

Workshop “Detection and measurement of RFI in radio astronomy”



Yebes Observatory (IGN, Spain), June 8-9, 2017



RFI Mitigation Project at Italian Radio Telescopes

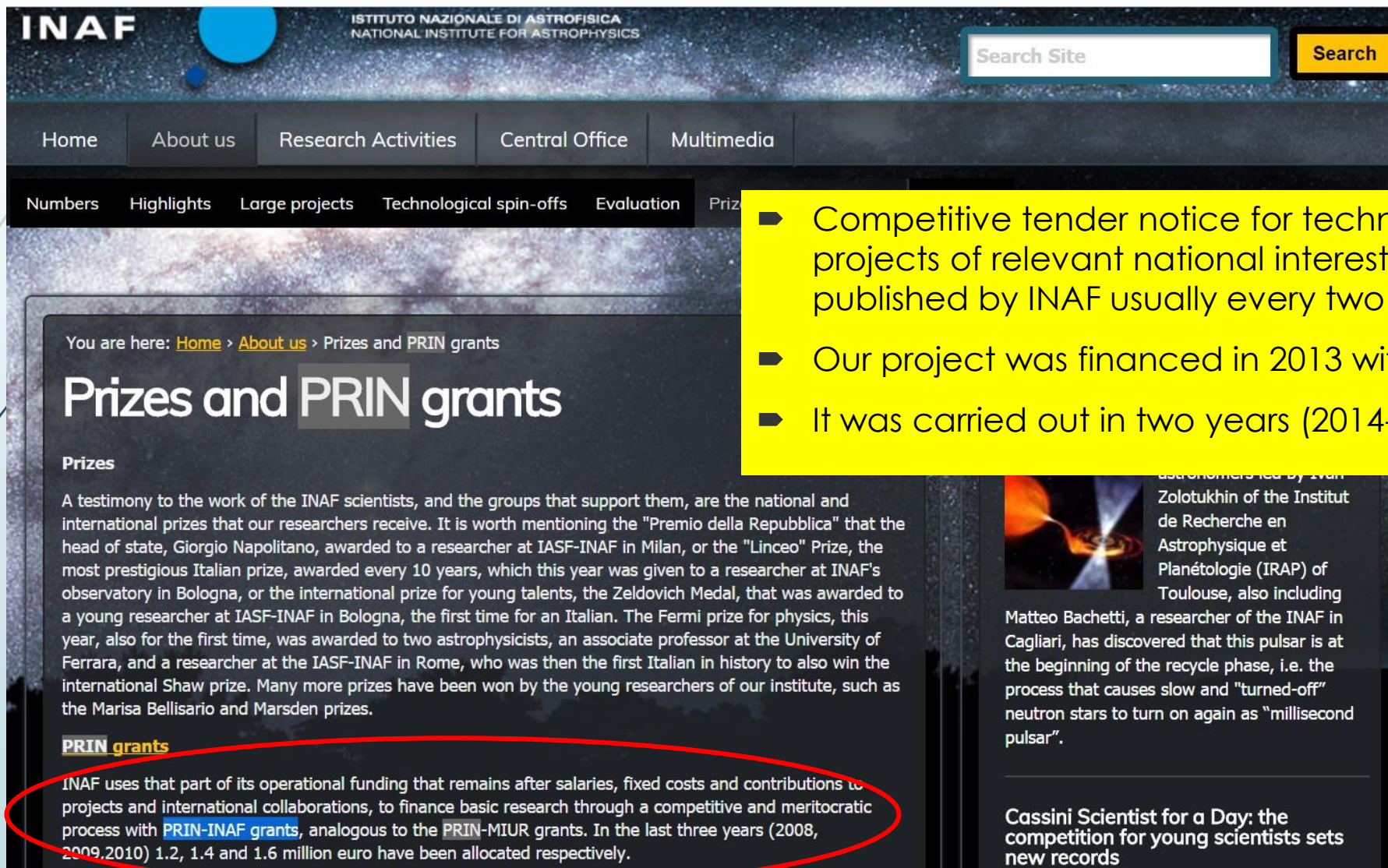
G. Serra and many other people
from Italian National Institute for Astrophysics
and other Institutes



Outline

- Introduction (Techno INAF-PRIN)
- RFI facilities at the Italian radio telescopes (IRTs)
- Project tasks for RFI monitoring: some examples of RFI at the IRTs
- Project tasks for RFI mitigation
 - WBLGB spectrometer
 - Off-line Dish Washer
- Tests of mitigation tools at the Sardinia Radio telescope
- Summary and ongoing developments

Intro: Techno INAF-PRIN



INAF ISTITUTO NAZIONALE DI ASTROFISICA
NATIONAL INSTITUTE FOR ASTROPHYSICS

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Prizes and PRIN grants

Prizes

A testimony to the work of the INAF scientists, and the groups that support them, are the national and international prizes that our researchers receive. It is worth mentioning the "Premio della Repubblica" that the head of state, Giorgio Napolitano, awarded to a researcher at IASF-INAF in Milan, or the "Linceo" Prize, the most prestigious Italian prize, awarded every 10 years, which this year was given to a researcher at INAF's observatory in Bologna, or the international prize for young talents, the Zeldovich Medal, that was awarded to a young researcher at IASF-INAF in Bologna, the first time for an Italian. The Fermi prize for physics, this year, also for the first time, was awarded to two astrophysicists, an associate professor at the University of Ferrara, and a researcher at the IASF-INAF in Rome, who was then the first Italian in history to also win the international Shaw prize. Many more prizes have been won by the young researchers of our institute, such as the Marisa Bellisario and Marsden prizes.

PRIN grants

INAF uses that part of its operational funding that remains after salaries, fixed costs and contributions to projects and international collaborations, to finance basic research through a competitive and meritocratic process with **PRIN-INAF grants**, analogous to the **PRIN-MIUR grants**. In the last three years (2008, 2009, 2010) 1.2, 1.4 and 1.6 million euro have been allocated respectively.

Astronomers led by Ivan Zlotukhin of the Institut de Recherche en Astrophysique et Planétologie (IRAP) of Toulouse, also including Matteo Bachetti, a researcher of the INAF in Cagliari, has discovered that this pulsar is at the beginning of the recycle phase, i.e. the process that causes slow and "turned-off" neutron stars to turn on again as "millisecond pulsar".

Cassini Scientist for a Day: the competition for young scientists sets new records

- Competitive tender notice for technological projects of relevant national interest (PRIN) published by INAF usually every two years
- Our project was financed in 2013 with 173 K€
- It was carried out in two years (2014-16)

Intro: Research units and people



IRA-Medicina staff

K-H. Mack (P.I.), A. Zanichelli,
M. Bartolini, C. Bortolotti, M.
Roma, A. Orlati, S. Righini, R.
Ambrosini

IRA-Noto staff

G. Nicotra, R. Platania

Non-staff people

M. De Biaggi, F. Cantini, E. Favero
(research grants with the project
funds)



OA-Cagliari staff

S. Poppi (Local Coordinator)
G. Serra, F. Gaudiomonte
M. Buttu, C. Migoni



OA-Catania staff

C. Trigillio (Local Coordinator)
P. Leto, C. Buemi

External collaborators

R. Prestage (NRAO), W. Baan (ASTRON), A. Jessner, A.
Kraus, P. Muller and B. Winkel (MPIfR), G. Busonera (CRS4)

Intro: Research Units and tasks

OA-Cagliari, IRA-Noto and –
IRA-Medicina (RFI teams)



Characterization of the RFI situation at
the three Italian radio-observatories

IRA-Medicina, IRA-Bologna
(backend team)



Development and implementation of
the FPGA firmware for on-line
mitigation (WBLGB spectrometer)

IRA-Bologna, IRA-Medicina,
OA-Cagliari
(software team)



Development of an off-line mitigation
tool (Dish Washer)

OA-Cagliari, IRA-Medicina
(RFI and backend teams)



Verification of the on-line mitigation
algorithms

INAF people and external
collaborators



Verification of the mitigation tools and
observational tests using telescopes

RFI facilities & ordinary activities at IRTS



32-m Medicina telescope & Northern Cross array



Fixed station: 0.3 - 12 GHz
Mobile lab: 0.3 - 40 GHz



64-m Sardinia Radio telescope
Mobile lab: 0.05 - 40 GHz



32-m Noto radio telescope
Fixed station: 0.1 - 3 GHz

Project tasks: RFI campaigns at IRTs

@ IRA- Medicina



Dedicated RFI meas. in the RA bands
by fixed station and IRA mobile lab)

@ IRA- Noto



2 RFI meas. campaigns in freq. range
0.05-40 GHz (in 2014 and 2016)

(by **OA-Cagliari mobile lab**)

@ SRT



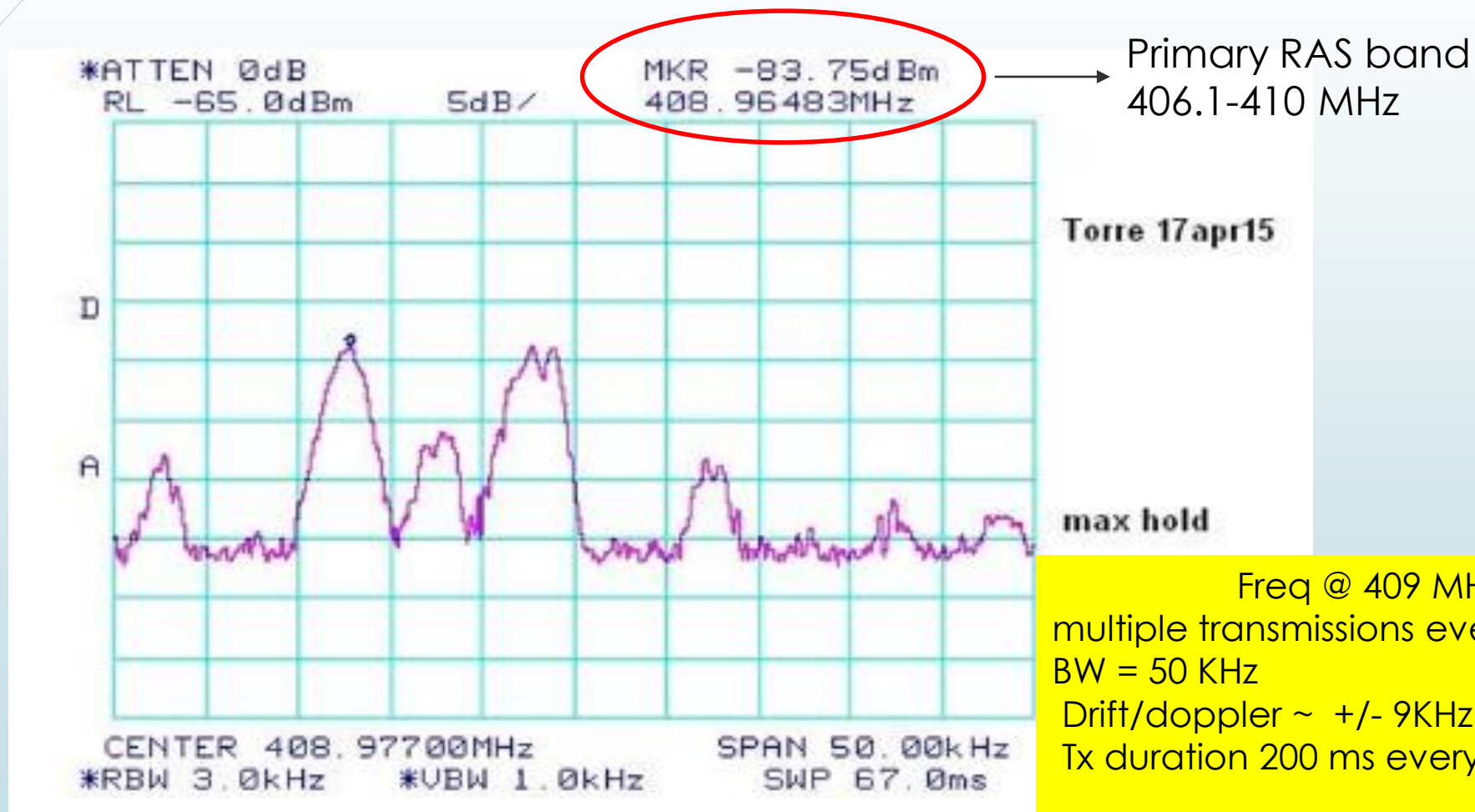
1 RFI meas. campaign (in 2015)

freq range 0.05-40 GHz

(by **OA-Cagliari mobile lab**)

Example of RFI @ Medicina

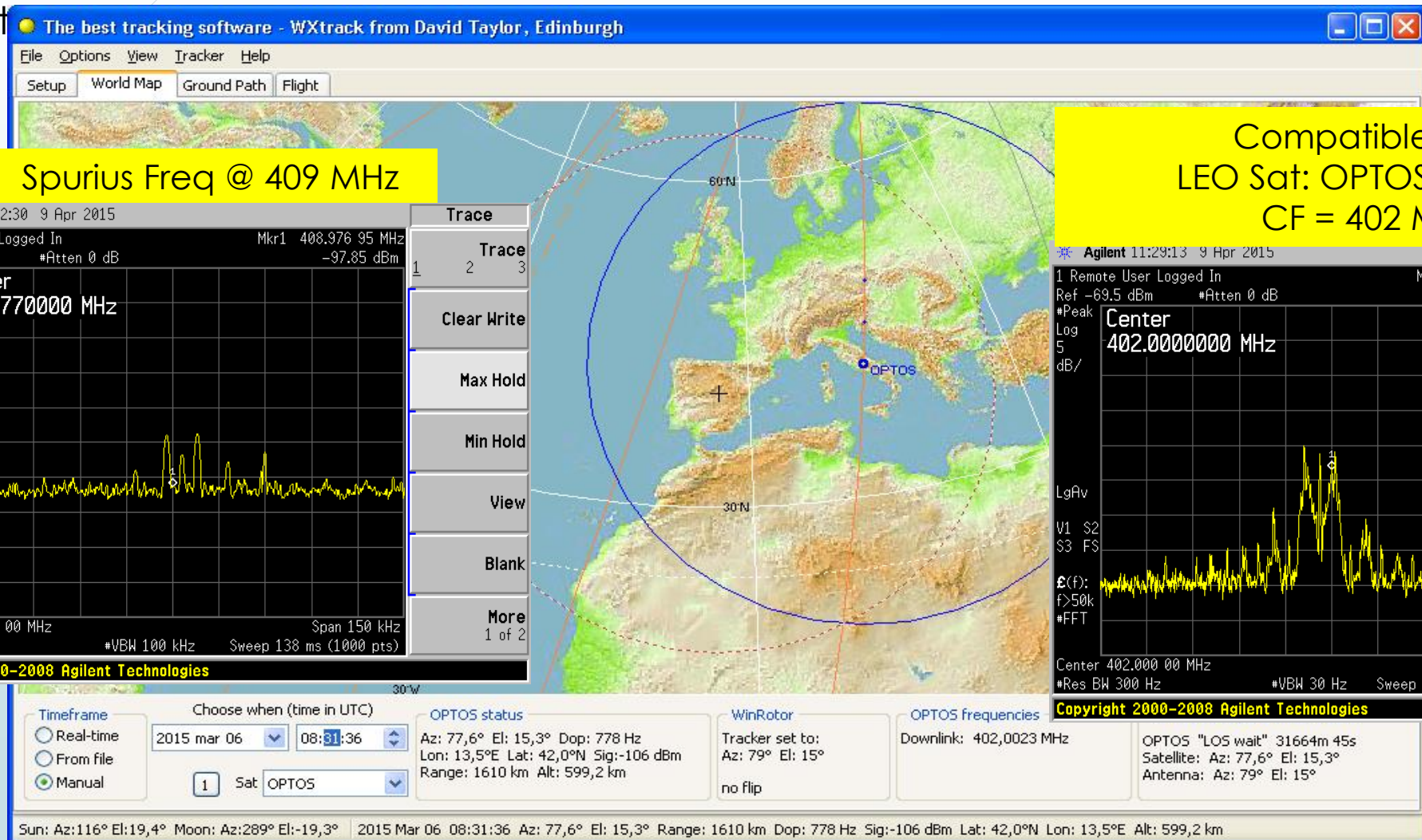
Satellite signal in the Best2-Northern Cross band (400-416 MHz)



Freq @ 409 MHz
multiple transmissions every 5 KHz
BW = 50 KHz
Drift/doppler ~ +/- 9KHz
Tx duration 200 ms every 10-20 s

Example of RFI @ Medicina (@SRT as well)

Satellit

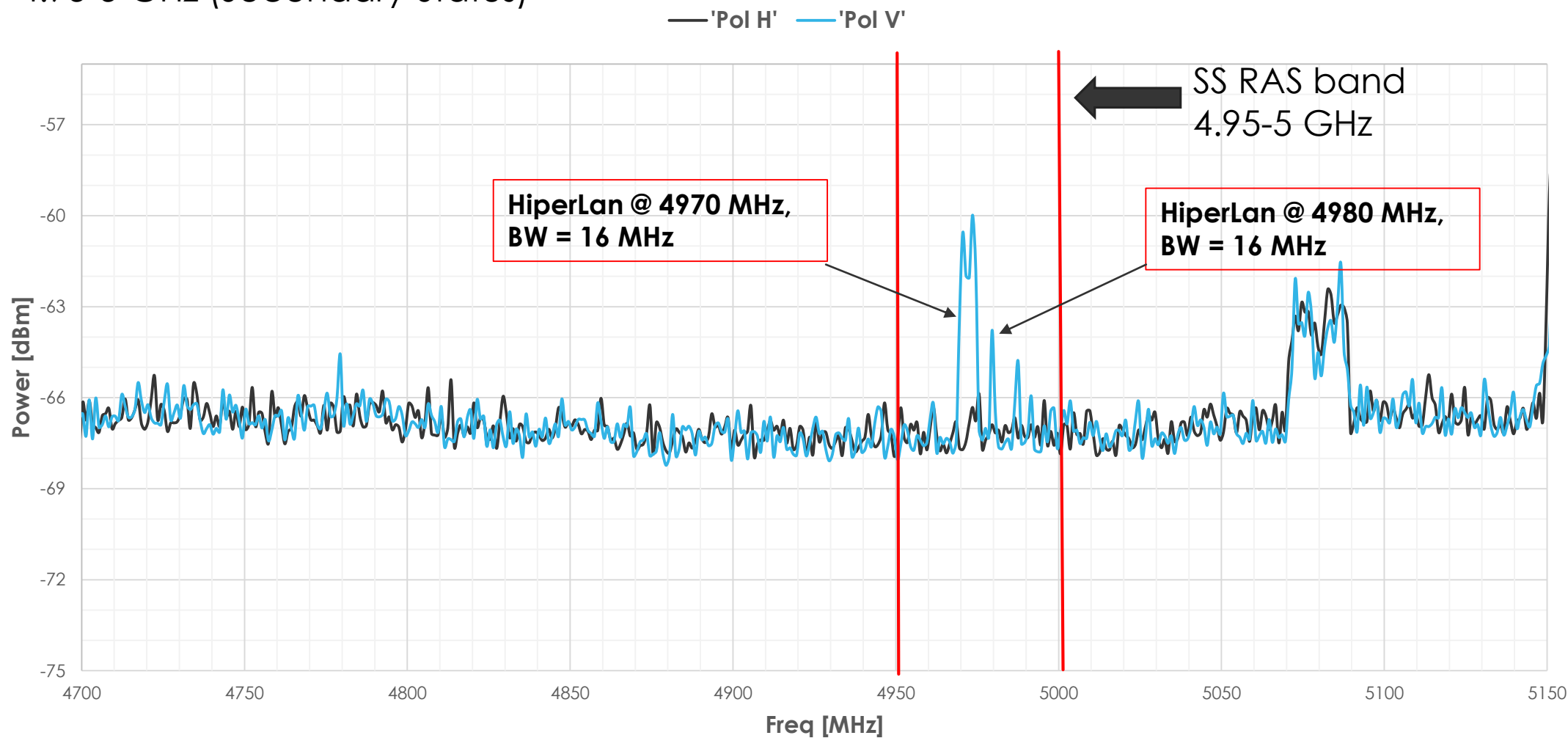


Spurious Freq @ 409 MHz

Compatible with
LEO Sat: OPTOS n.39420
CF = 402 MHz

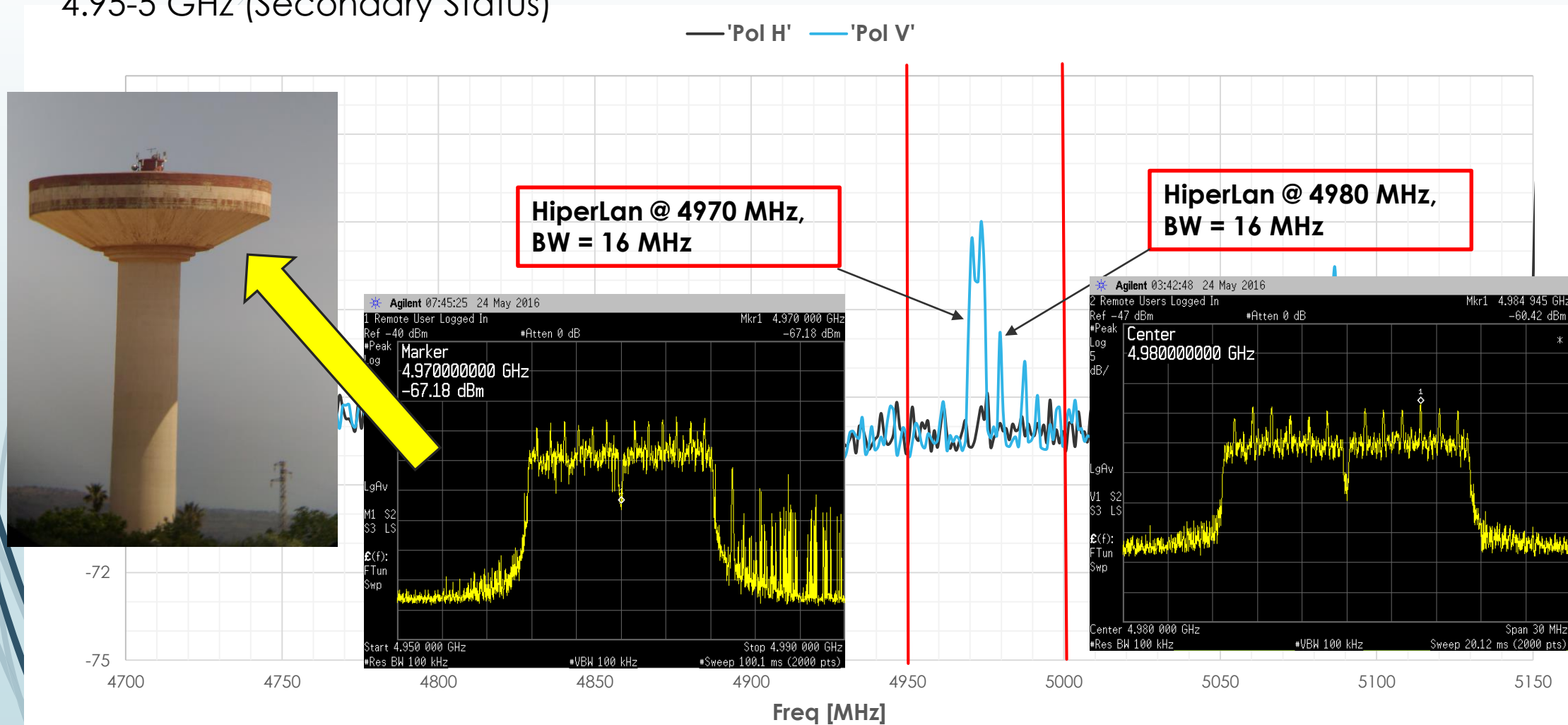
Example of RFI @ Noto

Low-C receiver band: two HiperLAN signals in a non-allocated band inside the RA band 4.95-5 GHz (Secondary Status)



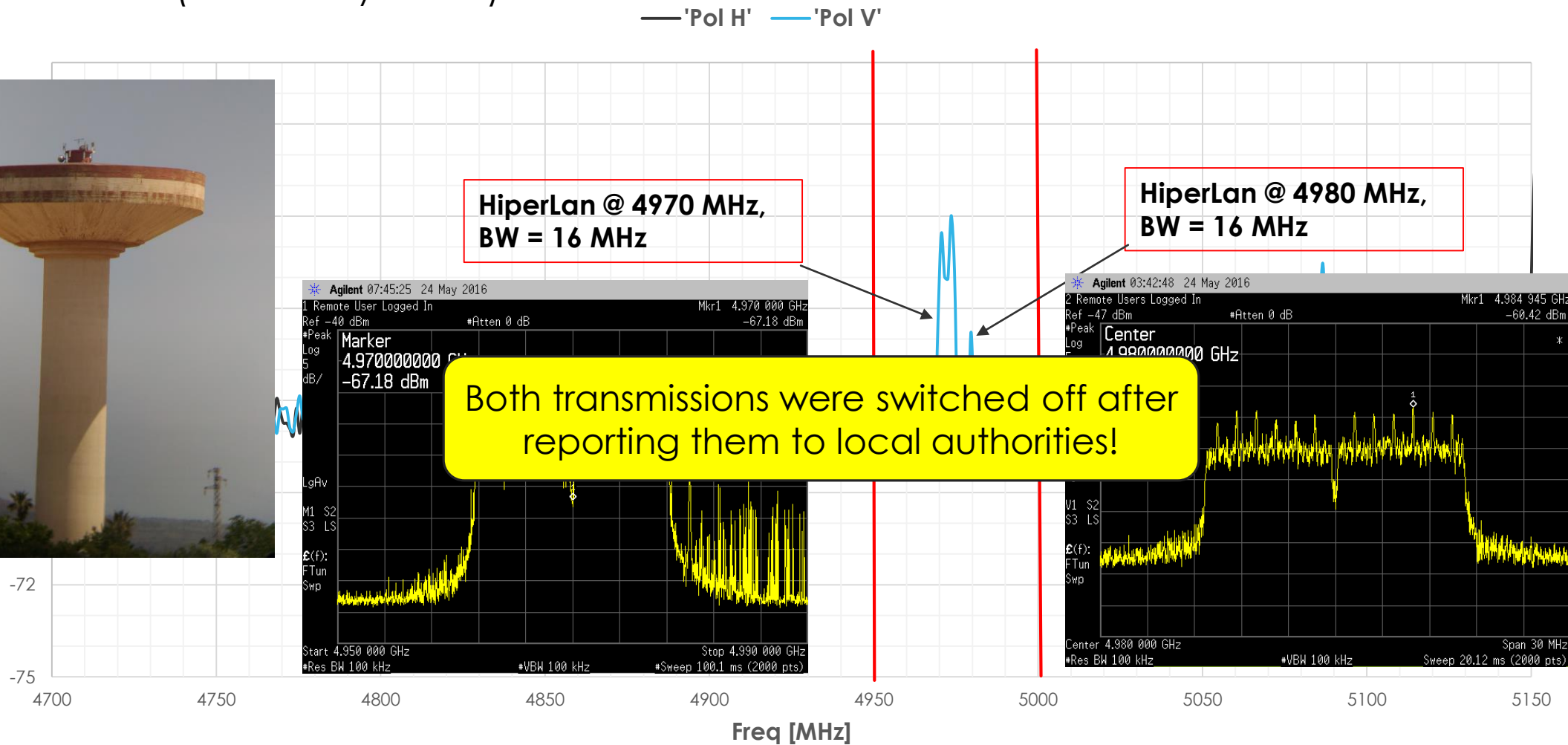
Example of RFI @ Noto

Low-C receiver band: two HiperLAN signals in a non-allocated band inside the RA band 4.95-5 GHz (Secondary Status)



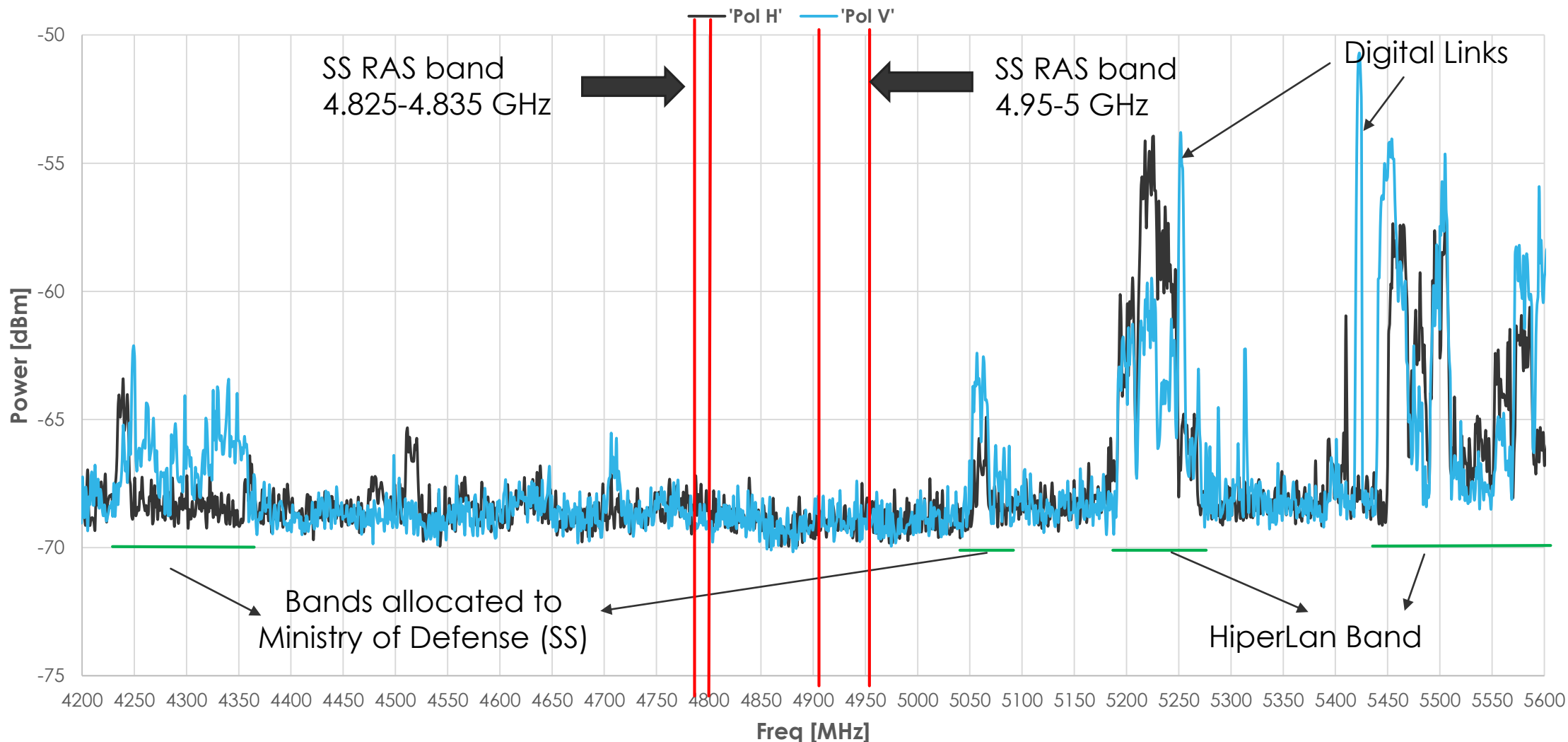
Example of RFI @ Noto

Low-C receiver band: two HiperLAN signals in a non-allocated band inside the RA band 4.95-5 GHz (Secondary Status)



Examples of RFI @ SRT

Low-C receiver (Under construction) band, no RFI in the SS RAS Bands but...



RFI occupancy* at Italian RTs (in current receiver bands)

Medicina

SRT

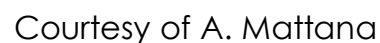
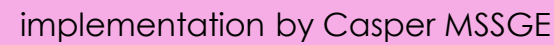
Noto

Receiver (focus)	Freq. Band [GHz]	Occupancy [%]	RFI notes	Receiver (focus)	Freq. Band [GHz]	Occupancy [%]	RFI notes	Receiver (focus)	Freq. Band [GHz]	Occupancy [%]	RFI notes
P BEST – N. Cross (primary)	.400-.416	35	Radio Links, Radiosondes power lines	Coaxial P-L (primary)	.305 - .410	52	Aeronautical digital links, self-RFI, power lines	low-C (secondary)	4.70-5.15	4	HiperLAN
low-L (primary)	1.35-1.45	15	Radar		1.3 - 1.8	57	Radar, satellite, cell phone, self-RFI	high-C (secondary)	6.5-6.8	0	-
high-L (primary)	1.595-1.715	35	Satellite		5.7 - 7.7	10	HiperLAN, digital links	Coaxial S/X (primary)	2.189-2.371	11	Digital links
low-C (secondary)	4.3-5.8	45	HiperLAN Radio links	C (tertiary)	18 – 26.5	7	Cell phone network digital links	K (secondary)	8.138-8.922	0	-
high-C (secondary)	5.9-7.1	25	Radio links		8.2- 8.6	3	Digital links		21.18-22.46	0	-
Coaxial S/X (primary)	2.20-2.36	45	Radio links	7 beam-K (Gregorian)	31.85 - 32.35	2	Surveillance radar	Q (secondary)	39.0-43.5	0	Only up to 40 GHz
	8.18-8.98	5	Radio Satellite								
2-beam-K (secondary)	18.0-26.5	5	Radio links	Coax. X-Ka (primary)							

* with respect to the whole bandwidth of each receiver

Wide Band Lowpass Giga Bit spectrometer

<http://casper.berkely.edu/wiki/ROACH/>



Project tasks: spectrometer for RFI mitigation

Wide Band Lowpass Giga Bit spectrometer

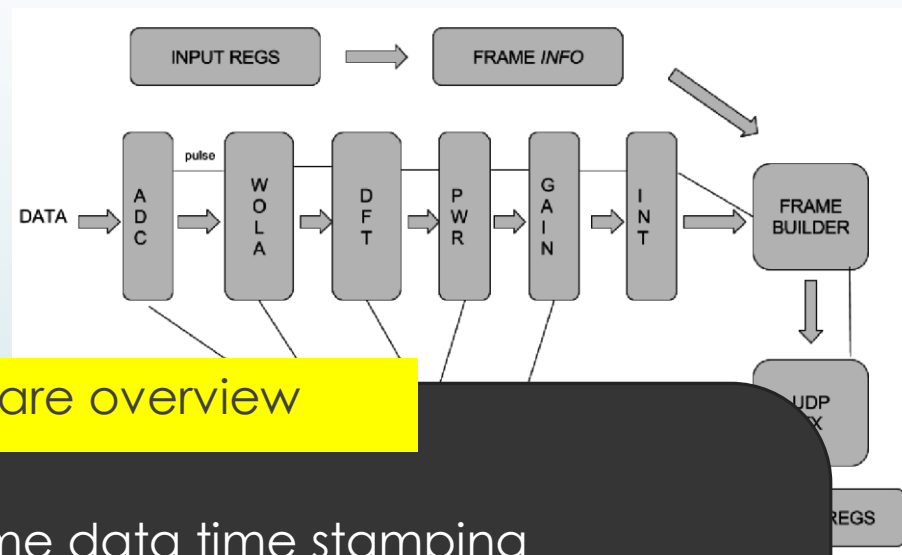
Roach Board 1

<http://casper.berkeley.edu/wiki/ROACH/>



Hardware features

- 8 bit, 1Gbps iADC
- External clock synchronization
- BW = 800 MHz
- integration time = 1 ms (minimum for fast RFI detection)
- Data output rate = 1.25 Gbps

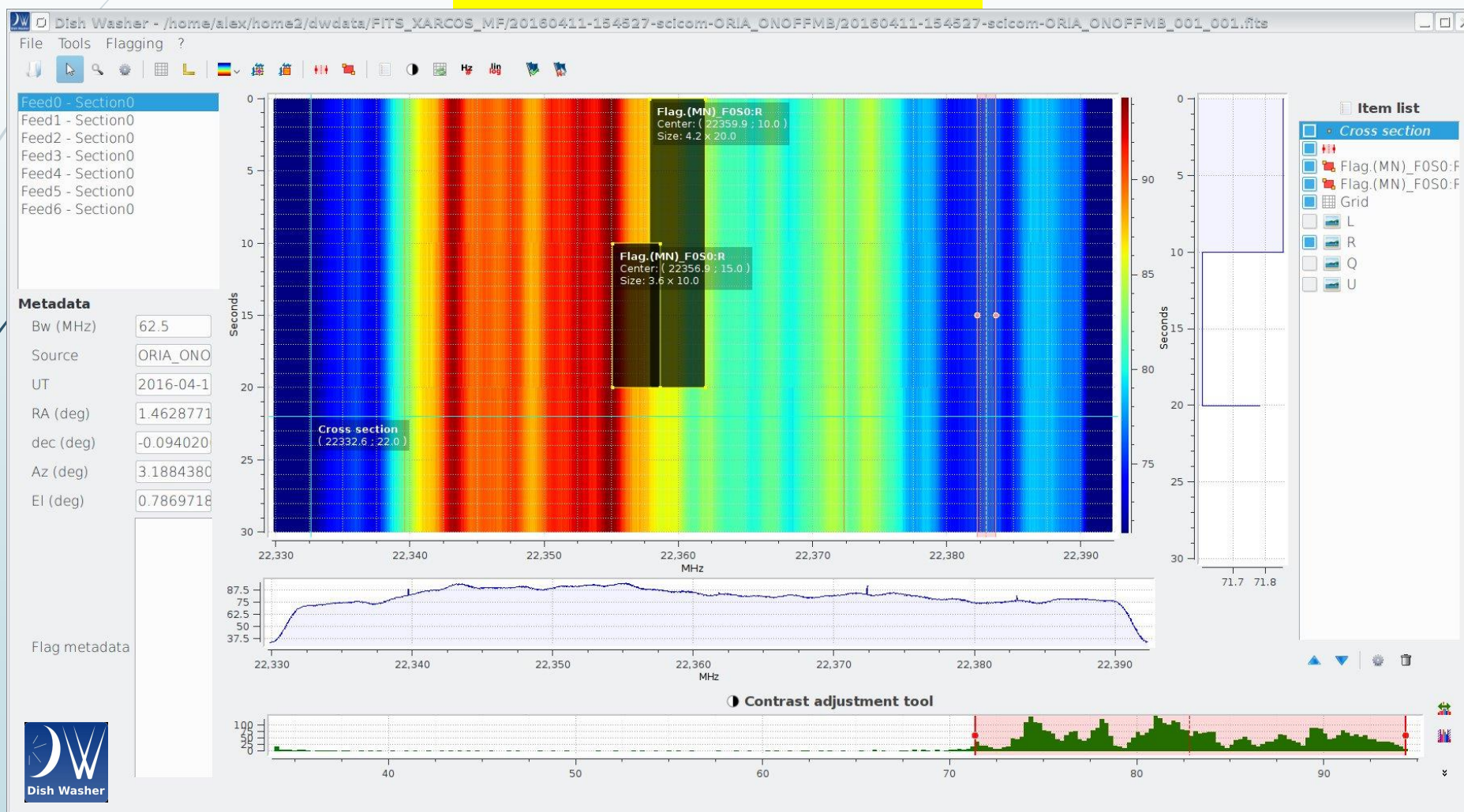


Firmware overview

- Realtime data time stamping
- Freq channels = 4096 (maximum via PFB and DFT)
- Configurable digital gain and DFT shift (robust to RFI signals)
- overflow monitoring
- every stage inspectable via ram blocks

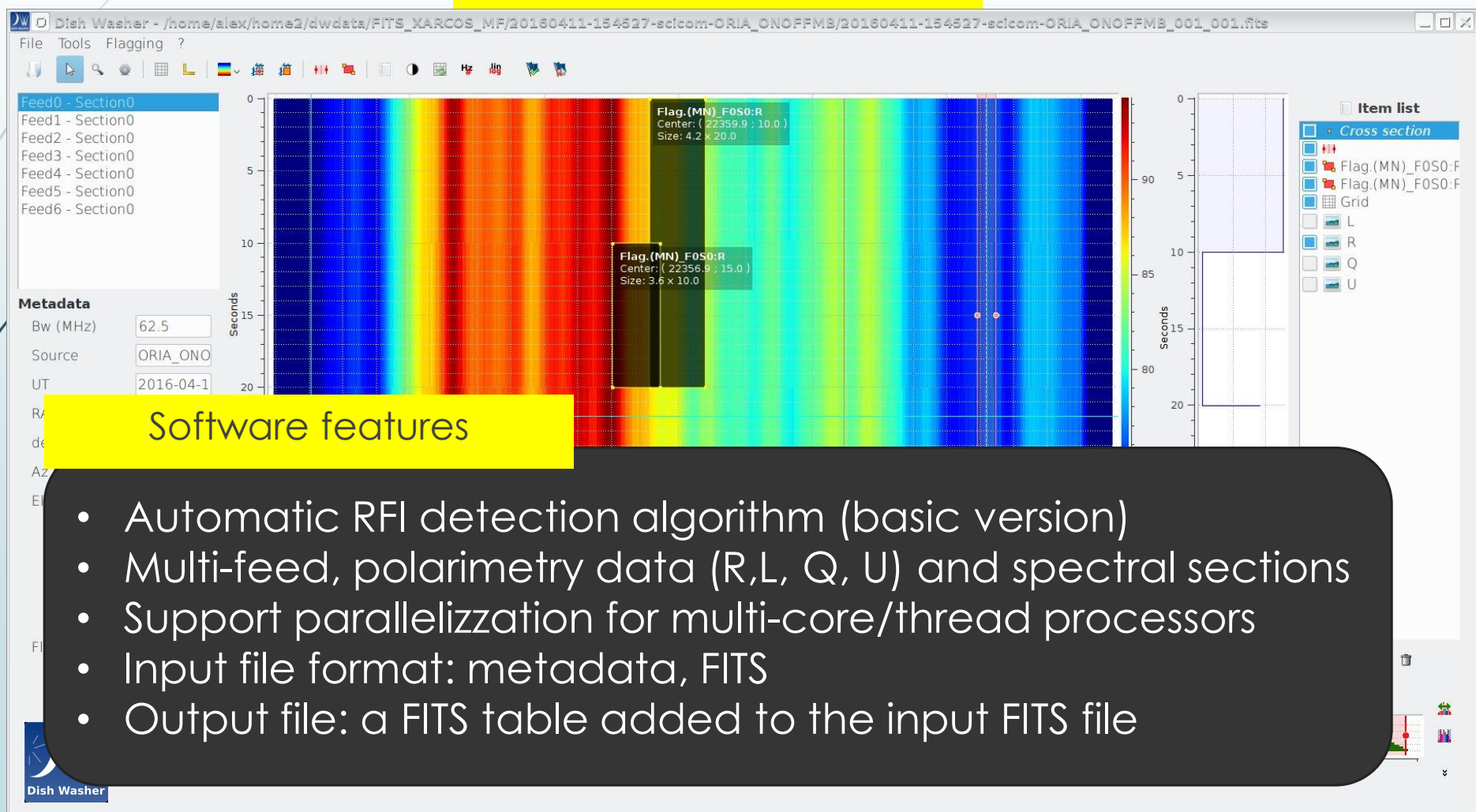
Project tasks: software for RFI mitigation

Off-line Dish Washer (DW)



Project tasks: software for RFI mitigation

Off-line Dish Washer (DW)



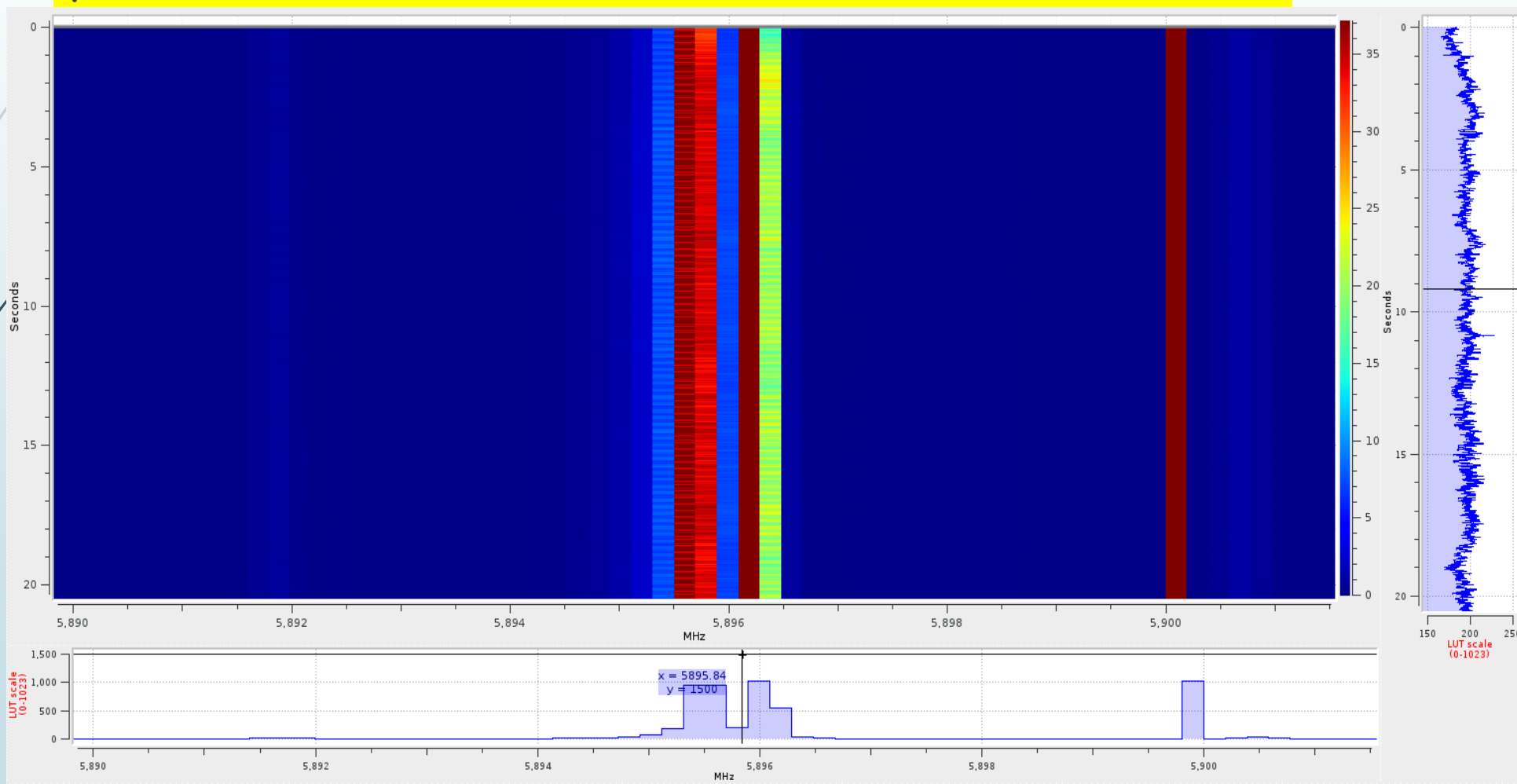
Mitigation tools: preliminary tests @ SRT

Set-up and adjustment

- WBLGB spectrometer firmware was installed on a SRT backend (ROACH 1) (DW not yet installed @ IRTs but tested off-line with Medicina and SRT dataset)
- Interfacing with the Antenna Control Software (Nuraghe) to get pointing coordinates, source name, UT and receiver setup
- the SRT IF baseband output was connected by a power splitter to:
 - the WBLGB spectrometer
 - spectrum analyzer with for the spectrometer digital gain adjustment (repeated for each SRT receiver band (L-P, C- and K-band))

Mitigation tools: preliminary tests @ SRT

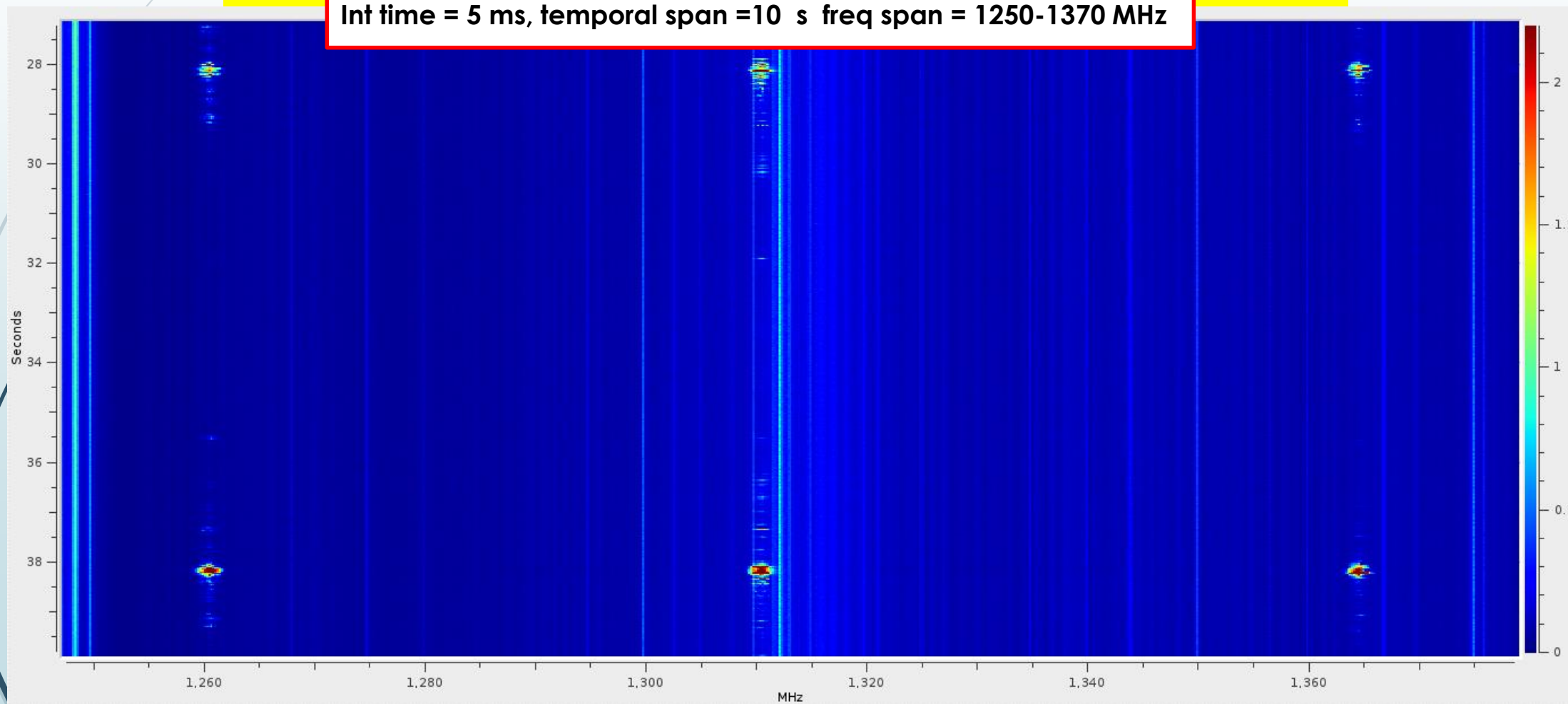
Strong known RFI (digital link) in the C-receiver band (5.7-7.7 GHz)



Mitigation tools: preliminary tests @ SRT

Strong known RFI (radar) in the L-receiver band (1.3-1.8 GHz) :

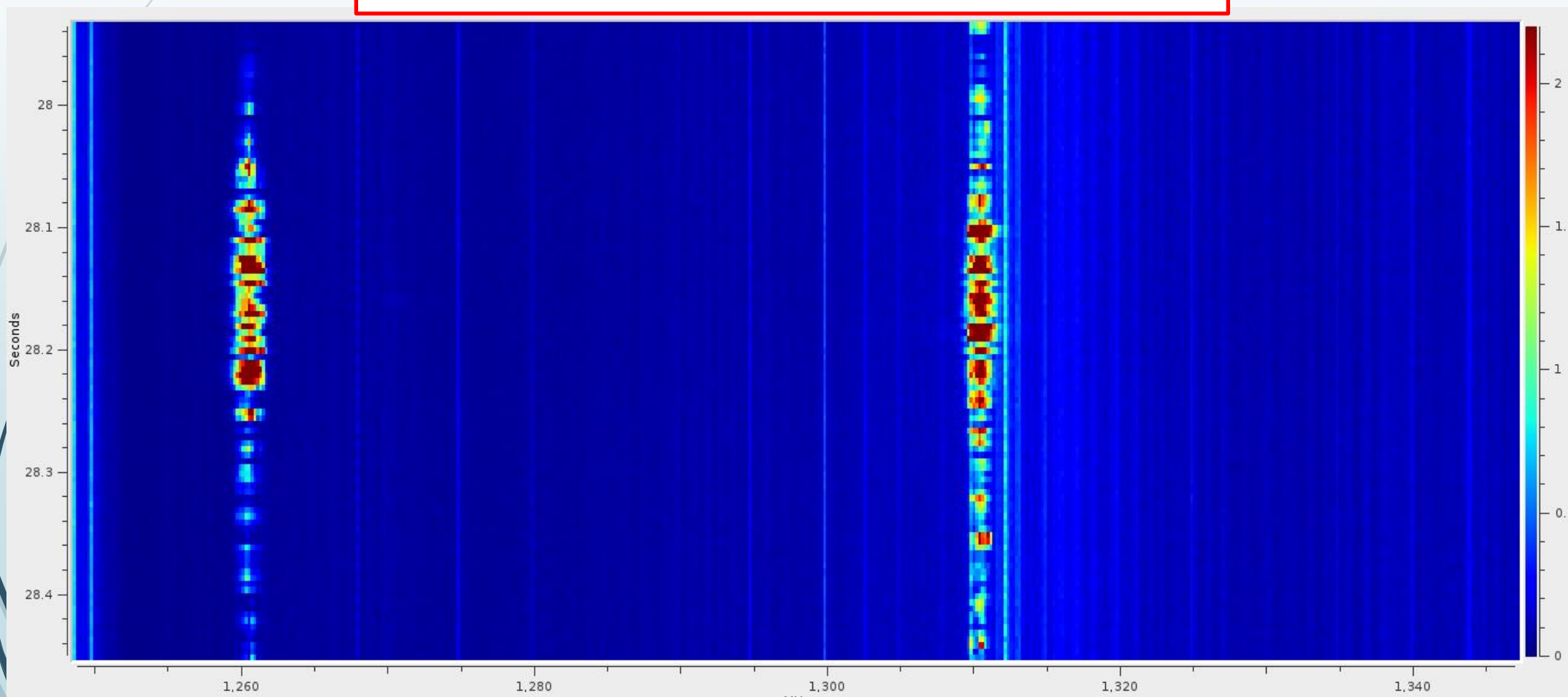
Int time = 5 ms, temporal span = 10 s freq span = 1250-1370 MHz



Mitigation tools: preliminary tests @ SRT

Strong known RFI (radar) in the L-receiver band (1.3-1.8 GHz) :

Int time = 5 ms, temporal span = 0.5 s freq span = 1260-1340 MHz



Mitigation tools: preliminary tests @ SRT

W3OH (astronomical calibrator) and strong RFI in the C-receiver band



Project Summary

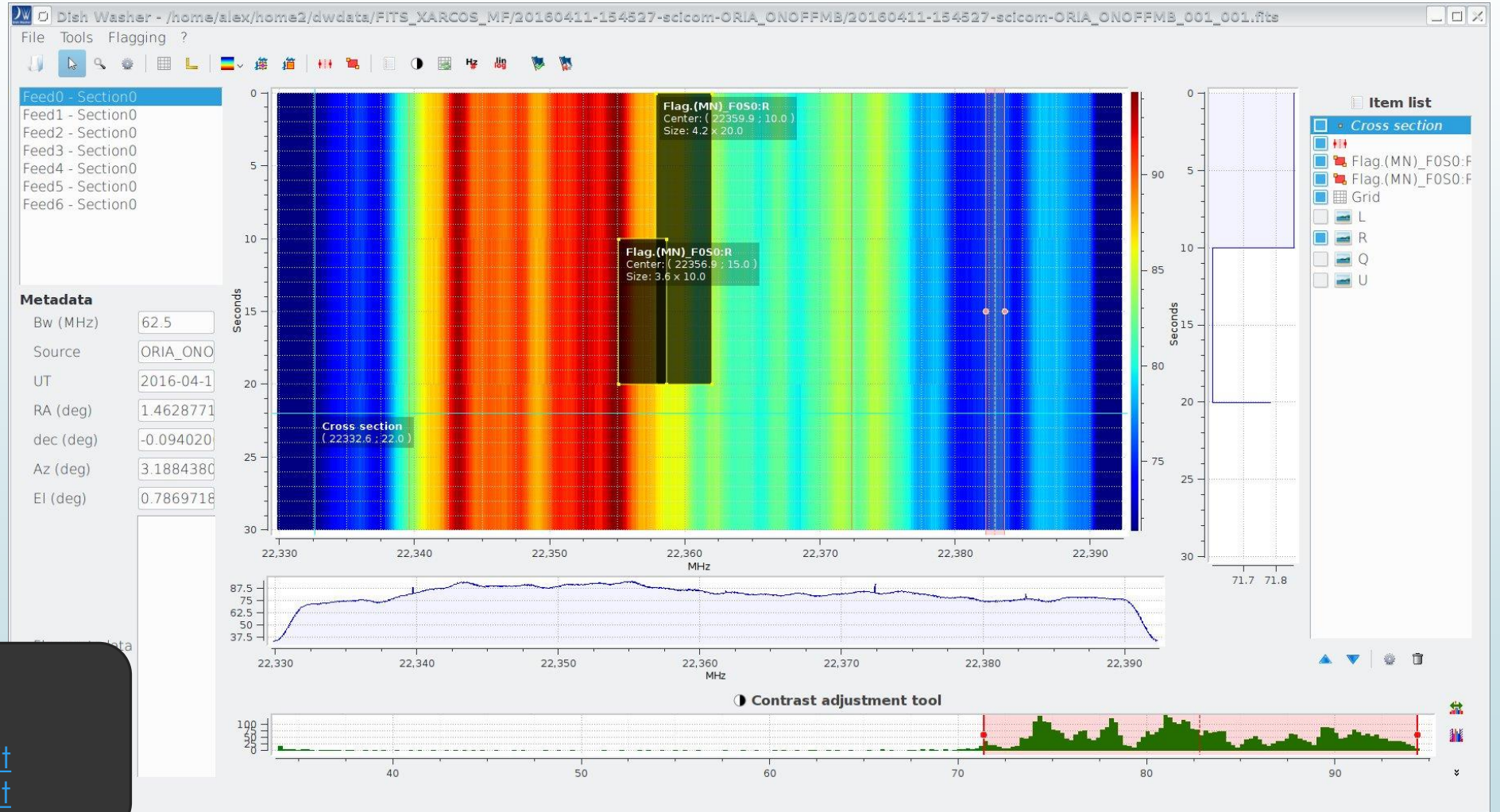
- In two years (2014-2016) the project has allowed to:
 - create a network of people from different observatories/facilities of the INAF and external partners to face together the RFI issue affecting the RA observations
 - support the local RFI groups in increasing the RFI monitoring at the national telescopes (RFI environment is constantly evolving)
 - develop HW/SW tools in common at IRT for a national RFI mitigation
 - WBLGB spectrometer by using the local know/How in using ROACH-FPGA technologies and Casper development environment
 - Dish Washer based on a python free software providing:
 - an user-friendly GUI
 - Manual RFI flagging
 - Automatic RFI flagging (at moment a basic sigma-clipping has implemented)
 - In general, a software platform where sophisticated automatic methods for RFI mitigation can be implemented

Ongoing developments

Since the end of project (July 2016) the same network of people have been still working on:

- Finish DW debugging and test with various types of data (spectropolarimeters, telescopes, also involving the international collaborators) → almost completed
- Public release: it requires improving/testing some graphical functionalities and updating the documentation → almost completed
- Implement more sophisticated RFI detection algorithms → to be done (collaboration with other people involved in RA community would be very welcome)
- Implement support for different input data formats → Start/continue collaboration with other facilities (GBT tests encouraging!)

Thanks for your attention. Questions?



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