

Coeval Star Formation and Black Hole Growth in the Most Massive Galaxies

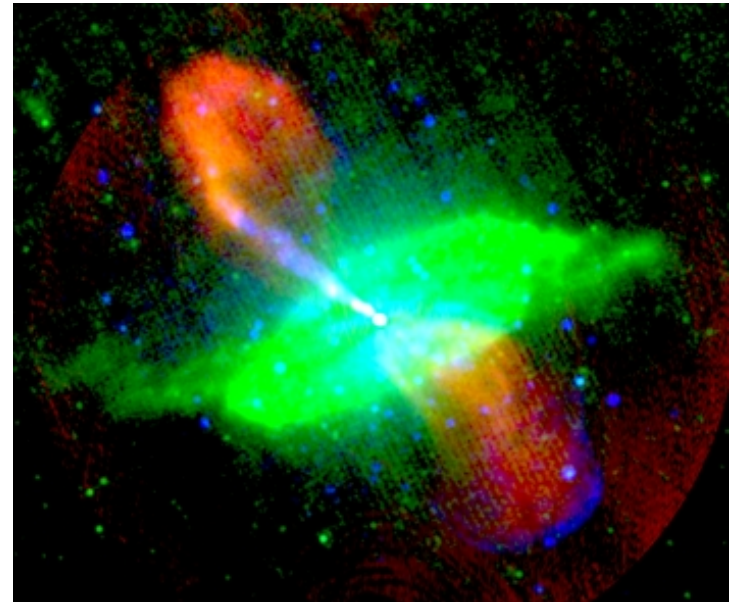
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High Redshift Radio Galaxies (HzRGs)

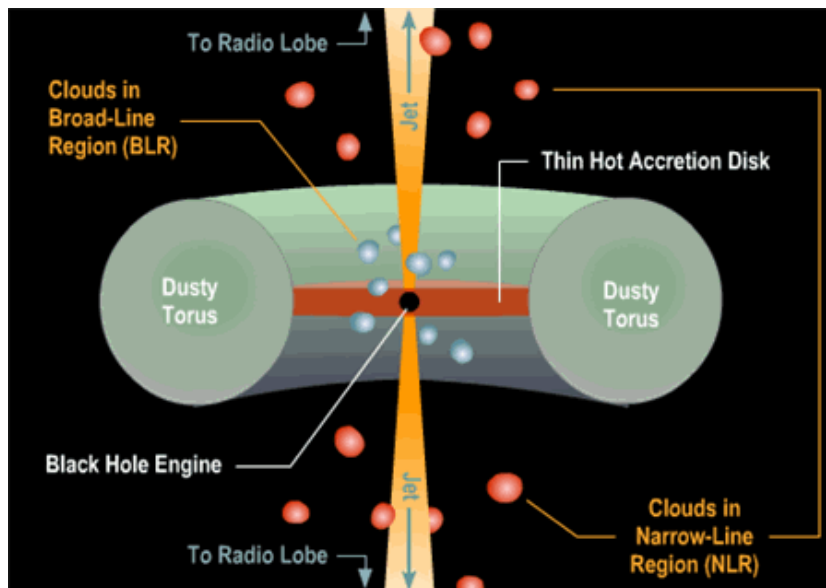
- Mostly hosted by elliptical galaxies
- Among most massive galaxies in Universe
- Contain a dust obscured radio-loud type II AGN and emit radio jets
- High accretion rates
- Host massive black holes due to bulge/black hole mass relation
- Rare objects as radio jets and star formation have very short periods ($<0.1\text{Gyr}$)



Hardcastle, 2006

Aims

- Using Spitzer-IRS MIR data to perform spectroscopy on high redshift radio galaxies (HzRGs)
- Looking for PAH emission and silicate absorption features



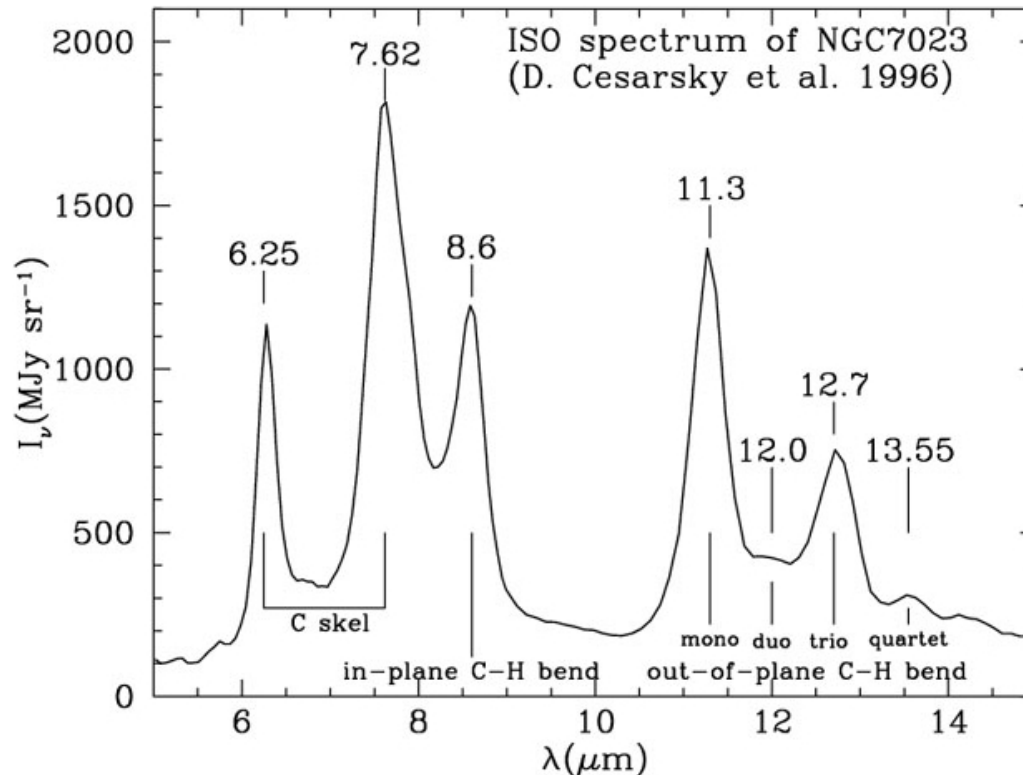
- To measure star-formation rates (SFR) and find out how compact the obscuring dust is around AGN

Motivation

- Find the link between star formation and AGN activity
- And see if central black hole and host galaxy grow coevally- probing their formation and evolution
- Silicate absorption (at 9-10 μ m) can tell us about dust distribution and may indicate dusty starbursts
- Rest-frame mid-IR luminosities will lead to a measure of the accretion rates in these AGN

Polycyclic Aromatic Hydrocarbons (PAHs)

- PAHs are an approximate indicator of galaxy SFRs

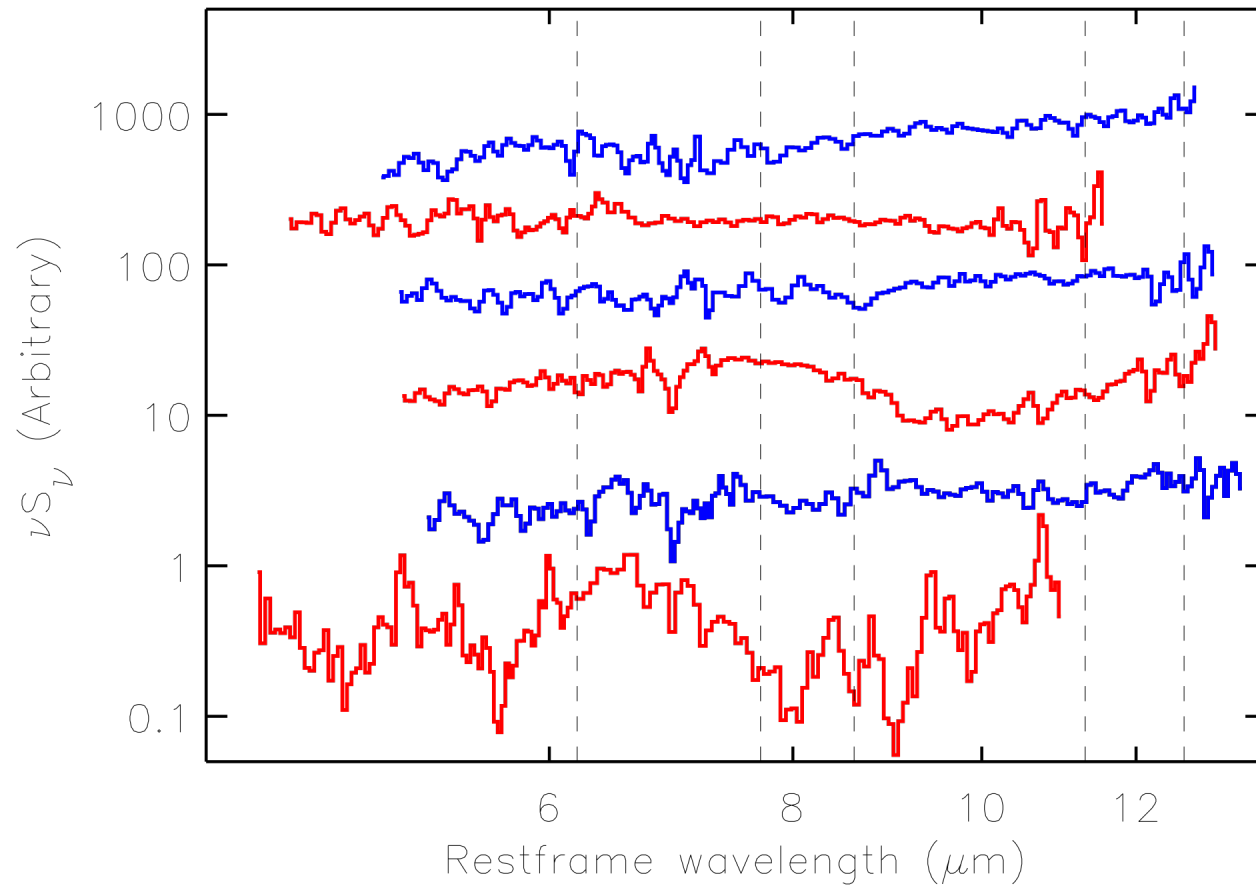


Three data sub-samples:

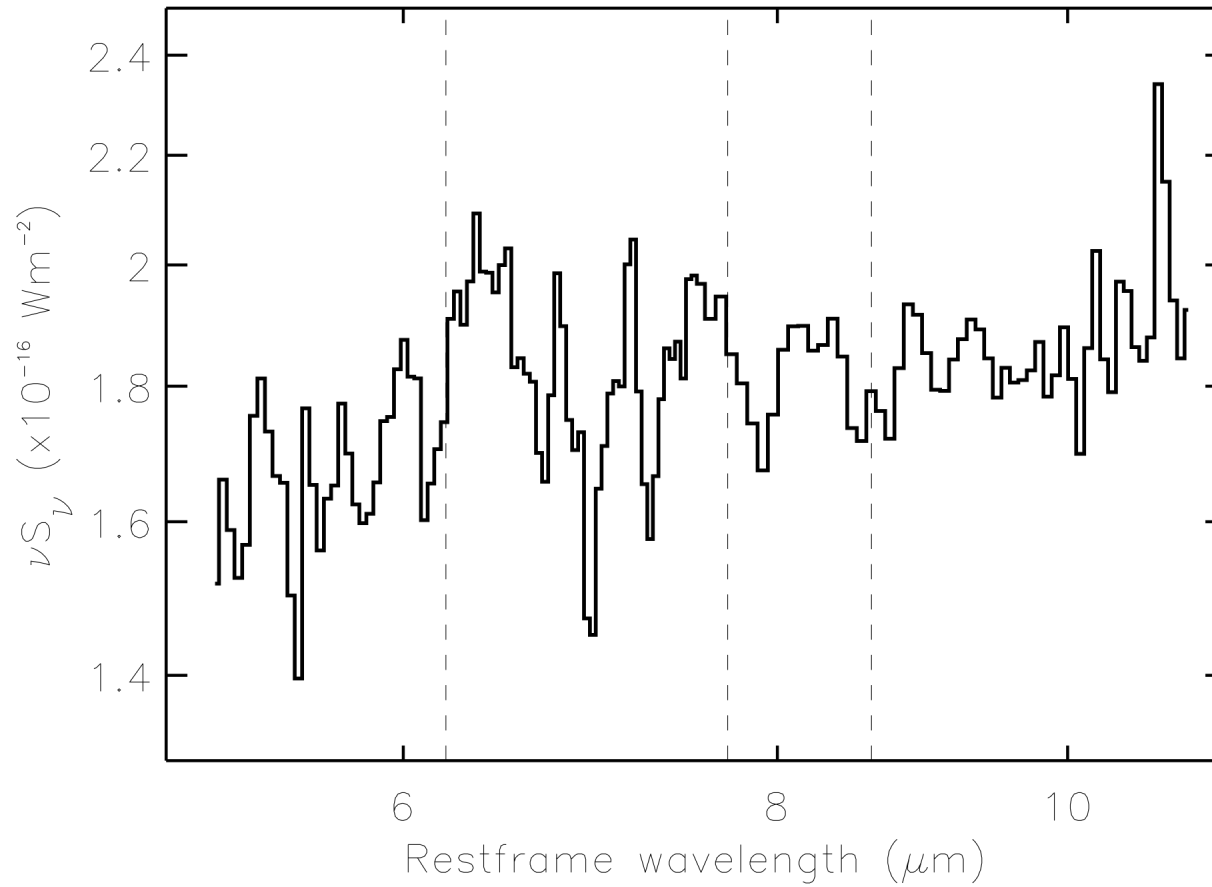
- High luminosity radio galaxies from All-sky surveys
- Lower luminosity radio galaxies from Boötes Field
- Radio luminous MIPS sources from First Look Field

Spectral Results

Individual spectra of the Bootes sub-sample

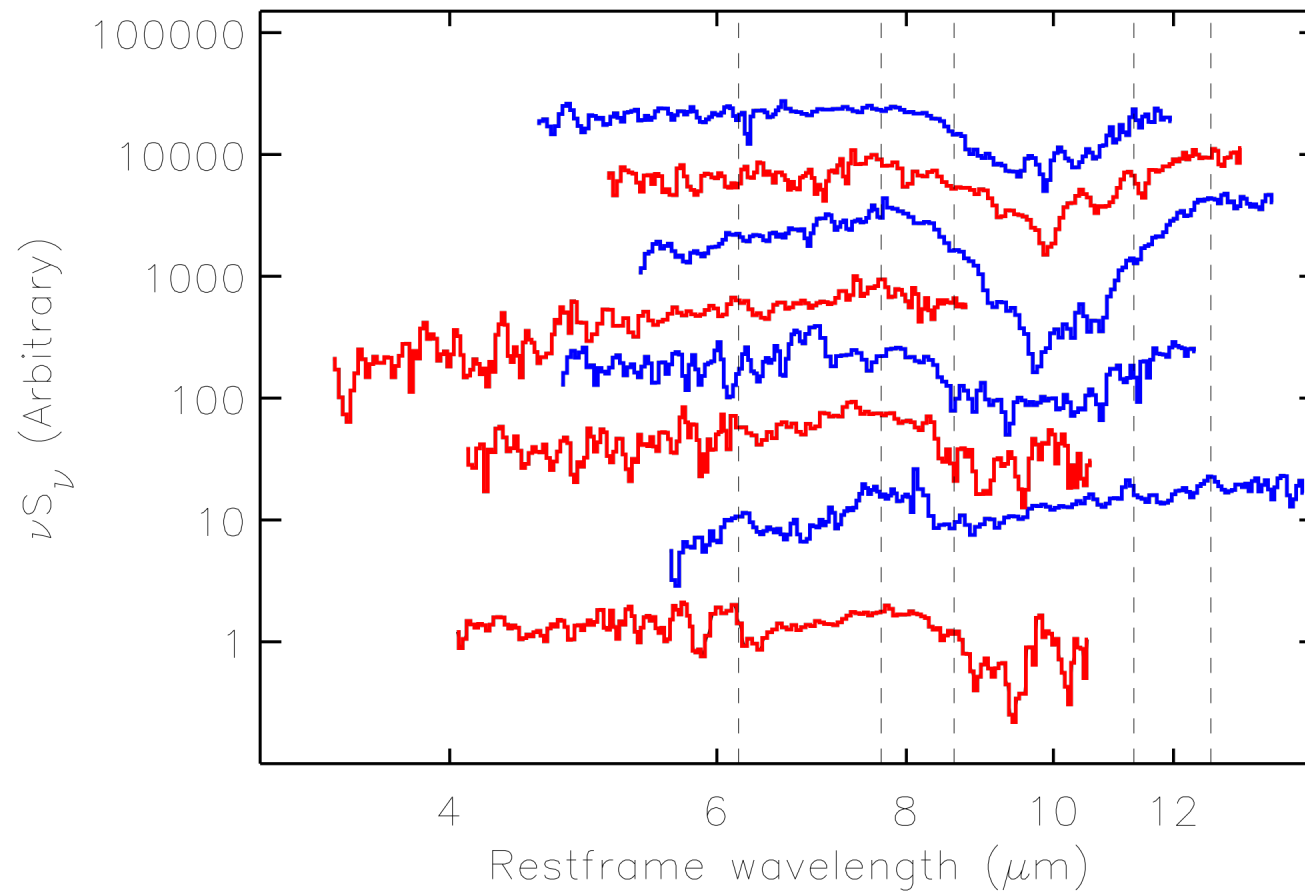


Stacked spectrum of the Bootes sub-sample

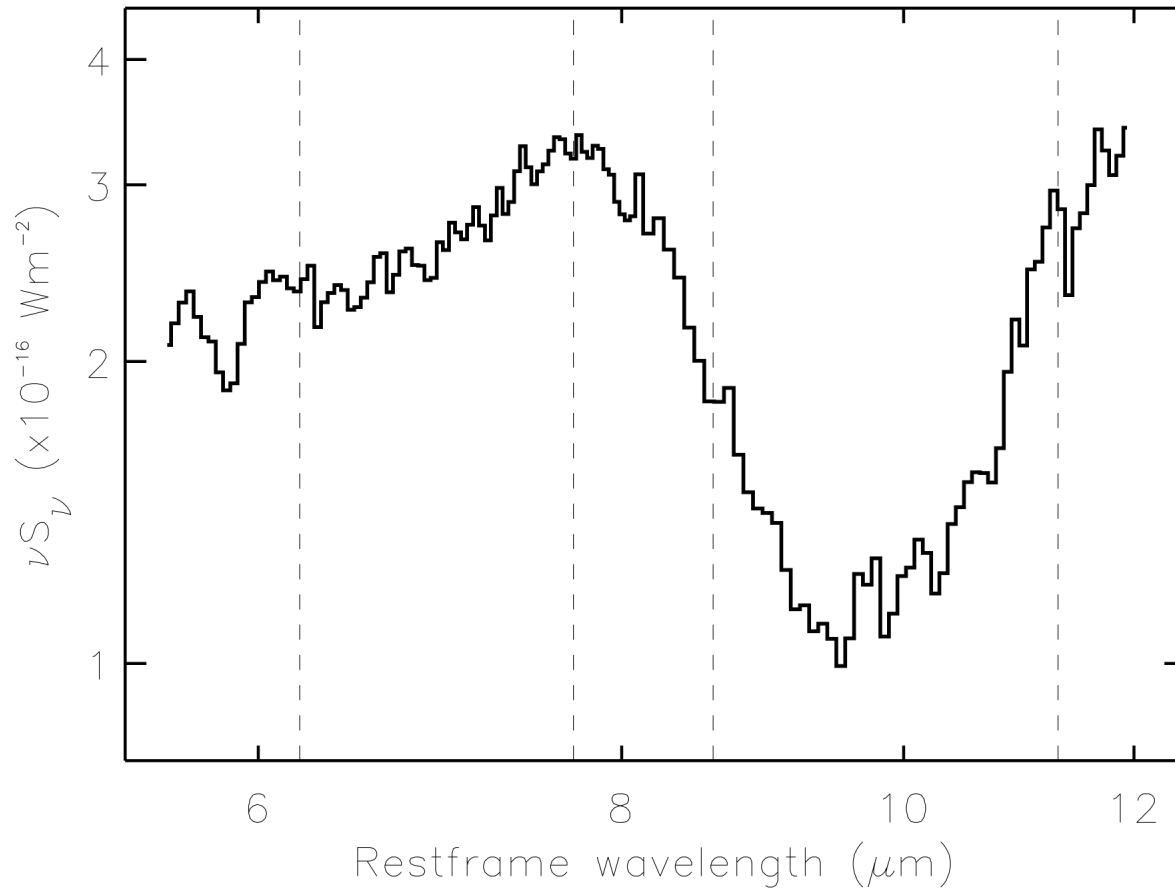


...deep silicate absorption

Individual spectra of the MIPS sub-sample

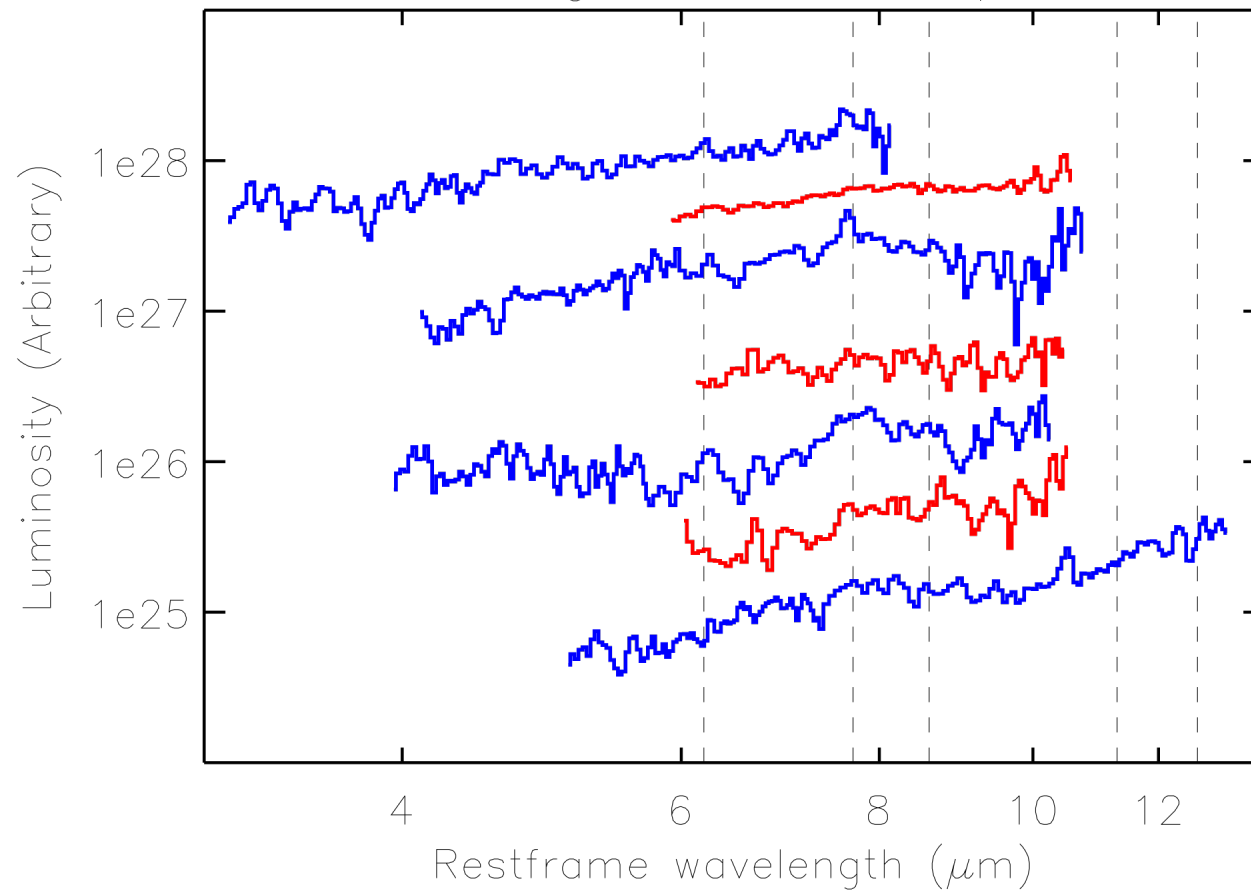


Stacked spectrum of the MIPS sub-sample

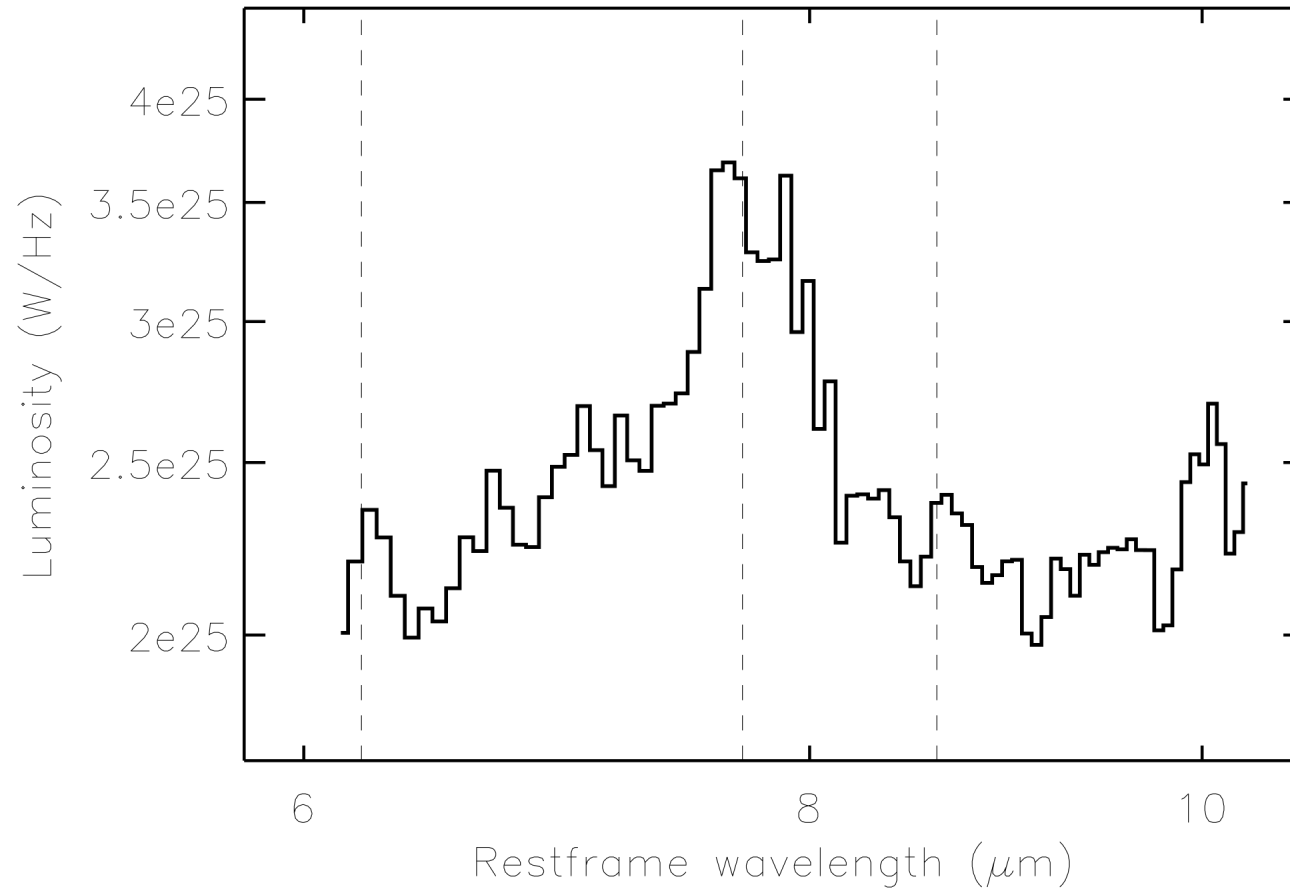


...PAH emission

Individual spectra of the high luminosity radio galaxies sub-sample



Stacked spectrum of the high luminosity radio galaxies sub-sample



Preliminary results

- PAH emission and silicate absorption detected in high luminosity galaxies
- Deep silicate absorption detected in MIPS sources
- No apparent detections in the Bootes field
- Further analysis still to be done:
 - Spectral Energy Distribution (SED) fitting