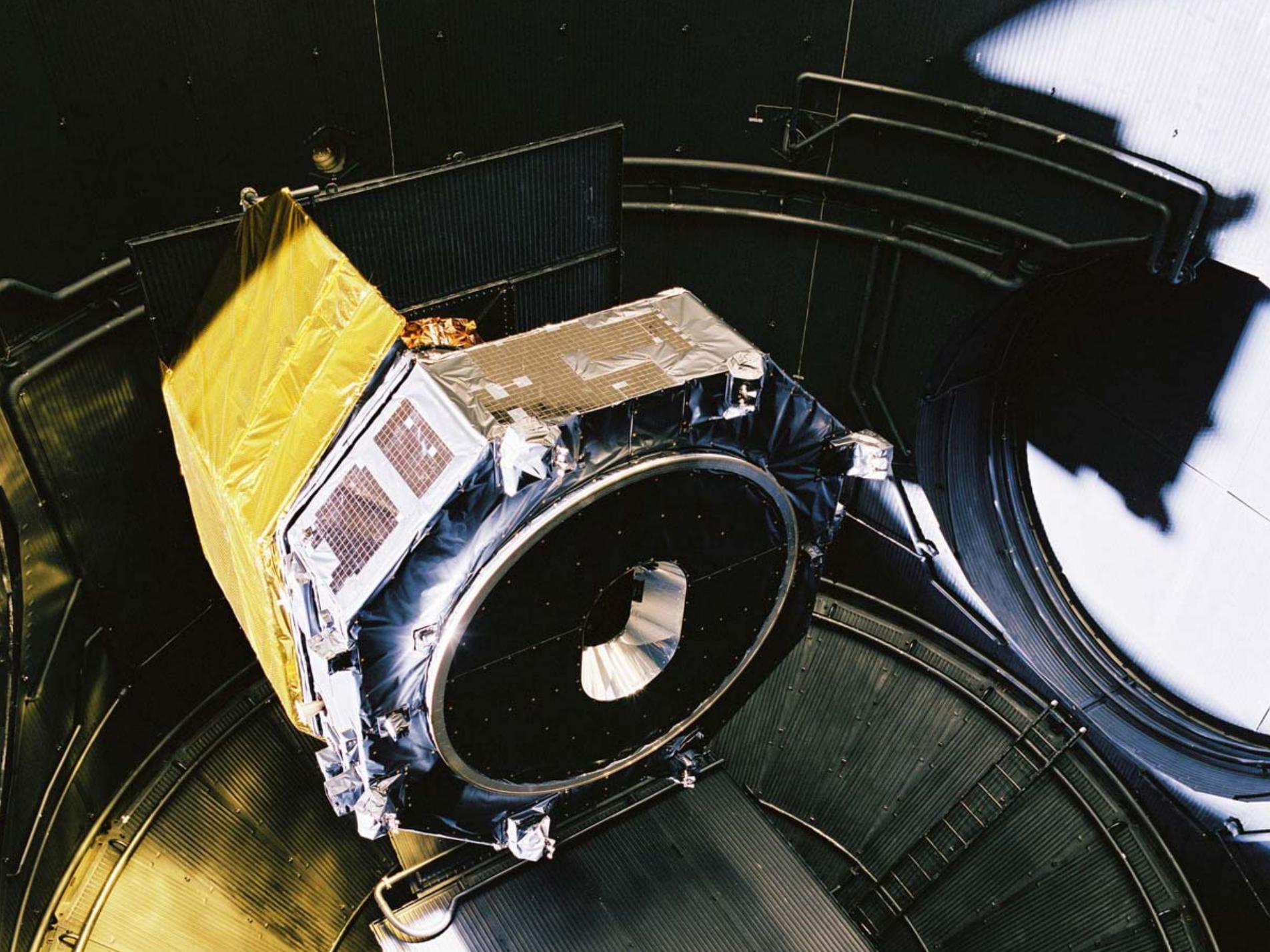
- **Satellite PFM programme**
 - two SVMs – QM and FM – with many common parts
 - only one cryostat/PLM is built
 - only one satellite is built – tested in STM configuration
 - FM cryostat
 - STM models of everything else
- **Telescope PFM programme**
 - telescope manufacture complete – has undergone lengthy cryo-testing and characterisation campaign in CSL
 - spare parts exist for FS
- **EQM programme – completed**
 - ISO hardware refurbished to provide ‘Herschel OB simulator’
 - science instrument CQMs
- **AVM programme**
 - one physical setup – configurable as either Herschel or Planck
 - test procedures and OBSW development

Herschel SVM/STM

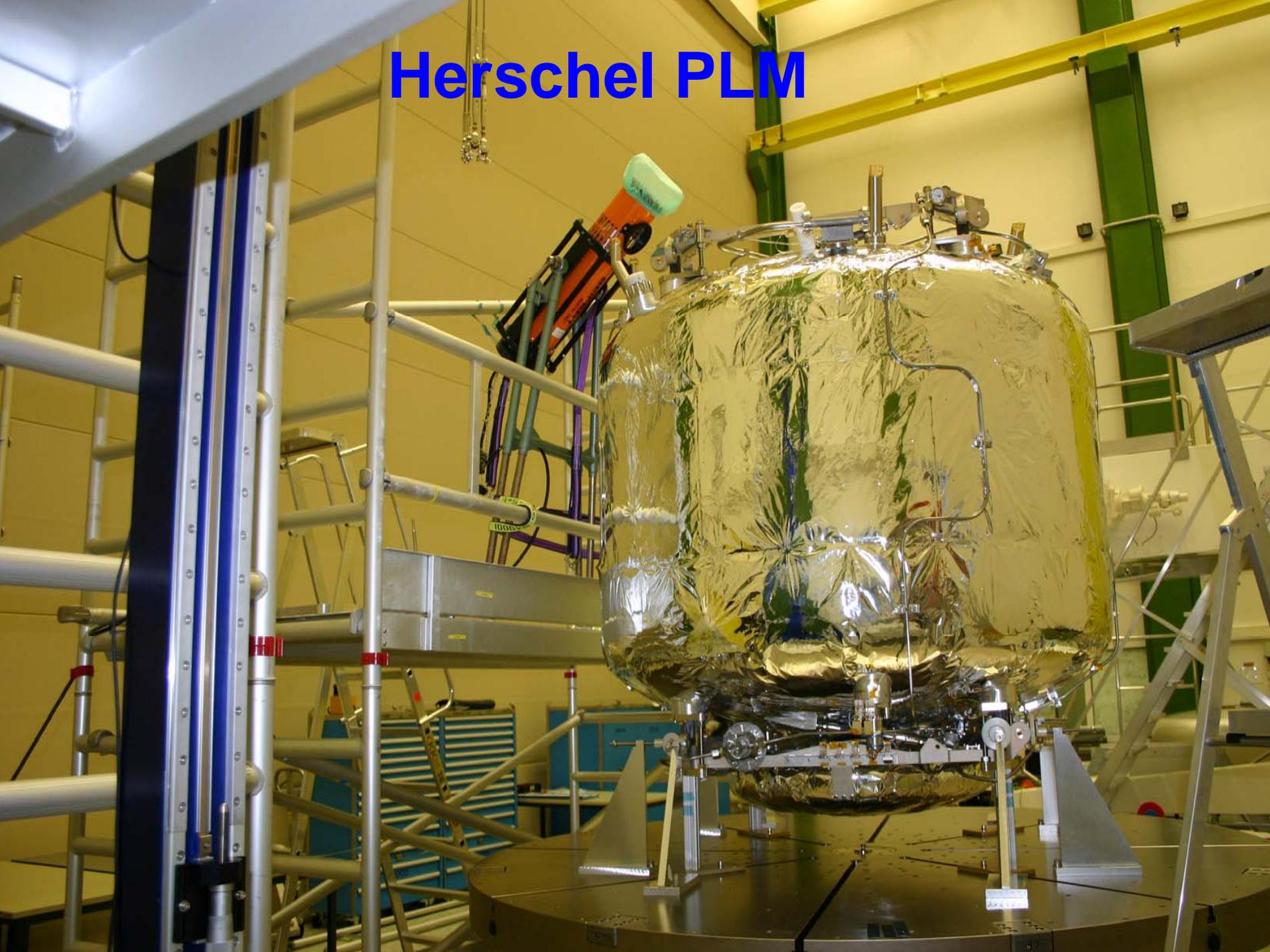






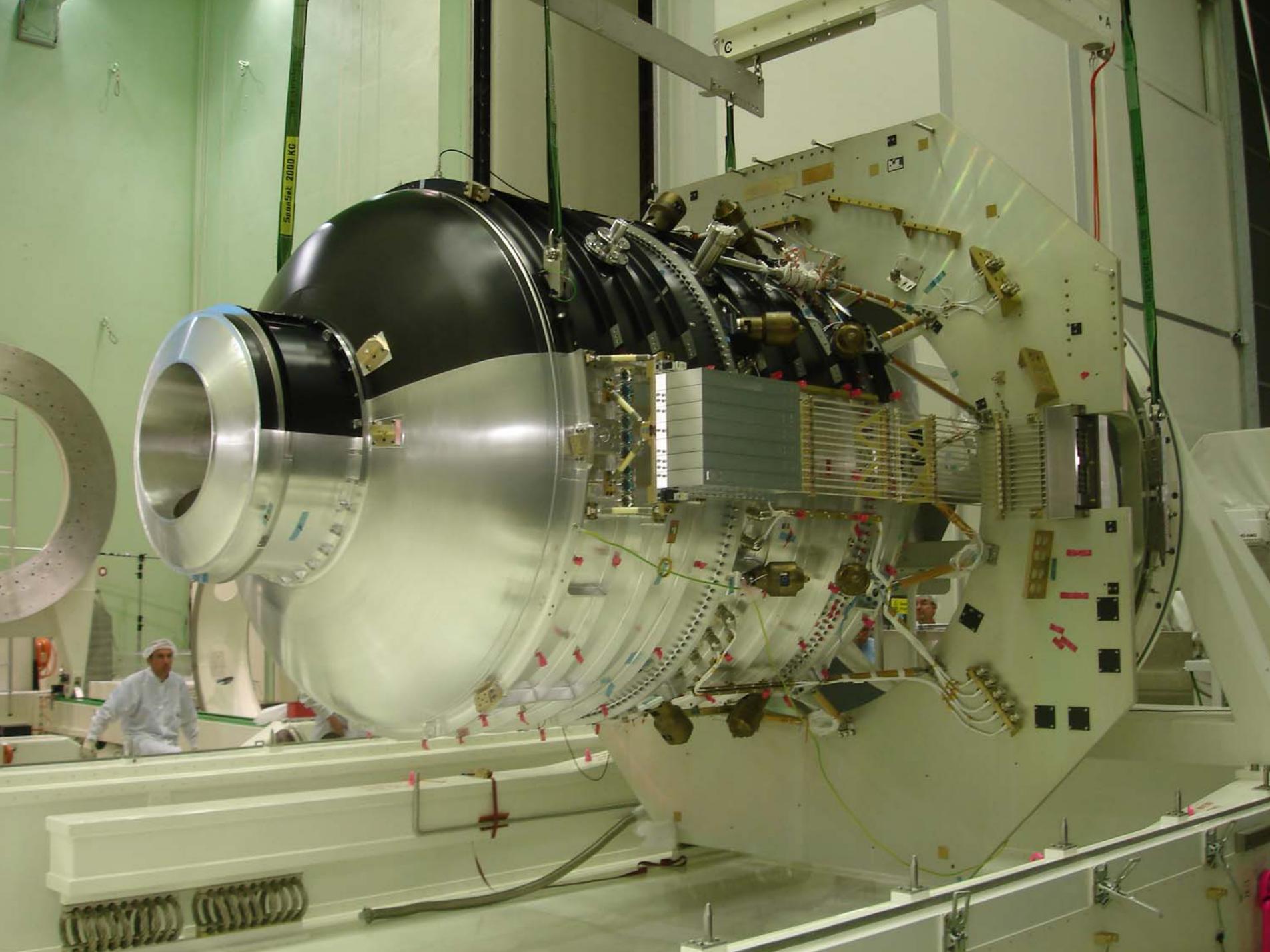


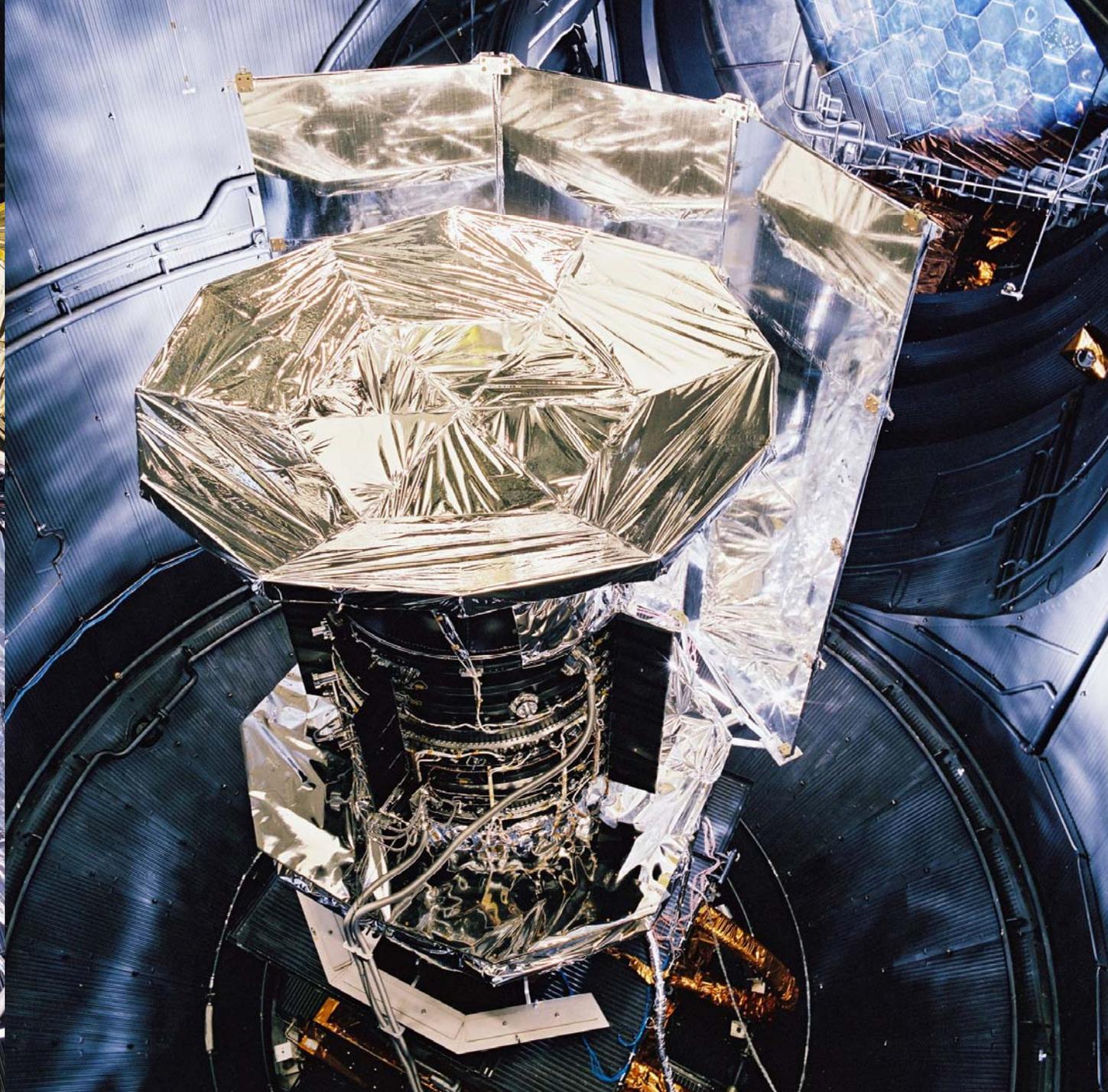
Herschel PLM













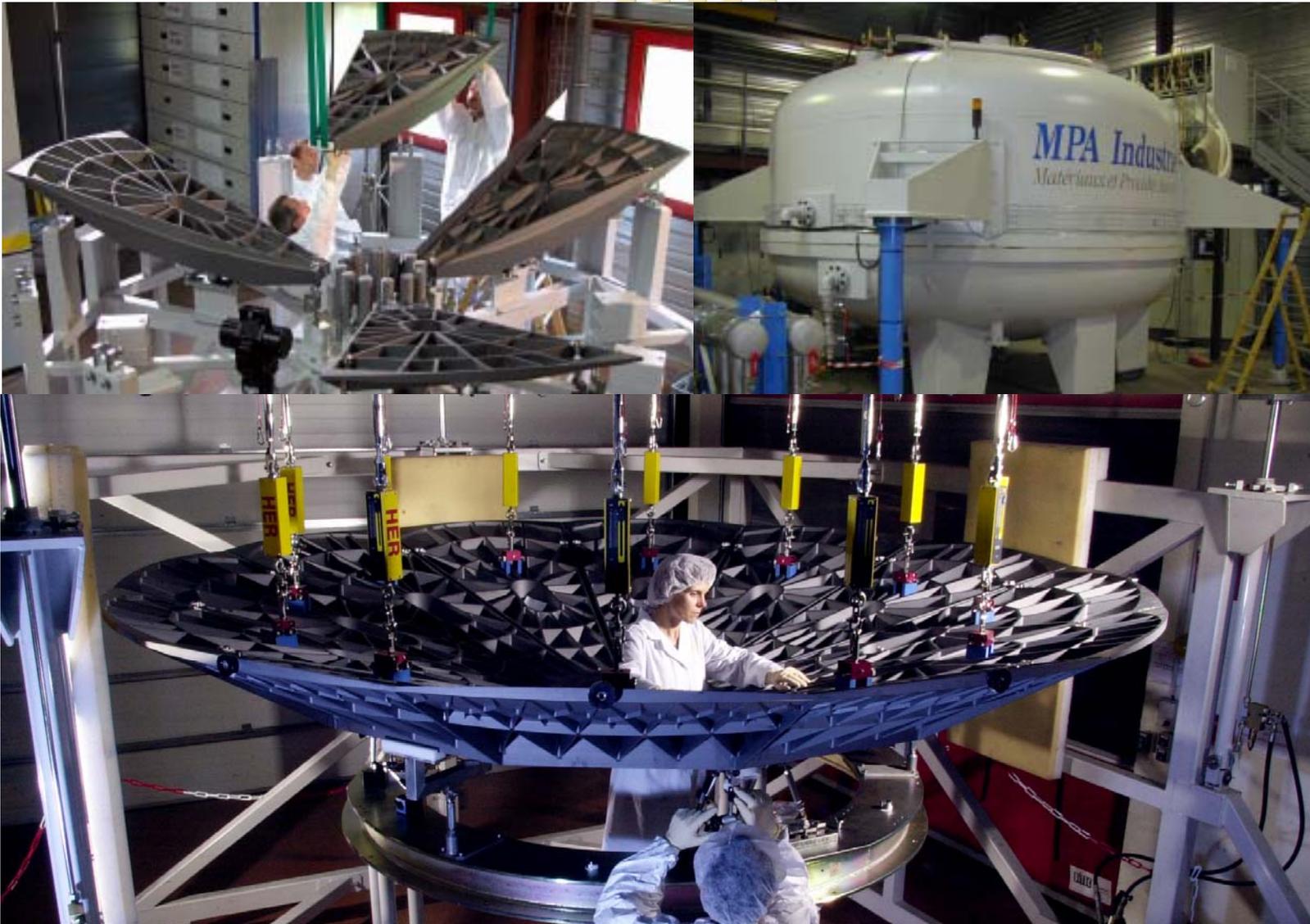
Herschel STM satellite

Herschel
Science Team
visited the
ESTEC Test
Centre to view
the Herschel
STM satellite
on 1 Feb 2006



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Brazing of primary mirror 'blank'

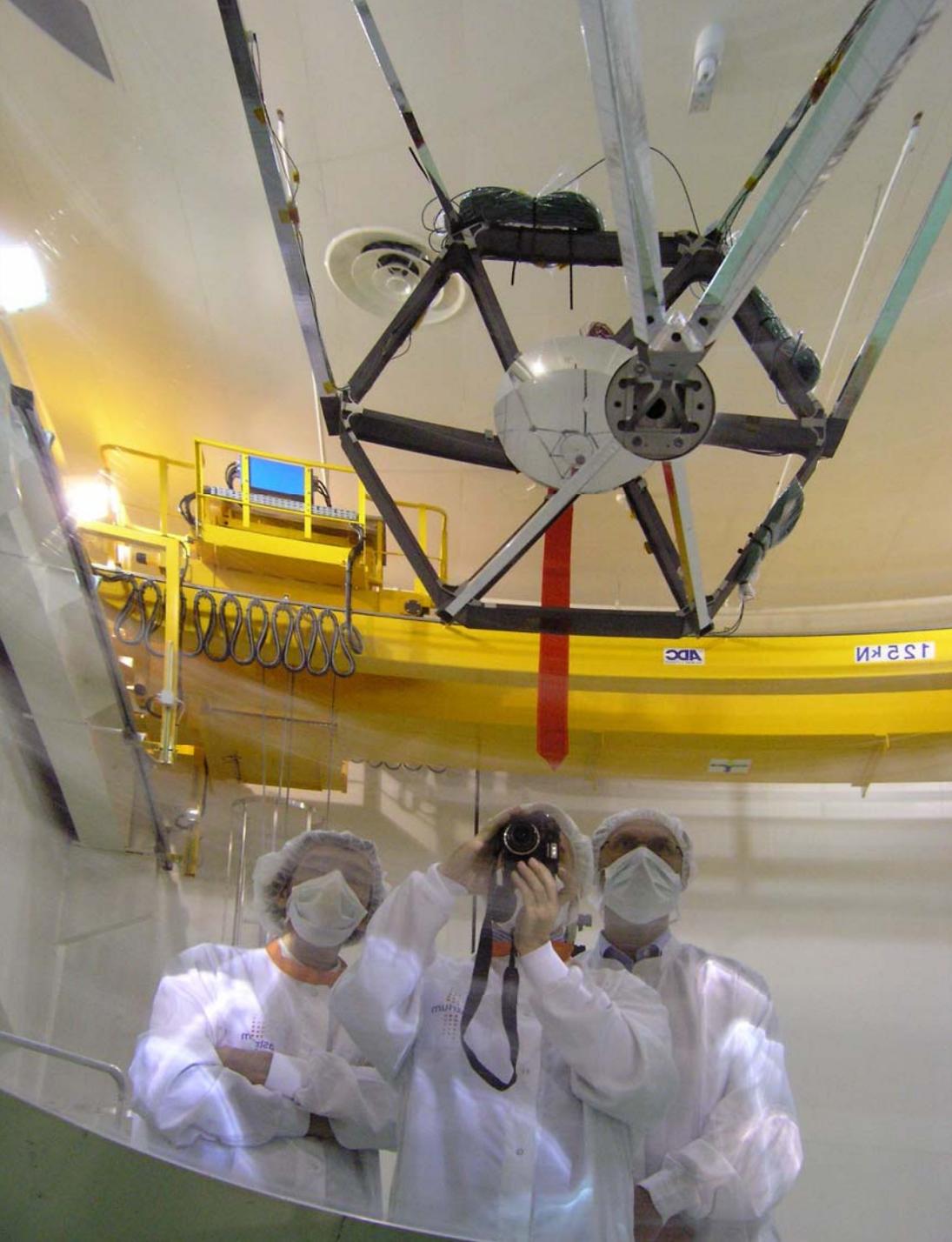
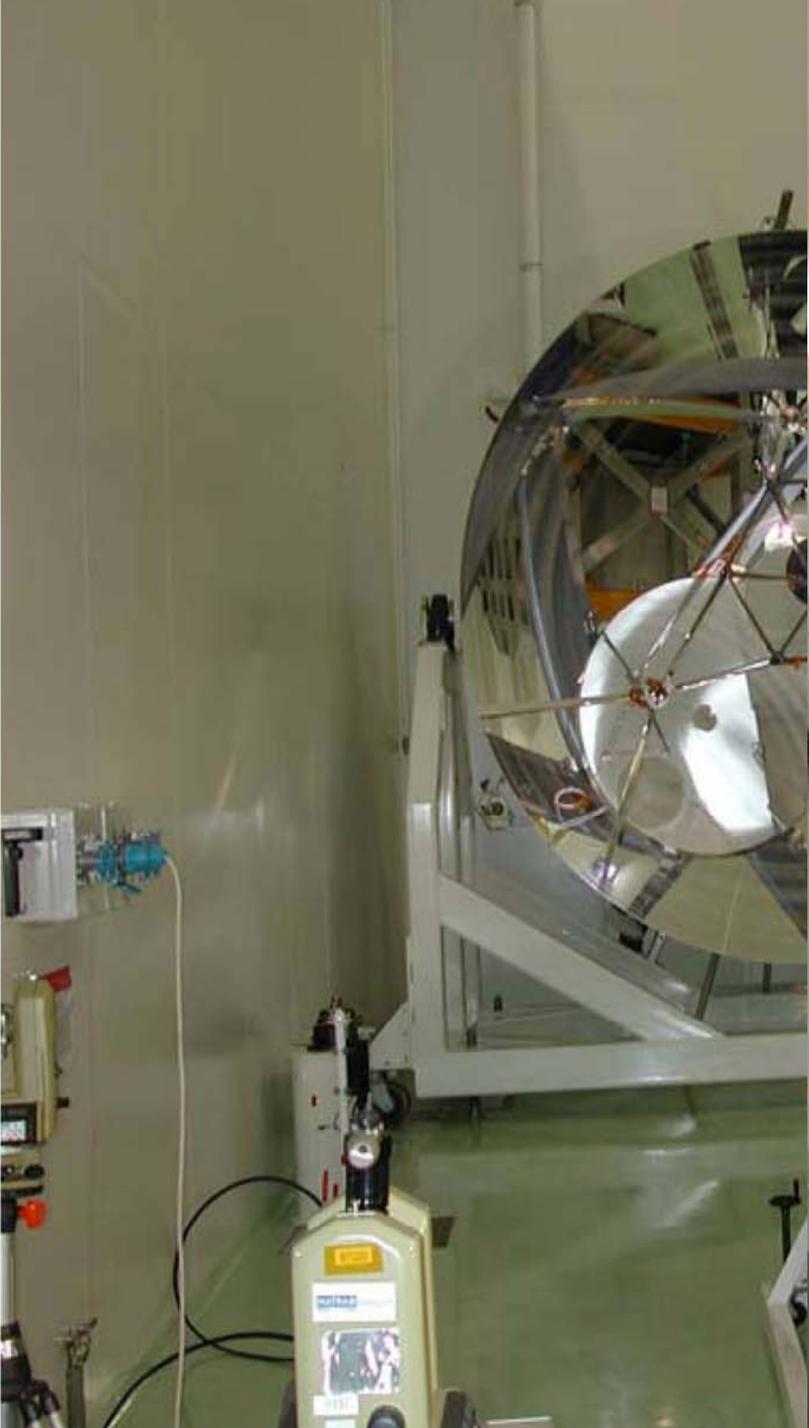


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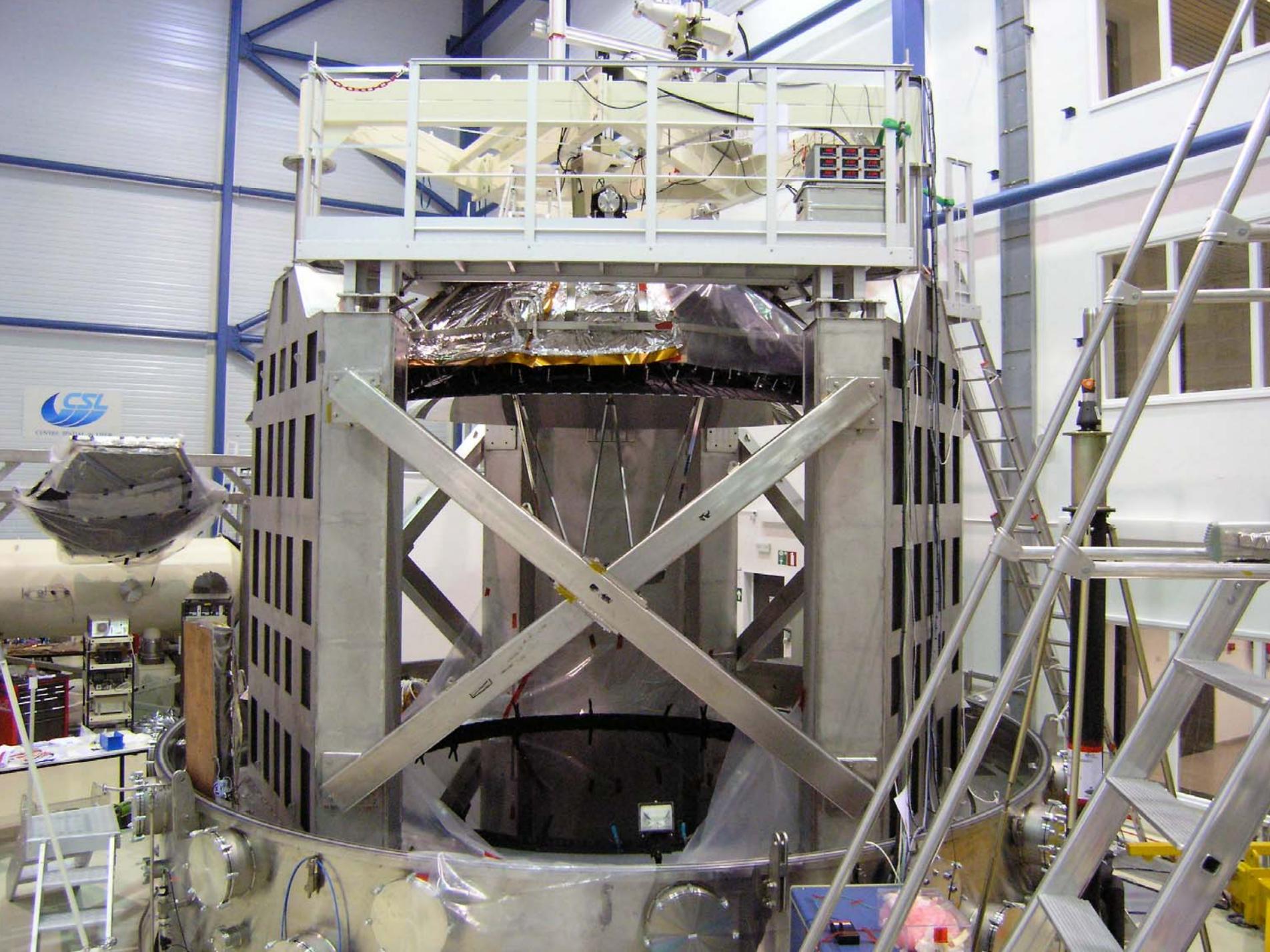


Grinding and lapping/polishing







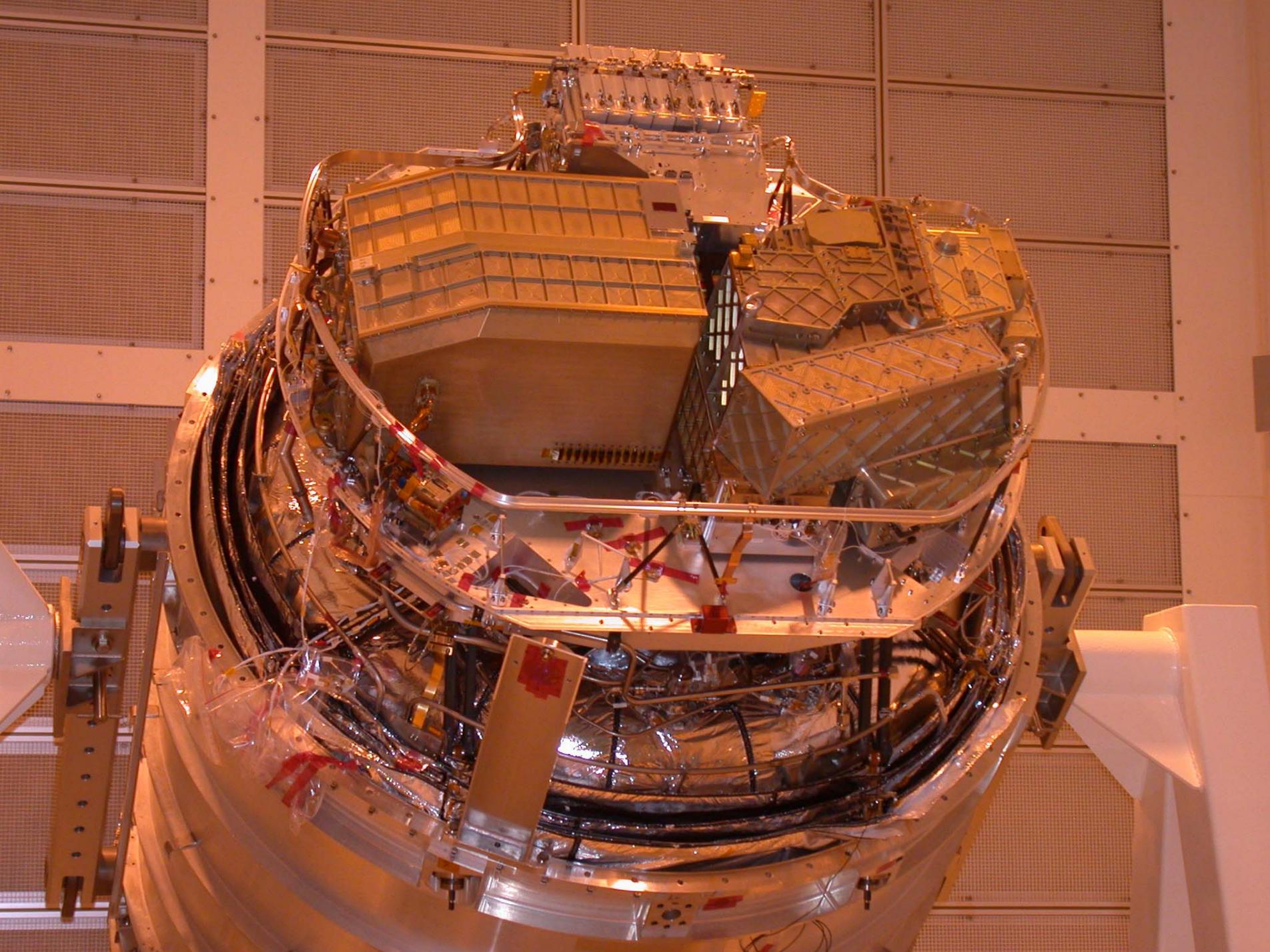


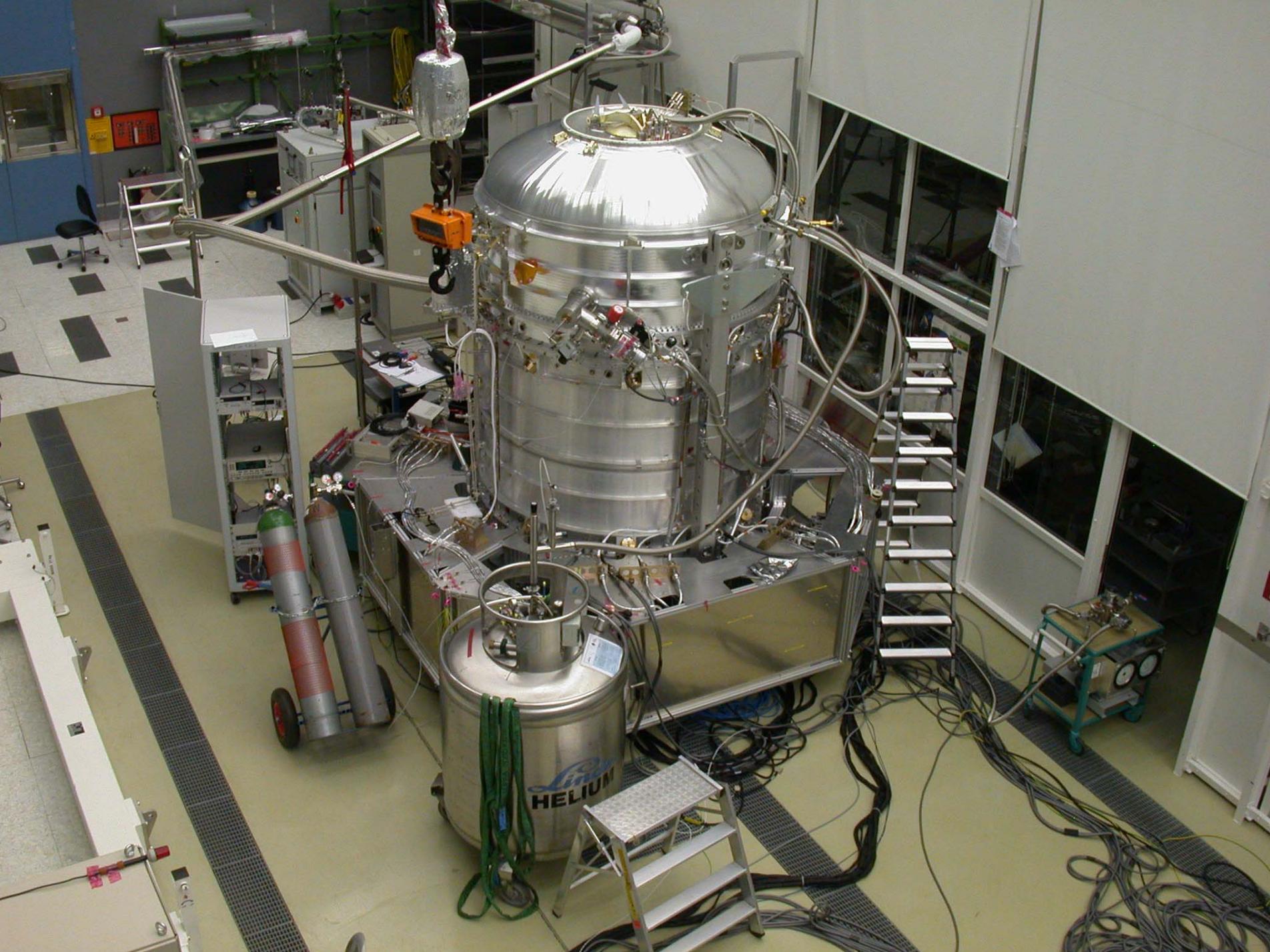
CSL
CENTRO SPAZIALE

Herschel EQM



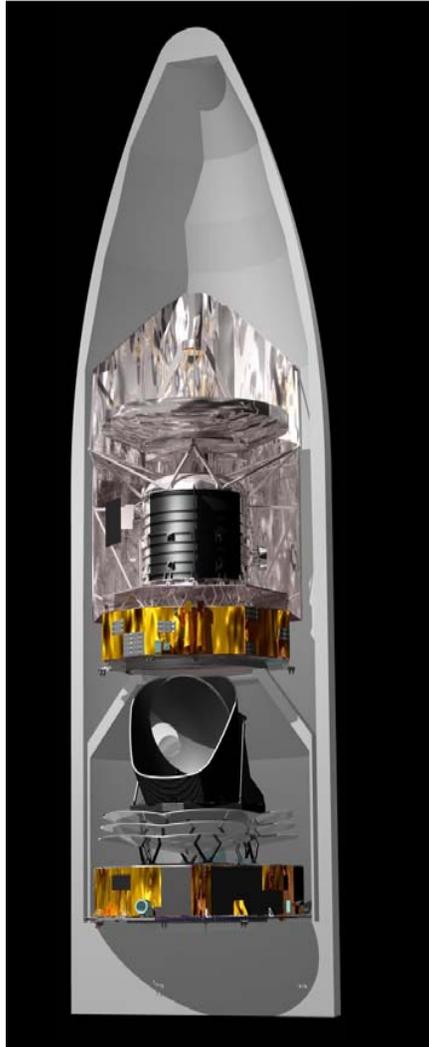
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Launcher



- Launcher version:
 - Ariane 5 - ECA (cryogenic upper stage)
 - qualification flight V164
- Payload configuration:
 - Planck in lower position
 - Herschel in upper position
 - Syllda5/ACU2624
 - Long fairing
- Launch Autonomy:
 - 25 hours
- L2 Injection strategy:
 - direct injection
 - 25 minutes powered phase
- L2 Injected mass capability:
 - ≥ 6273 kg including adaptors



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©CNES-septembre 2002 / illust. D.Ducros





1.5 Mkm

385 Tkm

150 Mkm





Herschel mission phases

- **Launch and early operations (LEOP)**
- **Commissioning and performance verification (SC + payload)**
- **Science demonstration phase**
- **Routine science operations phase (36 months)**
 - **Approx 1000 days / 20000 hours of schedulable science time**
 - **Guaranteed time programmes – GT (32%)**
 - open for GT holders only
 - **Open time programmes – OT (68%)**
 - including discretionary time and targets of opportunity
 - open for all – including GT holders
- **Three ‘Call for proposals’ (AO) cycles are foreseen**
 - one Call for ‘Key Projects’ programmes only (GT and OT)
 - two Calls for regular programmes (GT and OT)
- **Each AO will be divided in two parts**
 - GT awarded first
 - OT awarded after GT in same cycle

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Overall AO schedule

- **Logic: Issue 'Call for Proposals' (AOs) as late as possible**
 - for pure scientific reasons
 - and for performance knowledge reasons
 - but early enough for observers and HSC to prepare
 - to have observations available for scheduling
 - and enable community support staff 'training on the job'
- **1 Feb 2007: Issue AO for 'Key Objects' proposals**
- **5 Apr 2007: Submission deadline for GT KP proposals**
- **5 Jul 2007: Selection & announcement of GT KP programmes**
 - 'effective' AO date for open time proposers
- **25 Oct 2007: Submission deadline for OT KP proposals**
- **28 Feb 2008: Selection & announcement of OT KP programmes**
- **28 Feb 2008: Issue AO for 'Cycle 1 Regular GT' proposals**
- **3 Apr 2008: Submission deadline for GT1 proposals**
- **5 Jun 2008: Selection & announcement of GT1 programme**

AO doc package



- **Cover letter – formal Call by D/SCI**
- **Executive summary**
 - duplicates as ‘AO manual’
 - points to online package with latest news
- **Policies and procedures**
 - politics / policies etc.
- **Herschel Observers’ manual**
 - spacecraft matters (pointing, visibility, etc)
- **3 x Instrument Observers’ manual**
 - PACS, SPIRE, HIFI – ultimately HSC responsibility
- **HSpot manual**
 - + additional tools (background noise estimator, confusion noise estimator, ...)
- **Guaranteed time programme description**
 - with reserved observations lists

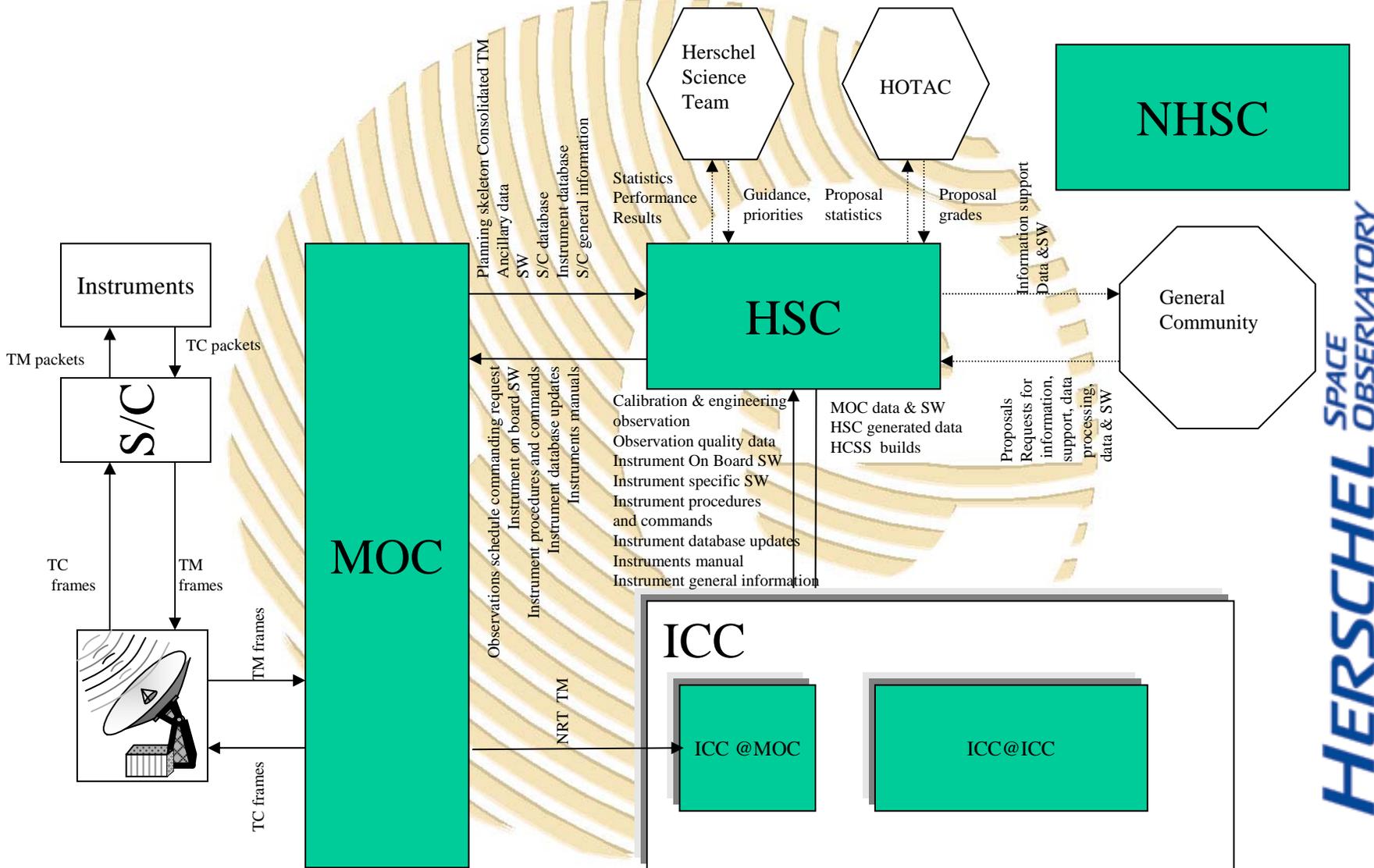
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Herschel 'Key Projects'

- **Foreseen to be important upfront (SMP/instrument AO)**
 - introduced to ensure that 'unusually large' observing programmes can be proposed, selected, and observed
 - need 'pre-identified' due to the nature of the foreseen science objectives and the lack of 'precursor' (IRAS-type) mission
- **Definition of a 'Key Project' programme - it must**
 - exploit unique Herschel capabilities address (an) important scientific issue(s) in a comprehensive manner
 - require a large amount of observing time to be used in a uniform and coherent fashion
 - produce a resulting well characterised dataset of high archival value
- **Data reduction**
 - it is recognised that there is a legitimate science return interest that
 - the data generated by the observations are timely reduced, and
 - the data products and tools are made public
 - therefore 'Key Project' consortia must demonstrate commitment and ability to perform data reduction, and must make data products and tools publicly available at the end of the proprietary time period

Ground segment



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PACS Photometry

Unique AOR Label: PACSPHOTO-0000

Target: None Specified

New Target

Number of visible stars for the target: 15
Star tracker target: Ra: 22.482 degrees Dec: -47.232 degrees

Instrument Settings

Blue channel

60-85 microns filter (Blue 1)
 85-130 microns filter (Blue 2)

For each pointing
On-source integration time (s): 60

Observing Mode Settings

Source

Map choice

Pointed
 Small Map
 Large Map

Pointing settings

Nodding
 On
 Off

Large Map
 On
 Off

Source Flux Estimation

Bright Moving Objects to Avoid

Earth+Moon
 Others

Chopping and Dithering settings

Chopping

On
 Off

Chopper avoidance angle (deg)

Angle from (degrees): 0.00
Angle to (degrees): 0.00

Dithering

On
 Off

Observation Est... Add Comments... Visibility... Star Tracker...

OK Cancel

PACS Photometry

Unique AOR Label: PACSPHOTO-0000

Target: m51 Type: Fixed Single
Position: 13h29m55.73000s,+47d13m53.4000s

New Target Modify Target Target List...

Number of visible stars for the target: 15
Star tracker target: Ra: 22.482 degrees Dec: -47.232 degrees

Instrument Settings

Blue channel

60-85 microns filter (Blue 1)
 85-130 microns filter (Blue 2)

For each pointing
On-source integration time (s): 60

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 Small Map
 Large Map

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 Off

Large Map
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On
 Off

Chopper avoidance angle (deg)

Angle from (degrees): 0.00
Angle to (degrees): 0.00

Dithering

On
 Off

Observation Est... Add Comments... Visibility... Star Tracker...

OK Cancel



SPACE OBSERVATORY

Target

Target Name (required): NED

M51

Fixed Moving

Coord Sys: Equatorial J2000

Proper Motion

Use P

PM RA (" /yr):
PM Dec (" /yr):

Epoch: 2000.00

Bright Moving Objects to Avoid

Earth+Moon
 Others

OK Cancel

PACS Time Estimation

Time

Observatory overhead (s): 180
On-source time (s): 48
Calibration time (s): 4
Total time (s): 283

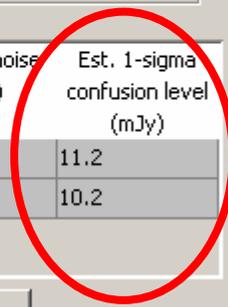
Band (um)	Est. 1-sigma noise (mJy)	Est. 1-sigma confusion level (mJy)
60-85	3.1	11.2
130-210	3.2	10.2

Ned Name Resolver- Working...

Ned Name Resolver
Searching for Name: M51

Click **Abort** to stop.

Abort





Data processing – HCSS/DP

- Ensemble of services in a single, coherent platform-independent system – HCSS/DP
- Offering the astronomical community the means (data, products, software) to do science without the need to buy licenses
- Source of software, calibration and documentation to reduce and do science with Herschel data ('Observer IA')
- Source of software and documentation to support the commissioning of the instruments, the validation of the observing modes and the calibration of satellite pointing and instruments ('Calibration IA')
- Toolbox to develop data processing algorithms
- Generation of standard observation products and quality information by using IA modules ('SPG' & 'QC' pipelines)
- On-line, context-sensitive and printable help documentation for end-users and developers based on XML/DocBook
- Data and services Virtual Observatory compliant + FITS

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Herschel data processing

ESA Home Research & Science Home
esa Herschel
 Astrophysics Planetary Exploration Solar System

HSCDT Team Resources
 HSCDT Public Home
 HSCDT Internal Home
 Contact List
 Contact List (HIFI)
 Contact List (PACS)
 Contact List (SPIRE)
 Distribution Lists (PACS)
 FTP area
 Installation Support
 Updates
 HSCDT Documentation
 Choose One
 HCSS Release Doc. Search
 HSCDT Software
 Choose One
 Problem Reporting
 Choose One
 Herschel Related Links
 Herschel Science Home
 HIFI Home
 PACS Home
 SPIRE Home
 RSSD Services
 My Portal
 Livelink
 Find Us
 Who's who
 RSSD Web Search
 a Search Scope
 Herschel
 Advanced Search

Average hot/cold frames for 242 Y-Factor

Relative counts Hot vs Cold (raw)

Max VSI spectrum values

Strip chart
 Dec Mec
 Colour range
 Instrument array
 Line summary
 Data selector

Linearity test
 Demodulated signal
 Script debugger for Interactive Analysis

```

def beam_atten_pipeline5.py
"""
"""
import numpy as np
import sys
import os

def main():
    """
    """
    # Get the wavelength
    lambda = float(sys.argv[1])

    # Get the beam size
    beam_size = float(sys.argv[2])

    # Get the offset
    offset = float(sys.argv[3])

    # Get the data
    data = np.loadtxt(sys.argv[4])

    # Calculate the beam attenuation
    beam_atten = 1 - np.exp(-lambda / beam_size)

    # Calculate the demodulated signal
    demodulated_signal = data * beam_atten

    # Save the demodulated signal
    np.savetxt(sys.argv[5], demodulated_signal)

if __name__ == '__main__':
    main()
  
```

- **DP-WG MoM**: DP Working Group Minutes of Meetings
- **Livelink documents**: Other related documents under Livelink, including DP CCB and User Group documentation.

Last updated: Friday, 26-Aug-05 13:02:20



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Herschel key dates

- FIRST feasibility study: 1982-83
- ...
- FIRST confirmed 4th Cornerstone: Dec 1993
- ...
- Selection of science instruments/PIs: 1997-98
- FIRST becomes Herschel: Dec 2000
- Start of spacecraft Phase B: mid 2001
- Start of spacecraft Phase C/D: early 2003
- SVM, PLM, telescope, CQM/EQM testing 2005
- SVM, PLM, instruments, telescope deliveries: 2006/07
- Integration & tests/verification: 2007
- Issue of first AO for proposals: 1 Feb 2007
- Launch: (not before 31 Jul) 2008
- Early operations: first 6 months
- Routine Science Operations: 2009 - 2012

HERSCHEL SPACE OBSERVATORY

Herschel Latest News



PACS FM instrument first light! PACS photometer first light was achieved on 15 November 2006 (left picture). The moment the beam was witnessed among other people (middle picture, left to right) Marc Sauvage, Koryo Okumura, and Roland Wiegand (standing). The photometer first light video is available here (avi, 11Mb). The PACS spectrometer first light was achieved only two days later, on 17 November 2006 (right picture). This was a spectrometer grating scan looking at a 30K optical ground support equipment black body source. In the picture 25 simultaneous spectra can be seen (confirming all 25 modules working!) and in the inset the time evolution of the signal in a single pixel. [November 2006]



Herschel FM telescope at Flight Qualification Review (FQR). The telescope being inspected in the Astrium, Toulouse facilities as part of the FQR on 9 November 2006. In left picture from left to right: Dominic Doyle, Dominique Pierot, Jacqueline Fischer, Göran Pilbratt, Yves Toulemont, Gerald Crone, Pierre Olivier, and Jean-Luc Varin. In middle picture M2 and its supporting structure is inspected, in right picture the 'scatter cone' can be seen. [November 2006]

www.rssd.esa.int/herschel
www.herschel.esac.esa.int
Initial AO foreseen 1 Feb 2007

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