



MeerKAT Simulation Studies

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Overview

“MeerKAT is South Africa’s 1% SKA demonstrator interferometer. The 80 dish telescope is designed to be a versatile instrument ideal for a wide range of science cases. I will provide some insights into the configuration design and the simulation study currently being started. From a scientific perspective, I will also provide an overview of some of the large HI science cases that are being planned for MeerKAT, as well as some results from my current work in the comparison of HI and CO dynamics of local galaxies.”

Overview

- MeerKAT Configuration
 - Background
 - Current Status
- Performance Simulations
 - Point Sources
 - Extended Emission

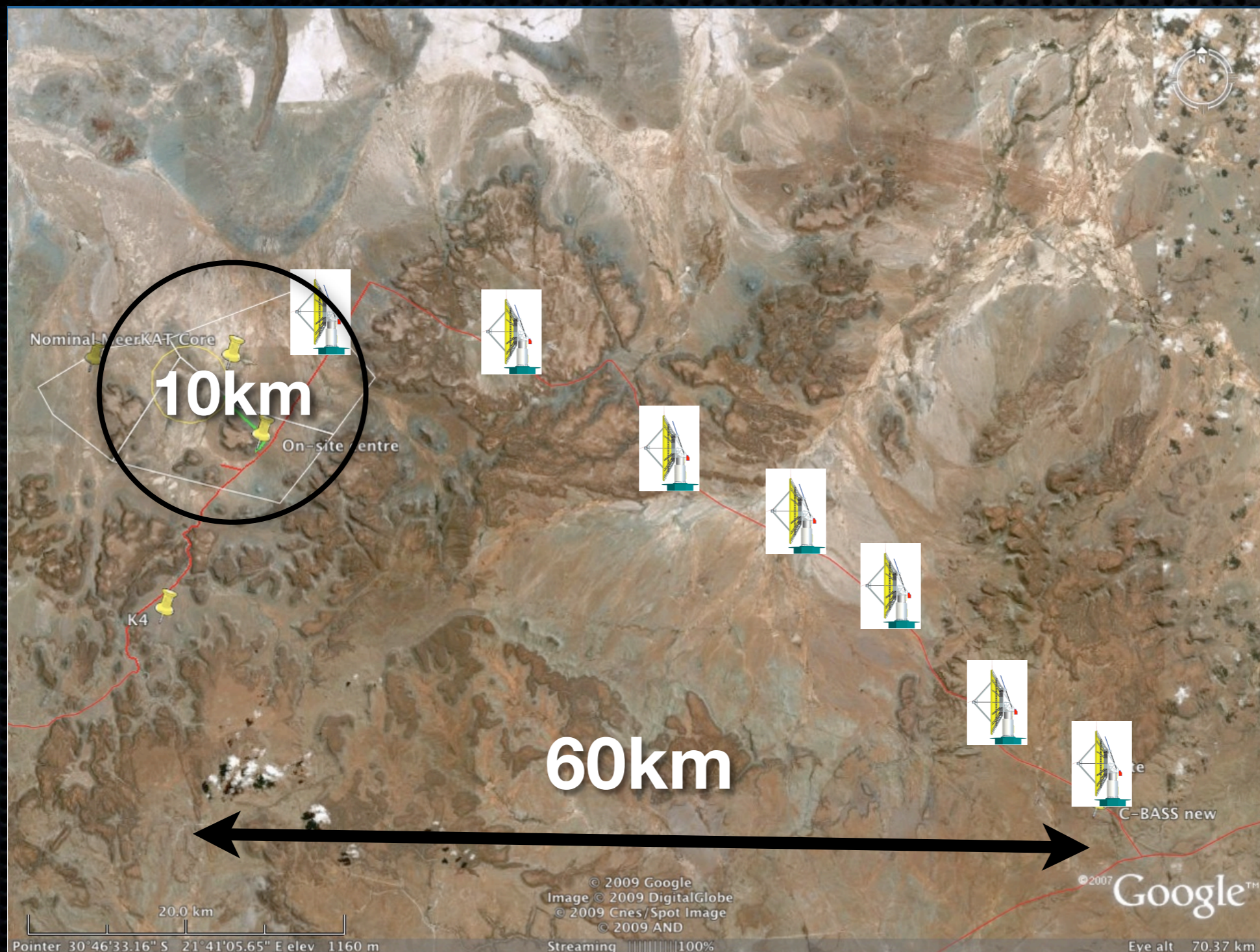
MeerKAT

Number of dishes ^a	80 (central array) + 7 (spur)
Dish diameter	12 m
Aperture efficiency	0.7
System temperature	30 K
Low frequency range ^a	0.58–2.5 GHz
High frequency range ^a	8–14.5 GHz
Field of view	1 deg ² at 1.4 GHz 6 deg ² at 580 MHz 0.5 deg ² at 2 GHz
A_e/T_{sys}	200 m ² /K
Continuum imaging dynamic range ^b	1:10 ⁵
Spectral dynamic range ^b	1:10 ⁵
Instrumental linear polarisation purity	–25 dB across field
Minimum and maximum bandwidth per polarization ^a	8 MHz–4 GHz
Number of channels	16384
Minimum sample time	0.1 ms
Minimum baseline	20 m
Maximum baseline	8 km (without spur) 60 km (with spur)

	KAT-7 2010	Phase 1 2013	Phase 2 2014	Phase 3 2015	Phase 4 2016
Number of dishes	7	80	80	87	87
Low freq. range (GHz)	1.2–1.95	0.9–1.75	0.9–1.75	0.9–1.75	0.58–2.5
High freq. range (GHz)	—	—	8–14.5	8–14.5	8–14.5
Maximum processed bandwidth (GHz)	0.256	0.850	2	2	4
Min. baseline (m)	20	20	20	20	20
Max. baseline (km)	0.2	8	8	60	60

MeerKAT RFP

MeerKAT V1

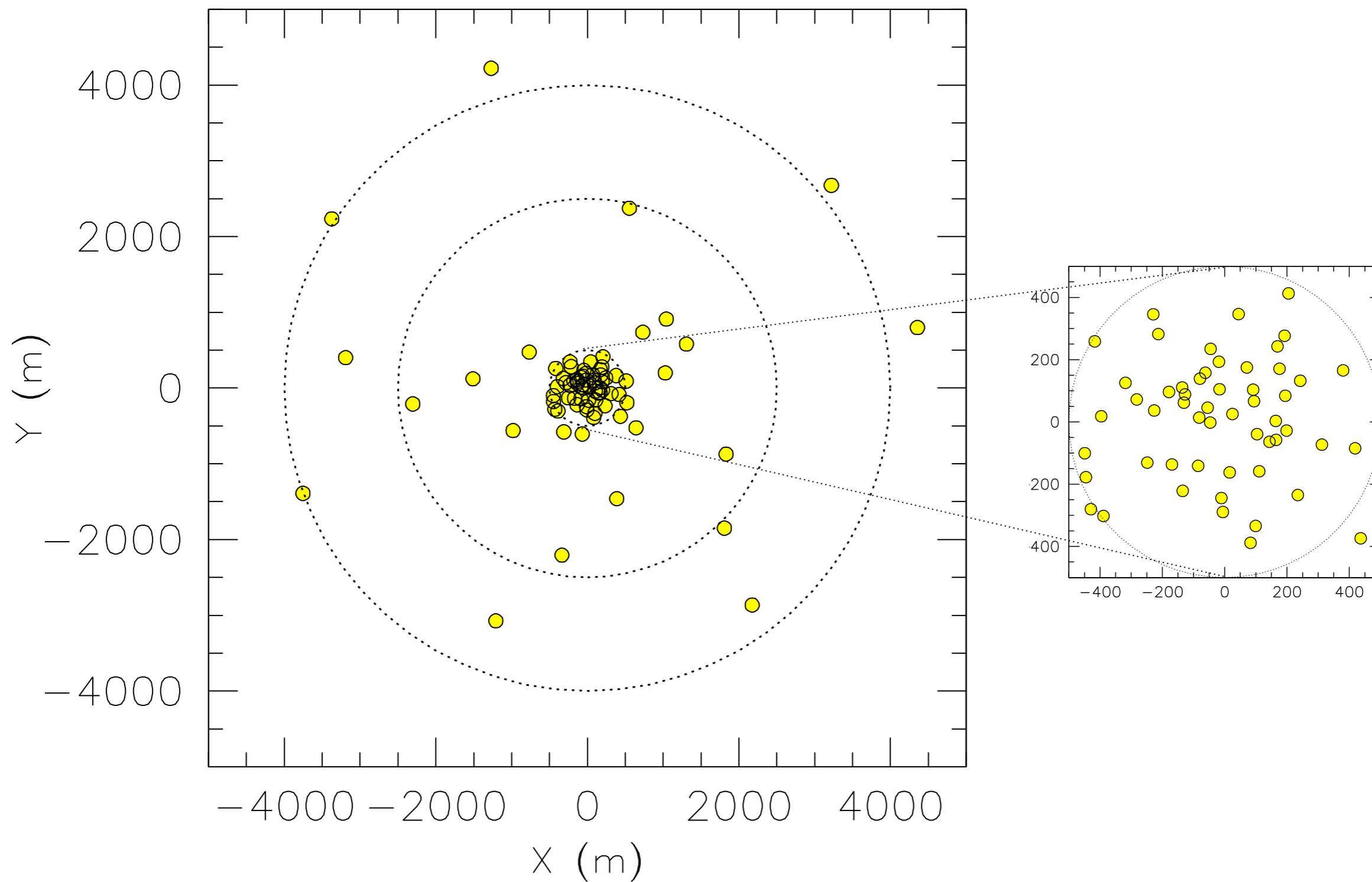


Justin Jonas

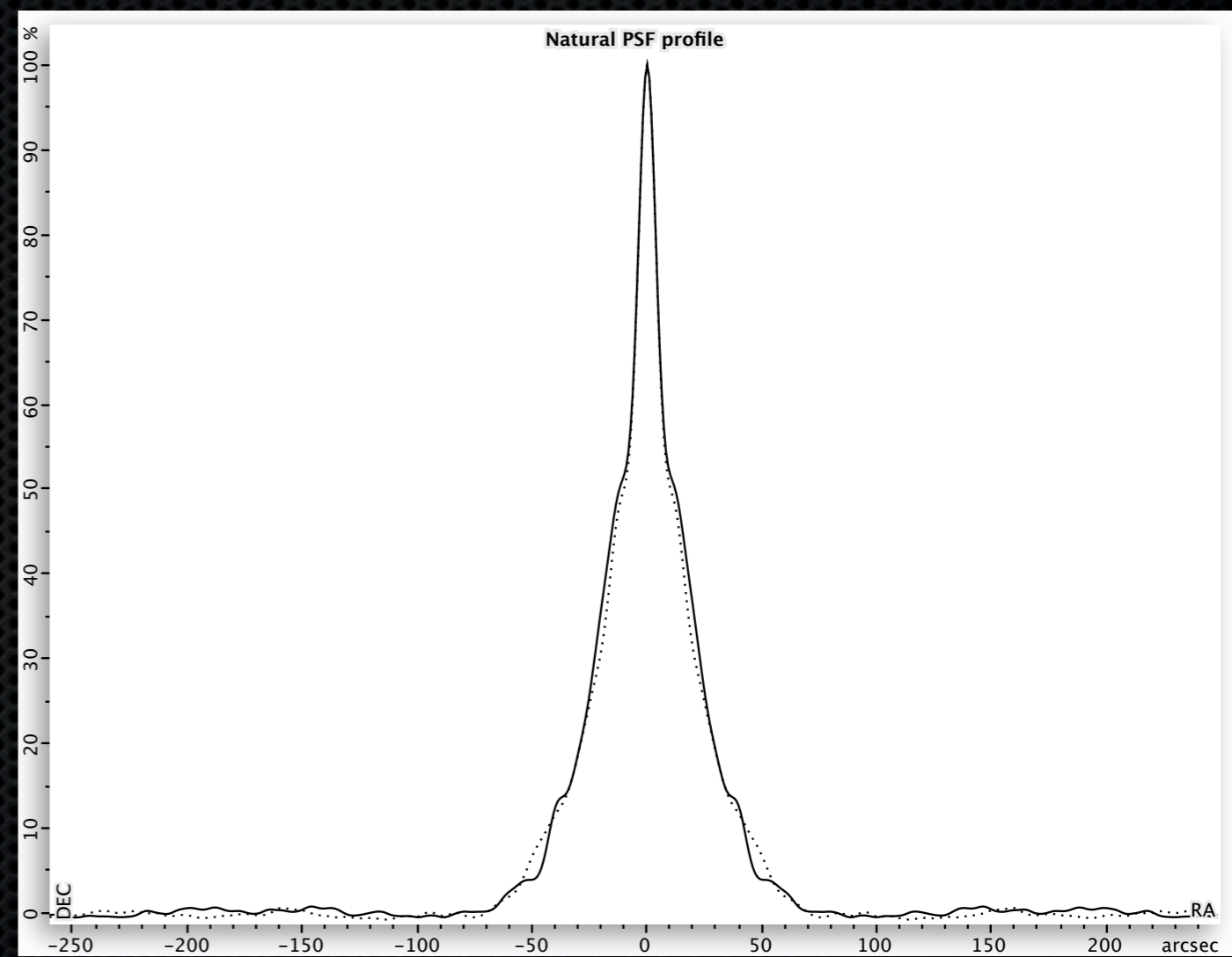
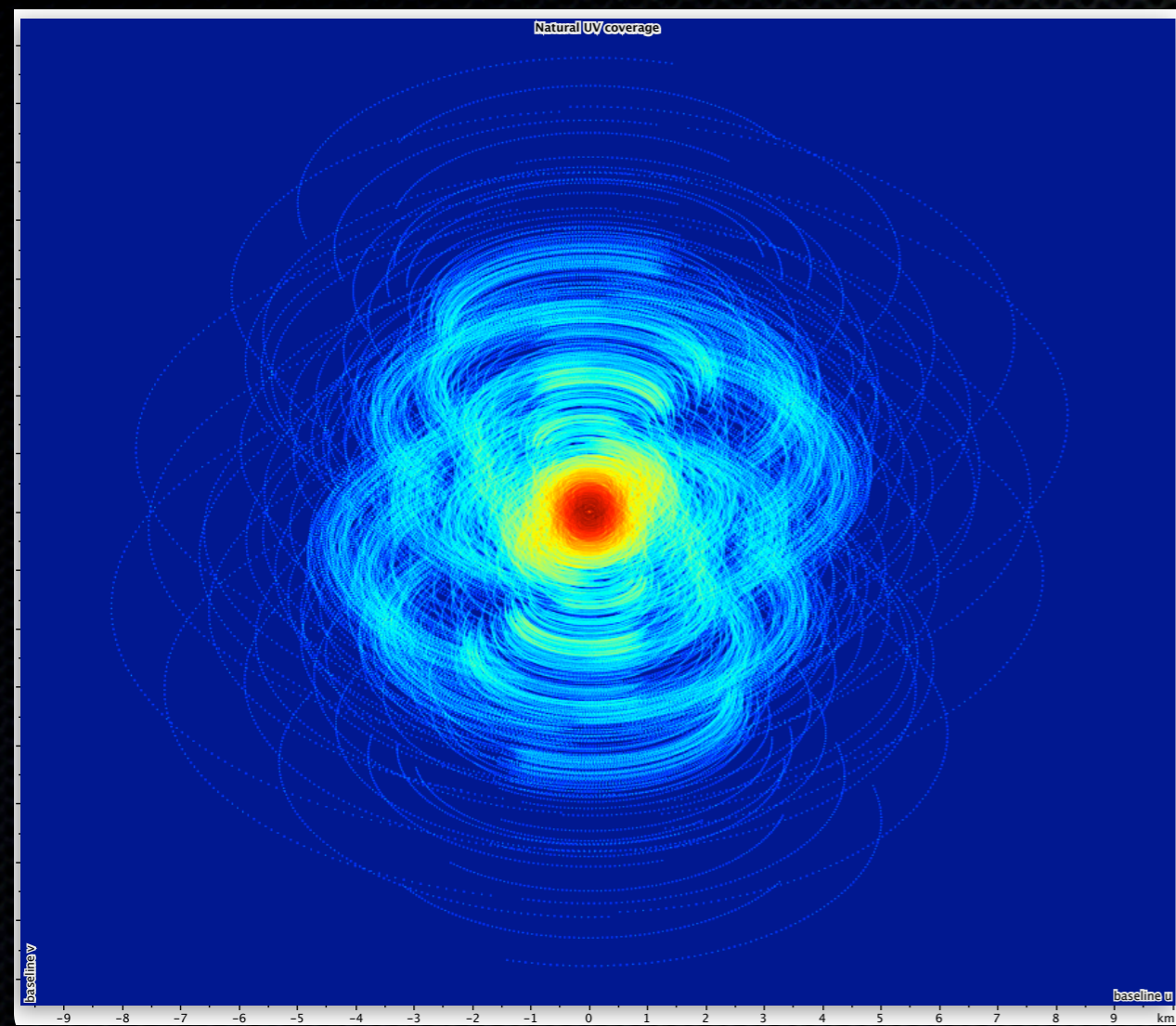
MeerKAT V1

- MeerKAT RFP
- 80 x 12m Dishes
- 2 Gaussian Components
 - 56 Dishes in $\sigma_{\text{inner}} = 300\text{m}$
 - 24 Dishes in $\sigma_{\text{outer}} = 2500\text{m}$
- Optimized for 8hr, $\delta = -30^\circ$ pointing
- Used Weighted PSS as Measure of Imaging Performance

MeerKAT V1

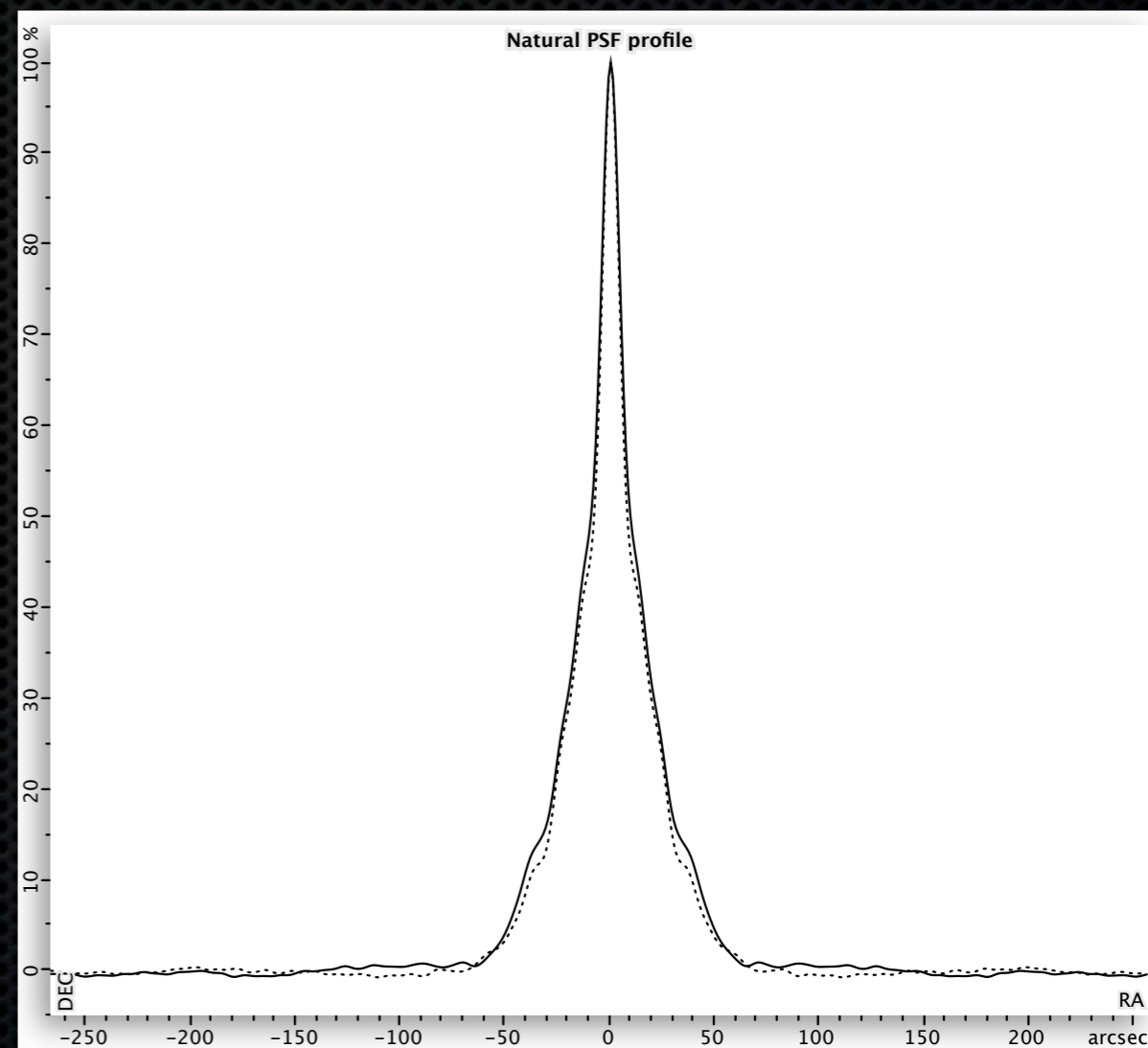
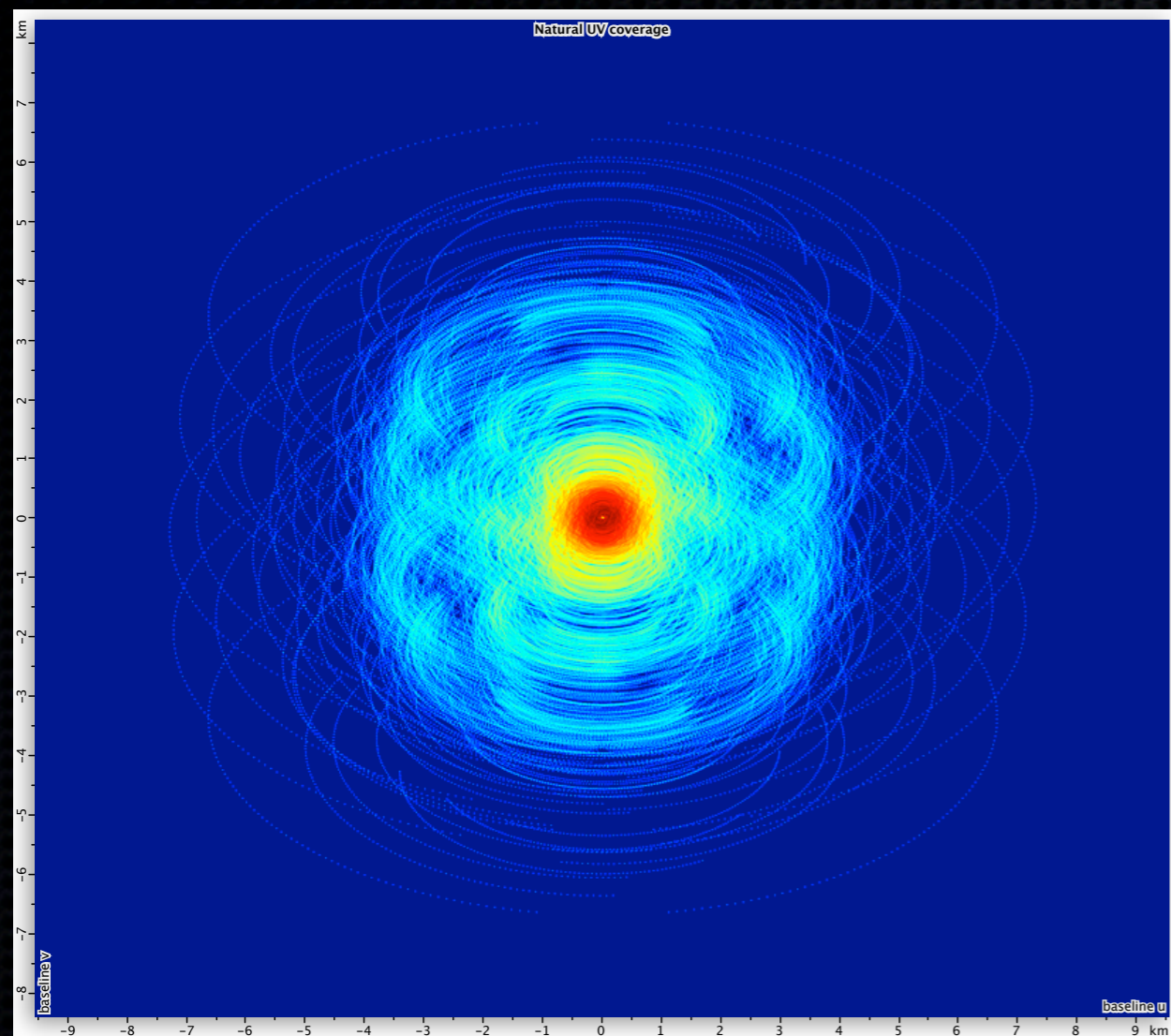


MeerKAT V1

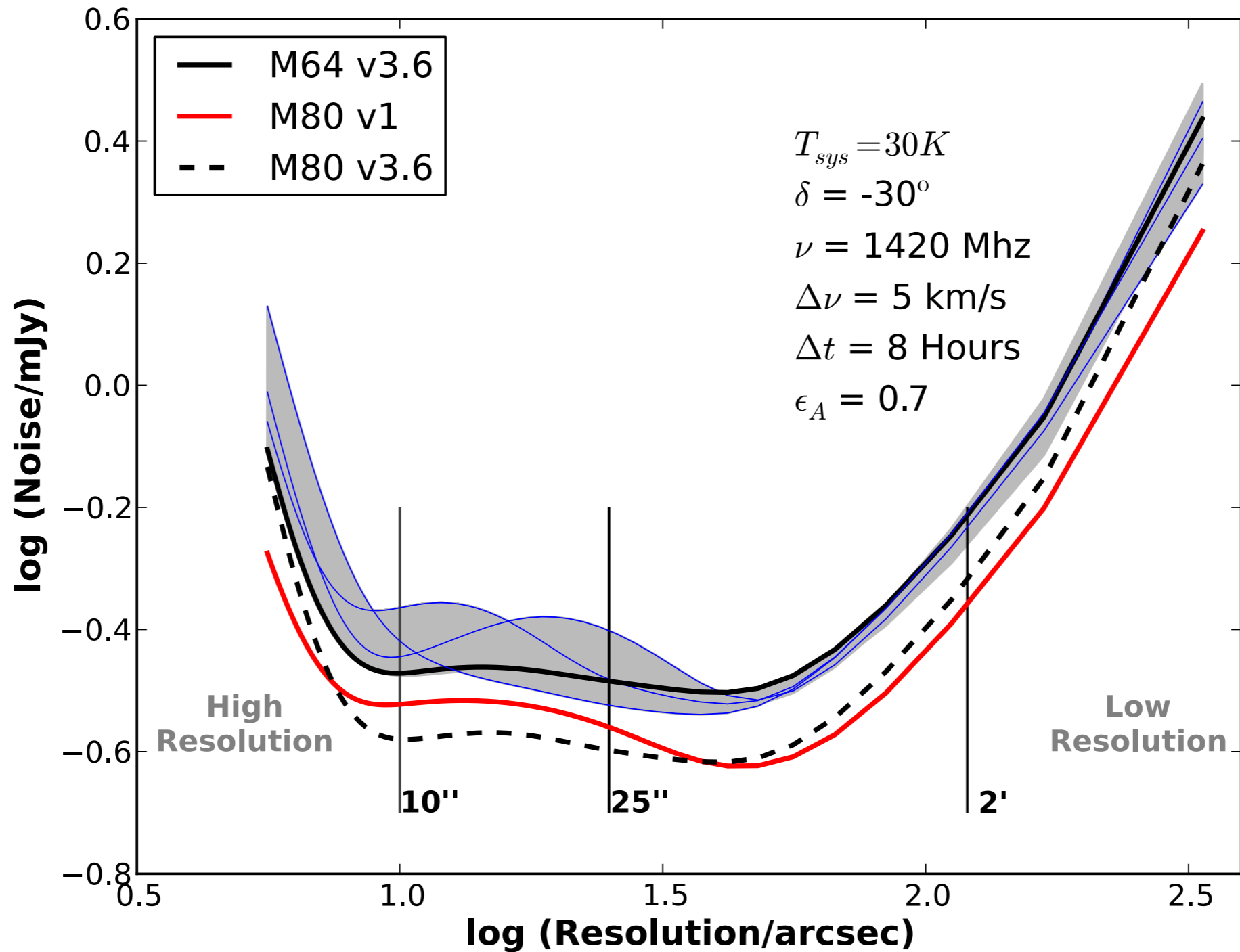


MeerKAT V3.6

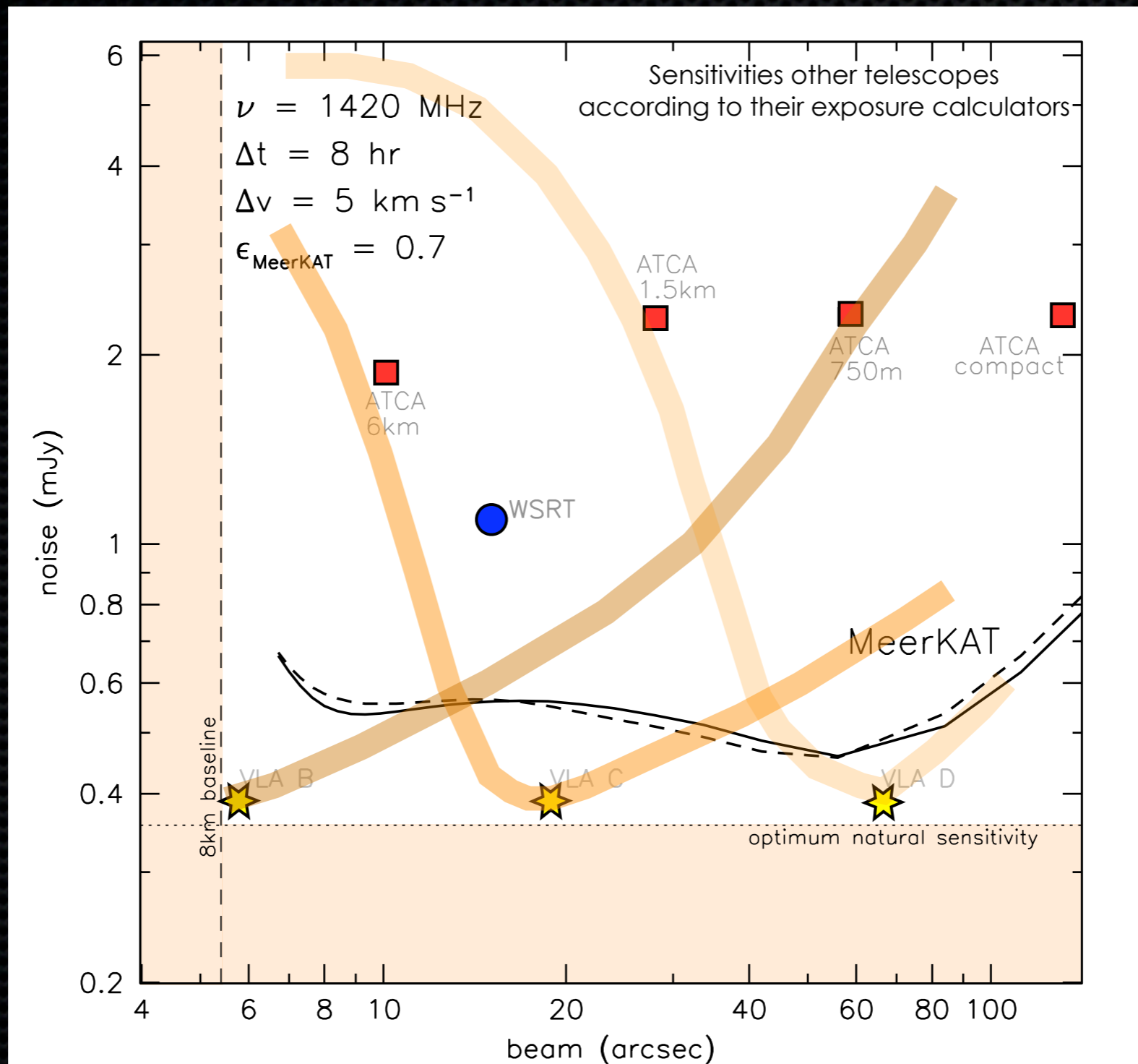
MeerKAT V3.6



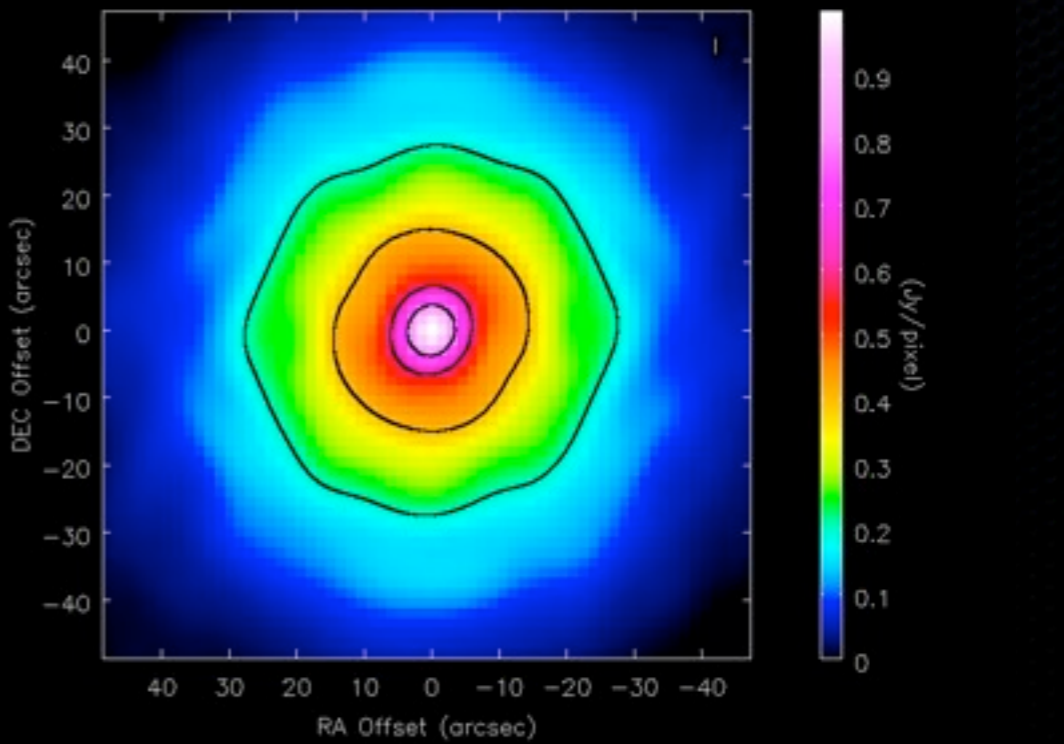
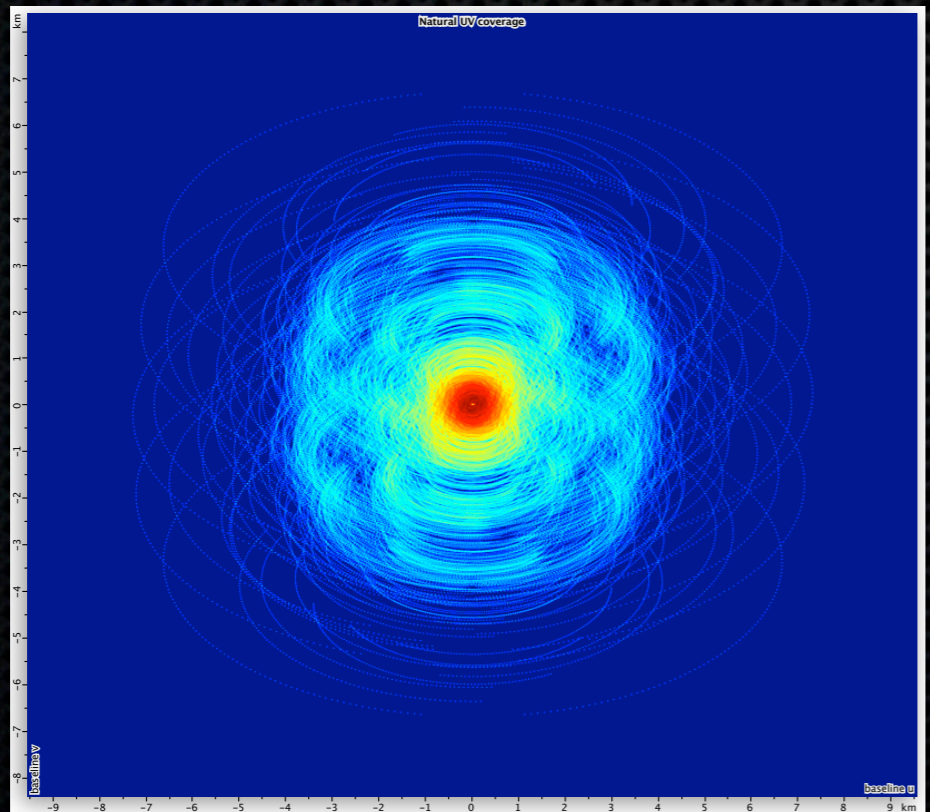
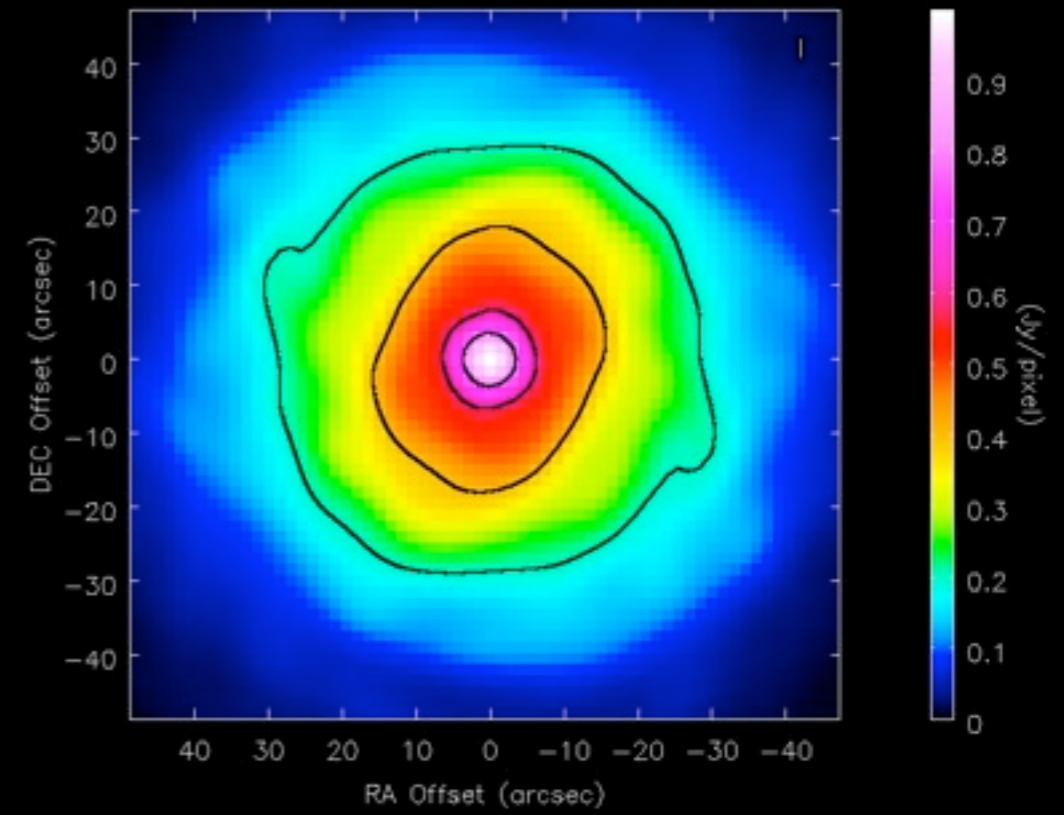
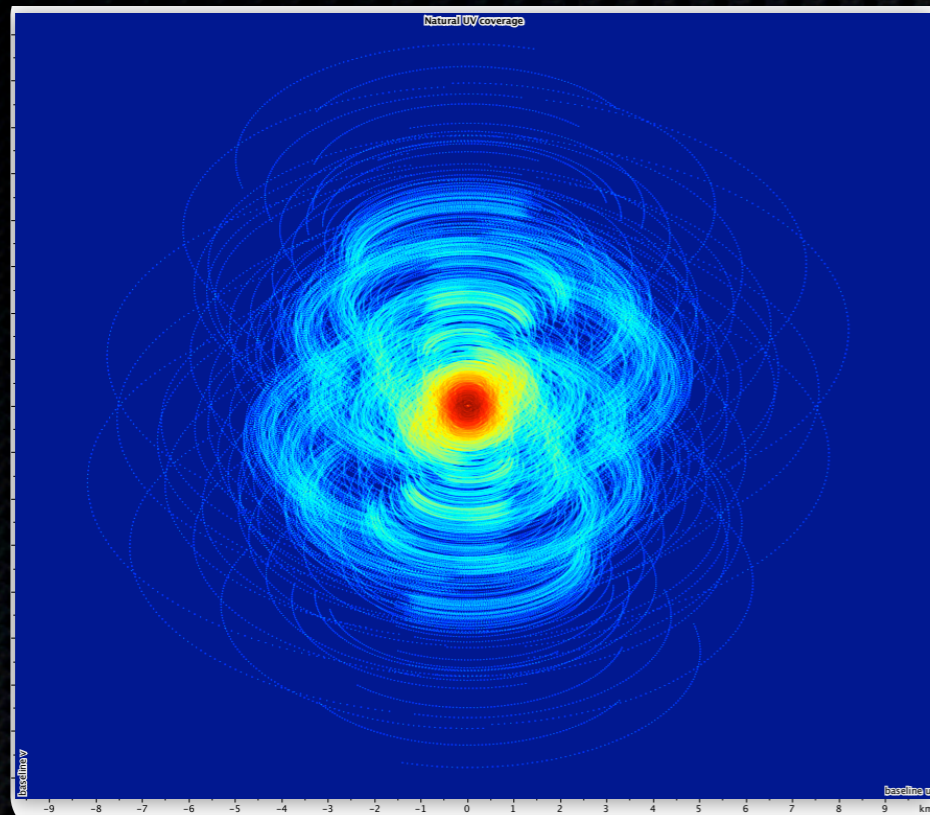
MeerKAT V3.6, 64-Dish



Sensitivities

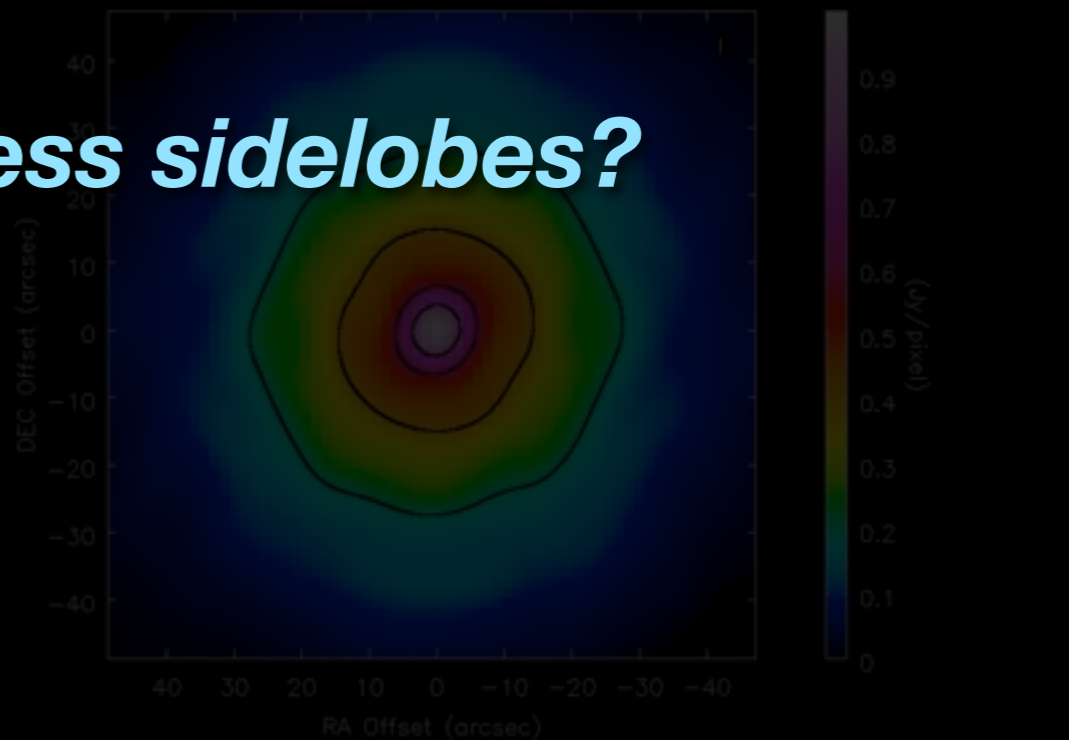
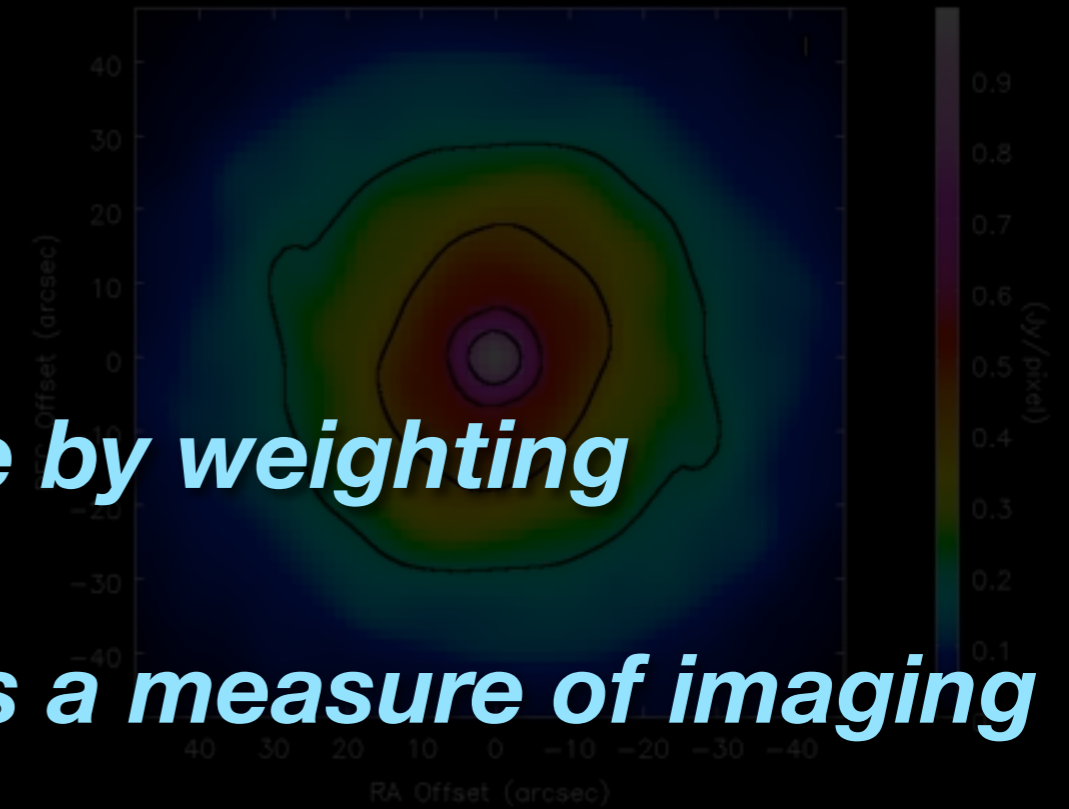
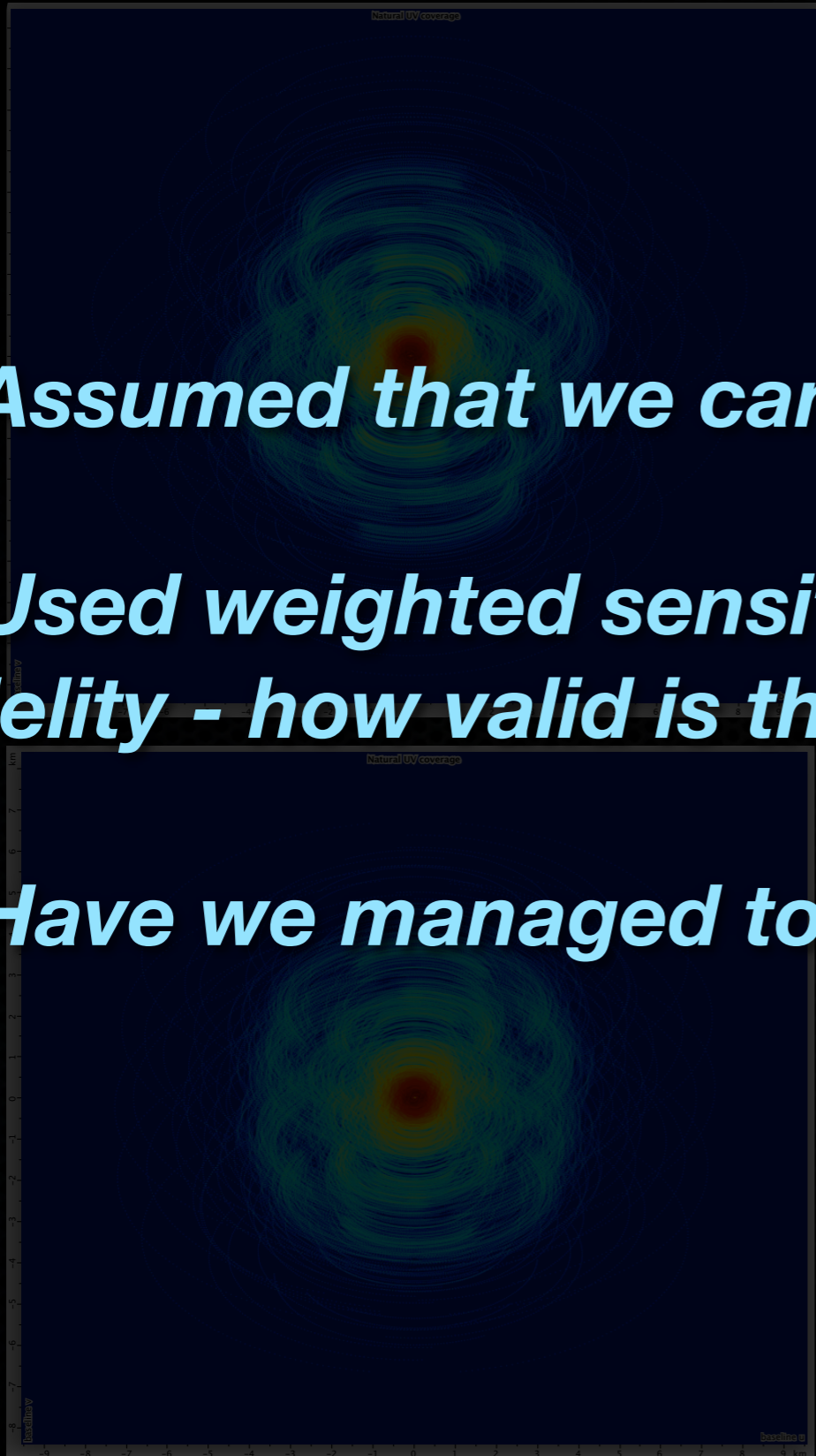


MeerKAT V1 and V3.6



MeerKAT V1 and V3.6

- *Assumed that we can image by weighting*
- *Used weighted sensitivity as a measure of imaging fidelity - how valid is this?*
- *Have we managed to suppress sidelobes?*



Simulation Studies

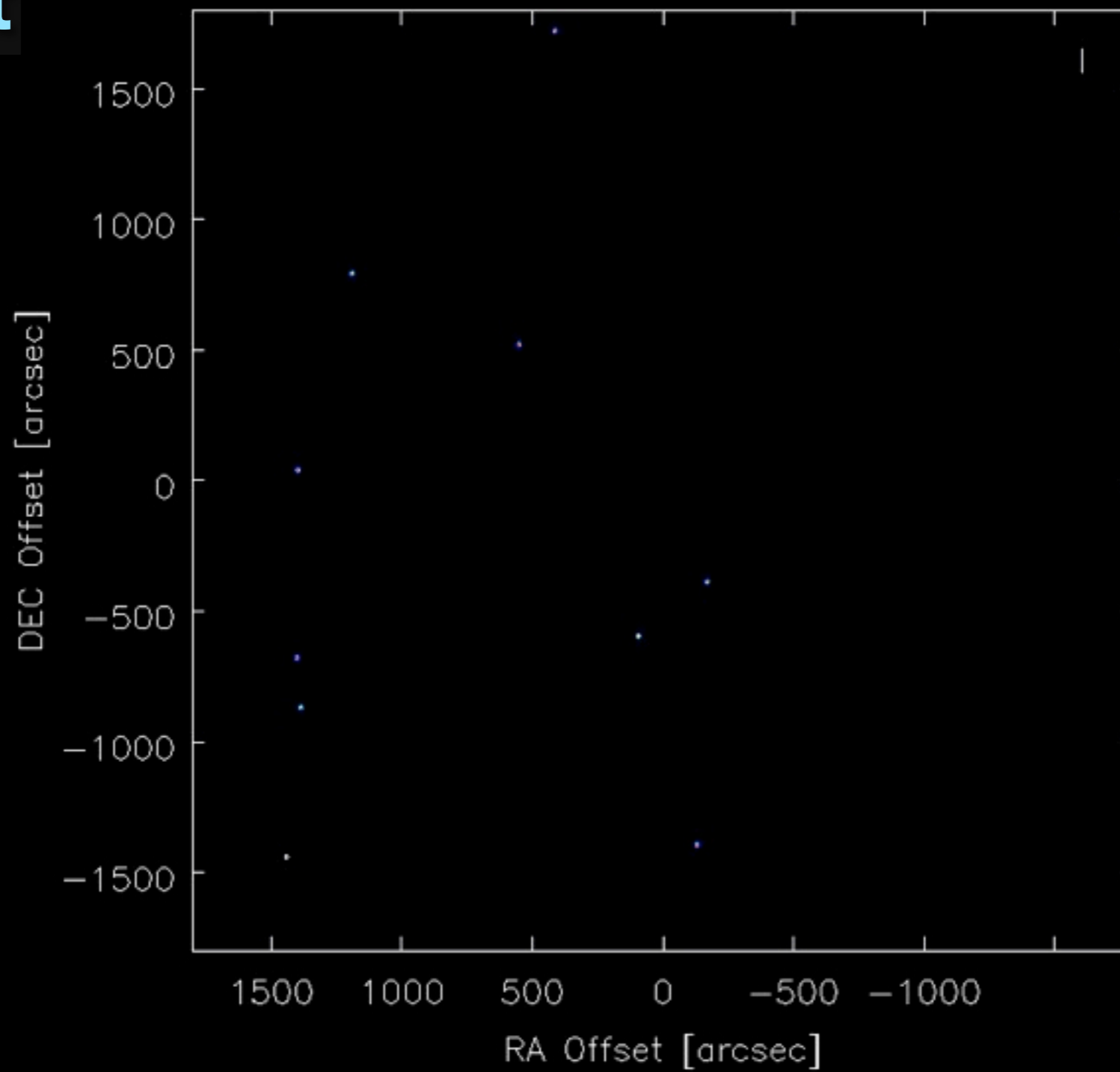
- In Support of Science Cases and Commissioning
- Uses CASA's simdata2 simulation task (hacked a little)
- Simulations
 - 10 Point Sources across FoV, Single Channel, Zenith Snapshot
 - NGC628 Clean Component Model (THINGS)
 - 58 Channel, 1.4GHz, $\Delta\nu = 12\text{kHz}$, Fitted across $0.5^\circ \times 0.5^\circ$, 8hr, $\delta = -30^\circ$

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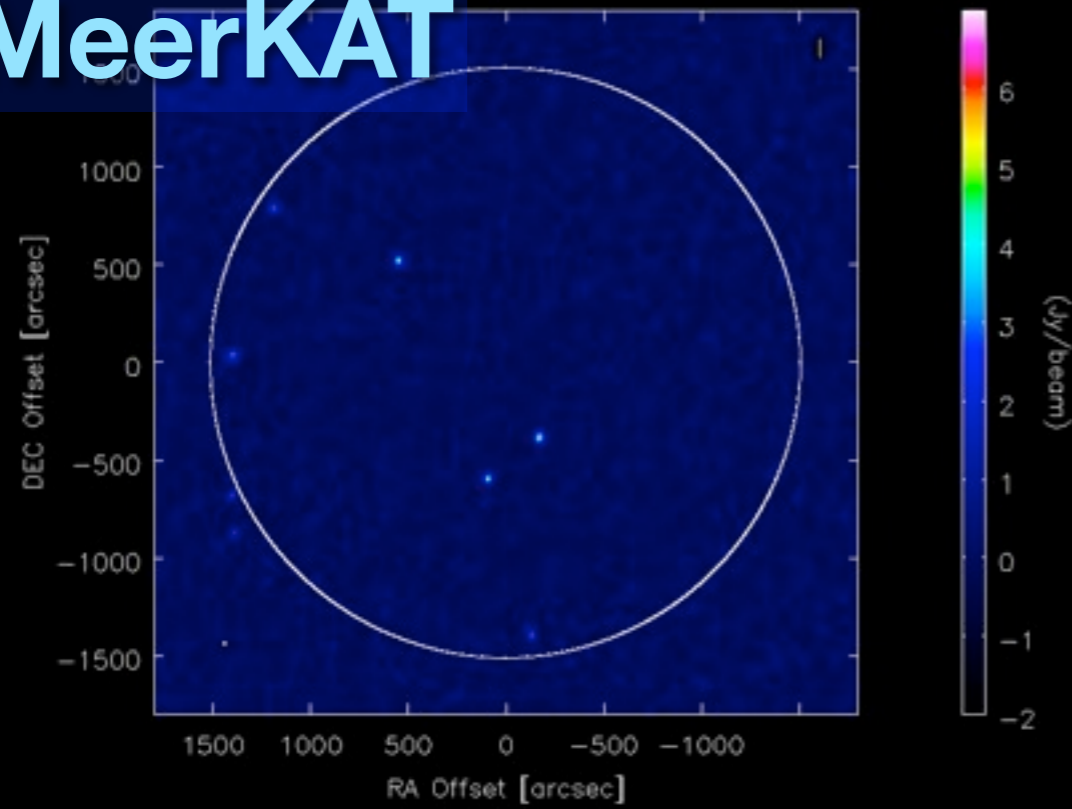
Point Sources

Input

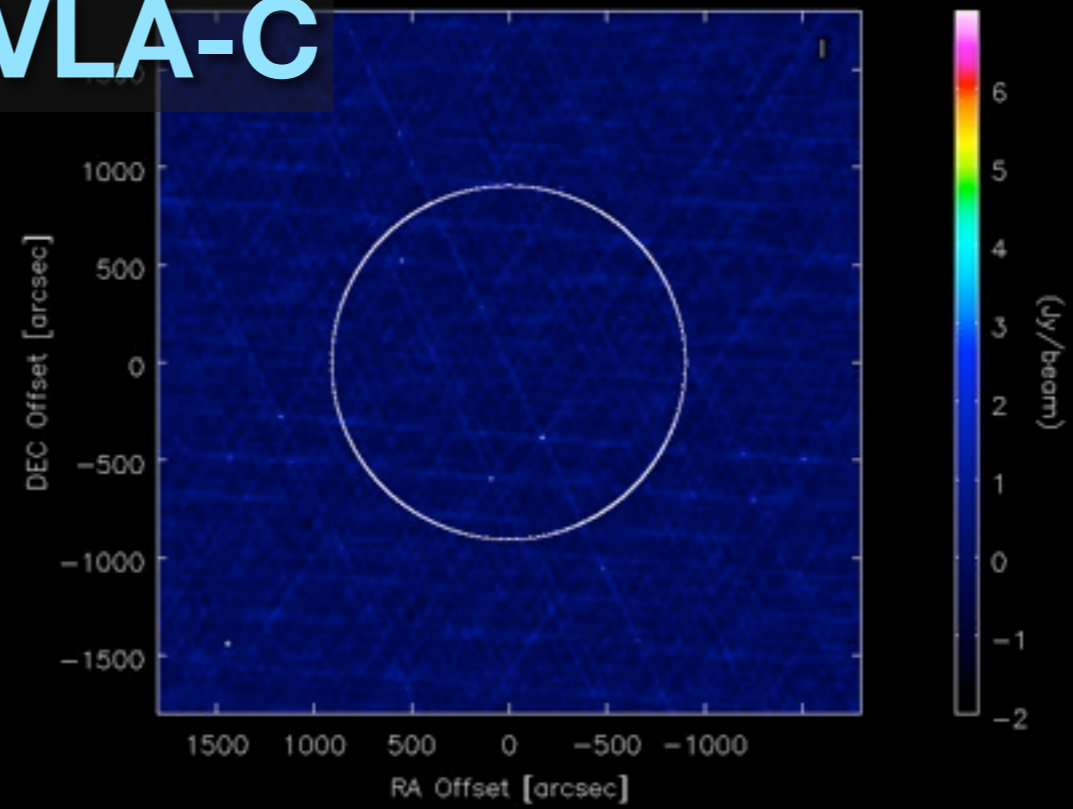


Point Sources

MeerKAT

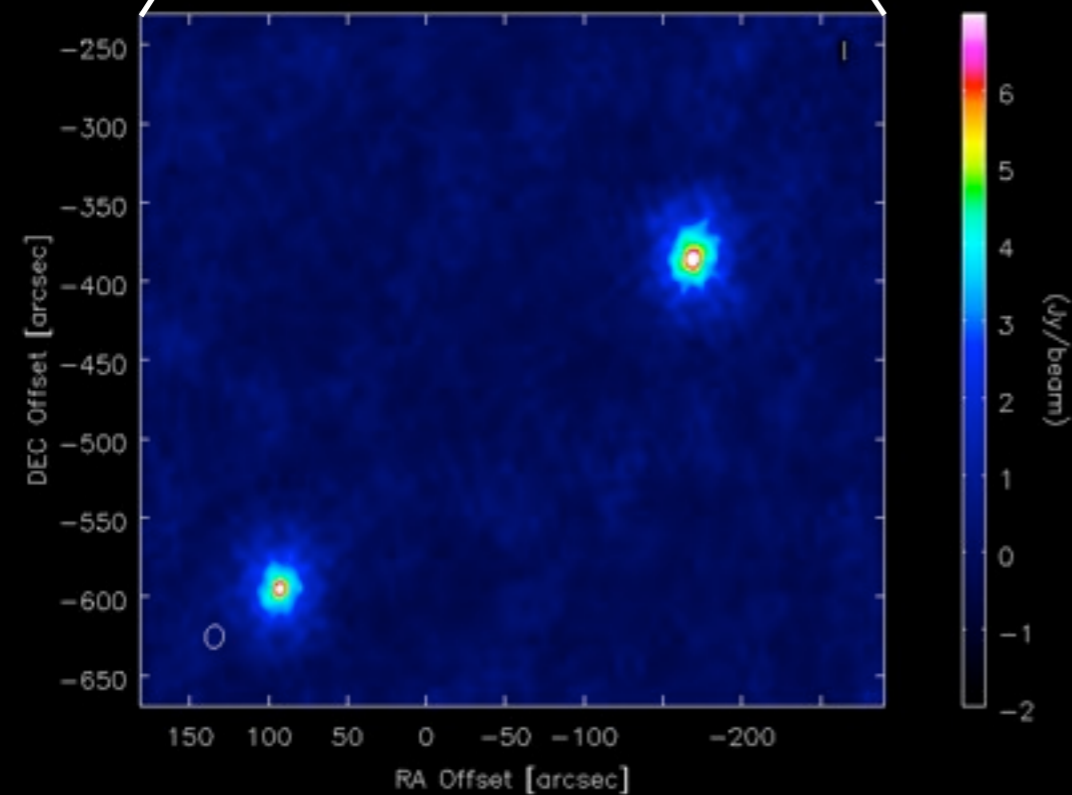
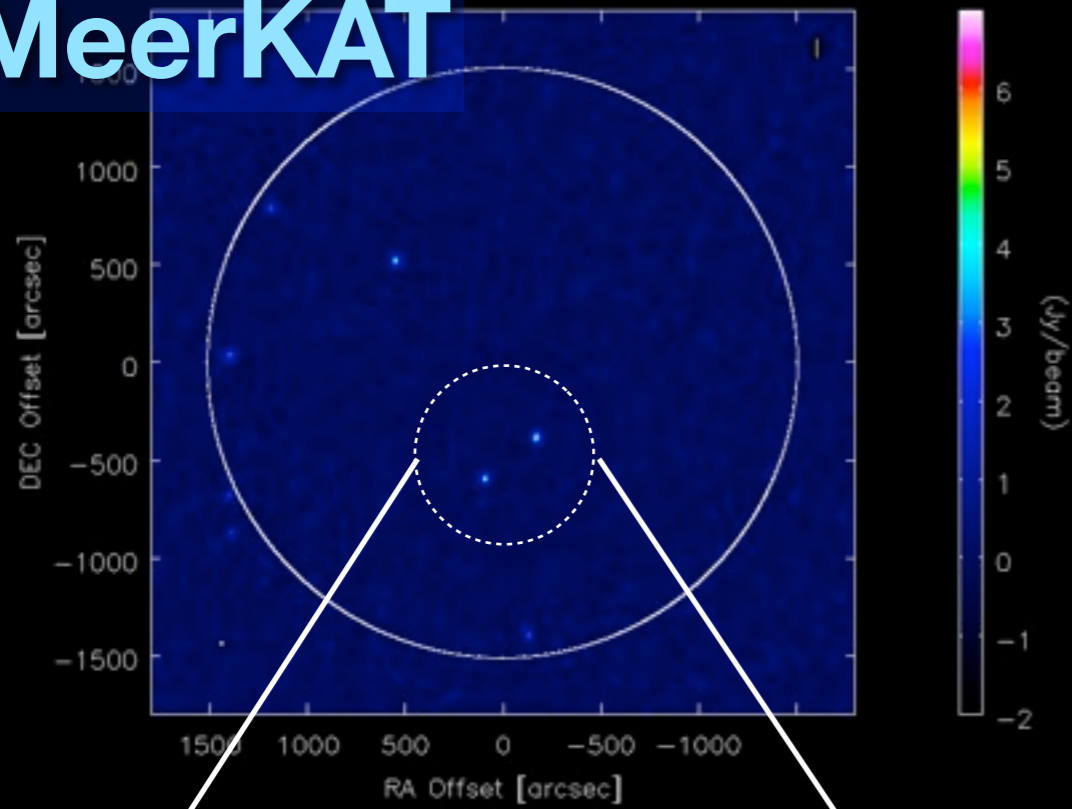


VLA-C

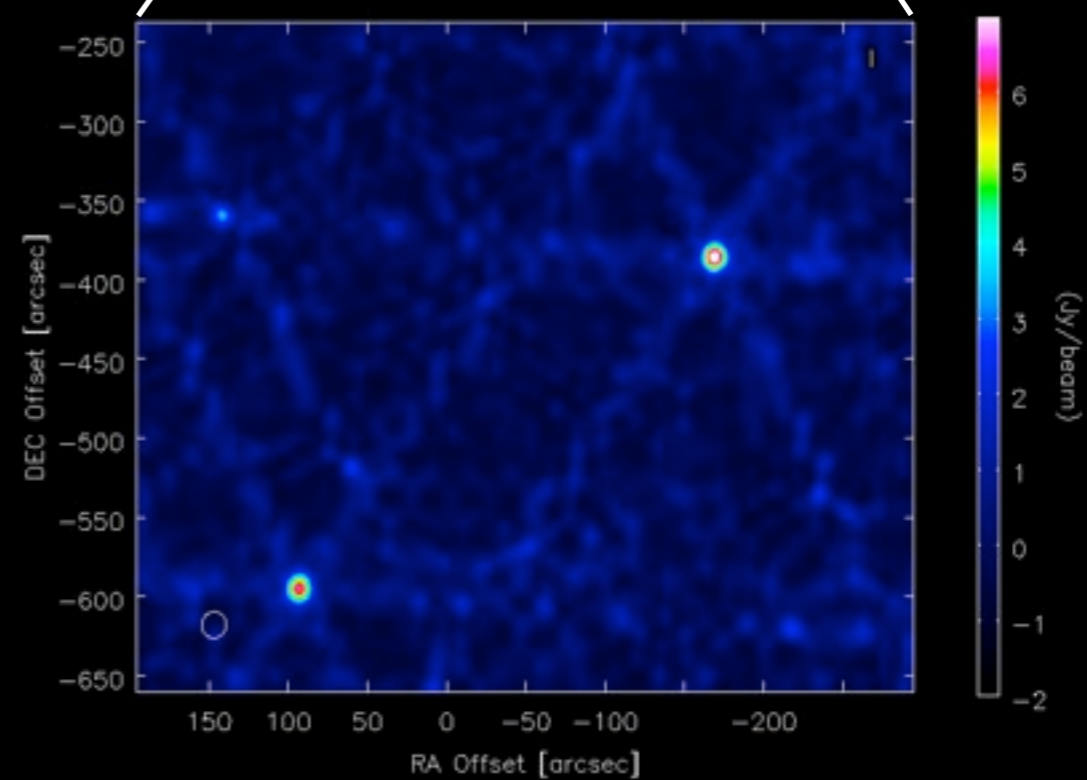
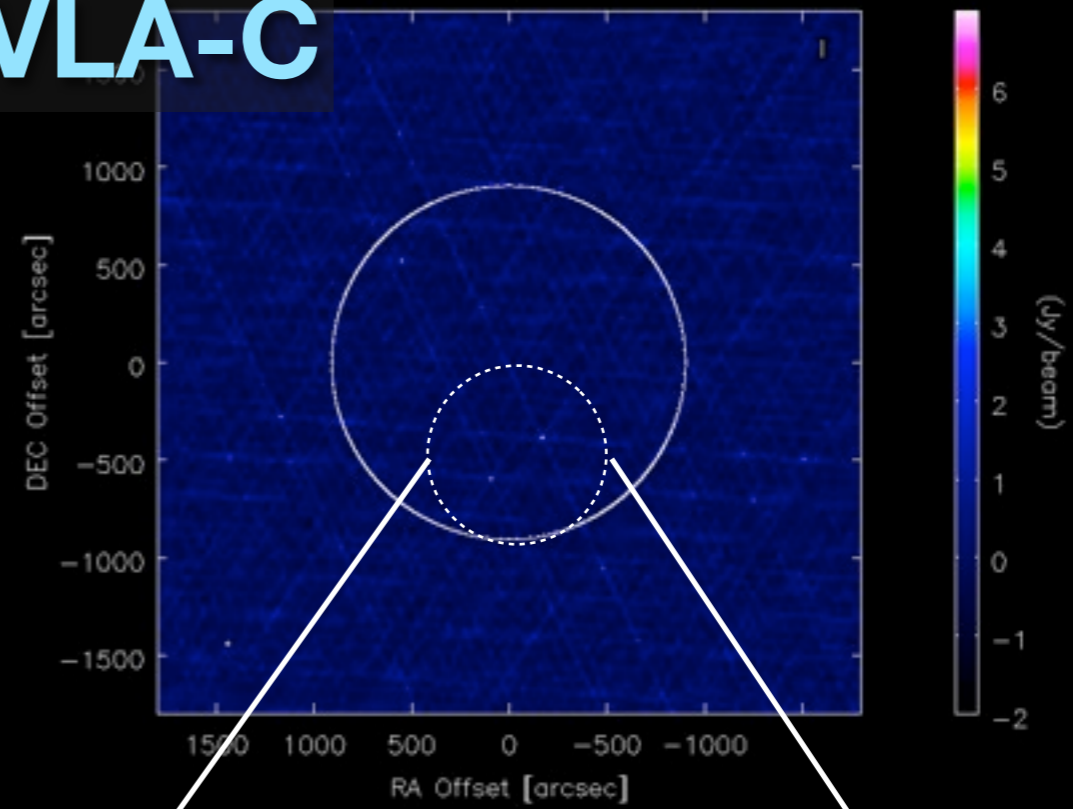


Point Sources

MeerKAT



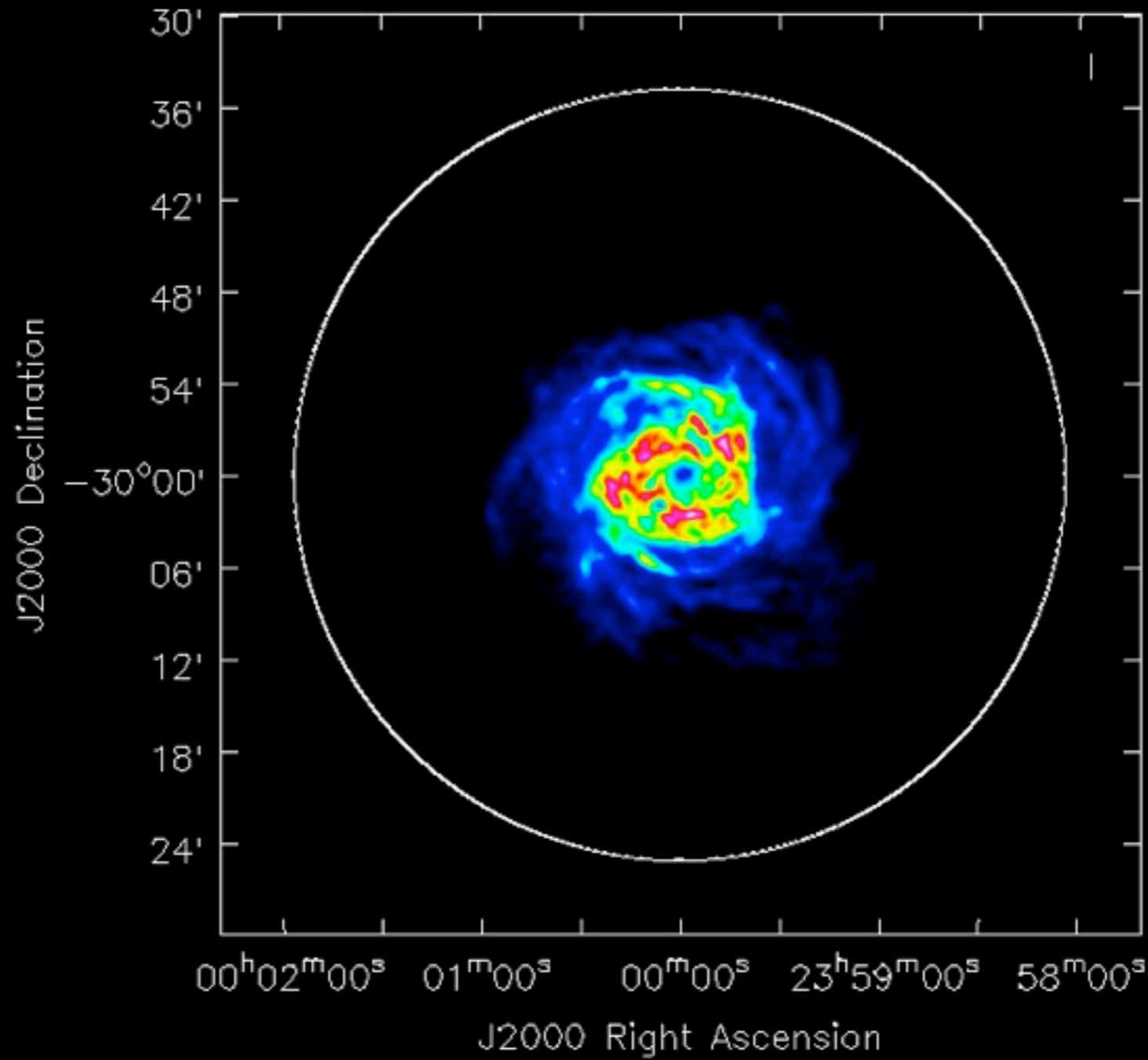
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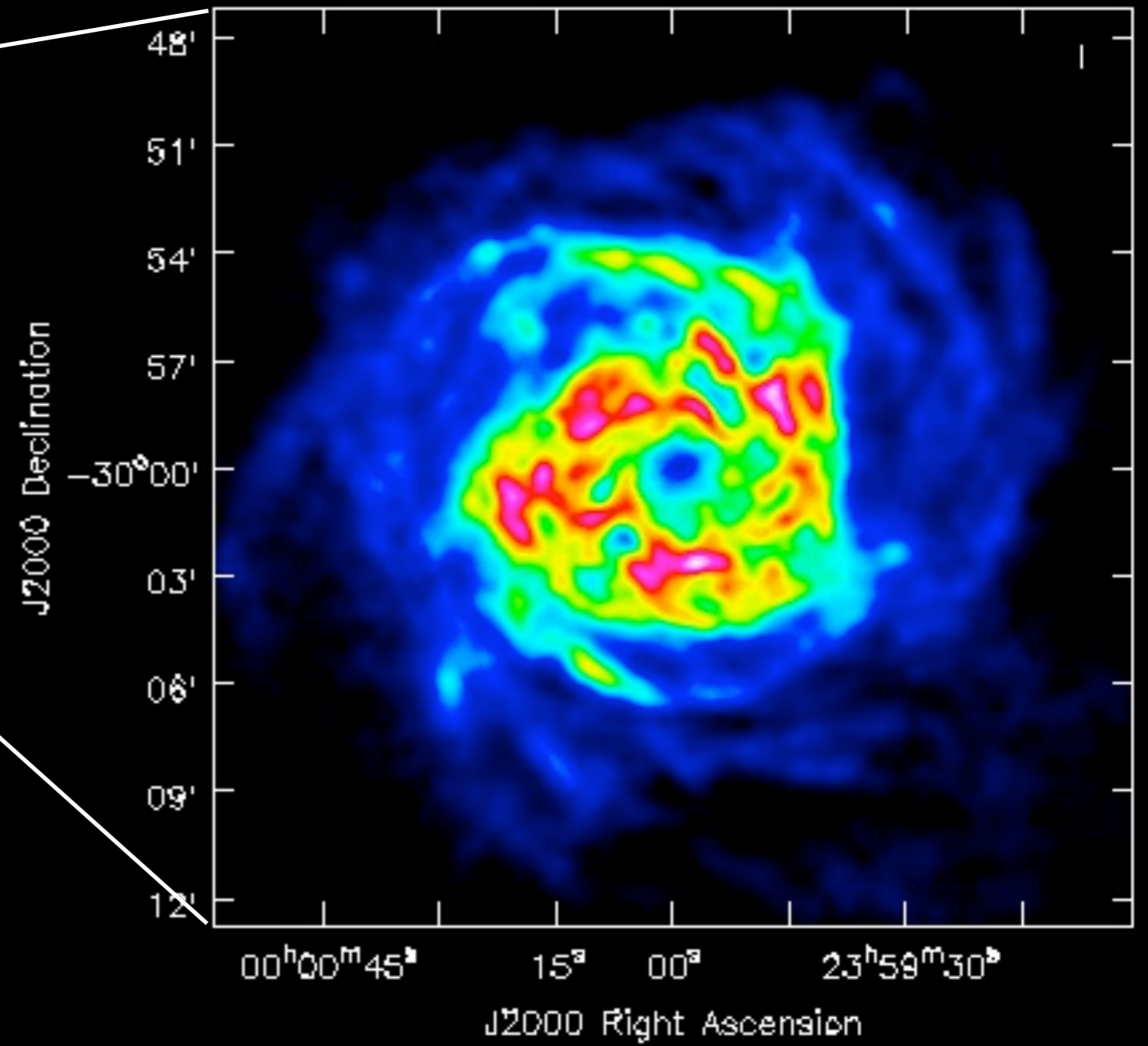
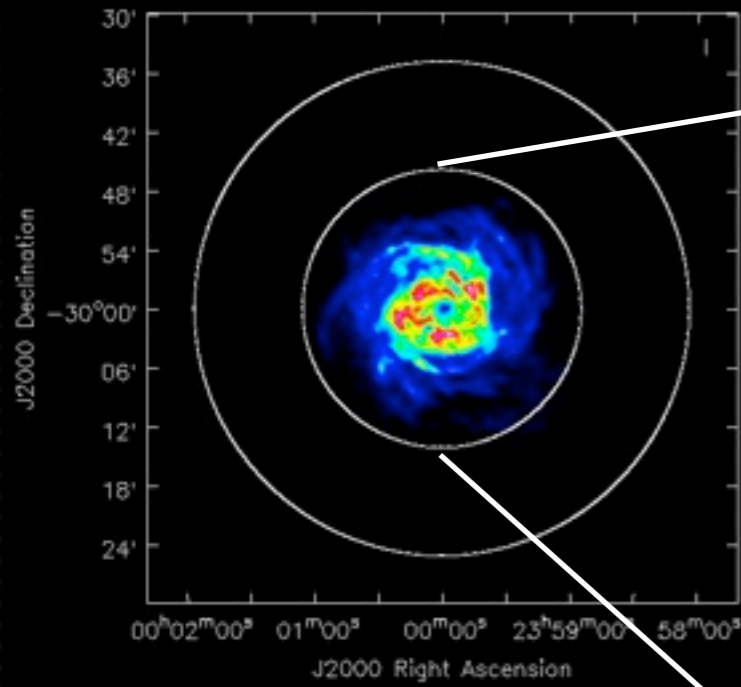
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NGC 628



NGC 628



Results

FLAT

Cleaned Image
Briggs
Robustness = 0
 $N_{\text{iter}}(\text{max}) = 1000$

Abs. Diff.

$\text{abs}[\text{Input} - \text{Cleaned}]$

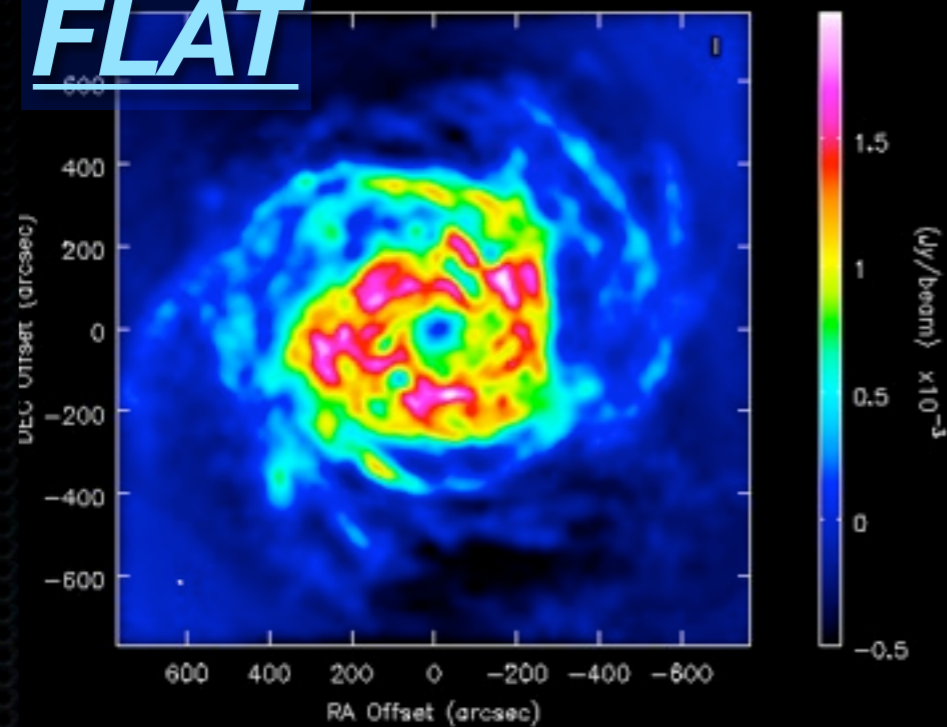
Fidelity

Input

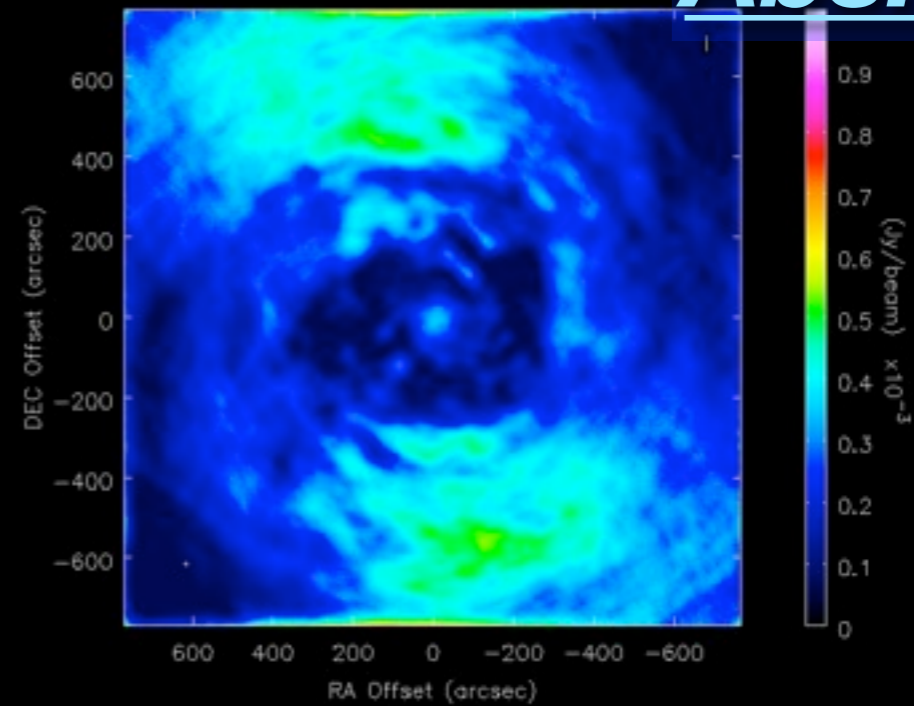
$\text{max}[\text{difference}, 0.7 * \text{rms}(\text{difference})]$

MeerKAT V1

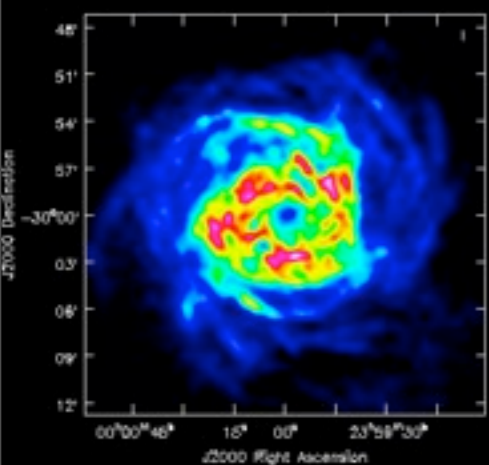
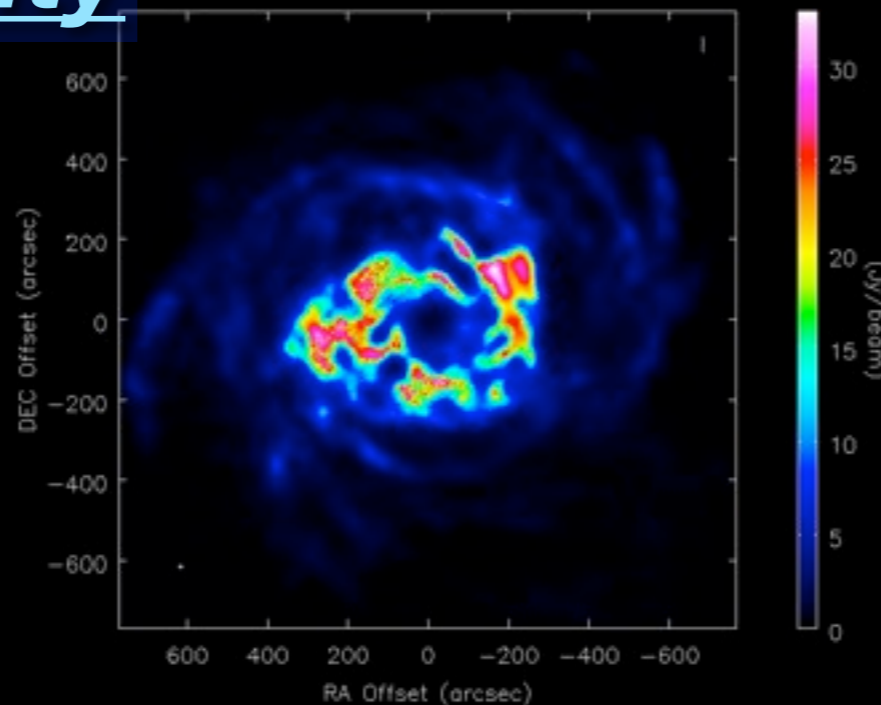
FLAT



Abs. Diff.

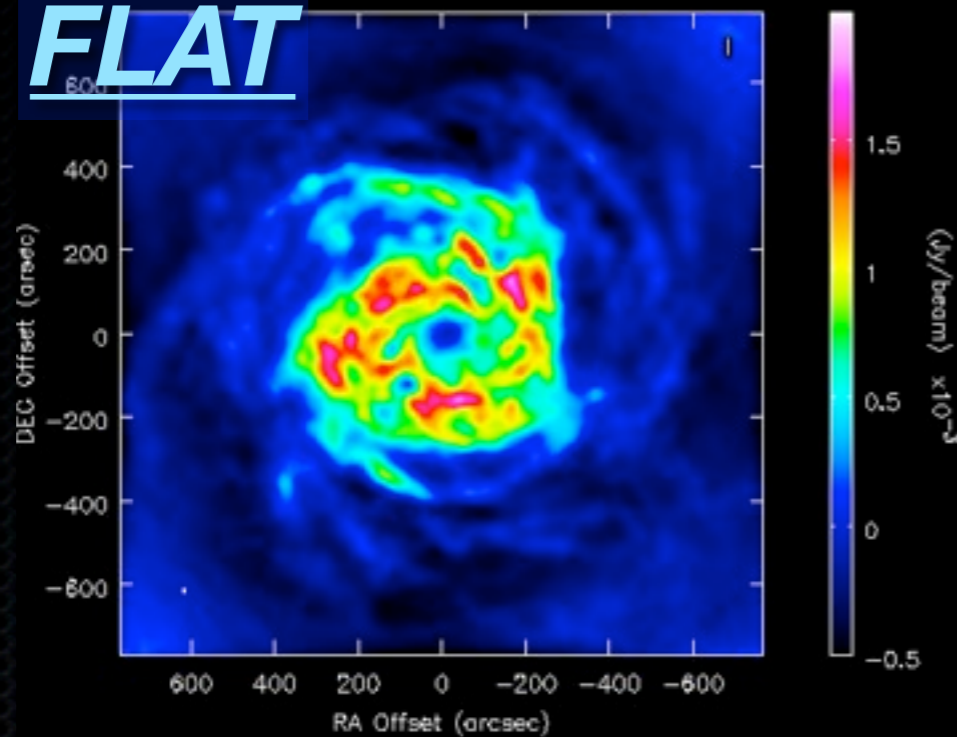


Fidelity

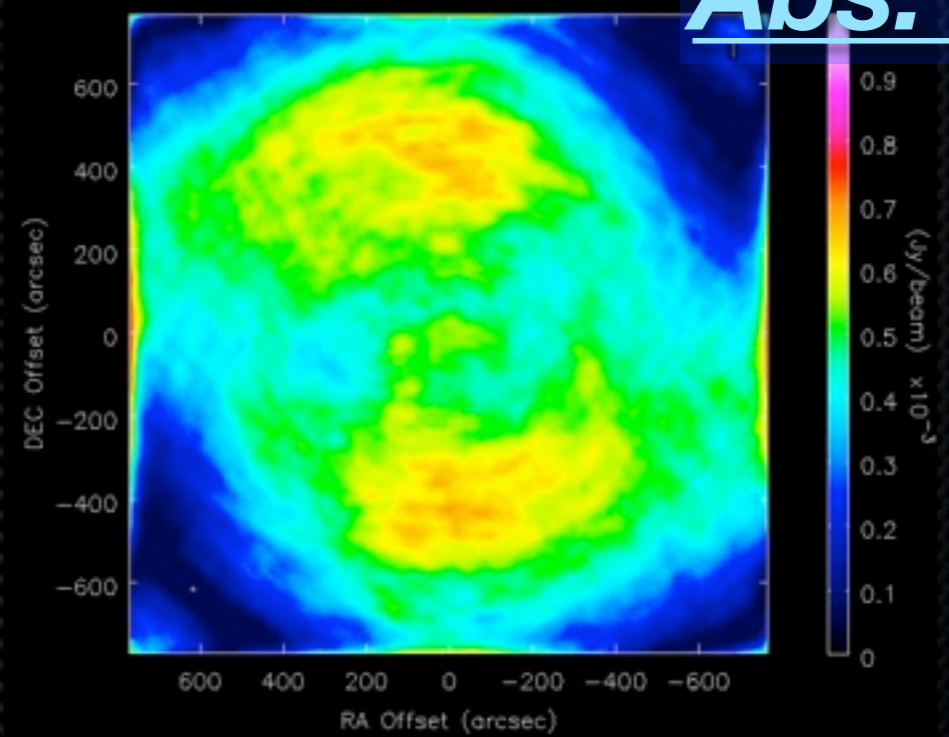


MeerKAT V3.6

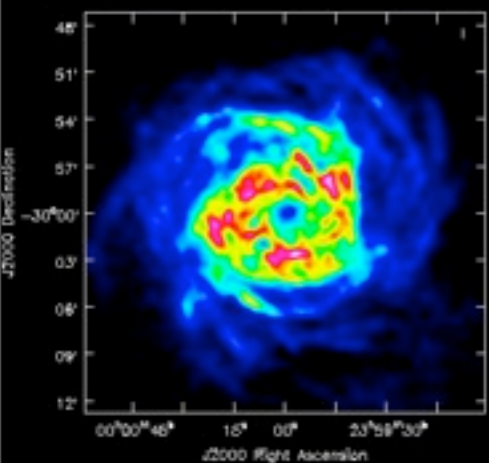
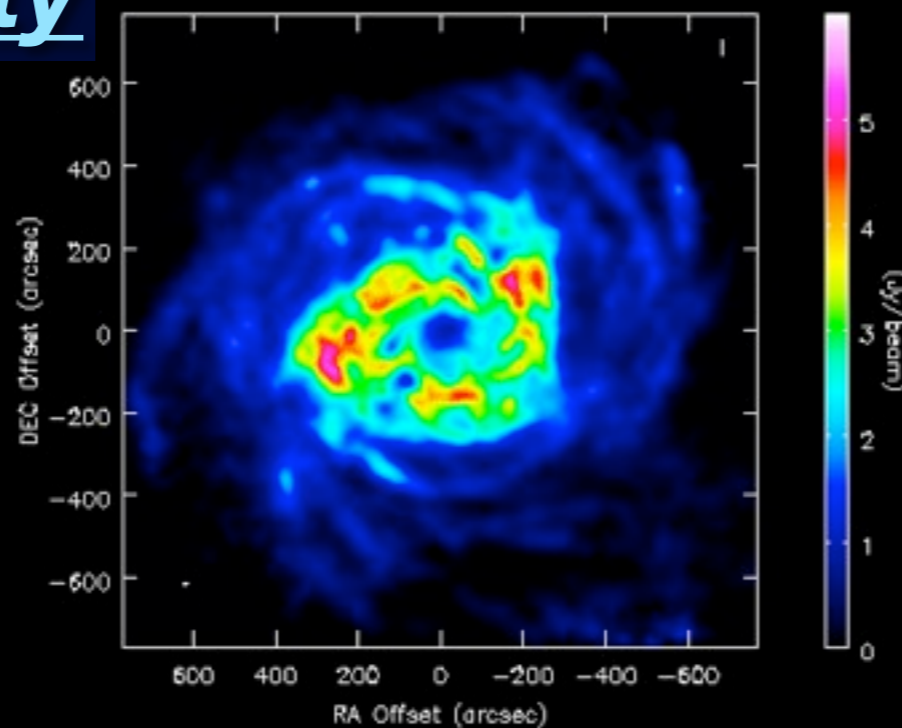
FLAT



Abs. Diff.

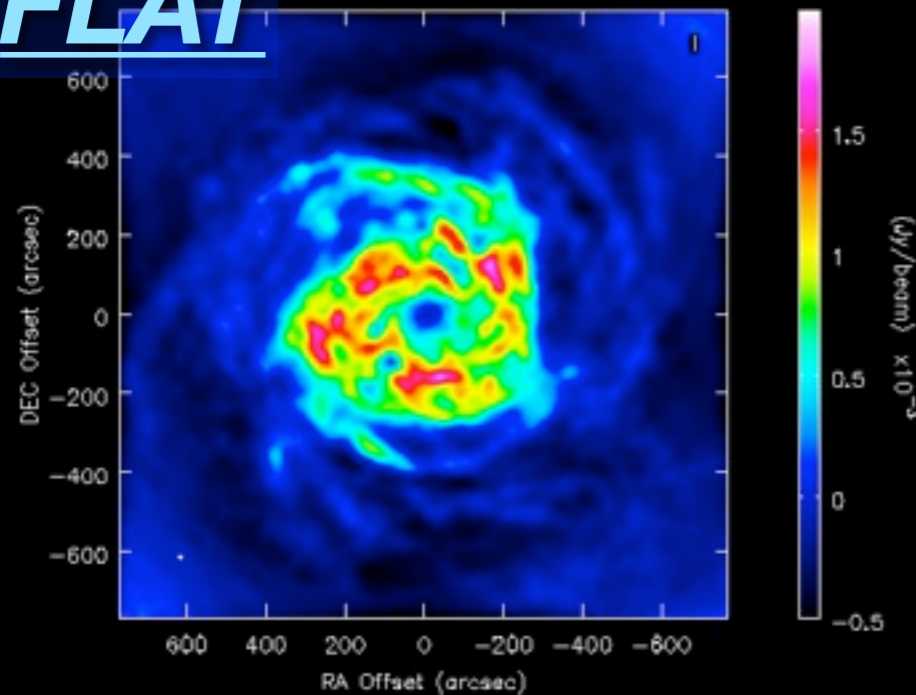


Fidelity

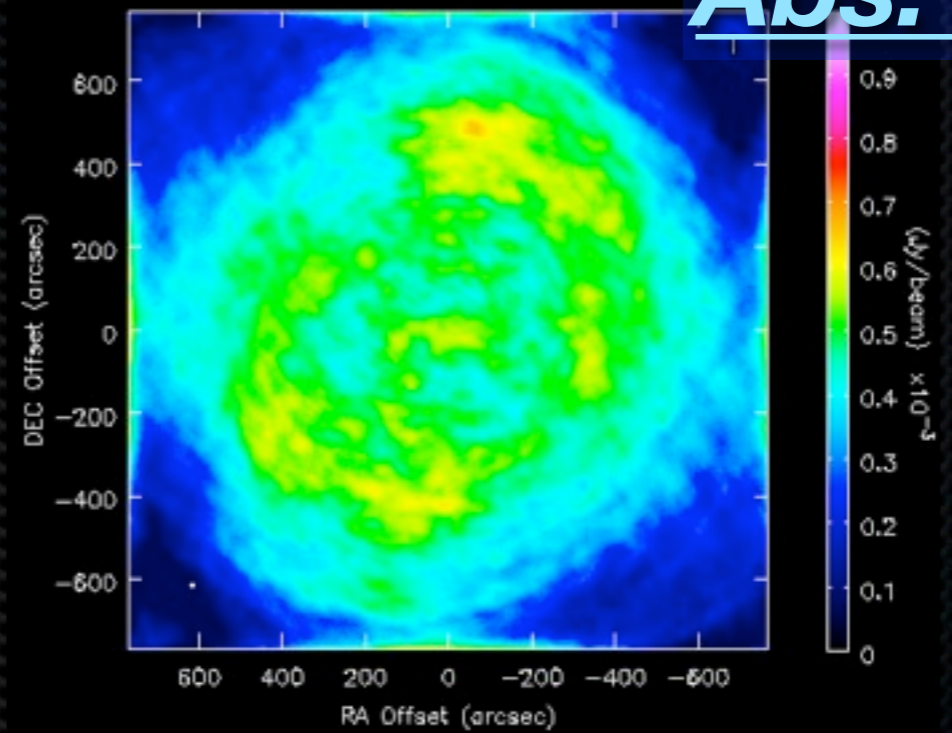


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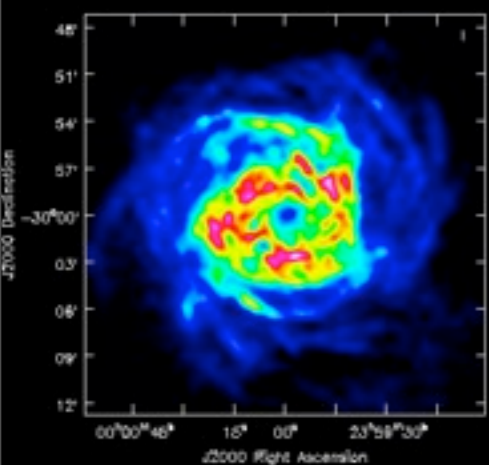
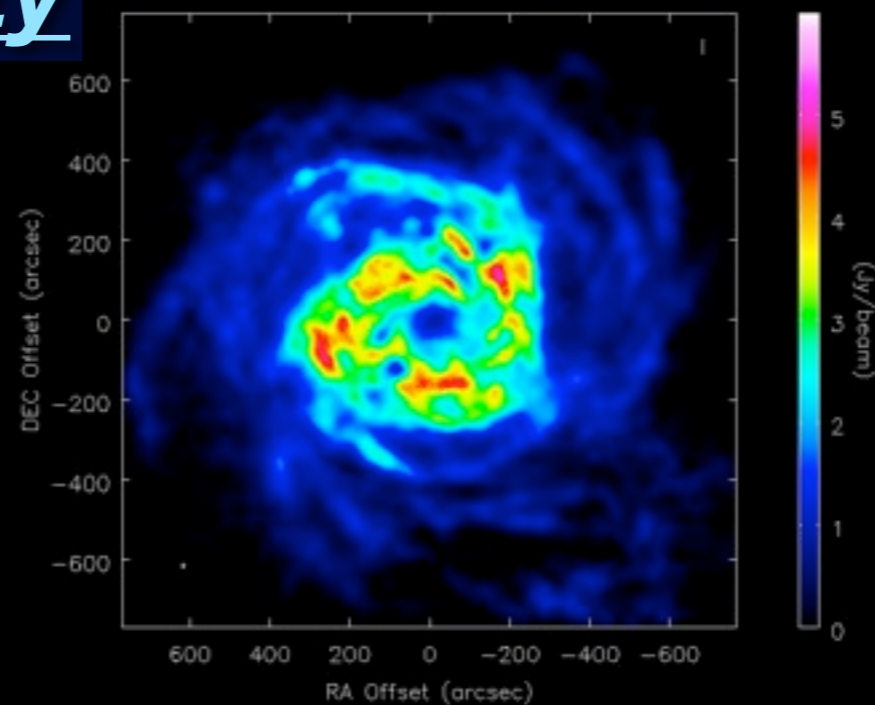
FLAT



Abs. Diff.



Fidelity



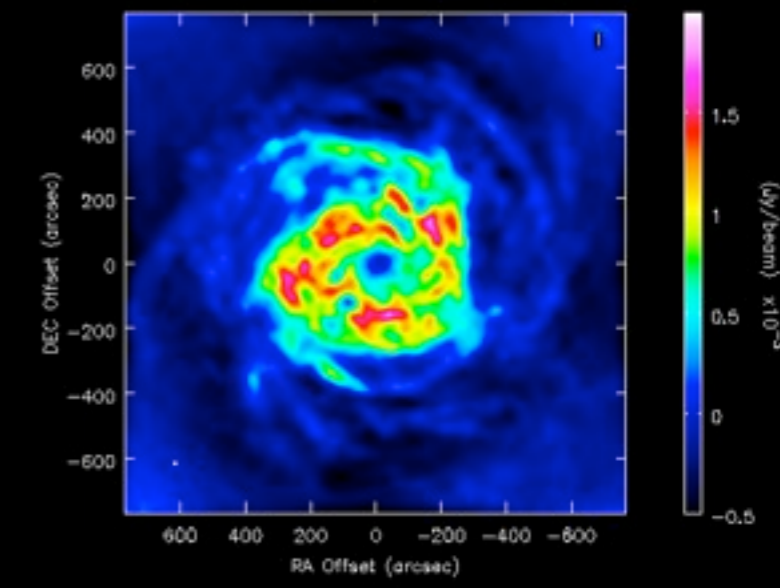
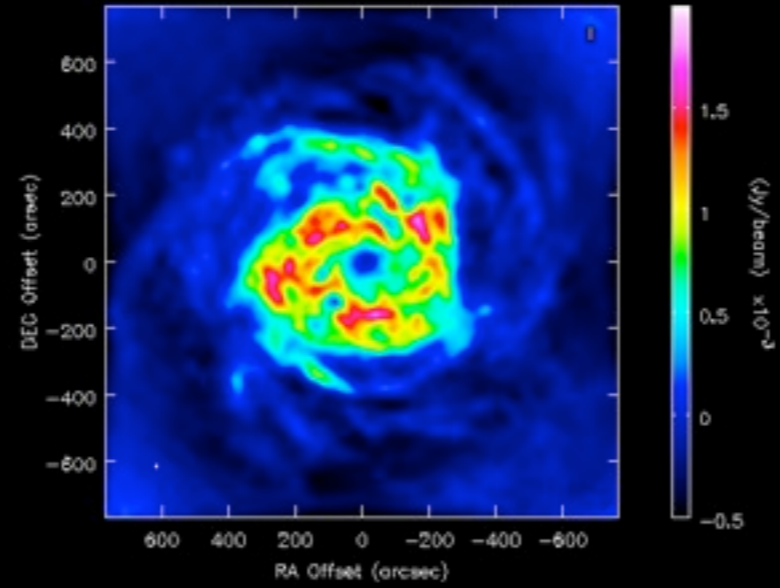
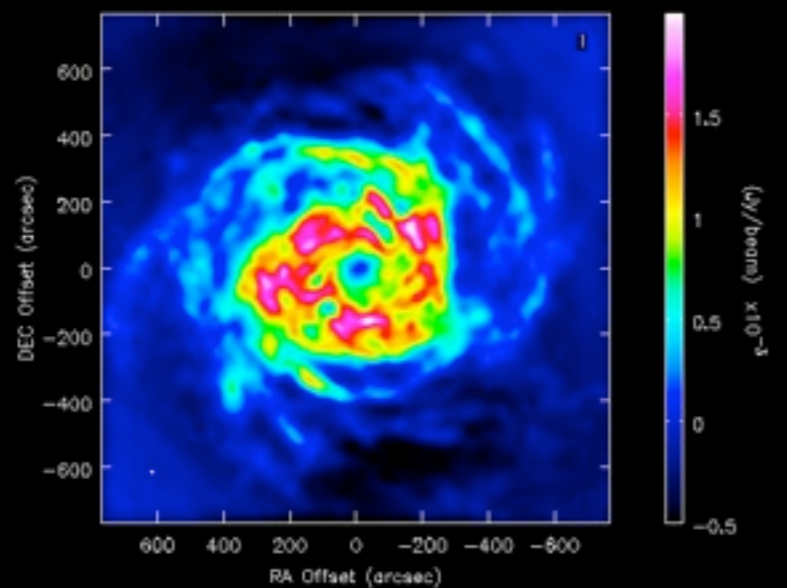
Comparisons

MeerKAT V1

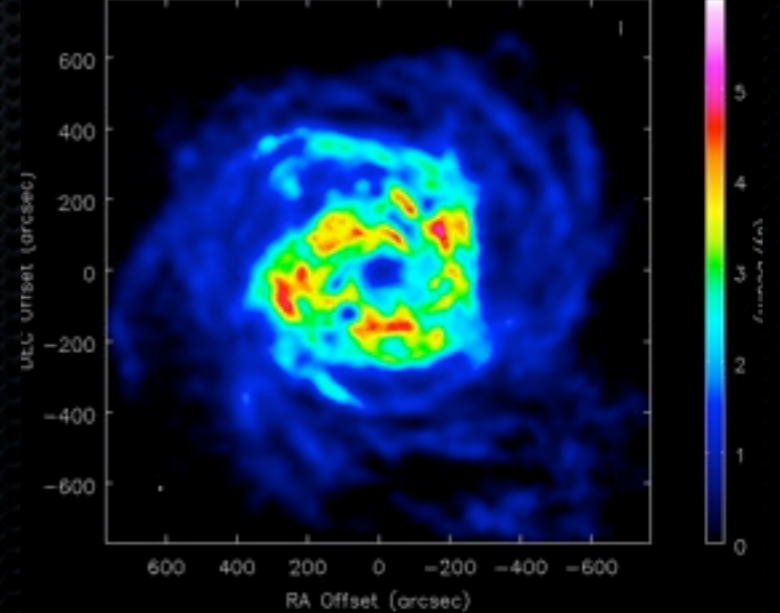
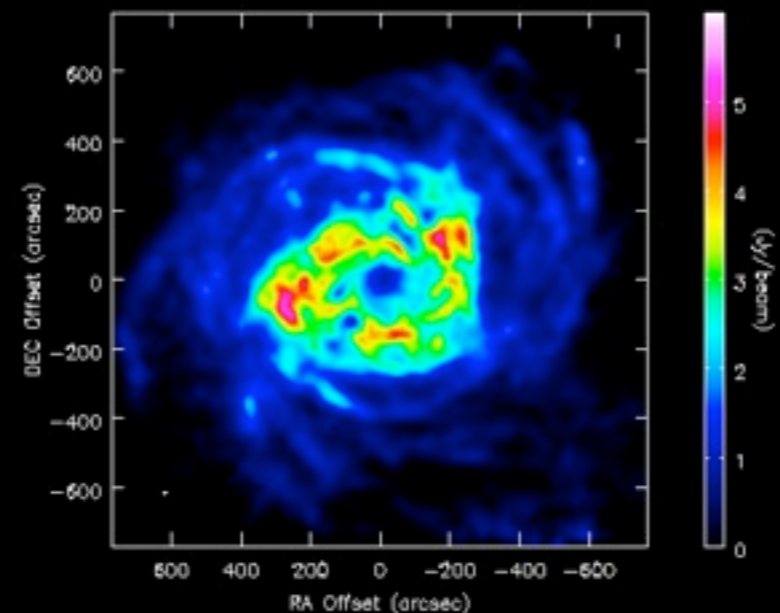
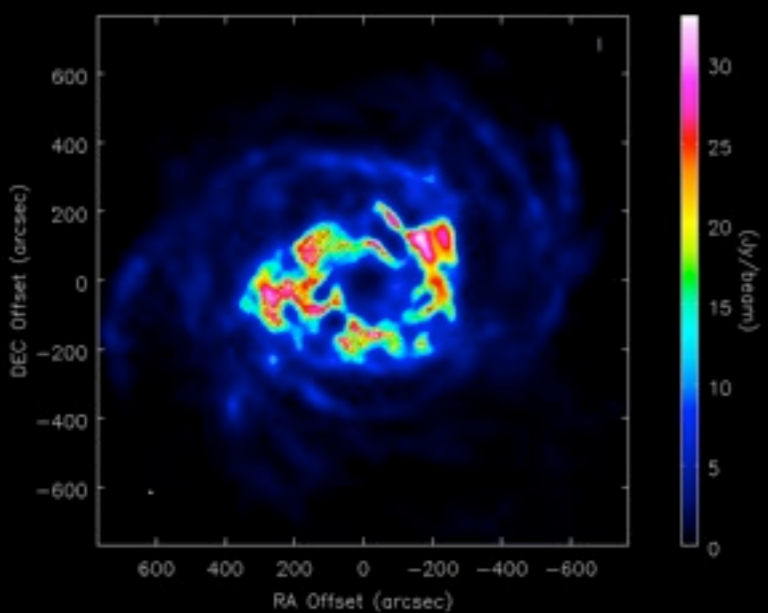
MeerKAT V3.6

MeerKAT-64

Cleaned



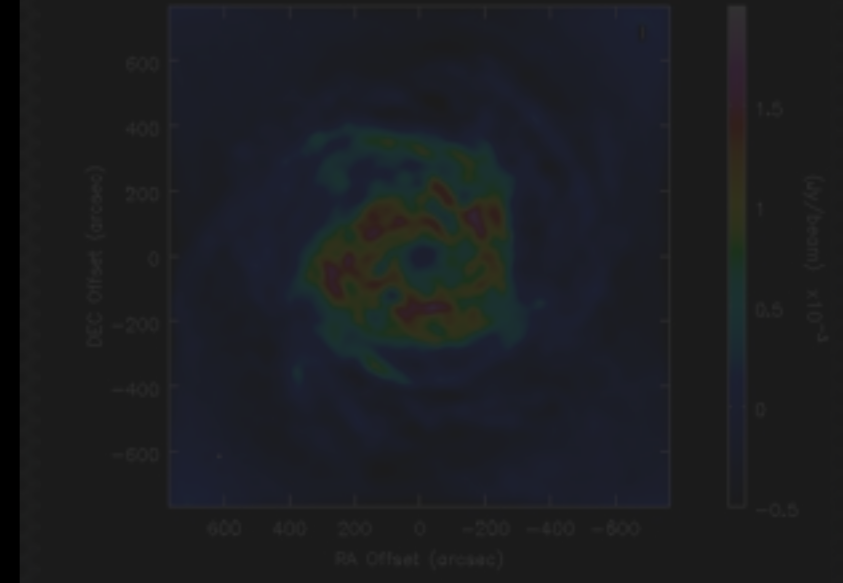
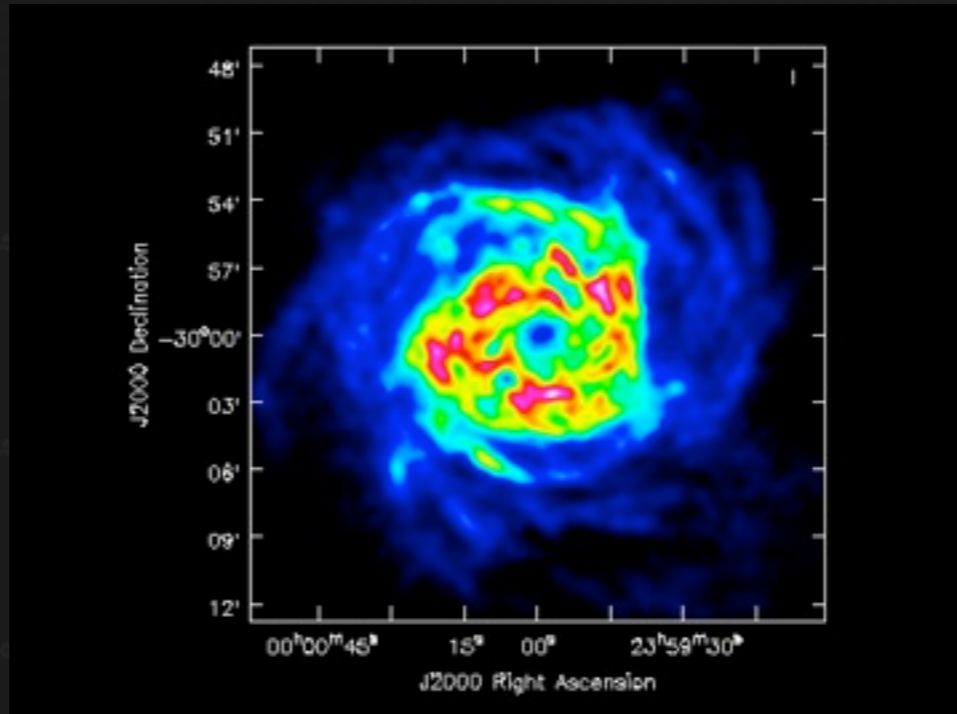
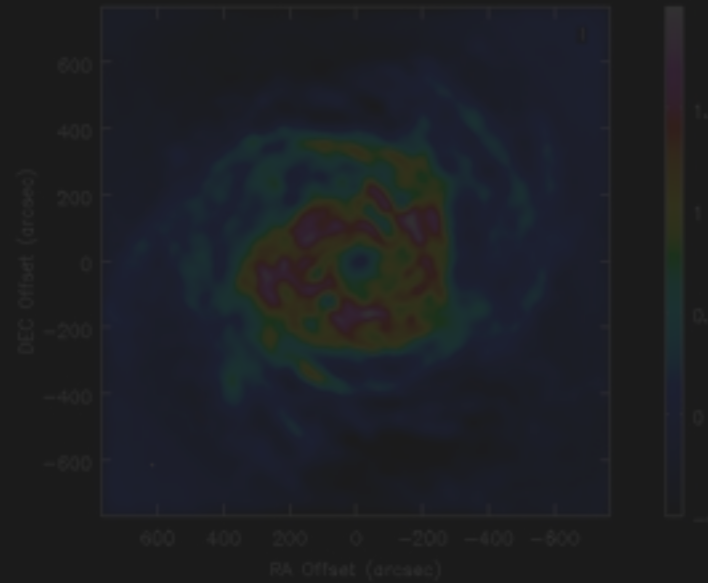
Fidelity



Comparisons

N628 Model

Cleaned

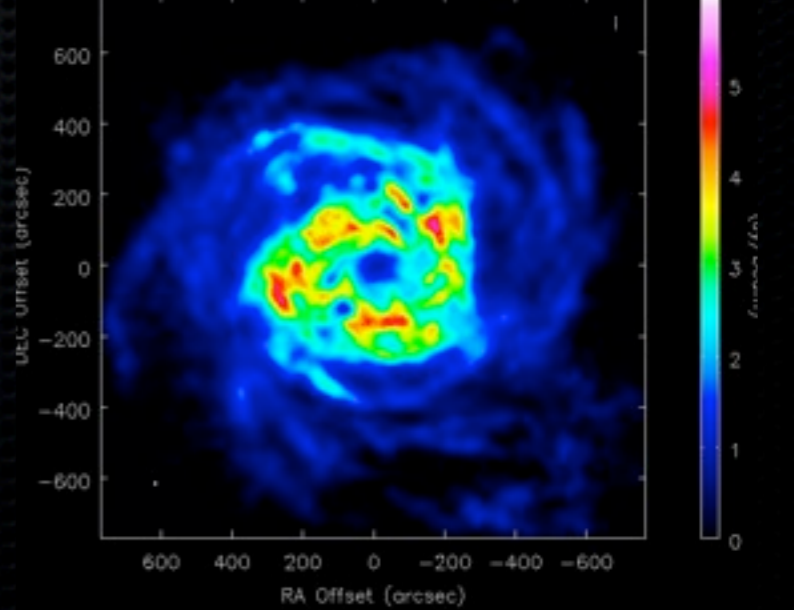
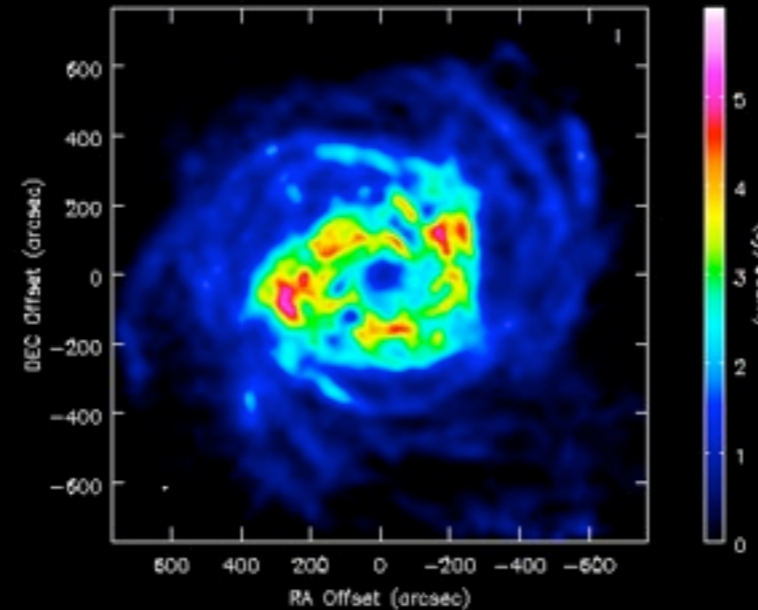
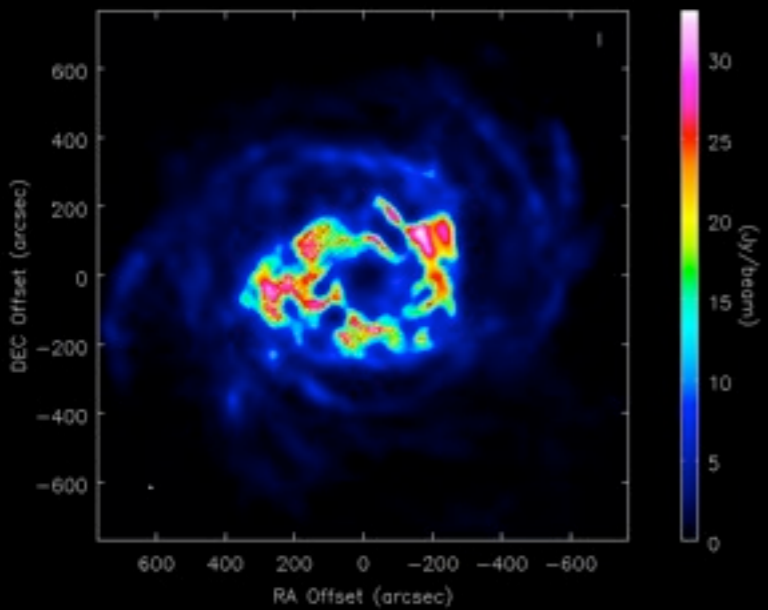


MeerKAT V1

MeerKAT V3.6

MeerKAT-64

Fidelity



Moving Forward...

- Use S^3 for Model Inputs for Deep HI and Continuum
- More Imaging Simulations
- Noisify (only works properly for ALMA)
- Import MeerKAT Specs into CASA
- Release Documentation and Use Cases to Local Community (end of 2010)

Moving Forward...

- First Light Spectral Line Imaging with MeerKAT Demonstrator KAT-7 (2011)
 - Commissioning later this year
- Independent Scientific Component
 - Comparisons of CO and HI Dynamics of THINGS Galaxies
- Member of SKA-SA Configurations Working Group, consultant slave at MeerKAT Project Office



Friday, 7 May 2010