



***Herschel* mission: Overview and observing opportunities**

Observing opportunities with *Herschel*: A window to the far-infrared universe – an ASTROCAM workshop

ESAC, Villafranca, 14-15 December 2006

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Research and Scientific Support Department

Herschel in a nutshell



- **ESA cornerstone observatory**
 - instruments 'nationally' funded, int'l - NASA, CSA, Poland – collaboration
 - ~1/3 guaranteed time, ~2/3 open time
- **FIR (57 - 670 μm) space facility**
 - large (3.5 m) monolithic low emissivity passively cooled telescope
 - 3 focal plane science instruments
 - 3 years routine operational lifetime
 - full spectral access
 - low and stable background
- **Unique and complementary**
 - for $\lambda < 200 \mu\text{m}$ larger aperture than cryogenically cooled telescopes (IRAS, ISO, Spitzer, Akari,...)
 - more observing time than balloon- and/or air-borne instruments
 - larger field of view than interferometers
- **Launch in 2008**
 - initial observing AO on 1 Feb 2007

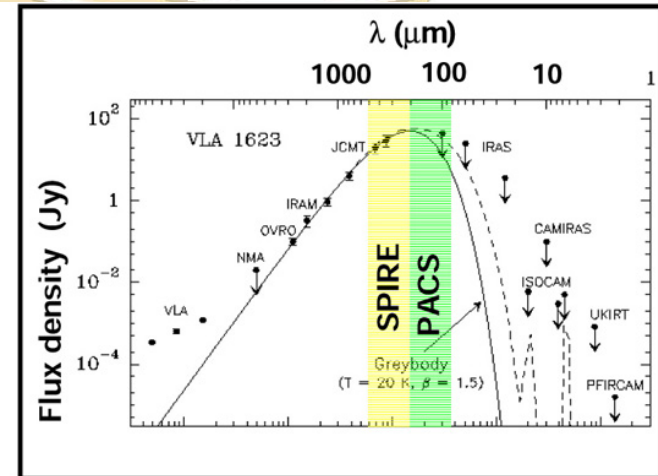
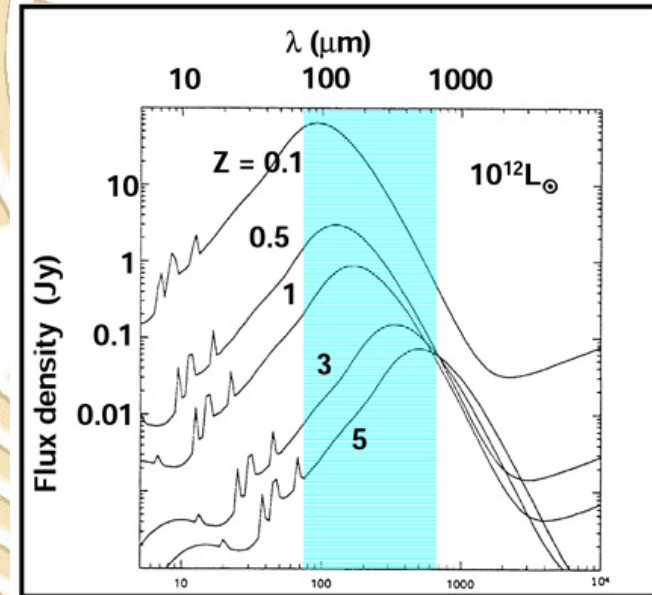


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The Cool Universe

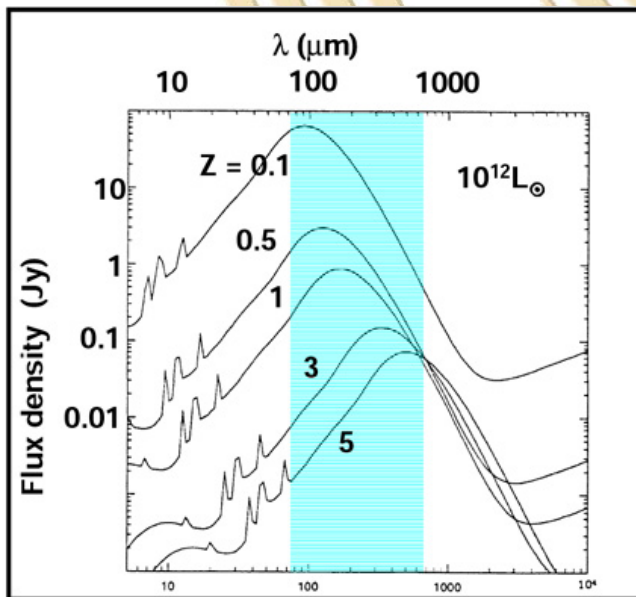
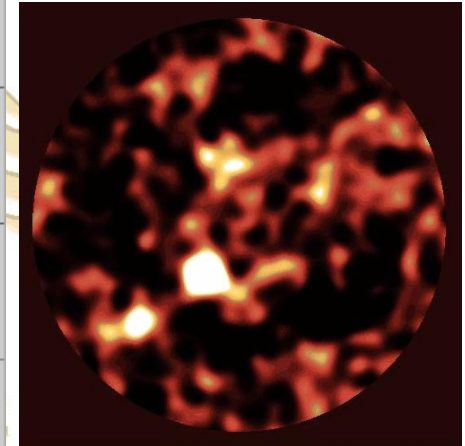
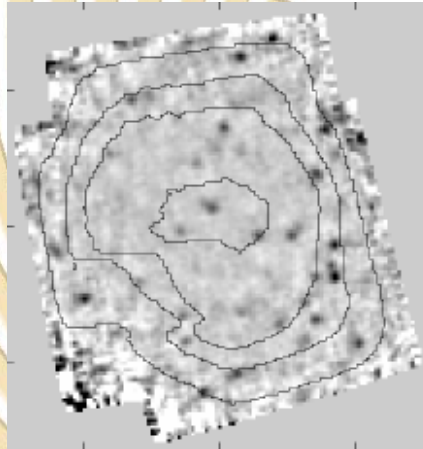
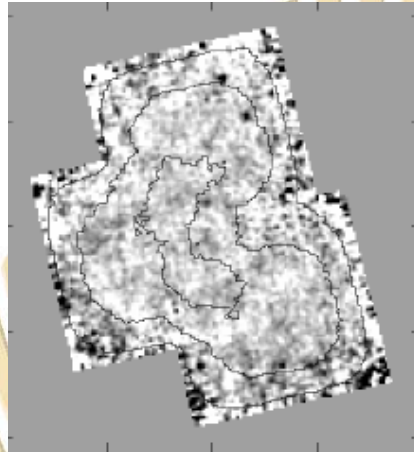


- **Herschel spectral coverage**
 - black-bodies 5-50 K
 - continuum radiation
 - dust grains (re-)radiating
 - gases 10-few100 K
 - brightest atomic/molecular lines
- **Herschel strengths**
 - covers IR dominated galaxies & protostar SED peaks
 - wide area mapping
 - full coverage spectral scans & particular (water) lines
- **Herschel emphasis**
 - formation and evolution of galaxies & stars
 - ISM physics & chemistry
 - solar system bodies



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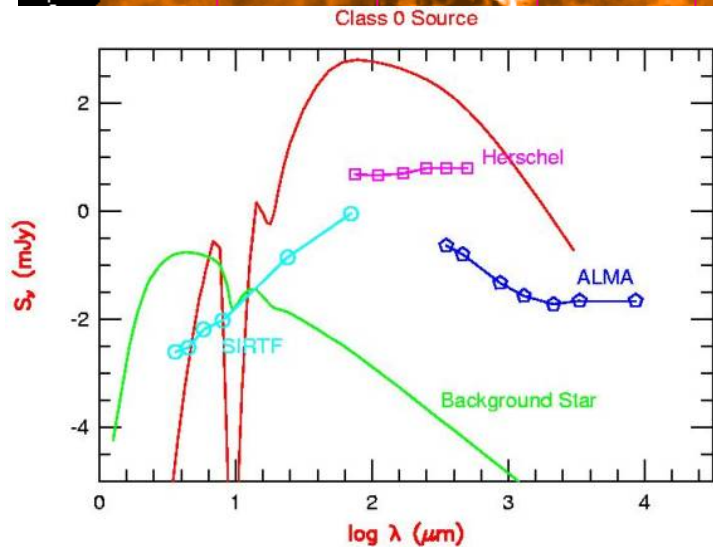
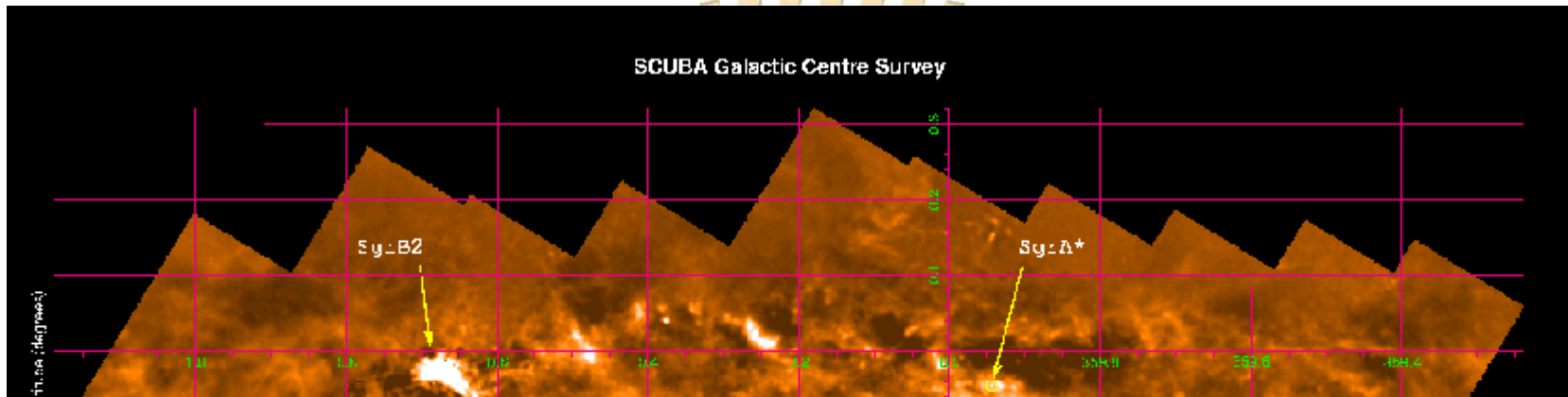
Galaxy formation & evolution



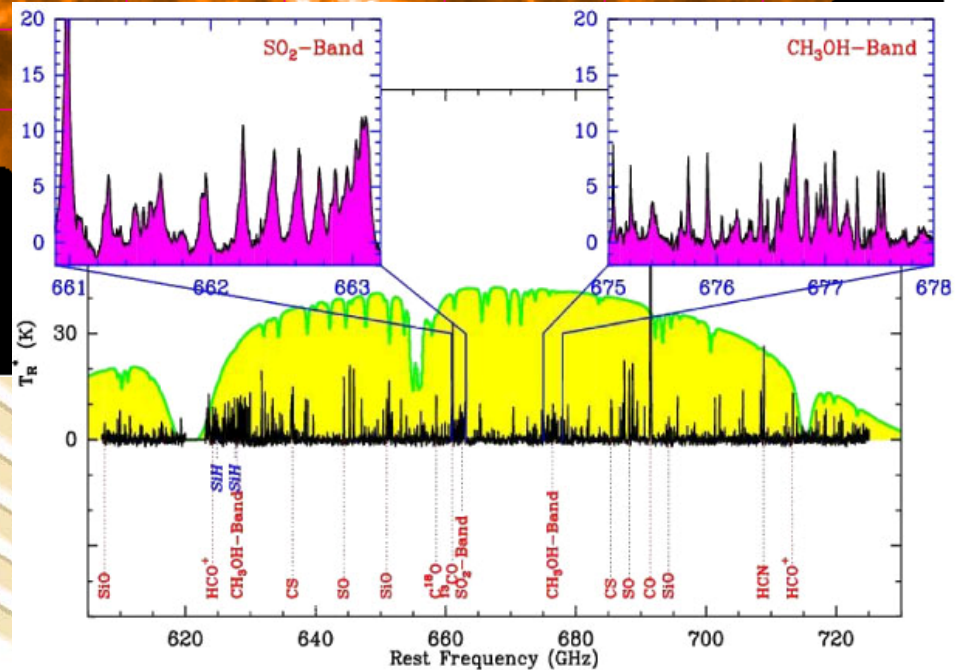
Early Spitzer results:

- 24 μm – 70%
- 70 μm – 23%
- 160 μm – 7%
- but now claiming all resolved!

Surveys & follow-up



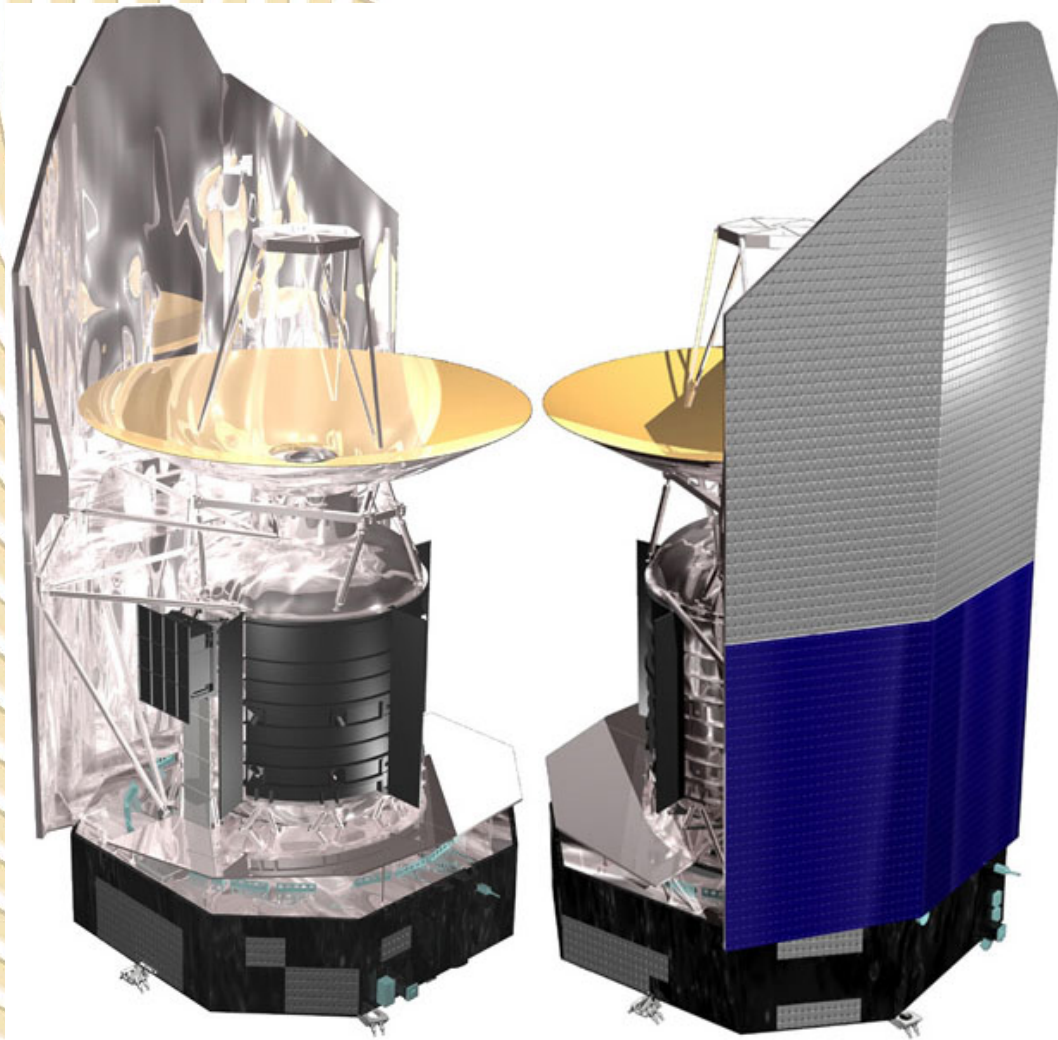
0.1 L_{sun} protostar at 300 pc



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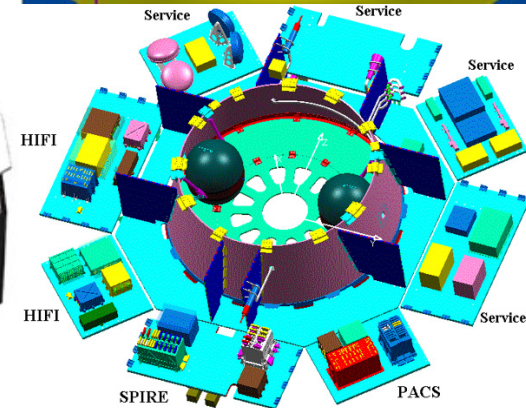
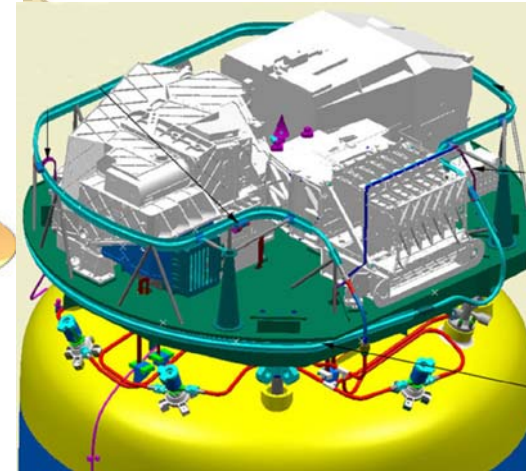
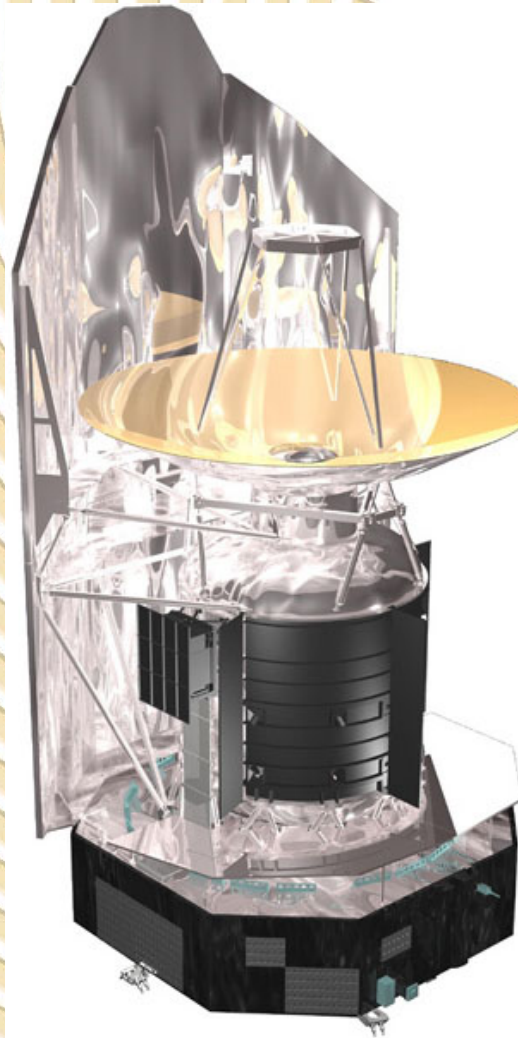
Herschel spacecraft specs

- telescope (eff) diam (3.3) 3.5 m
- *telescope WFE* $< 6 \mu\text{m}$
- *telescope temp* $< 90 \text{ K}$
- *telescope emissivity* $< 4\%$
- *abs/rel pointg (68%)* $< 3.7'' / 0.3''$
- science instruments 3
- science data rate 130 kbps
- *cryostat lifetime* $> 3.5 \text{ years}$
- height / width $\sim 7.5 / 4 \text{ m}$
- launch mass $\sim 3200 \text{ kg}$
- power $\sim 1500 \text{ W}$
- orbit 'large' Lissajous around L2
- solar aspect angle 60-120 deg
- launcher (w Planck) Ariane 5 ECA

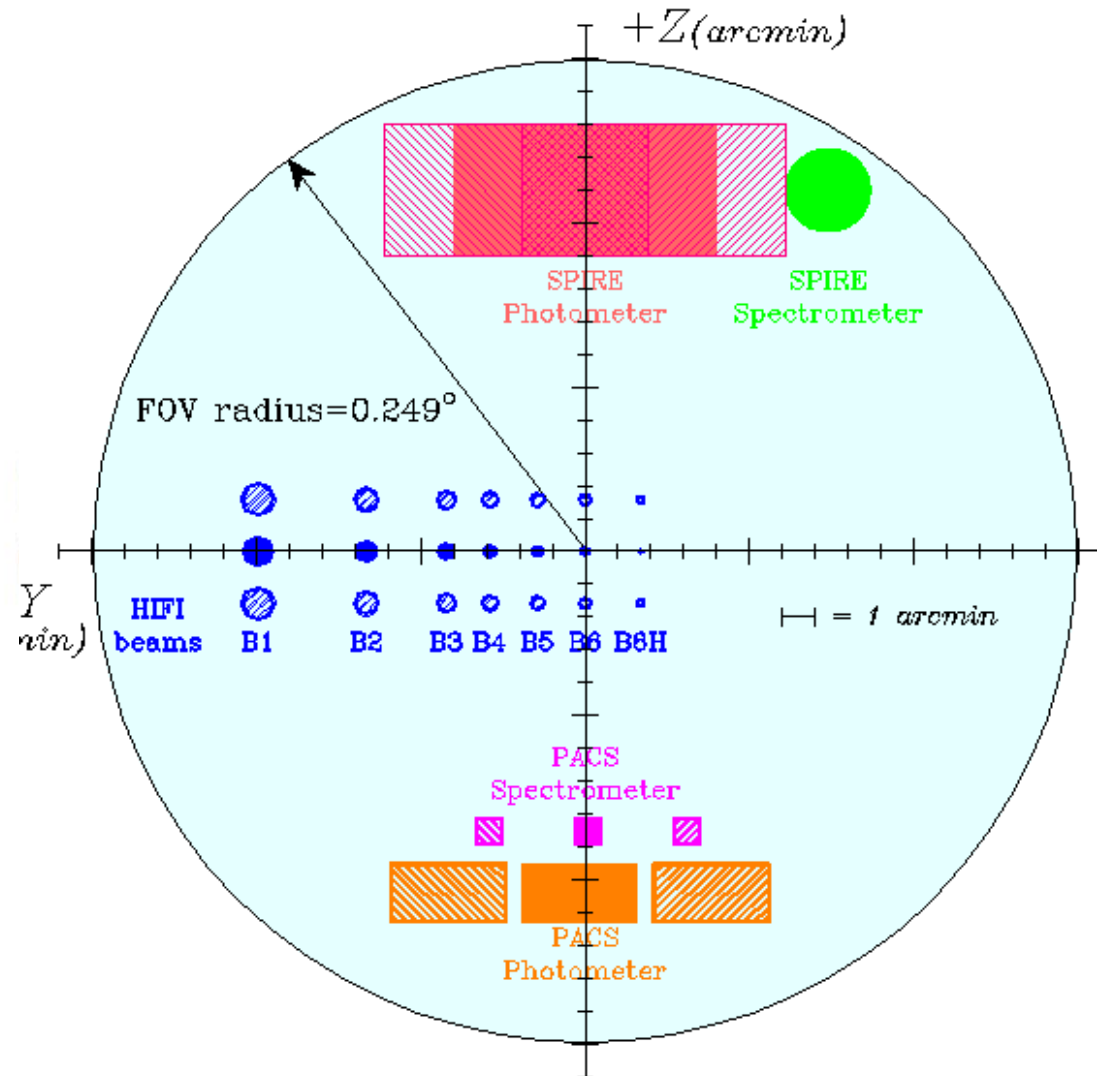


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Herschel focal plane

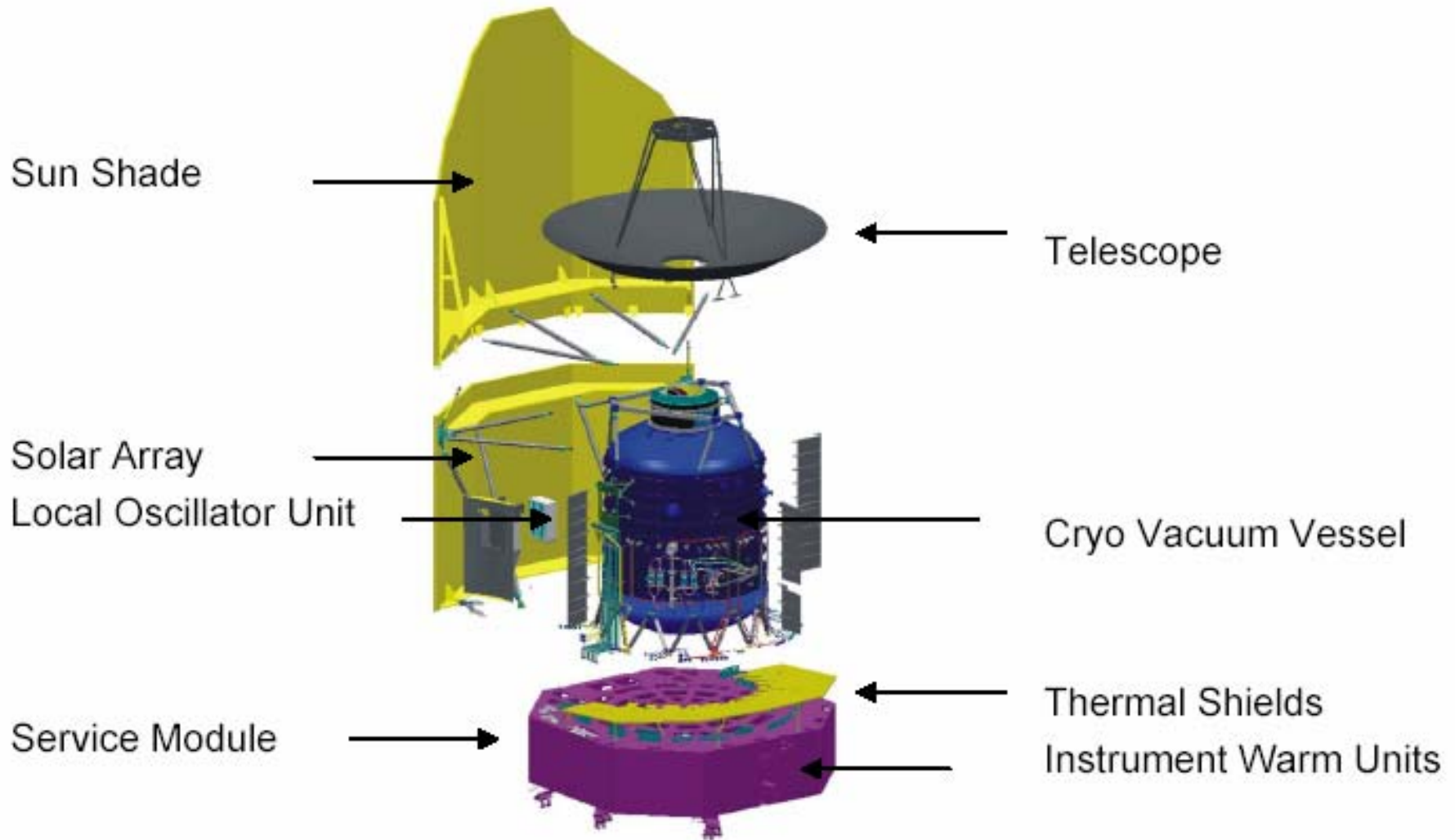




Herschel observatory capabilities

- **Photometry - imaging, 6 broad bands in 75-500 μm range**
 - **PACS** – simultaneous 2 colour fully-sampled ($0.5F\lambda$) imaging with FOV 1.75×3.5 arcmin and $R \sim 2.5$ centred at 75/110 and 170 μm
 - **SPIRE** – simultaneous 3 colour $2F\lambda$ imaging with FOV 4×4 arcmin and $R \sim 3$ centred at 250, 363, and 517 μm
 - for larger fields 'on-the-fly' mapping, mosaicing
 - sensitivity is somewhat wavelength and observing mode dependant, very roughly for point sources $1\text{mJy} - 1\sigma - 1$ hour; for mapping confusion limit is important
- **Spectroscopy - in 57-670 μm range, varying R in $20-10^7$ range**
 - **PACS** – grating spectrometer, 5×5 spatial $\times 16$ spectral pixels, FOV 0.8 arcmin, $R \sim 1500-4000$, $\lambda \sim 57-210 \mu\text{m}$
 - **SPIRE** – FTS spectrometer, $R \sim 20-100+$, FOV 2.6 arcmin, $\lambda \sim 200-670 \mu\text{m}$
 - **HIFI** – interferodyne spectroscopy with R up to 10^7 , $\lambda \sim 157-212$ and $240-625 \mu\text{m}$, 2 orthogonal polarisations, 4000 spectral channels per polarisation, single pixel on the sky, mapping by 'on-the fly' or mosaicing observations

Herschel spacecraft





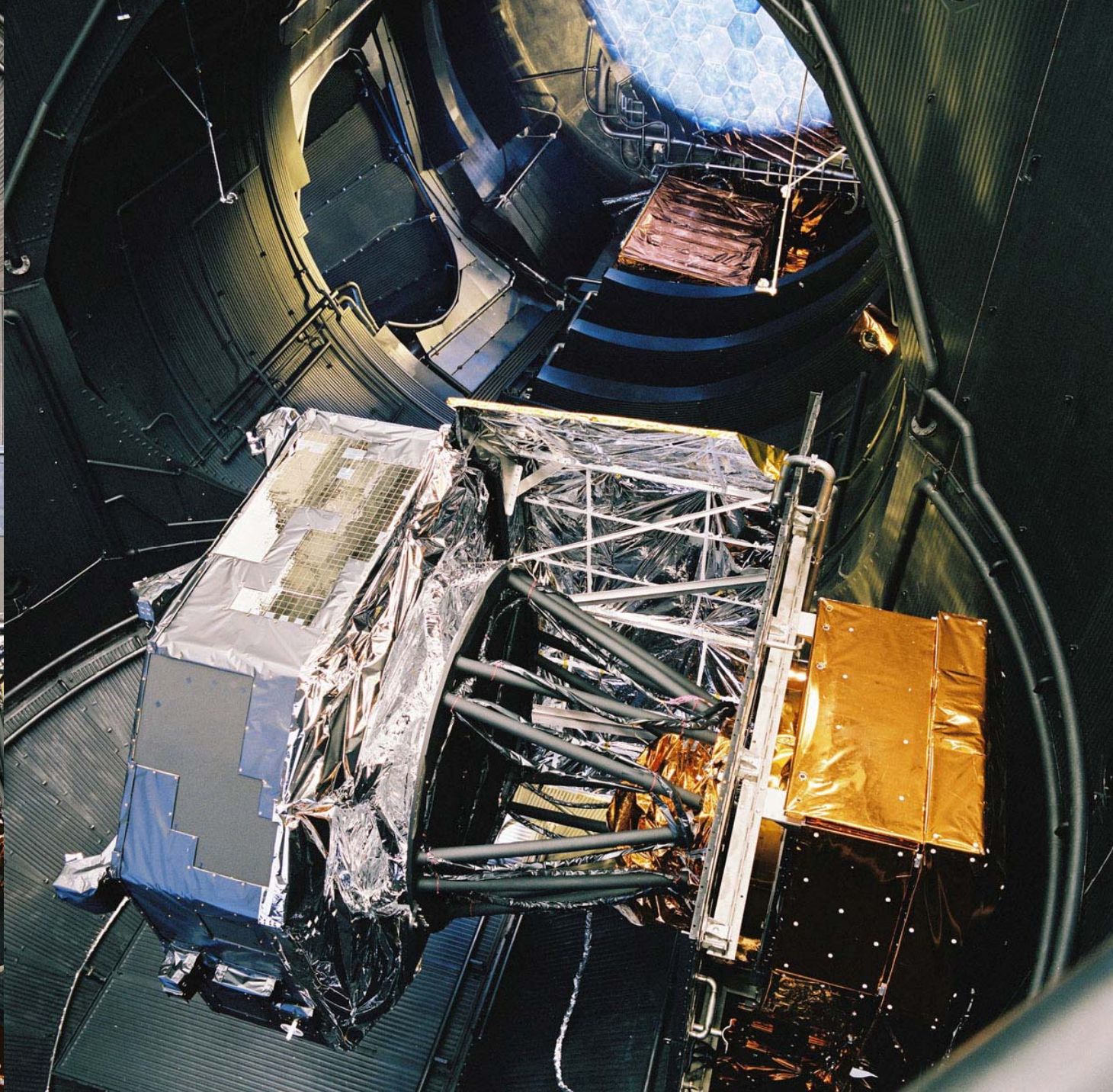
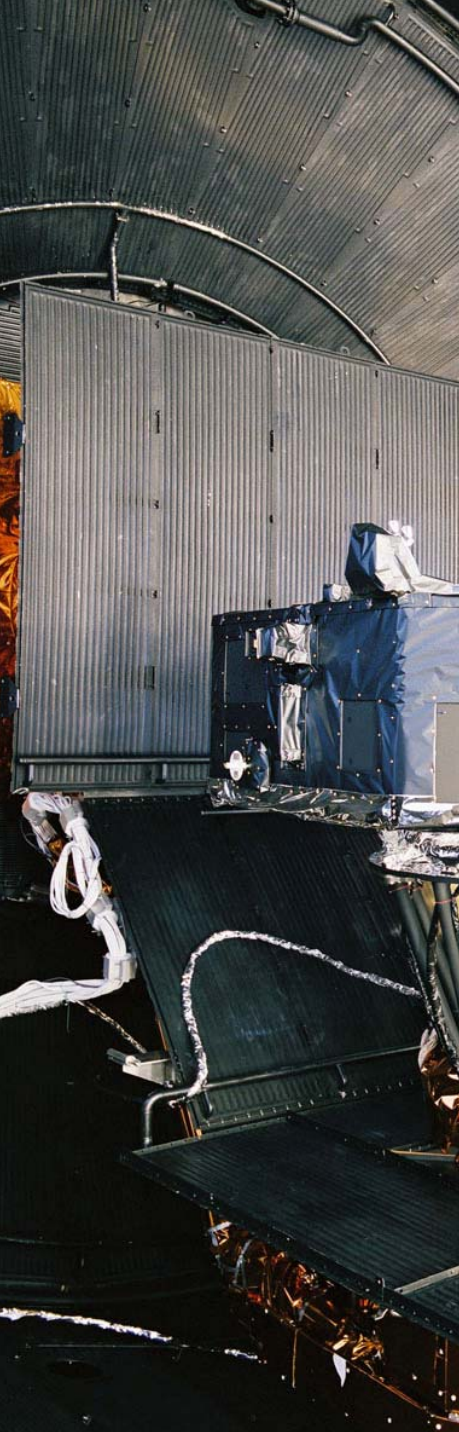
Herschel spacecraft programmes

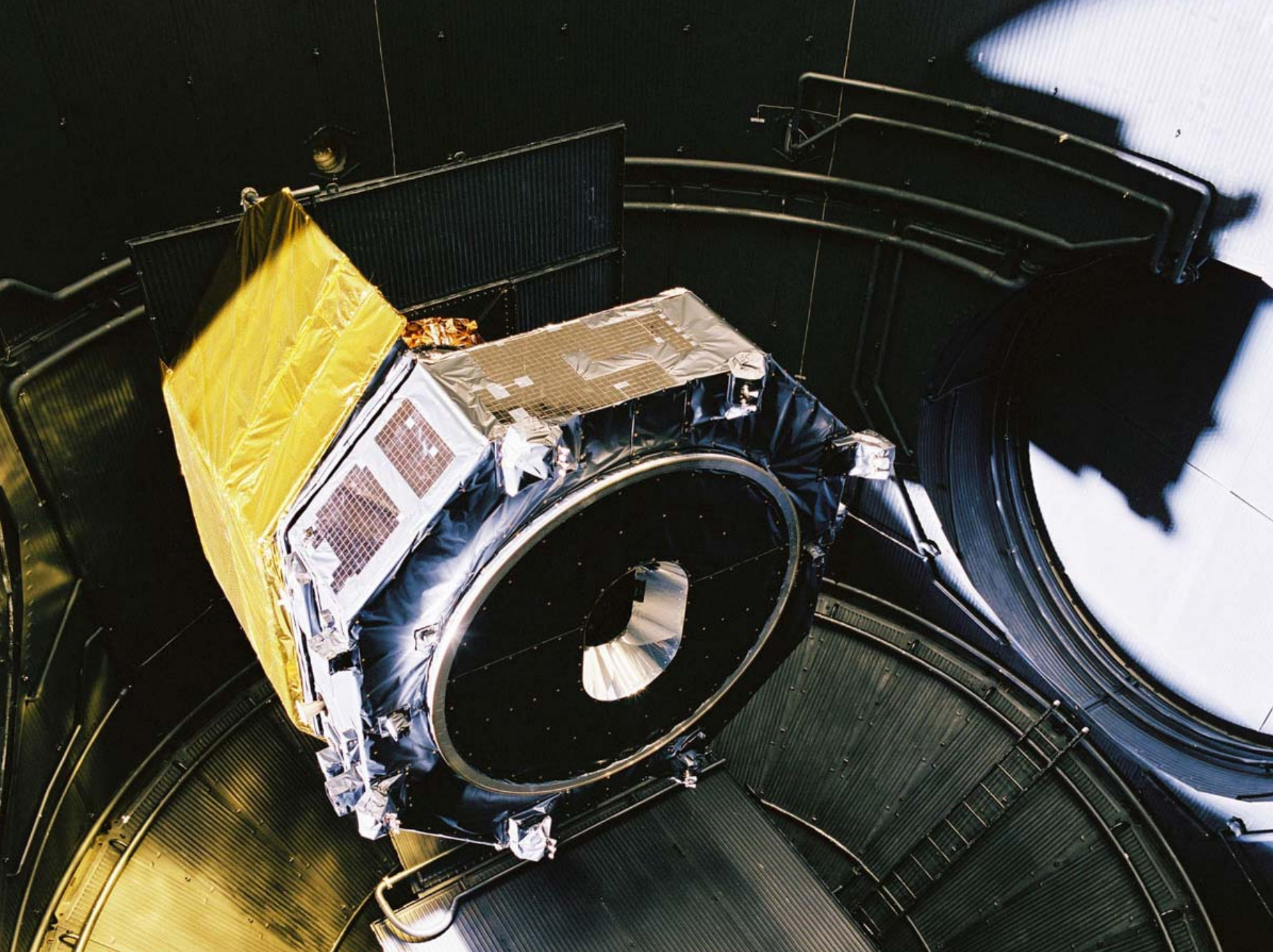
- **Satellite PFM programme**
 - two SVMs – QM and FM – with many common parts
 - only one cryostat/PLM is built
 - only one satellite is built – tested in STM configuration
 - FM cryostat
 - STM models of everything else
- **Telescope PFM programme**
 - telescope manufacture complete – has undergone lengthy cryo-testing and characterisation campaign in CSL
 - spare parts exist for FS
- **EQM programme – completed**
 - ISO hardware refurbished to provide ‘Herschel OB simulator’
 - science instrument CQMs
- **AVM programme**
 - one physical setup – configurable as either Herschel or Planck
 - test procedures and OBSW development

Herschel SVM/STM

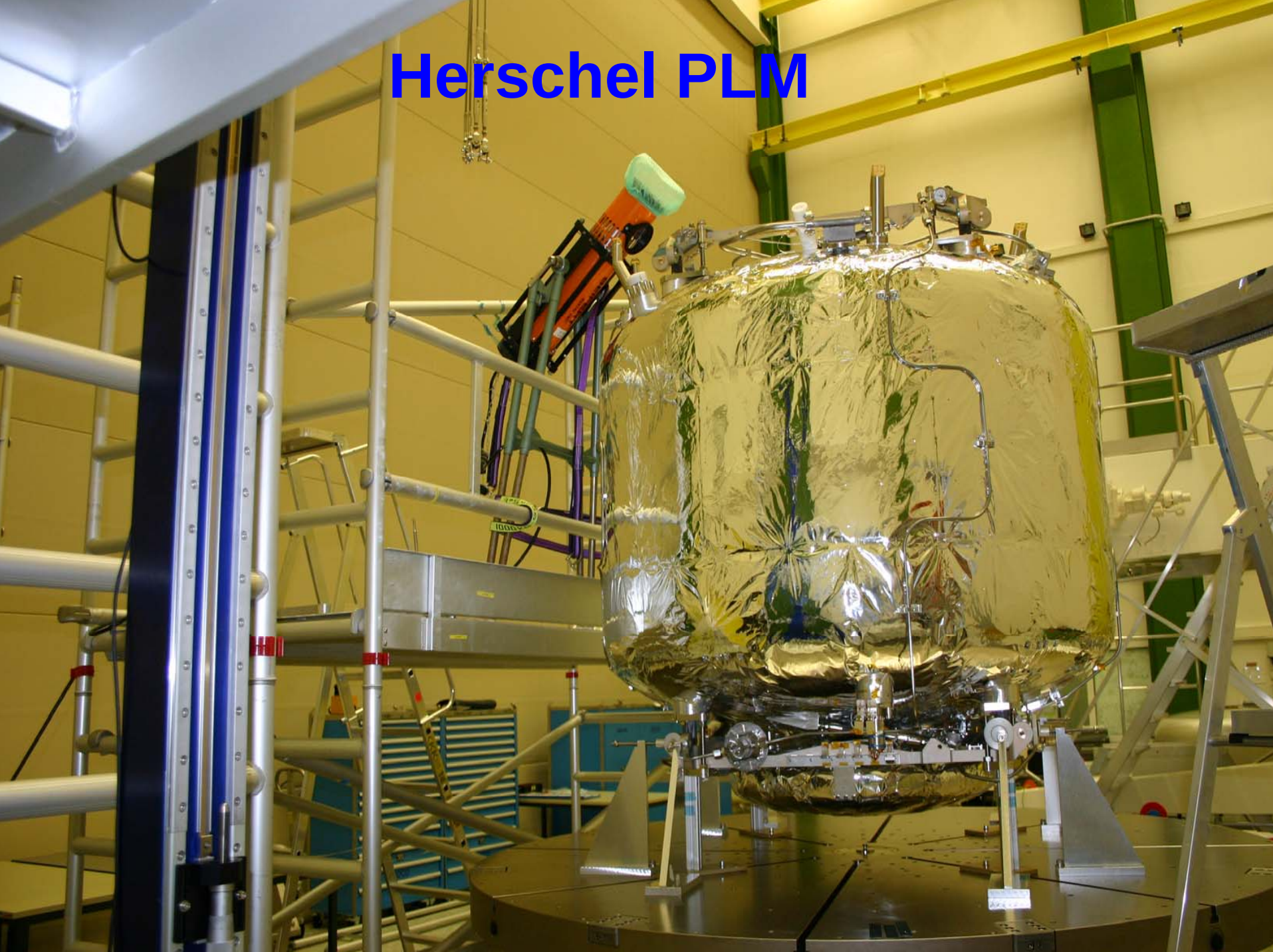






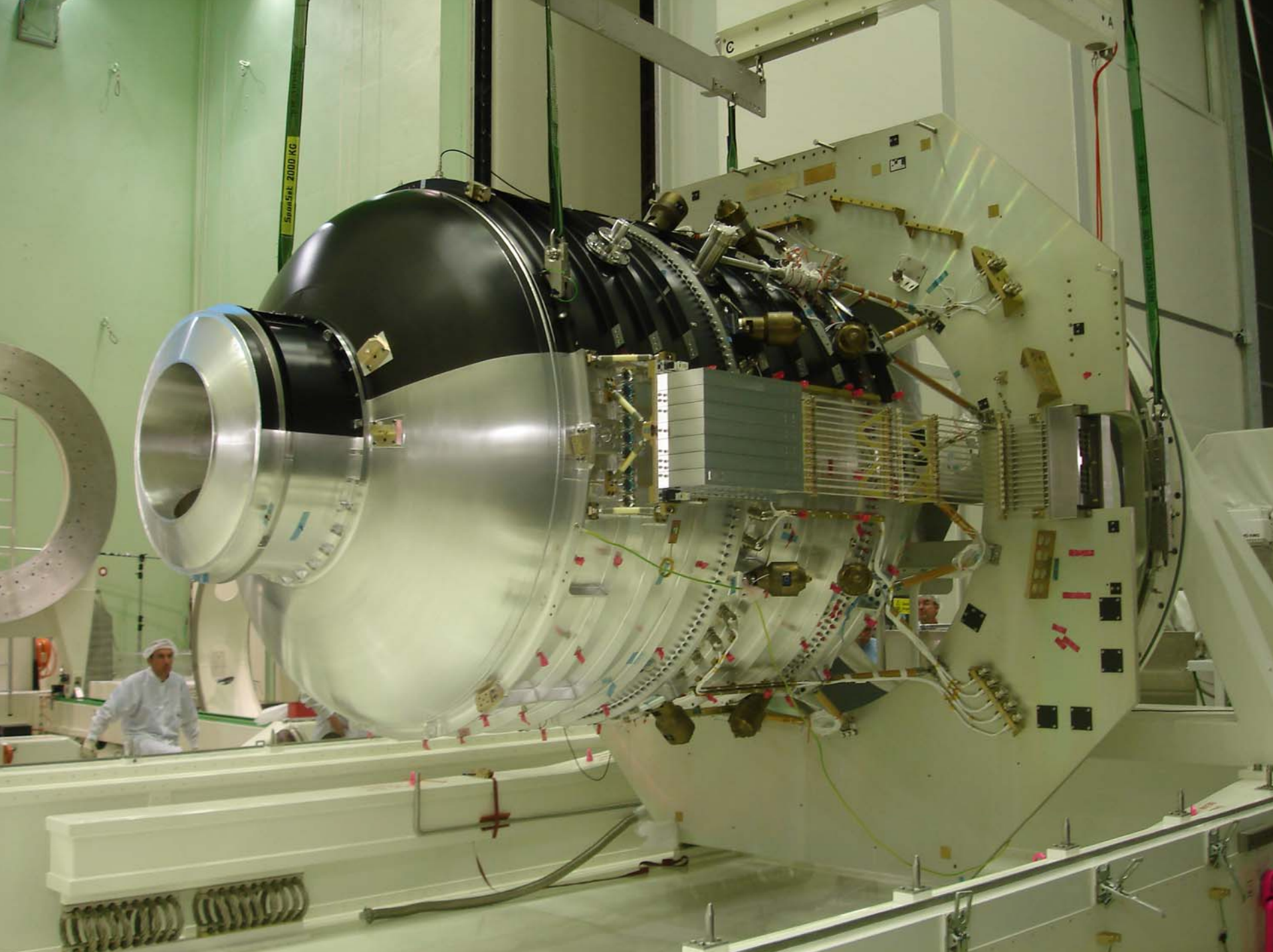


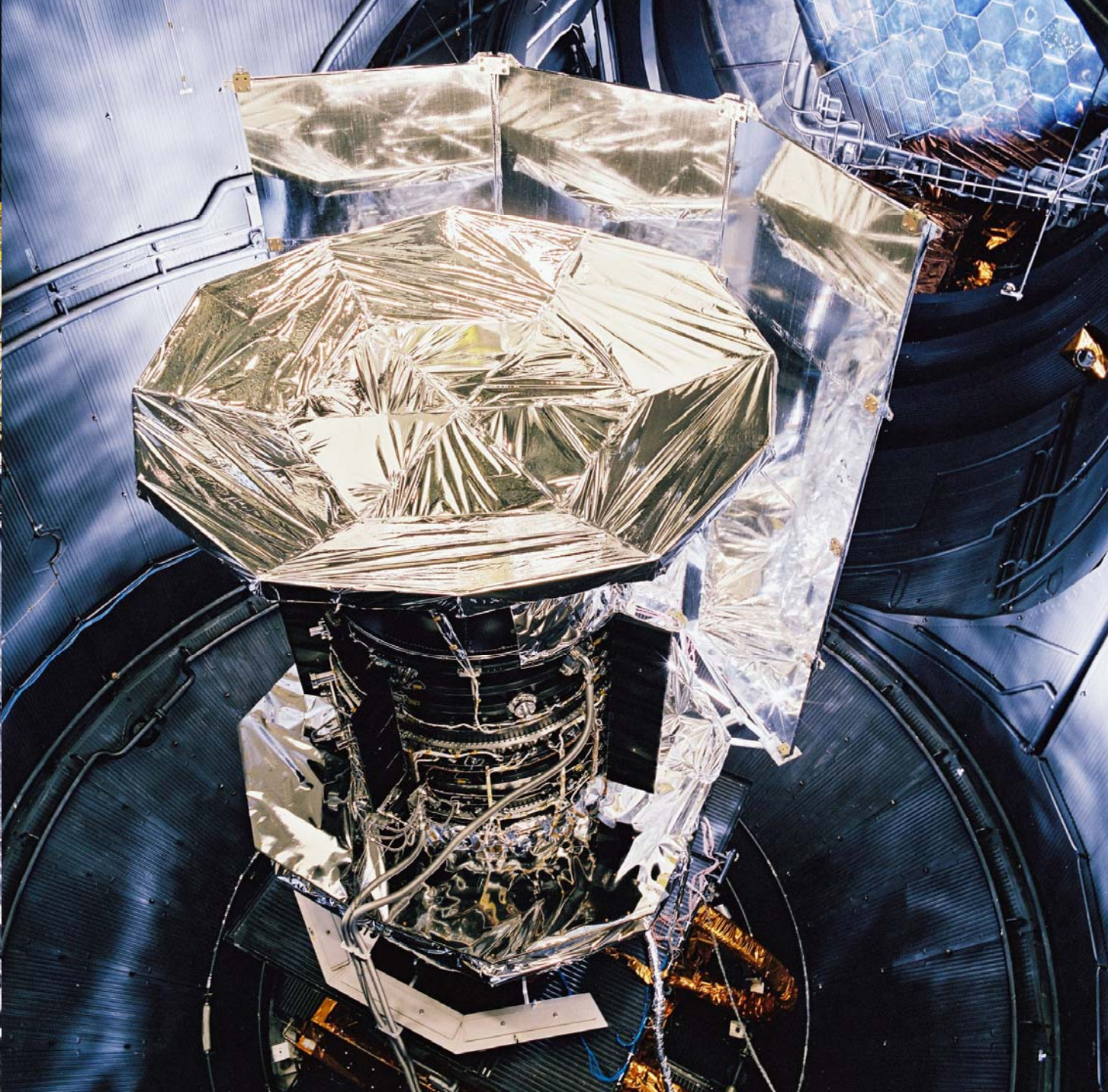
Herschel PLM













Herschel STM satellite

Herschel
Science Team
visited the
ESTEC Test
Centre to view
the Herschel
STM satellite
on 1 Feb 2006

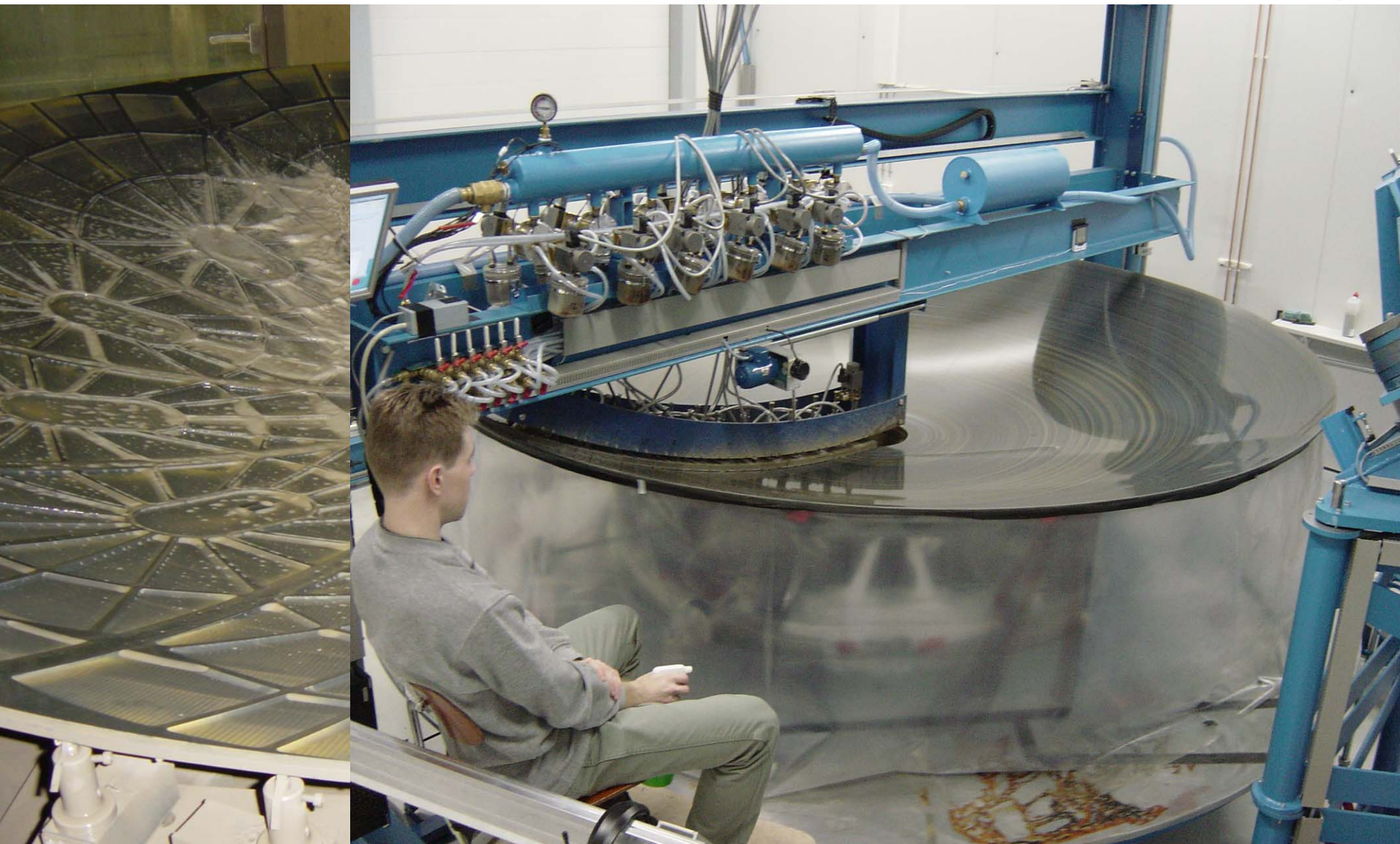


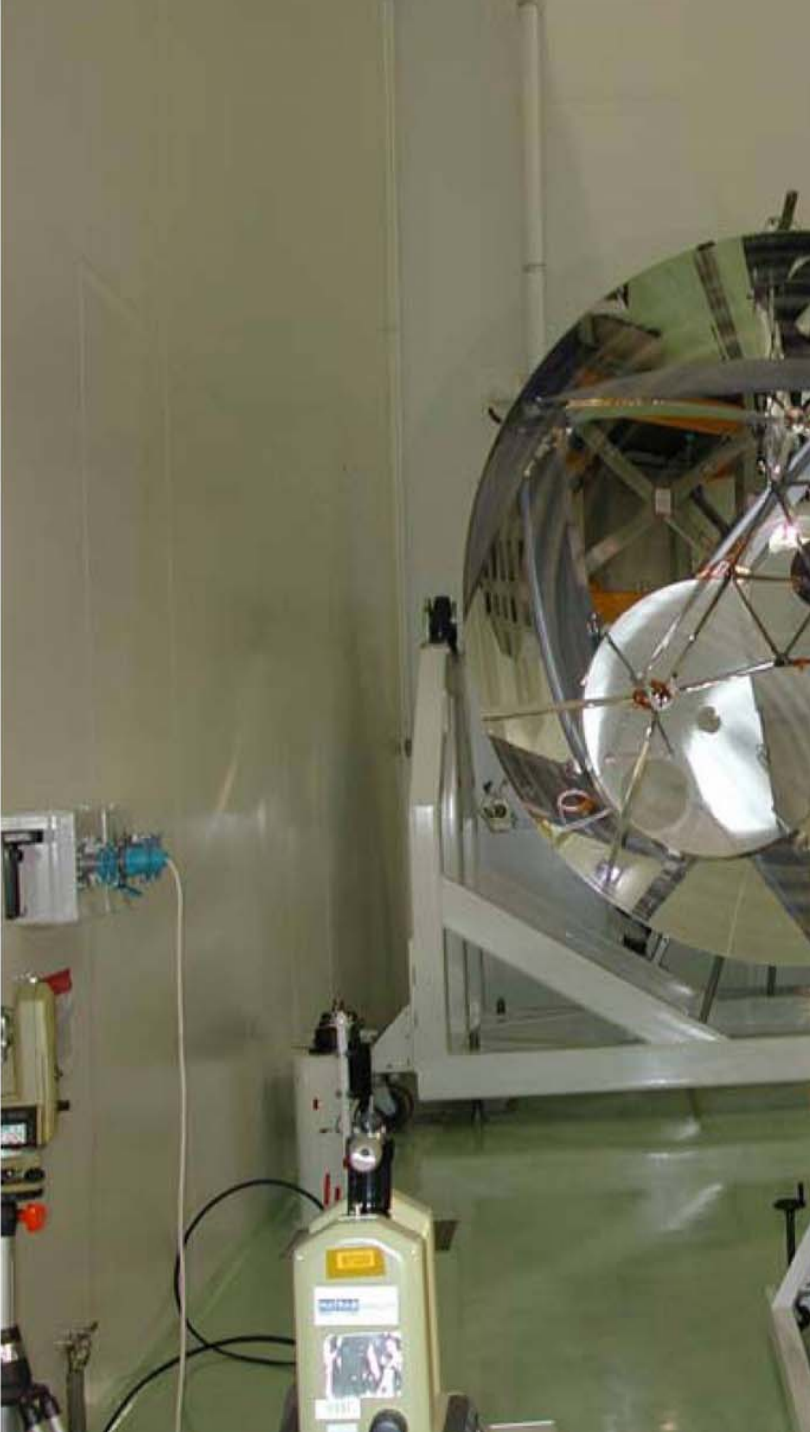
Brazing of primary mirror 'blank'



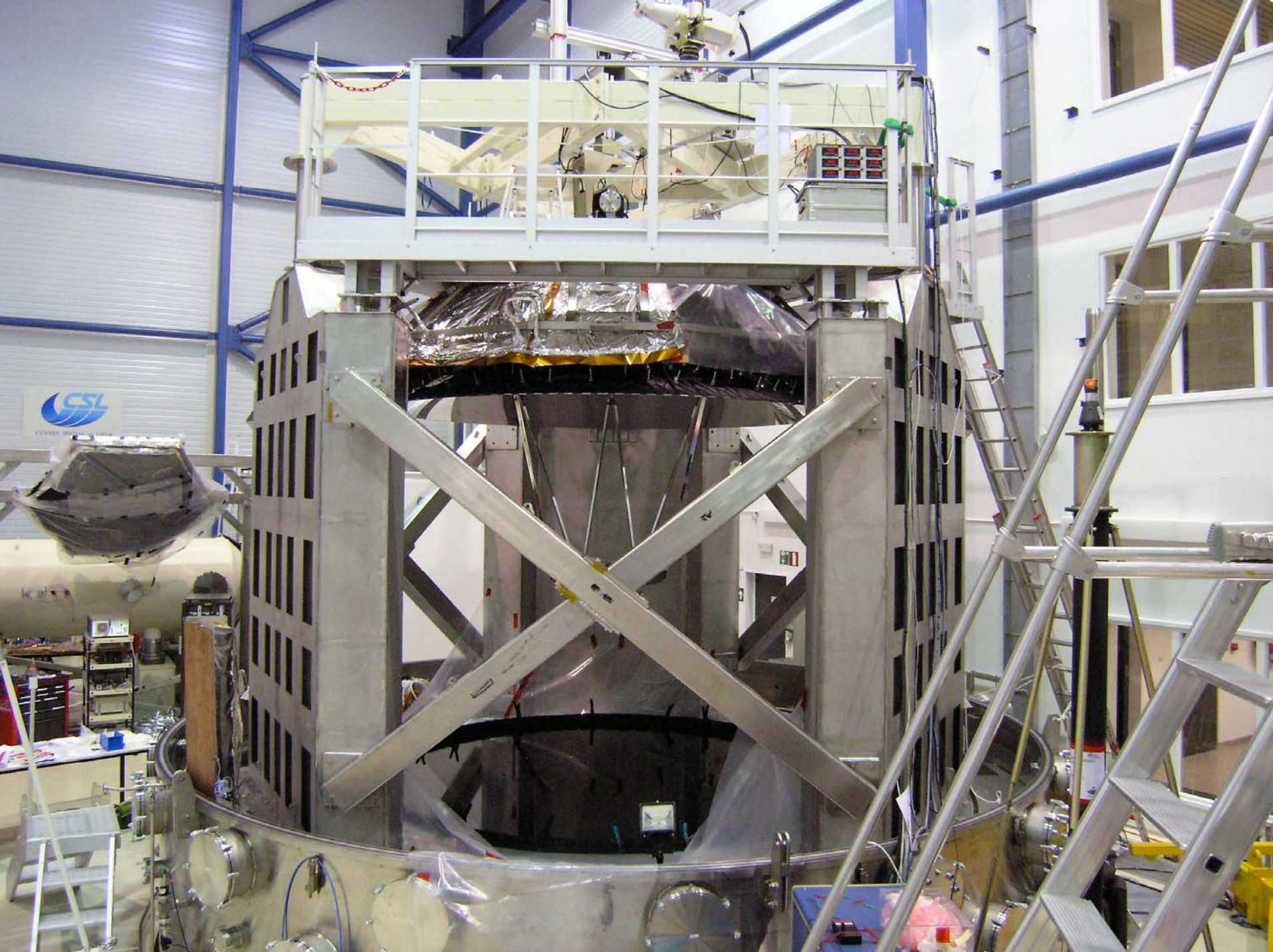
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Grinding and lapping/polishing





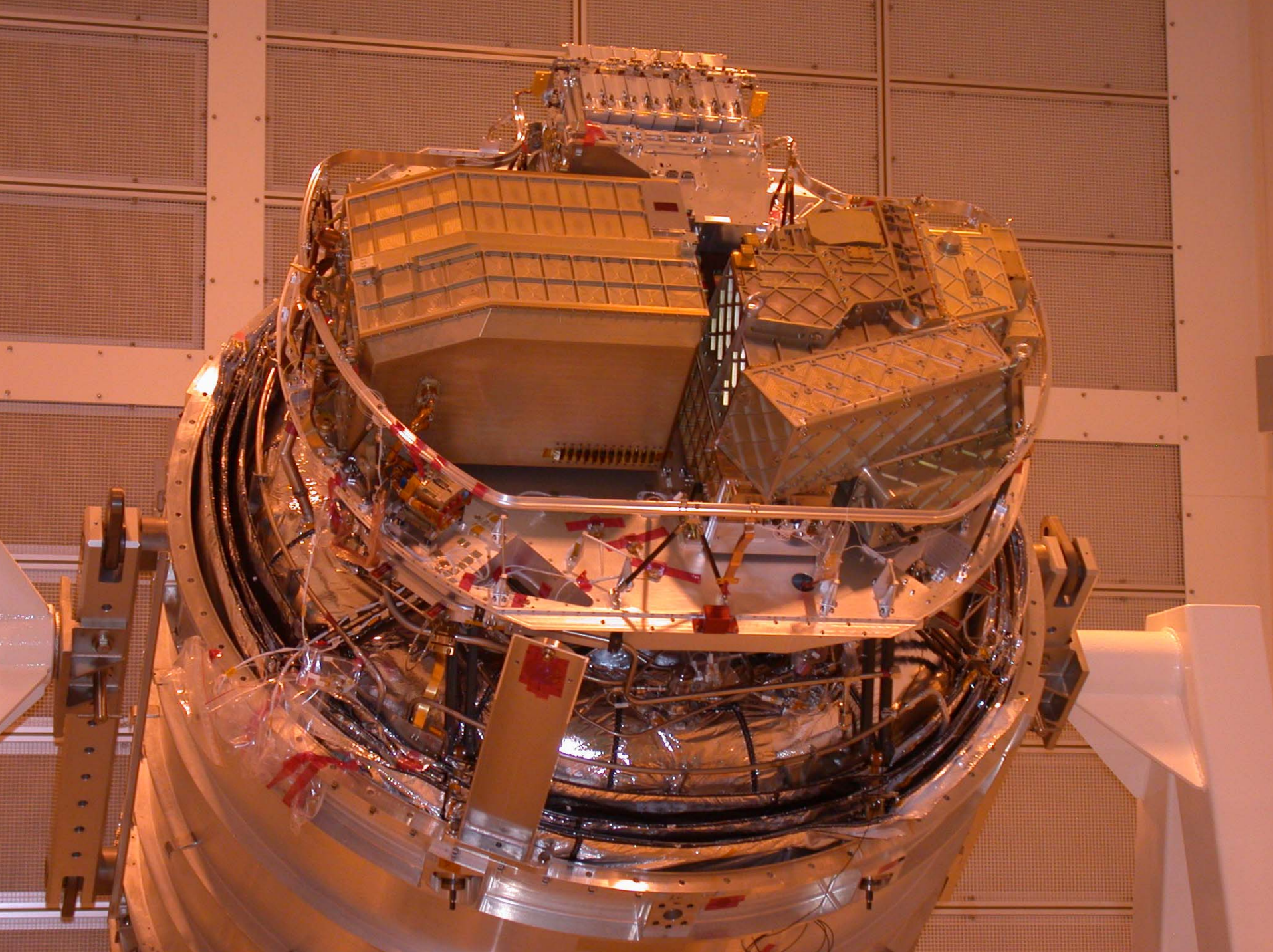


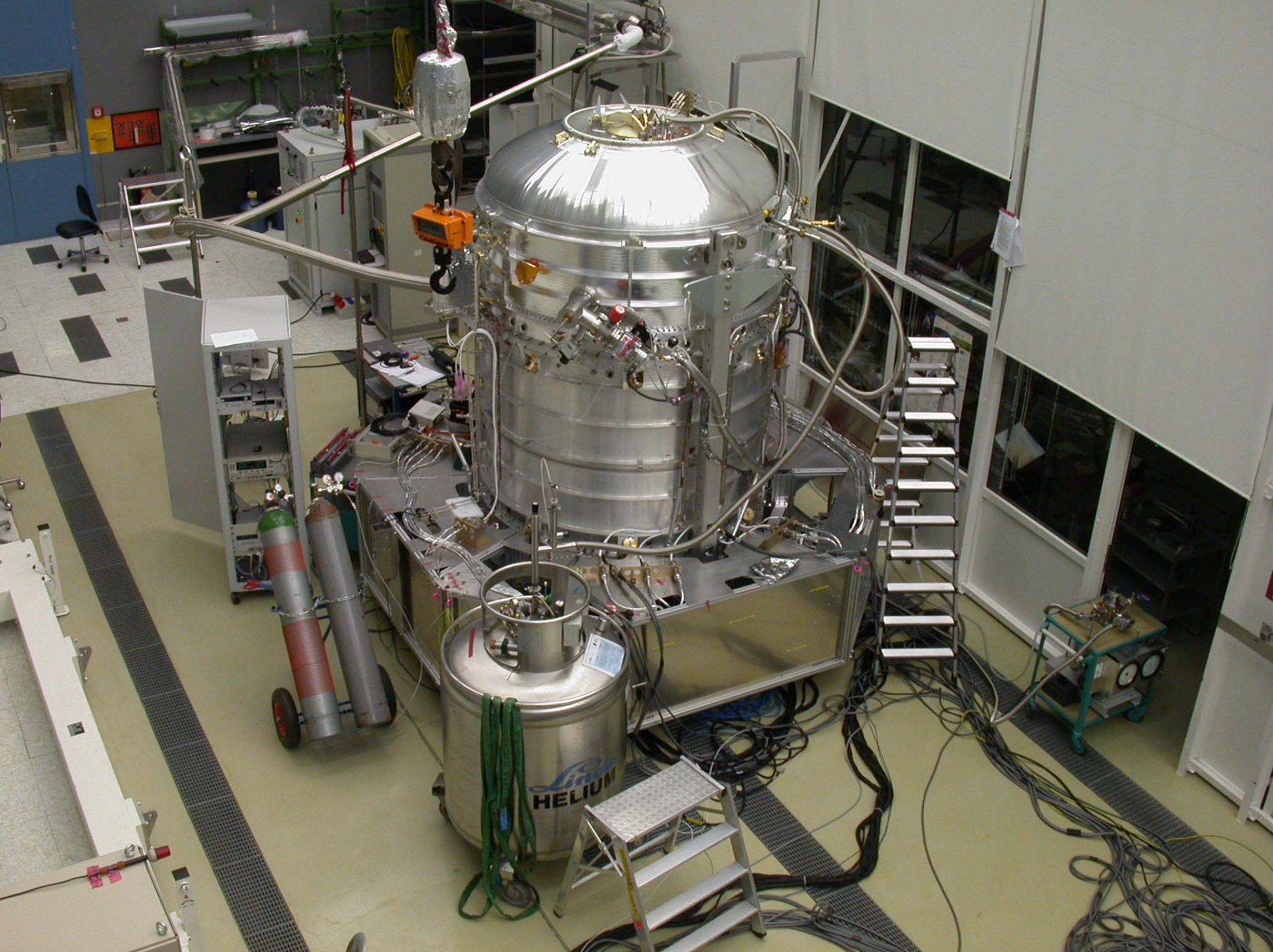


Herschel EQM



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Launcher



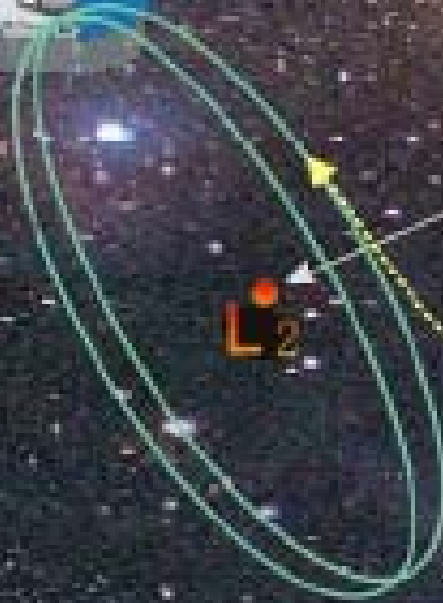
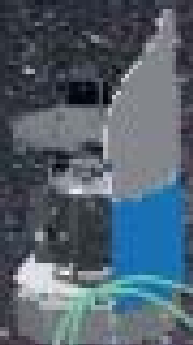
- Launcher version:
 - Ariane 5 - ECA (cryogenic upper stage)
 - qualification flight V164
- Payload configuration:
 - Planck in lower position
 - Herschel in upper position
 - Sylدا5/ACU2624
 - Long fairing
- Launch Autonomy:
 - 25 hours
- L2 Injection strategy:
 - direct injection
 - 25 minutes powered phase
- L2 Injected mass capability:
 - ≥ 6273 kg including adaptors



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L₂

1.5 Mkm

383 Tkm

150 Mkm





Herschel mission phases

- **Launch and early operations (LEOP)**
- **Commissioning and performance verification (SC + payload)**
- **Science demonstration phase**
- **Routine science operations phase (36 months)**
 - Approx 1000 days / 20000 hours of schedulable science time
 - **Guaranteed time programmes – GT (32%)**
 - open for GT holders only
 - **Open time programmes – OT (68%)**
 - including discretionary time and targets of opportunity
 - open for all – including GT holders
- **Three ‘Call for proposals’ (AO) cycles are foreseen**
 - one Call for ‘Key Projects’ programmes only (GT and OT)
 - two Calls for regular programmes (GT and OT)
- **Each AO will be divided in two parts**
 - GT awarded first
 - OT awarded after GT in same cycle

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Overall AO schedule

- **Logic: Issue 'Call for Proposals' (AOs) as late as possible**
 - for pure scientific reasons
 - and for performance knowledge reasons
 - but early enough for observers and HSC to prepare
 - to have observations available for scheduling
 - and enable community support staff 'training on the job'
- **1 Feb 2007: Issue AO for 'Key Projects' proposals**
- **5 Apr 2007: Submission deadline for GT KP proposals**
- **5 Jul 2007: Selection & announcement of GT KP programmes**
 - 'effective' AO date for open time proposers
- **25 Oct 2007: Submission deadline for OT KP proposals**
- **28 Feb 2008: Selection & announcement of OT KP programmes**
- **28 Feb 2008: Issue AO for 'Cycle 1 Regular GT' proposals**
- **3 Apr 2008: Submission deadline for GT1 proposals**
- **5 Jun 2008: Selection & announcement of GT1 programme**

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AO doc package



- **Cover letter – formal Call by D/SCI**
- **Executive summary**
 - duplicates as ‘AO manual’
 - points to online package with latest news
- **Policies and procedures**
 - politics / policies etc.
- **Herschel Observers’ manual**
 - spacecraft matters (pointing, visibility, etc)
- **3 x Instrument Observers’ manual**
 - PACS, SPIRE, HIFI – ultimately HSC responsibility
- **HSpot manual**
 - + additional tools (background noise estimator, confusion noise estimator, ...)
- **Guaranteed time programme description**
 - with reserved observations lists

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Herschel 'Key Projects'

- **Foreseen to be important upfront (SMP/instrument AO)**
 - introduced to ensure that 'unusually large' observing programmes can be proposed, selected, and observed
 - need 'pre-identified' due to the nature of the foreseen science objectives and the lack of 'precursor' (IRAS-type) mission
- **Definition of a 'Key Project' programme - it must**
 - exploit unique Herschel capabilities address (an) important scientific issue(s) in a comprehensive manner
 - require a large amount of observing time to be used in a uniform and coherent fashion
 - produce a resulting well characterised dataset of high archival value
- **Data reduction**
 - it is recognised that there is a legitimate science return interest that
 - the data generated by the observations are timely reduced, and
 - the data products and tools are made public
 - therefore 'Key Project' consortia must demonstrate commitment and ability to perform data reduction, and must make data products and tools publicly available at the end of the proprietary time period

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PACS Photometry

Unique AOR Label: PACSPHOTO-0000

Target: None Specified

New Target

Number of visible stars for the target: 15
Star tracker target: Ra: 22.482 degrees Dec: -47.232 degrees

Target

Target Name (required): NED
M51

Fixed Moving

Coord Sys: Equatorial J2000
RA:
Dec:
Epoch: 2000.00

Proper Motion
☒ Use P
PM RA (" /yr):
PM Dec (" /yr):

Blue channel
☒ 60-85 microns filter (Blue 1)
☐ 85-130 microns filter (Blue 2)

Source
Map choice
☒ Pointed
☐ Small Map
☐ Large Map

Pointing
Nodding
☒ On
☐ Off

Bright Moving Objects to Avoid
☒ Earth+Moon
☒ Others

Chopping and Dithering settings
Chopping
☒ On
☐ Off
Chopper avoidance angle (deg)
Angle from (degrees) 0.00
Angle to (degrees) 0.00

Dithering
☐ On
☒ Off

Observation Est... Add Comments... Visibility... Star Tracker...

OK Cancel

PACS Photometry

Unique AOR Label: PACSPHOTO-0000

Target: m51 Type: Fixed Single
Position: 13h29m55.73000s, +47d13m53.4000s

New Target Modify Target Target List...

Number of visible stars for the target: 15
Star tracker target: Ra: 22.482 degrees Dec: -47.232 degrees

Instrument Settings

Blue channel
☒ 60-85 microns filter (Blue 1)
☐ 85-130 microns filter (Blue 2)

For each pointing
On-source integration time (s): 60

Observing Mode Settings

Source
Map choice
☒ Pointed
☐ Small Map
☐ Large Map

Pointing settings
Nodding
☒ On
☐ Off
Large
☒ On
☐ Off

Source Flux Est

Chopping and Dithering settings
Chopping
☒ On
☐ Off
Chopper avoidance angle (deg)
Angle from (degrees) 0.00
Angle to (degrees) 0.00

Dithering
☐ On
☒ Off

Observation Est... Add Comments... Visibility... Star Tracker...

OK Cancel

PACS Time Estimation

Time
Observatory overhead (s) 180
On-source time (s) 48
Calibration time (s) 4
Total time (s) 283

Band (um)	Est. 1-sigma noise (mJy)	Est. 1-sigma confusion level (mJy)
60-85	3.1	11.2
130-210	3.2	10.2

Confusion Messages

Done Messages

Ned Name Resolver- Working...



Ned Name Resolver
Searching for Name: M51

Click Abort to stop.

Abort



Data processing – HCSS/DP

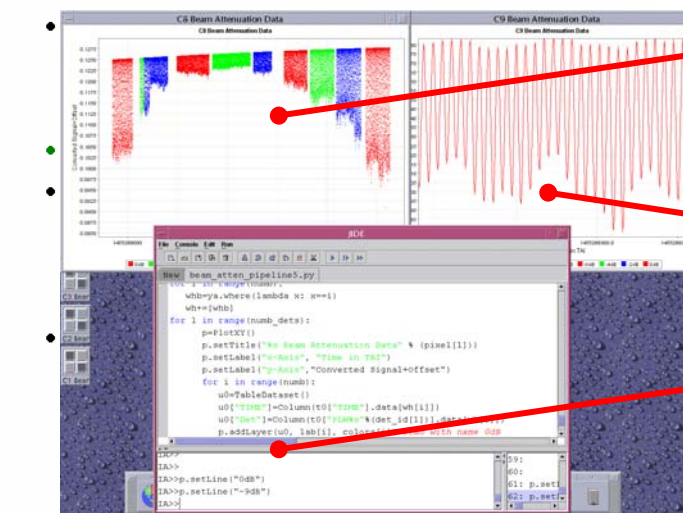
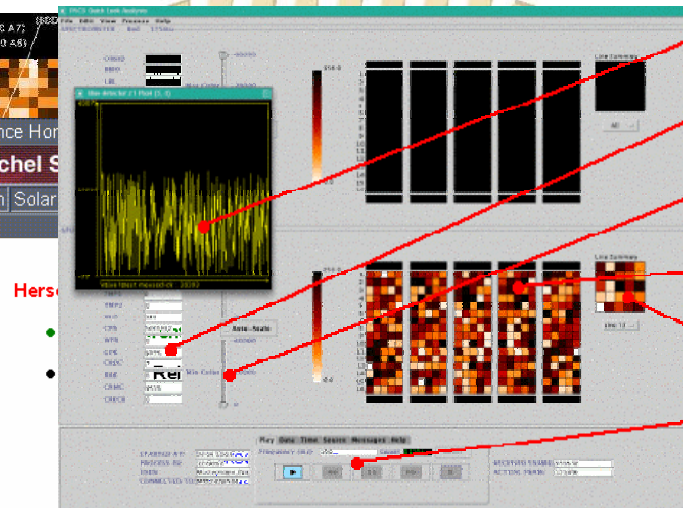
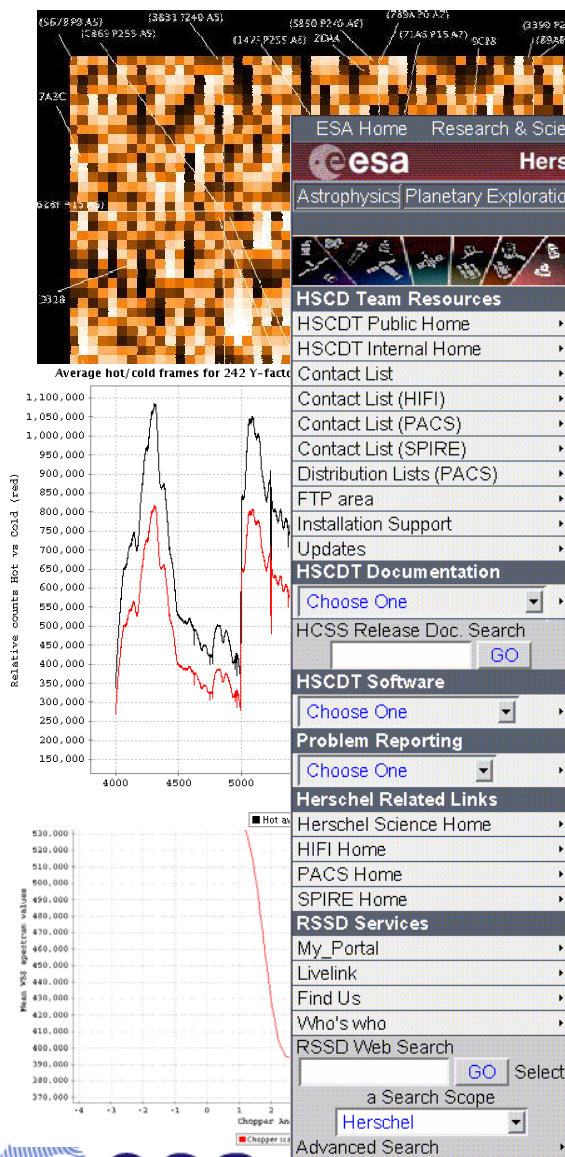
- Ensemble of services in a single, coherent platform-independent system – HCSS/DP
- Offering the astronomical community the means (data, products, software) to do science without the need to buy licenses
- Source of software, calibration and documentation to reduce and do science with Herschel data ('Observer IA')
- Source of software and documentation to support the commissioning of the instruments, the validation of the observing modes and the calibration of satellite pointing and instruments ('Calibration IA')
- Toolbox to develop data processing algorithms
- Generation of standard observation products and quality information by using IA modules ('SPG' & 'QC' pipelines)
- On-line, context-sensitive and printable help documentation for end-users and developers based on XML/DocBook
- Data and services Virtual Observatory compliant + FITS

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Herschel data processing

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- [DP-WG MoM](#): DP Working Group Minutes of Meetings
- [Livelihood documents](#): Other related documents under Livelihood, including DP CCB and User Group documentation.

Last updated: Friday, 26-Aug-05 13:02:20





Herschel key dates

- FIRST feasibility study: 1982-83
- ...
- FIRST confirmed 4th Cornerstone: Dec 1993
- ...
- Selection of science instruments/PIs: 1997-98
- FIRST becomes Herschel: Dec 2000
- Start of spacecraft Phase B: mid 2001
- Start of spacecraft Phase C/D: early 2003
- SVM, PLM, telescope, CQM/EQM testing 2005
- SVM, PLM, instruments, telescope deliveries: 2006/07
- Integration & tests/verification: 2007
- Issue of first AO for proposals: 1 Feb 2007
- Launch: (not before 31 Jul) 2008
- Early operations: first 6 months
- Routine Science Operations: 2009 - 2012

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